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**Reacting Fluids Laboratory  
Department of Chemical Engineering  
Louisiana State University**

**COUPLED RADIATING SHOCK LAYERS  
WITH FINITE RATE CHEMISTRY EFFECTS**

(NASA-CR-132735) COUPLED RADIATING SHOCK  
LAYERS WITH FINITE RATE CHEMISTRY EFFECTS.  
PART 2: ANALYSIS OF THE INTERACTION OF  
ABLATING PROTECTION SYSTEMS AND STAGNATION  
REGION HEATING (Louisiana State Univ.)

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**Analysis of the Interaction of Ablating  
Protection Systems and Stagnation Region Heating**

by

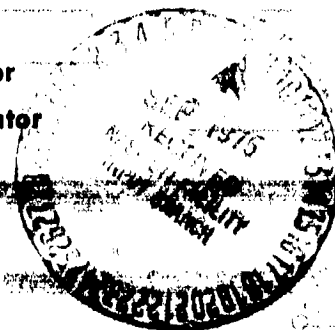
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**Langley Research Center  
National Aeronautics and Space Administration**

**June 1, 1975**

**Baton Rouge, Louisiana**



# PREFACE

This is the fourth in a series of final reports and completes the documentation of research on NASA Grant NGR 19-001-059. It also served as the Ph.D. dissertation for John F. Balhoff.

COUPLED RADIATING SHOCK LAYERS  
WITH FINITE RATE CHEMISTRY EFFECTS

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Doctor of Philosophy

in

The Department of Chemical Engineering

by

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## ABSTRACT

The conservation equations that describe the mass, momentum, and energy transfer in the shock layer around an ablating blunt body were developed and placed into numerical solution form. Effects of line and continuum radiation along with finite rate chemistry were considered. Sixteen chemical reactions with 19 chemical species were included in the analysis, and 12 of these species were employed in the radiative flux divergence calculation. The shock layer was coupled to a phenolic-nylon charring ablator through a surface energy balance.

The resulting conservation equations were a set of parabolic integro-partial differential equations which were quasilinearized and solved using the Crank-Nicolson implicit numerical technique. The partial differential equations required two sets of boundary conditions and one set of initial conditions. The boundary conditions were specified at the shock using the Rankine-Hugoniot relations and conditions at the ablator were determined by employing the nonequilibrium Hertz-Knudsen equation for sublimation and surface energy balance. The stagnation line equations are a set of ordinary integro-differential equations which supply the initial conditions for the analysis. The same boundary conditions were employed at the stagnation line as for the shock layer solution.

Results are presented for coupled solutions at six key points along a reentry trajectory for the stagnation line. Four of these six

solutions were continued around-the-body. A study was also performed to determine the uncertainty in radiative heating due to the uncertainty in chemical kinetic data.

On the basis of this research, it was concluded that the numerical solution of the quasilinearized conservation equations yielded stable and accurate solutions to the nonequilibrium, radiating shock layer surrounding a blunt reentry body. These nonequilibrium solutions, when compared to equilibrium solutions at comparable conditions yielded consistently lower wall radiative heating rates. The difference was most pronounced at low heating rates (e.g. 121 BTU/sq. ft-sec for nonequilibrium) where the equilibrium results were 190% higher. At higher heating rates (~2000 BTU/sq. ft-sec) the equilibrium results were found to be generally less than 20% higher than nonequilibrium. The species profiles for low heating rates differed significantly especially in the air layer from the nonequilibrium and equilibrium analyses whereas the species profiles for the high heating rate cases were very similar.

For the around-the-body analyses, it was observed that one nondimensional blowing profile along the ablator surface was sufficient for all four cases to adequately couple the shock layer with the ablator response. The coupling between ablator and shock layer was found to differ by only 16% for the worst case.

The choice of species boundary conditions at the ablator wall does not greatly affect the shock layer solution. The surface heating rate was found to differ by only 4% for the case of a nonequilibrium wall versus an equilibrium wall. However, the

choice of species boundary conditions at the shock does appear to affect the solution. The species composition at the shock were assumed to be equilibrium. In light of the results of this investigation, the assumption of equilibrium appears to be inaccurate (especially at low heating rates) and further study is recommended in this area.

One of the major problems encountered in this study was the tremendous amount of computer time necessary. Over 100 hours of computer time on an IBM 360/65 computer was required. The specification of a correct shock standoff distance did not alter the heating rate analysis, but the amount of computer time was found to decrease significantly when the converged shock standoff distance was used. It is felt that more study is needed in the area of shock standoff calculations.

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## CHAPTER I

### INTRODUCTION AND BACKGROUND

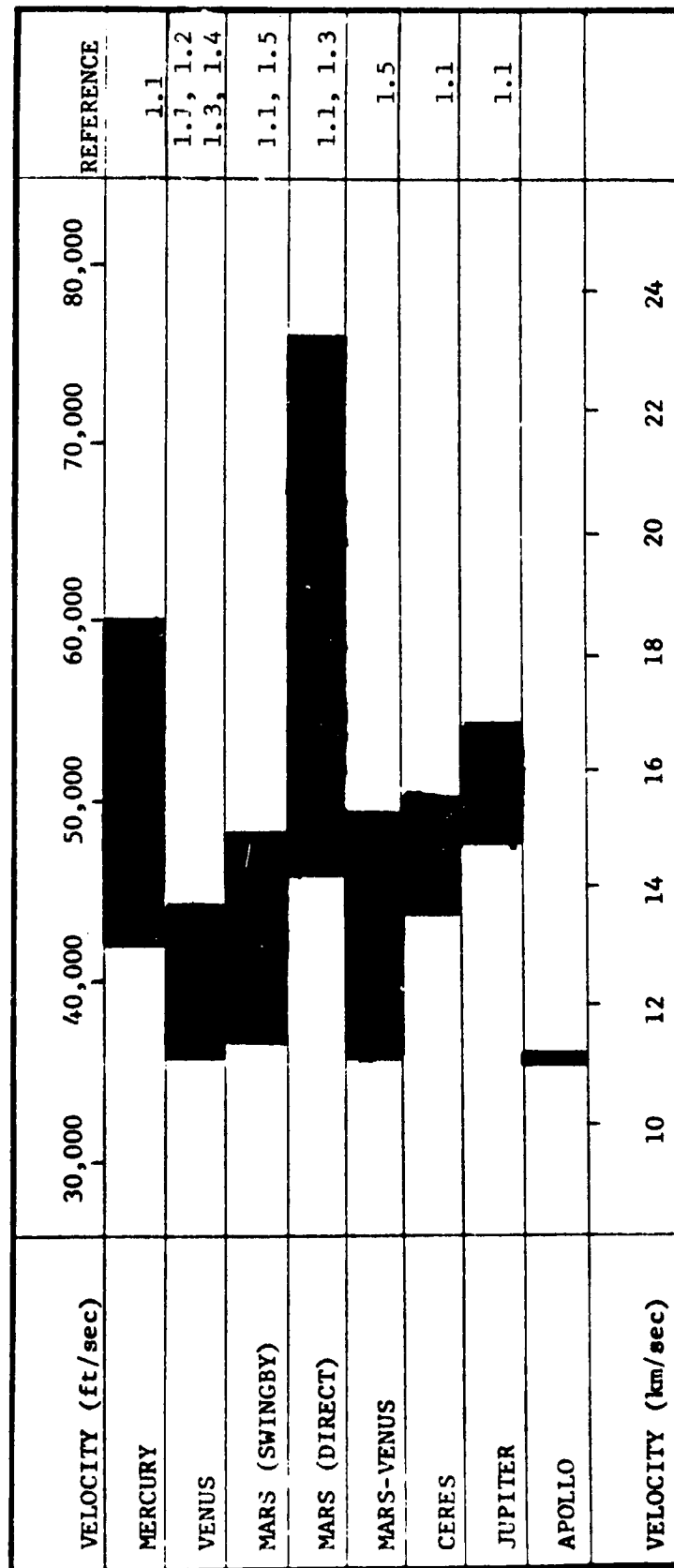
#### Introduction

Aerodynamic heating is one of the most severe problems encountered by a reentry vehicle when returning from interplanetary missions. A reentry vehicle with a blunt body configuration and having a charring ablative heat shield has been found to be the most effective design for combatting this problem. This chapter discusses the advantages of the blunt body configuration over other alternatives and presents a detailed description of the processes which occur in a charring ablator and its surrounding flow field environment. The description includes a comprehensive overview of the intricate heat and mass transport mechanisms involved in the ablative heat shield and the shock heated air layer flowing over the reentry capsule.

#### The Problem of Atmospheric Entry

A reentry vehicle encounters extreme heating during hyperbolic atmospheric entry when utilizing aerodynamic braking. This heating is a result of applying the frictional resistance of the atmosphere to decrease the speed of the vehicle. Since kinetic energy ( $KE = MV^2/2$ ) is a function of the square of the velocity, it is necessary to know the ranges of relevant velocities for manned missions. In Figure 1.1 these ranges for flights to the planets and and the ranges encountered by Apollo are given. Notice that Apollo

Figure 1.1  
Expected Earth Entry Velocities for Several Planetary Missions



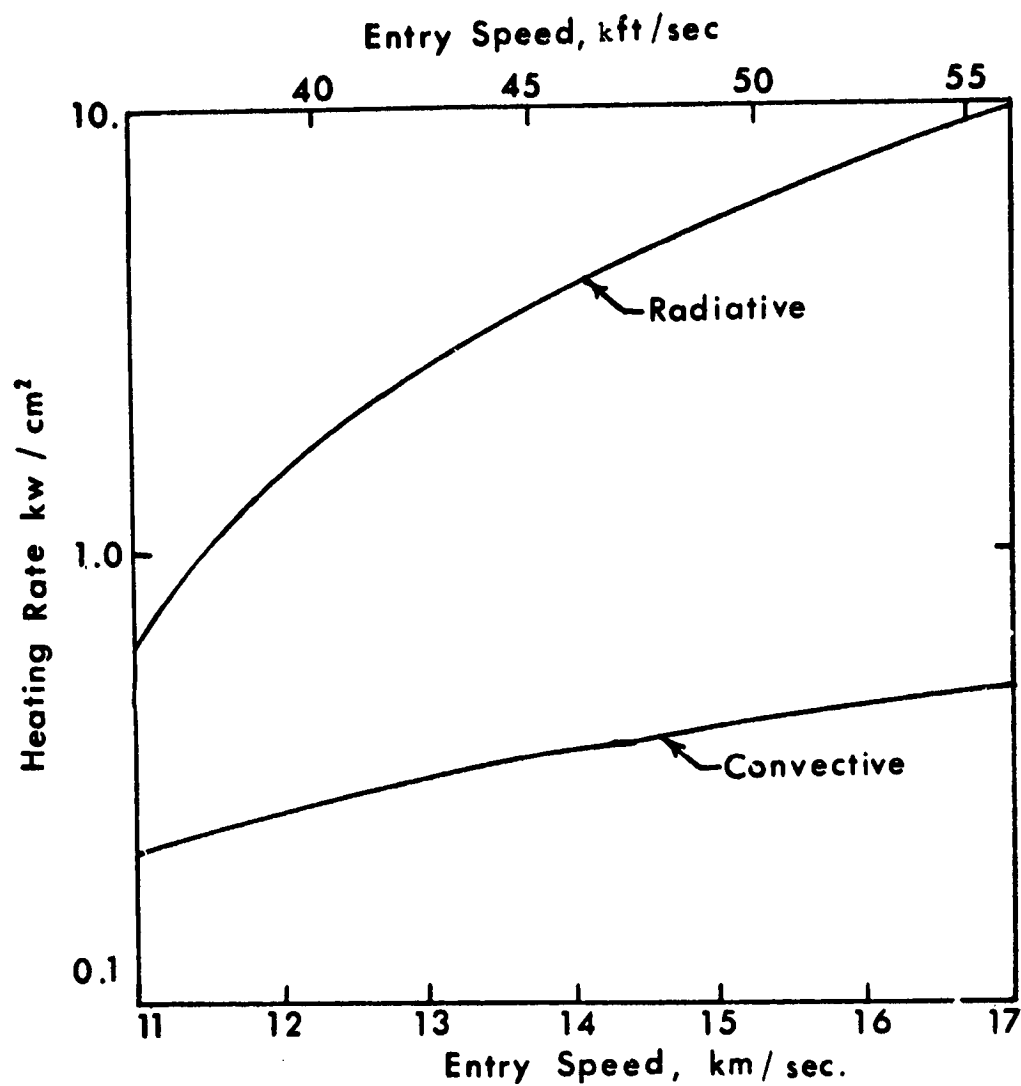


Figure 1.2. Peak Stagnation Point Heating Rates  
(Ref. 1.1).

has the lowest return velocity range. This kinetic energy upon reentry is transformed into thermal energy and a bow shock wave is formed. Passing through the shock wave, the air is heated as it flows over the heat shield of the reentry vehicle. The air layer then transfers a portion of this heat to the surface of the space capsule primarily by convection and radiation. The relative importance of convective and radiative heating can be seen in Figure 1.2. Note that radiation heat transfer becomes increasingly important as the reentry velocity increases. Thus, it is extremely important to have a reasonable description of the radiative transport mechanisms for return from planetary missions ~~where the velocities are much~~ higher than Apollo. In Figure 1.3, the region where radiative heat transfer is significant is shown.

The alternative to aerodynamic braking would be to slow down the vehicle by using reverse thrust. Since the return velocity is of the same magnitude as the launch velocity, the reverse thrust method requires the same amount of fuel for reentry, therefore doubling the fuel requirement and adding unnecessary weight. In terms of weight efficiency for entry deceleration, an ablation protection system requires 10 to 50 times less entry vehicle weight than would be required by the reverse thrust method (Ref. 1.7). Excessive weight on a return capsule is quite prohibitive since it has been estimated that each kilogram carried throughout a manned planetary mission can represent between 300 and 1000 kg on the launch pad (Ref. 1.1). On this basis, excess weight cannot be tolerated especially in the return phase of the mission and reverse thrust becomes unsuitable.



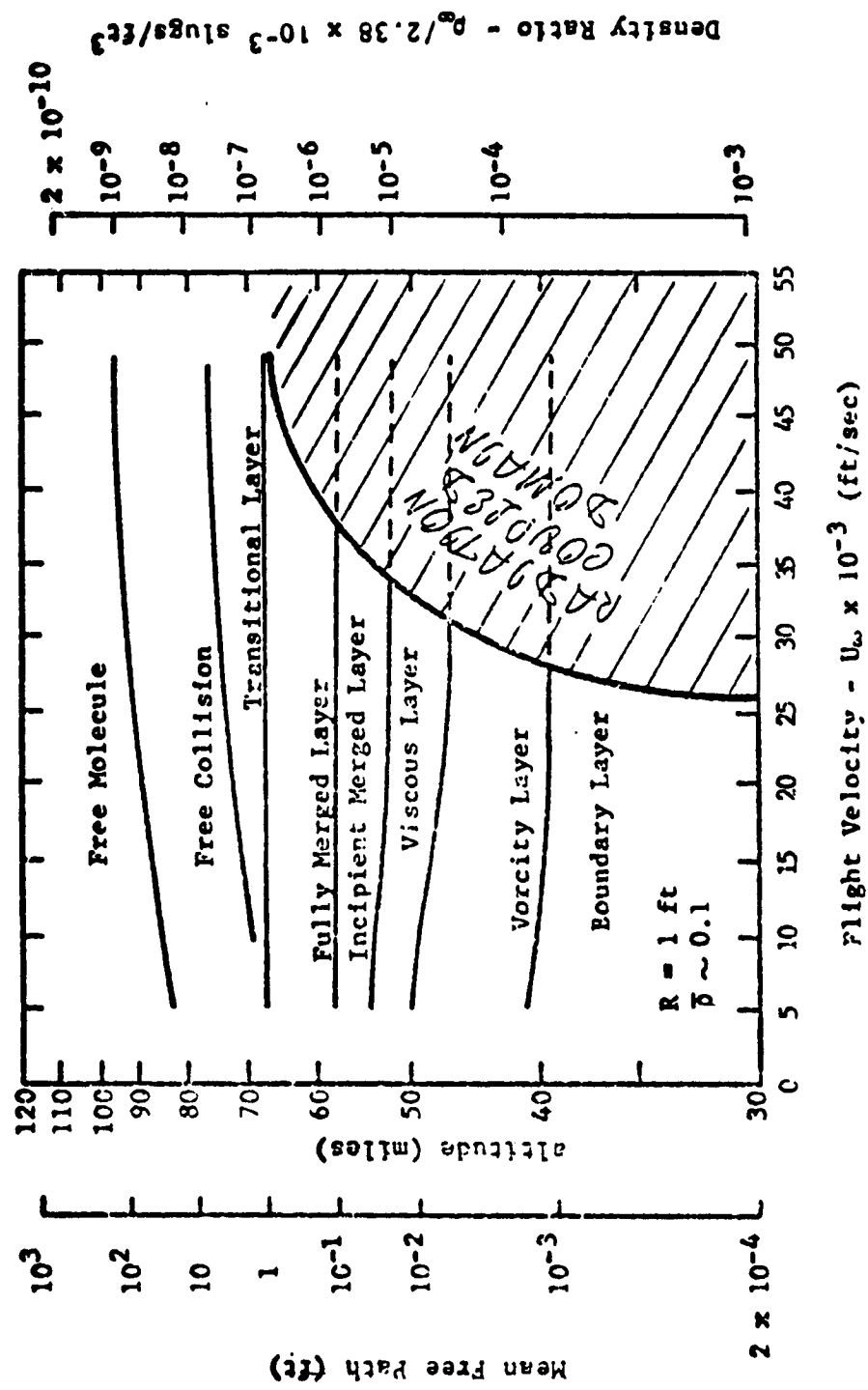


Figure 1.3. Flight Regimes (Modified from Ref. 1.6)

First it was thought that a slender body would be the logical design. The slender body does have the smallest aerodynamic drag. However, this shape produces a weak, attached shock wave and a large percentage of the heat generated on reentry is absorbed by the body. Slender body configurations are best suited for ballistic missiles and supersonic flow applications where low heating loads are experienced for short durations (Ref. 1.8).

The blunt or high drag configuration is more applicable to reentry of space vehicles since these have a much larger fraction of the total energy transfer to the atmosphere. An extremely strong, detached shock wave is formed around the blunted vehicle causing a major portion of the energy to be absorbed by the air flowing between the shock wave and the vehicle surface. This energy is then carried away in the wake behind the craft. The shock heated layer surrounding the vehicle becomes progressively hotter during the course of reentry causing dissociation and ionization of the air. Heat is transferred by convection and radiation to the capsule surface from the shock layer. The amount of heat transferred to the space craft is enough to cause surface temperatures of the order of  $3600^{\circ}\text{K}$ . This high temperature occurs despite the fact that most of the heat has been diverted away from the craft by using the blunt body configuration. The space capsule cannot survive these temperatures unless an efficient thermal protection system is used. The answer to heat shielding of reentry vehicles was the charring ablator which is discussed subsequently.

### The Charring Ablator

To dissipate the tremendous amount of heat that is generated at reentry, several methods were suggested (Ref. 1.9, 1.10, 1.11). These are (1) Heat Sink, (2) Radiation Cooling, (3) Transpiration Cooling, (4) Convection Cooling, (5) Film Cooling, and (6) Ablative Cooling. The most important of these is ablative cooling since it combines the advantages of all of the other methods. The combination of wide range of application with high reliability made the system extremely successful.

The charring ablator is a special type of ablator with qualities which make it a superb heat shielding material. The charring ablator utilized several mechanisms to reduce the heat transferred to the interior of the space craft. This type of ablator is normally made up of a composite virgin plastic such as phenolic resin and carbon or phenolic resin and nylon. As shown in Figure 1.4, the virgin material of these composites decomposes when subjected to intense heating and forms a porous carbon matrix along with pyrolysis gases which flow through and react in this porous carbon matrix. The carbon surface is removed by chemical reactions, sublimation and is eroded away by the shearing force in the shock layer (spallation). The pyrolysis gases flow through the char, reacting with themselves and the char layer and finally are injected into the boundary layer at the surface.

The ablative process is an intricate, complex, and yet orderly transfer of heat and mass. The important mechanisms involved are (a) heat conduction into the material and storage by its effective heat capacity, (b) heat absorption by the heat capacity of the

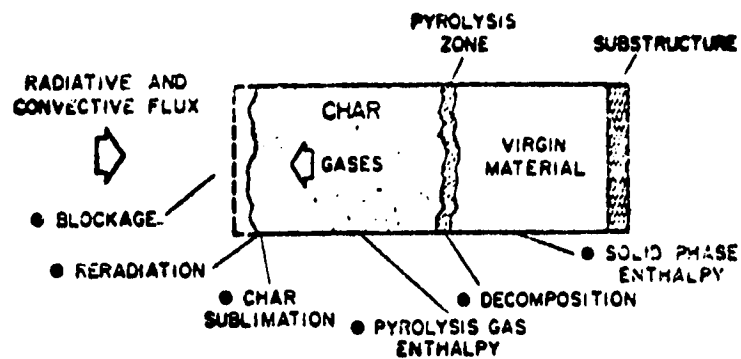


Figure 1.4. Schematic of Ablator Energy Absorption Mechanisms.

PHENOLIC NYLON  
BLUNT BODY

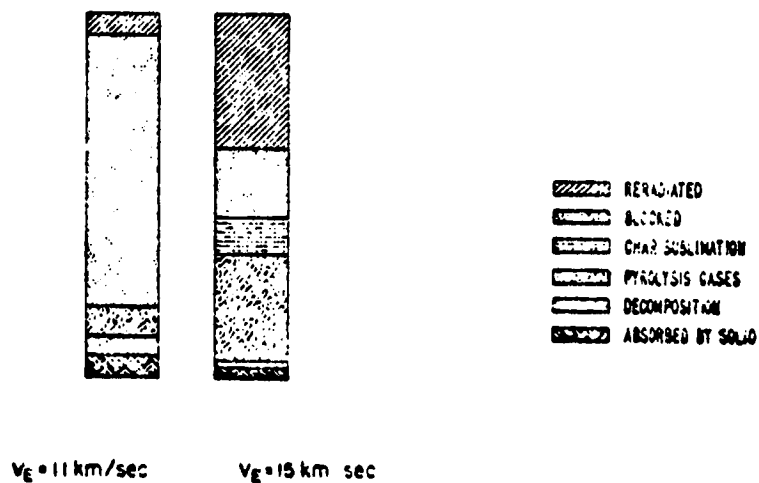


Figure 1.5. Heat Rejection Mechanisms for a Phenolic Nylon, Blunt Body Ablator.

pyrolysis gases, (c) primarily endothermic reactions of the gases (homogeneous and heterogeneous), (d) sublimation and surface reactions, (e) surface reradiation, (f) blowing into the shock layer by the pyrolysis gases, and (g) blockage of shock layer radiation by ablation species. This flow of pyrolysis gases through the char layer gives the charring ablator the added effect of transpiration cooling.

The degradation of the primary pyrolysis gases when flowing through the char into smaller molecules has two very important effects (1) decomposition is endothermic, because heat is consumed in the breaking of chemical bonds, and (2) the overall heat capacity of the smaller molecules is larger than that of the original decomposition products.

It is essential for the ablator to have a low thermal conductivity, a high thermal capacity, and a high heat of sublimation. The superiority of the charring ablator over subliming and melting ablators is largely due to its ability to reradiate substantial quantities of energy from the surface. The surface of the char possesses a high-emissivity; and, at high temperatures, reradiation is one of the most important energy absorbing mechanisms for charring ablators. For most carbonaceous materials, a reasonable assumption is constant emissivity and constant absorptivity (absorption is important when considering radiating shock layers). The values of emissivity and absorptivity are not necessarily equal, though, due to the fact that the radiation and reradiation normally involve different wave-length ranges (Ref. 1.12).

One of the outstanding features of an ablator is that it not only absorbs heat, but through the injection of gaseous ablation

products, it also modifies the adjacent boundary layer and greatly reduces the level of aerodynamic convective heating. These injected gases have the additional property of being good radiation absorbers, thus reducing the radiative heat transferred to the surface. More will be said about these two phenomena in the following section when discussing the flow field interaction.

The relative importance of the various heat rejection mechanisms is shown in Figure 1.5 for two entry velocities of 11 km/sec and 15 km/sec. Only the energy which reaches the ablator surface is considered, therefore the reduction of convective and radiative heating by gaseous injection is not included. At the higher velocity of 15 km/sec, the heat rejection modes of reradiation, char sublimation, and pyrolysis gas heat capacity and reactions become increasingly significant.

#### Flow Field Interaction

The ablator must be described in its proper environment in order to visualize the physical phenomena occurring. In the case of a reentry vehicle with an ablative heat shield, this environment is a shock heated shock layer shown schematically in Figure 1.6. At the outer edge of the shock layer, shock heated atmospheric gases ( $\sim 15,000^\circ\text{K}$ ) move primarily by convection toward the stagnation point, where diffusive and viscous mechanisms become predominant. The region in the immediate vicinity of the stagnation point is thus referred to as the diffusion or viscous layer. The region between the bow shock and the diffusion layer is called the air layer and the ablation layer is located between the diffusion zone and the

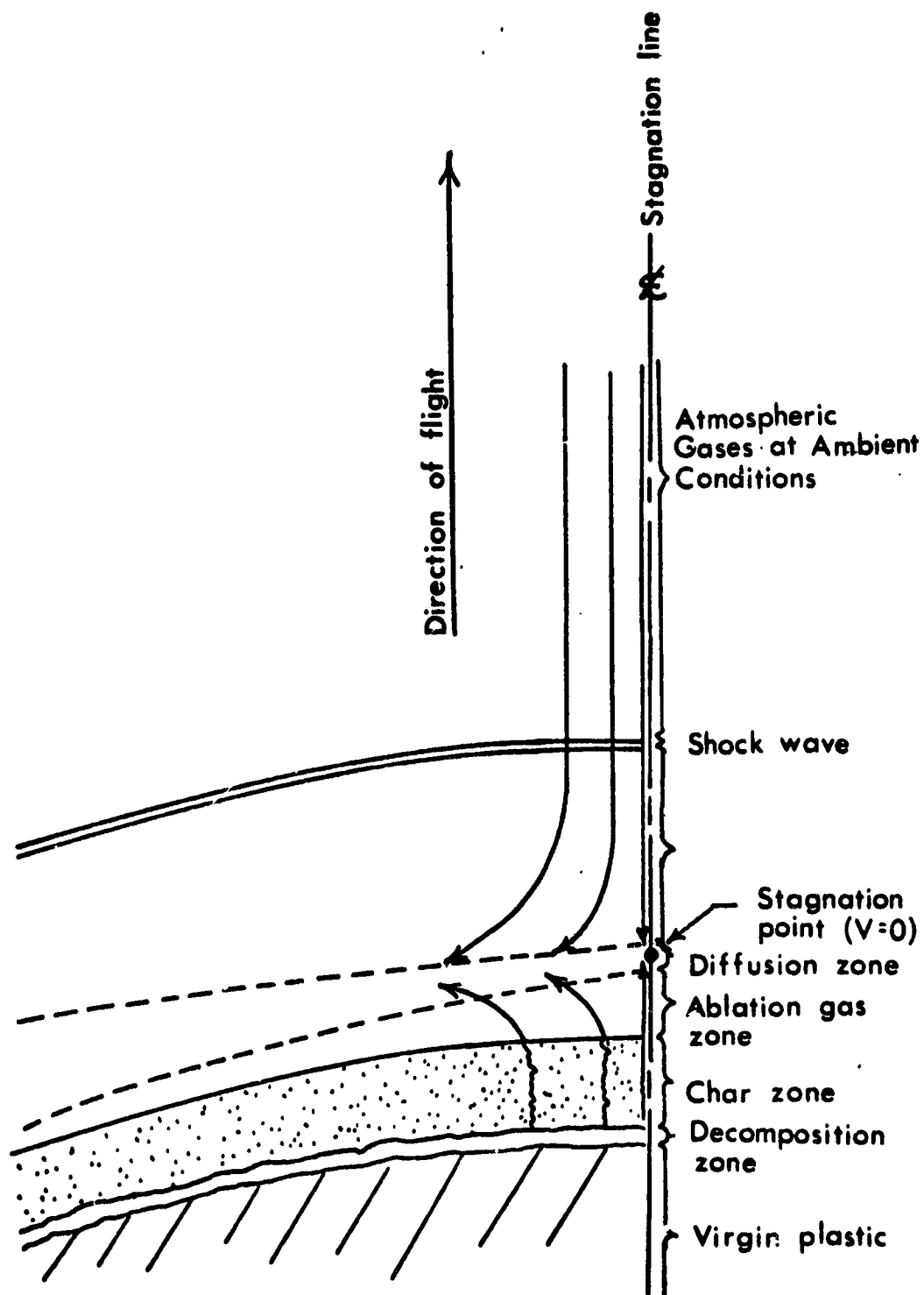


Figure 1.6. Illustration of Charring-Ablator and Flowfield Interaction.

ablator. As in the air layer, convective effects also predominate in the ablation layer.

Although simply depicted in Figure 1.6, the overall process is extremely complex and presents a significant challenge to an accurate mathematical analysis. The major difficulties are describing the radiative heating, the chemical reactions, and the interaction between the flow field and the ablator. In Figure 1.7 the modes of heat transfer in the shock layer are illustrated along with the interactions with the ablator surface. The following discussion describes these heat transfer modes and the coupling that exists between the flow field and the ablator.

In the air layer the predominant radiation effect is emission which occurs in all directions. The radiant heat emitted away from the body through the bow shock, is absorbed by the ambient air and is called precursor radiation. This effect increases the free stream enthalpy, thus elevating the temperatures in the shock layer. The radiative flux toward the body is partially absorbed in the ablation layer, increasing the temperature of the gases in this region. The radiant energy that is transmitted through this layer is then absorbed by the solid char surface where part is reradiated back into the shock layer and the remaining is conducted through the char to the decomposition zone. In the ablator the energy is absorbed by primarily endothermic reactions in the char zone and the decomposition of the polymer. The remaining fraction of energy is finally conducted through the virgin material to the substructure.

Heat is also transported to the vehicle by convection and conduction, but radiation is by far the most important mode of heat



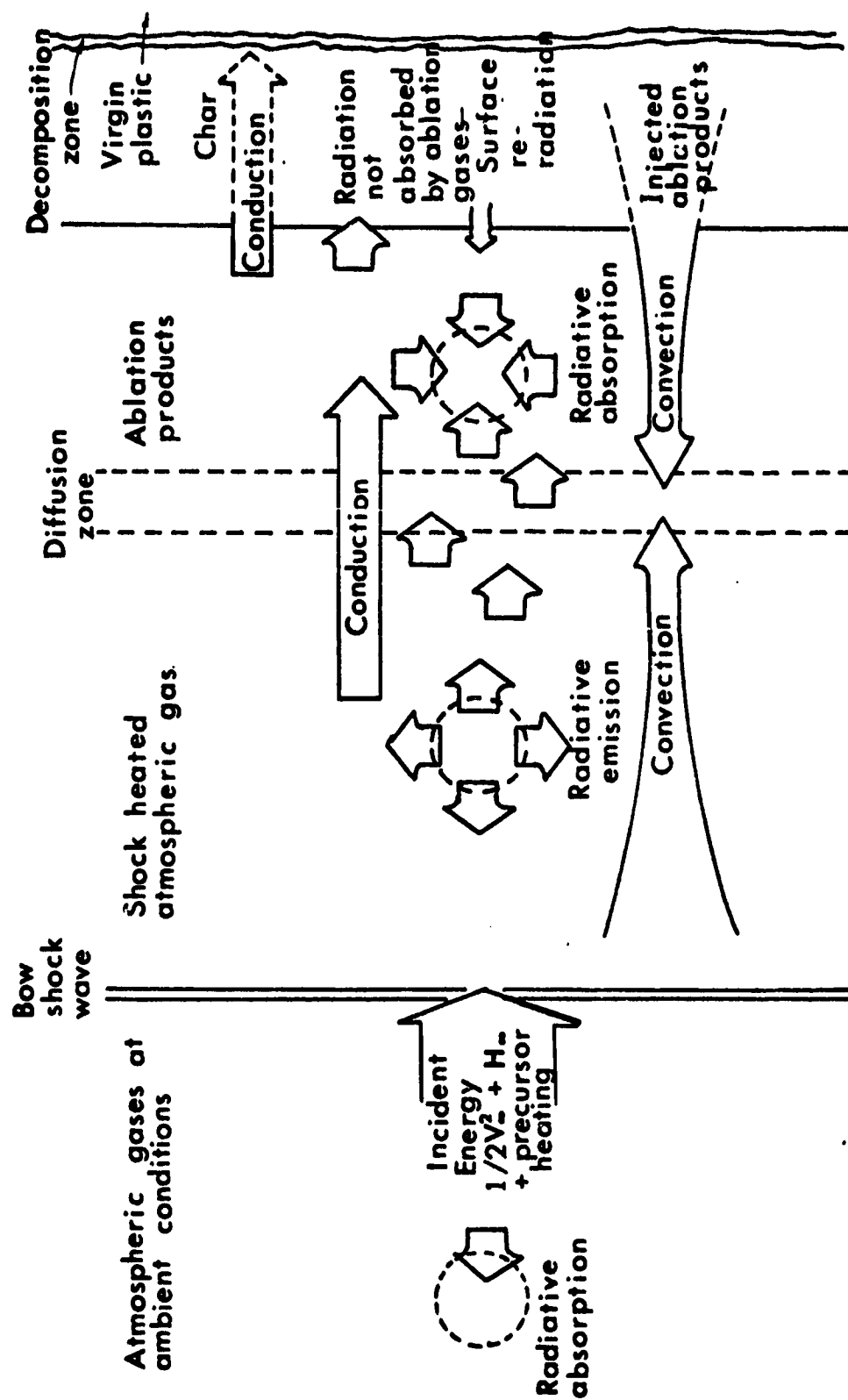


Figure 1.7. Illustration of Heat Transfer Mechanisms in Ablating Thermal Protection Systems.

transfer. The foregoing discussion has described the intimate coupling of ablator and flow field. The necessity exists, therefore, to accurately describe the major transport mechanisms in the shock layer and ablator including their coupling effects.

#### Summary

The ablator and its interaction with the surrounding flow field has been shown to be a very complicated process of mass and heat transport. The ablator and its associated flow field have been studied by many investigators who have made various assumptions about chemistry, heat transfer, and momentum transfer. The state of the art has reached a point where most of the analyses have assumed either equilibrium chemistry or no radiation heat transfer. Very few have actually coupled the response of the ablator to the flow field and those that do have used equilibrium chemistry models.

This research was undertaken to develop an analysis which incorporates the effects of finite rate chemistry, coupling of the flow field to a charring phenolic nylon ablator, and detailed line and continuum radiation heat transfer. The analysis is formulated to describe a reentering blunt vehicle typical of manned interplanetary missions.

A comprehensive overview of the work that has been done in the area of the shock heated flow field is given in the following chapter. The discussion will cover all of the recent pertinent work on non-equilibrium and radiation blunt body flows.

## REFERENCES

- 1.1 Syvertson, C. A., "Research Problems in Atmospheric Entry and Landing for Manned Planetary Missions," NASA TM D-4977 (January 1969).
- 1.2 Stewart, J. D. and D. H. Greenshields, "Entry Vehicles for Space Programs," Journal of Spacecraft and Rockets, Vol. 6, No. 10 (October 1969), pp. 1089-1102.
- 1.3 Allen, Julian H., "Hypersonic Aerodynamic Problems of the Future," in The High Temperature Aspects of Hypersonic Flow, ed. by W. C. Nelson (Pergamon Press, MacMillan Co., New York, New York), 1964, pp. 1-42.
- 1.4 Jaworski, W. and R. G. Nagler, "A Parametric Analysis of Venus Entry Heat Shield Requirements," Jet Propulsion Laboratory Tech. Report No. 32-1468, (California Inst. of Tech., Pasadena, Calif. 1970).
- 1.5 Willis, E. A., Jr. and J. A. Padruitt, "Round-Trip Trajectories with Stopovers at Both Mars and Venus," NASA TN D-5758 (April 1970).
- 1.6 Edelman, R. K. and D. Fortune, "Mixing and Combustion in the Exhaust Plumes of Rocket Engines Burning RP1 and Liquid Oxygen," TR-631 General Applied Science Laboratories, Inc., Nov. 1966.
- 1.7 Eggers, A. J. and N. B. Cohen, "Progress and Problems in Atmospheric Entry," Proceedings of the XVI th International Aero. Congress, Gauthier-Villars, Paris, 1965.
- 1.8 April, Gary C., "Evaluation of the Energy Transfer in the Char Zone During Ablation," Ph.D Dissertation, Louisiana State University, Baton Rouge, La. (1969).
- 1.9 Penner, S. S. and O. P. Sharma, "Ablation with Radiant Heating," Zeitschrift fur Flugwissen-schafter, 14, pp. 169-179, 1966.
- 1.10 Kelley, J. B. and M. R. L'Ecuyer, "Transpiration Cooling - Its Theory and Application," Report No. TM-66-5, Jet Propulsion Center, Purdue University, Contract NSG-592, N66-30856, (June, 1966).
- 1.11 Steverding, B. and V. A. Nieberlein, "Ablation for Heat Shielding," Chemical Engineering, 72, pp. 163-168, (July 19, 1965).

- 1.12 Walbert, G. D. and E. M. Sullivan, "Ablative Heat Shields for Planetary Entries - A Technology Review," Presented at ASTM/IES/AIAA Space Simulation Conference, September 14-16, 1970.

## CHAPTER II

### A REVIEW OF RADIATING AND NONEQUILIBRIUM FLOW FIELDS

#### Introduction

In the past decade, numerous investigations of flow fields surrounding blunt reentry vehicles have been performed to determine the heating rates encountered by such vehicles. Various degrees of detail have been incorporated into the modeling of the flow field processes. The momentum, energy, and species equations have undergone assumptions pertaining to the viscous, conducting, radiating, and chemical nature of the flow. Radiating solutions have developed from the initial gray emission simplifications to continuum models and further to detailed line and continuum models. Chemical models of frozen, equilibrium, and nonequilibrium chemistry have emerged with nonequilibrium being the realistic and most difficult analysis. At each stage of development, more understanding was gained and accordingly computation time increased. In this chapter the best currently available studies involving radiating and nonequilibrium flow will be discussed -- first the radiating analysis, then the nonequilibrium analysis.

#### Radiating Solutions

Radiative energy transport becomes an exceedingly important mode for heat transfer at high velocity entries and low altitudes. Many investigators have attempted to define the important mechanisms

involved in this radiative transport. A review is given below of the significant investigations performed in contributing to the current state-of-art prediction of radiative heat transfer. The analyses are split into direct methods which calculate a shock shape from a given body shape and inverse methods which use a specified shock shape and determine the body shape.

Direct Methods: One of the first efforts by Lockheed Missiles & Space Company to model the shock layer processes was by Wilson and Hoshizaki (Ref. 2.1) who attempted the relatively simple problem of inviscid and radiating flow using the optically thin approximation. The major effort by this group has been in assessing the important mechanisms of radiative transport. This first effort used an integral method to solve the conservation equations and the analysis applied away from the stagnation point. Later, Hoshizaki and Wilson (Ref. 2.2) extended their work to include viscous effects. This analysis was for a viscous, radiating but nonabsorbing gas in the shock layer. Hoshizaki and Wilson studied the effect of radiation cooling for a  $30^\circ$  hemisphere-cone and showed that the loss of energy in the shock layer by radiation reduced both the radiative and convective heat transfer. Solutions were presented for heating around the body which showed that the reduction of heating rate persists far from the stagnation line. The momentum and energy equations were solved using an integral method in which the tangential velocity and enthalpy profiles were represented by fifth and sixth-order polynomials respectively. Solutions were obtained for air only which eliminated the need to solve the species equations. The numerical solution proceeded by first obtaining a solution at

the stagnation point. The coefficients of the polynomials of tangential velocity and enthalpy were determined by using the boundary conditions at the wall and shock and the assumption that the shock was concentric. The solution proceeded downstream with an initial estimate for shock shape. This was compared to the output shock shape and a new estimate was then made. Iterations were performed until convergence was attained for the shock shape.

The third effort by Hoshizaki and Wilson (Ref. 2.3) was a major step forward in complexity by eliminating the assumption of an optically thin shock layer and accounted for the spectral behavior of the continuum absorption and emission radiative processes. The numerical method differed from the previous investigation by employing a combined finite difference-integral method. The solutions to the momentum and species continuity equations were obtained in an approximate manner by means of an integral method. The velocity and concentration profiles resulting from the approximate solution were then used to solve the energy equation by means of a modified finite difference method (Ref. 2.4). This analysis included the effects of injected ablation products which were assumed to have thermodynamic and transport properties identical to those of air. The injected gas was assumed to be inert within the context that no chemical reactions between ablation products and air were allowed. Solutions were obtained at the stagnation line and around the body. Iterations were made around the body as in Ref. 2.2 until the shock shape had converged. The primary purpose of this investigation was to determine the effect of self-absorption on the radiative and convective heating. Numerical results obtained showed that self-

absorption reduced the radiative heating by an order of magnitude but had little effect on convective heating.

The complex chemical equilibrium which results from injection of realistic ablation products into the shock layer was considered by Hoshizaki and Lasher (Ref. 2.5). In this analysis the shock layer chemical equilibrium composition was calculated using the FEMP computer program (from Ref. 2.4) which minimizes free energy given a specified temperature and pressure. According to Wilson (Ref. 2.6), "the use of FEMP, along with an even more detailed accounting of spectral radiative transport, raised a computational cost barrier which has since prevented the application of the direct finite difference analysis away from the stagnation point," Hoshizaki and Lasher (Ref. 2.5) state, therefore, that although around the body analyses can, in principle, be performed, only stagnation point solutions were obtained. The solution of the momentum equation was obtained using the Karman-Pohlhausen integral method and splitting the flow field in two parts - ablation layer and air layer. The ablation layer was terminated when ablation injected species comprised approximately 1% of the gas mixture. The velocity at the air-ablation interface was matched to combine the two solutions.

The analysis of Hoshizaki and Lasher included radiative emission and absorption, radiation cooling, and coupling between convection and radiation. Only continuum radiation was considered - lines for atomic and ionic species were not considered. The infinite slab, one dimensional approximation to the radiative flux was used which is tantamount to saying that gradients of temperature and density in the stream-wise direction are negligible. Wilson (Ref. 2.7) later studied



the effects of the infinite slab prediction and found it to be quite reasonable. Numerical results showed that absorption by ablation products reduced the radiative surface flux by almost a factor of two. This result was in contrast to an earlier paper by Howe and Sheaffer (Ref. 2.8) who found that mass additions enhanced, rather than reduced, the radiative flux to the surface. Hoshizaki and Lasher claimed that this result was due to Howe and Sheaffer's gray gas approximation. It was also shown that atomic carbon was the principal absorber and that the ablation product molecules play a minor role in determining the radiative flux.

Chin (Ref. 2.9) investigated radiation transport in the stagnation region of an ablating body with an inviscid, nonconducting flow model. The gas in the shock layer was assumed to be in thermodynamic equilibrium and at quasi-steady state. The flow field was assumed to have an interface between the ablation products and the shock-layer air. The stagnation region air layer and ablation layer were analyzed by means of similarity transformations with a continuity of pressure across the interface. The wall was assumed to be at the equilibrium sublimation temperature corresponding to the surface pressure. The ablator is coupled to the flow field by performing the following surface energy balance.

$$v_w = (q_w - \sigma T_w^4) / \rho_w \Delta H_v \quad [2.1]$$

where:  $v_w$  = velocity of injected gases into shock layer  
 $q_w$  = heat flux toward surface from shock layer  
 $\sigma T_w^4$  = energy reradiated by surface

$\rho_w$  = density of injected gases

$\Delta H_v$  = heat of ablation

A heating rate,  $q_w$ , was assumed and then updated until the flow field solution corresponded to the surface energy balance.

The radiation transport in the analysis attempted to introduce the effect of lines on the radiative surface heat flux. Only nitrogen lines were included and the ablation layer was modeled with a continuum approximation. Chin also made perturbation studies of the radiative properties to examine radiative uncertainties. Chin's results indicated that the ablation layer was very effective in reducing wall heat flux (41%) and that self-absorption and energy-loss effects reduced the sensitivity of the wall radiative flux to environmental variables and to uncertainties in radiative properties.

In order to remove major deficiencies in the treatment of the momentum equation due to the integral technique used in Refs. 2.5 and 2.10, Wilson (Ref. 2.11) solved the momentum equation in finite difference form. This method permitted solutions to "massive" blowing problems and placed all three conservation equations in the same solution form. The opacity data in the radiative transport model was extended to include atomic line transitions of H and C species. Thus, transport through the ablation product gases was more complete than in previous published studies. Finally, the inclusion of continuum and line opacity data for the He atoms permitted the solution to be applied to both Earth and Jovian entry (RATRAP from Ref. 2.12).

In comparing his results with Chin (Ref. 2.9), Wilson found that Chin's solution overpredicted the effectiveness of radiative

blocking. This result was due to the omission of C and H lines by Chin (only N lines were included in Ref. 2.9). Wilson showed that Chin's inviscid solution compared closer to a continuum only calculation as shown in Figure 2.1. The result of this research demonstrated that, unlike the situation with convective heating, massive blowing was ineffective in achieving a major reduction in the surface radiative flux. These conclusions held not only for earth entry but particularly for Jovian entry when predominately carbon species comprised the ablation product gases.

The end result of the work of Hoshizaki, Wilson and Lasher is the "VISC Code" reported in Ref. 2.6. In this investigation, the simplifying but nonessential gas dynamic approximations that  $\rho\mu = \text{constant}$  and the construction of an inner inviscid layer were removed. The results used a completely variable  $\rho\mu$  as well as a technique for integrating the conservation equations toward both the shock and wall from the interior shock layer point where the stream function passes through zero. The radiative transport subroutine was further extended to include wall emission/reflection terms. The previously employed analytic method (Ref. 2.13) for obtaining the flux divergence due to lines was dropped and replaced by a direct numerical differencing of the line flux. The code was used to obtain heating predictions for reentry into the earth's atmosphere and for probe entry into the atmospheres of Jupiter and Venus.

Another group to study radiation transport in the shock layer (stagnation line only) was that of Rigdon, Dirling, and Thomas of McDonnell Douglas Corporation. Rigdon et al. (Ref. 2.14 and 2.15) solved the viscous stagnation line equations using an initial value

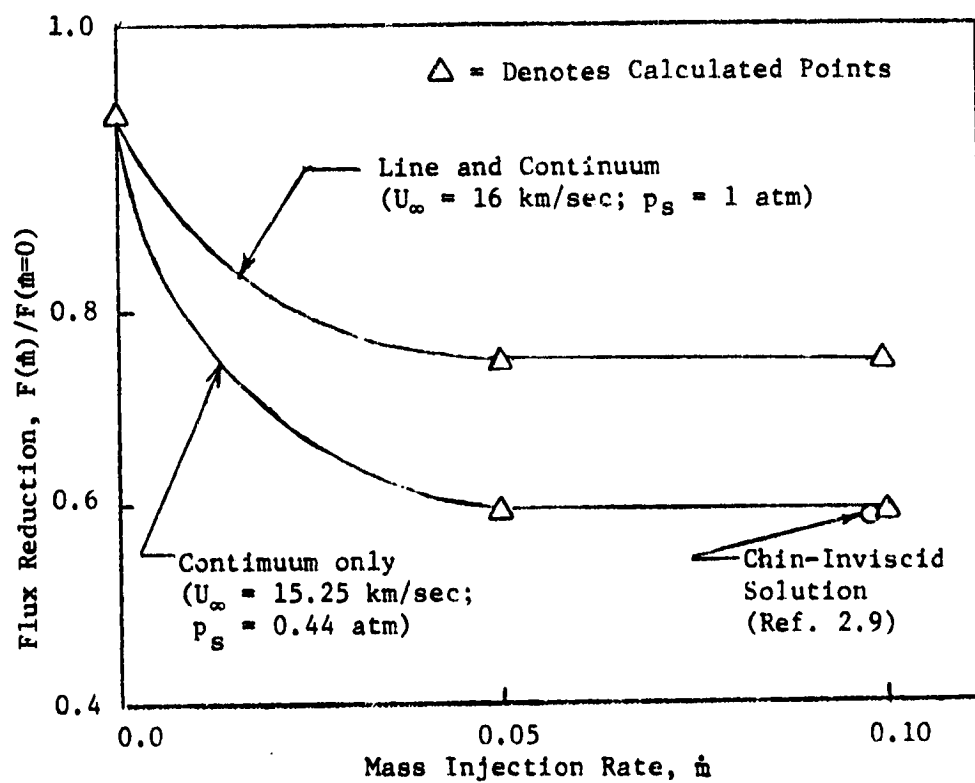


Figure 2.1. The Effect of Mass Injection on Surface Flux Reduction (Ref. 2.11)

(marching) finite difference technique. The solution was started at the stagnation point and proceeded in both directions (toward the wall and toward the shock), and the initial slope (at the stagnation point) was iterated on until the boundary conditions at both the wall and shock were satisfied. This technique was used in order to avoid the numerical problems associated with marching techniques that start the solution at a wall with mass injection where the slope is close to zero. Difficulties arose from trying to maintain numerical precision when taking differences between numbers of the same size.

Rigdon, et al. (Ref. 2.15) used a binary diffusion coefficient based on the C-N interaction and demonstrated that the choice of binary diffusion coefficient was not critical in the prediction of wall heating rates. The species equations were solved by assuming that ablation species were diffusing into air species and the mixture was in chemical equilibrium. In Ref. 2.15 the effect of injection of ablation species were considered. This was an extension of previous work of Rigdon et al. (Ref. 2.14) where only air blowing was studied. The effects of the thermodynamic and transport properties of ablation products on the heat flux at the wall was investigated, and radiation preheating of ambient air (precursor radiation) was considered. For these solutions the radiative transfer was computed using the detailed spectral radiative properties of both air and ablation products including atomic line transitions.

Comparison of the detailed solutions of Rigdon, et al. (Ref. 2.15) with other less exact methods, e.g., line grouping, showed a necessity for retaining full spectral detail in the description of the radiative properties. Furthermore, the thermodynamic properties

of the ablation products must be properly described. Results showed that precursor radiation increases the wall heat flux by less than 10% for entry velocities up to 55,000 feet per second and shock layer pressures of approximately one atmosphere.

NASA's Langley Research Center has done some extensive investigation of the radiating flow around blunt reentry bodies (Refs. 2.16 - 2.20). All of these investigations are restricted to the chemical equilibrium flow.

The inviscid, radiating, nonadiabatic flow field equations were solved by Suttles (Ref. 2.19) using an approximate computational scheme in which a method of integral relations was applied. The method was used for a one-strip approximation with an equilibrium-air gas model. A frequency dependent, self-absorbing radiation model was employed which accounted for atomic-line radiation as well as the important continuum radiation processes. This work was an extension of the work of Garrett, et al. (Ref. 2.21) who used the integral relations approach to a nonradiating flow field. Mass injection at the surface was not included. Comparison were made to the stagnation solutions of Olstad (Ref. 2.22) and Callis (Ref. 2.23). Suttles showed that the nonadiabatic solution resulted in significantly reduced enthalpies and temperatures, increased densities, and decreased shock standoff distance. Suttles concluded that although the approximate method did not produce all details of the flow, it did yield a good description of the flow field through quantities such as shock standoff distance and flow properties at the body surface.

A combined chemical equilibrium chemistry flow field and ablation study was performed by Smith et al. (Ref. 2.16) for a blunt body reentering earth's atmosphere. Three interacting regions were distinguished: inviscid air layer, ablation boundary layer, and charring carbon phenolic heat shield. The outer air layer was analyzed as an inviscid region using a one strip integral method. The inner ablation layer was analyzed using two techniques. For small ablation rates a boundary layer solution was used; whereas for large ablation rates (where the boundary layer solution was unstable) an integral method was used. No radiation coupling between air and ablation layers was considered. The radiation model used was that of Wilson and Hoshizaki (Ref. 2.10) which included various line and continuum mechanisms for atoms, ions and molecules excluding line mechanisms for C and H atoms. Heating rates for the stagnation region and around the body along a trajectory were presented. Coupling of the ablator to the flow field was accomplished in a similar manner to Chin (Ref. 2.9) in which the mass loss rate was determined by a surface energy balance. An equilibrium enthalpy of ablation was used along with surface reradiation of energy (the surface was assumed to be at the equilibrium sublimation temperature). Results were compared to Chin and indicated a smaller blockage of radiation (22% compared to 46%) due to the ablation layer. A similar result was found when compared to Rigdon et al. (Ref. 2.15) who predicted 55% blockage by the ablation layer. The conclusion was that the differences occurred due to the differences in radiative transport models of the three investigations.

In a recent study by Sutton (Ref. 2.18), the fully coupled radiating gas flow about an ablating planetary entry body was applied to Venusian entries. The method of solution coupled an inviscid flow solution and a boundary layer solution (laminar or turbulent) in which the divergence of the radiative flux was included in the energy equation for the solution of each gas layer. The treatment of radiation included molecular band, continuum, and line transitions with a detailed frequency dependence of the absorption coefficient. The species equations were solved in elemental form with chemical equilibrium being assumed. Results showed that the radiative flux toward the body was attenuated in the boundary layer at downstream regions of the body as well as at the stagnation point. The results were the same even when radiation absorption by ablation products was taken into account. The radiative heating rates along the downstream regions were shown to exceed stagnation point values under certain conditions. Blockage of radiation was found to be 10% to 20% for nominal entry conditions but radiation blockage as large as 30% resulted at higher velocities.

In an investigation by Garrett, et al. (Ref. 2.21), the viscous shock layer equations in the stagnation region were solved using an implicit finite difference scheme. The flow field was assumed to be in thermodynamic and chemical equilibrium. The radiation model used was the computer code RATRAP (Ref. 2.24) which accounted for both continuum and line radiation exchange. The species equation was solved by assuming ablation products diffusing into air species as a binary system. The assumption of the binary diffusion model was evaluated by comparing heating rate results for two cases. One



case assumed the binary diffusion coefficient to be that of hydrogen-nitrogen diffusion and the other to be that of carbon-nitrogen diffusion. The results showed that a maximum difference of 3 percent occurred in the radiative heat flux which indicated that the assumption of a binary diffusion model was valid for the conditions examined.

For massive blowing cases, damping of the enthalpy profile was required to keep the solution stable. A damping factor of 0.5 was found to be the most efficient when considering the trade-off of computer time and stability. Garrett et al. improved on the stability of their solution by using a windward differencing technique instead of central differencing. Oscillations in the enthalpy and species solutions were removed by use of windward differencing.

Comparison by Garret (Ref. 2.21) with other investigations (Ref. 2.9, 2.15, 2.16, 2.19, 2.25) demonstrated lower wall radiative heating rates for carbon phenolic ablation. The results indicated that the ablation products were highly effective in blocking the incident radiation from the high temperature outer layer of the shock. For blowing rates of 0.1 and 0.2, typical reductions in radiative heat flux at the wall ranged from 34 to 39 percent of the values for no blowing.

Engel et al. (Ref. 2.26) concluded from the results of Refs. 2.27 and 2.28 that the energy and species equations could be simplified by setting the diffusion term equal to zero. This assumption yielded a step function for the elemental mass fractions where ablation elemental mass fractions were used from the wall to the stagnation point and air elemental compositions were used in the region from stagnation point to the shock. The species compositions

were then calculated assuming thermo-chemical equilibrium.

The thin shock layer, stagnation line equations were solved including line and continuum radiation from air and ablation species. The flow field was coupled to a phenolic nylon ablator by assuming the ablator surface to be at its sublimation temperature and the heat absorbed to be that by quasi-steady ablation - similar to the coupling by Chin (Ref. 2.19) and Smith et al. (Ref. 2.16). An implicit finite difference scheme was used to solve the energy and momentum quasi-linearized equations.

Engel's results indicated that the continuum contribution to the surface radiative heat flux is essentially unchanged by increasing the ablation rate above about 5% of the freestream blowing rate. Molecular absorption of radiant energy in the ablation layer was found to contribute significantly to the reduction of wall heating rates (52% for the case studied).

Contemporary with the study performed by Engel was an investigation by Esch (Ref. 2.28) to determine, primarily, the effect of assuming binary diffusion. Esch's investigation differed from Engel in that he solved the species equations with the diffusion term included. The radiation model was identical to Engel's. Comparison of binary diffusion results with multicomponent diffusion results led to the conclusion that binary diffusion was sufficient to describe the shock layer process if an appropriate value of the diffusion coefficient was selected. He used the binary coefficient based on  $H-C_2H_2$  interaction. The effect of using air thermodynamic and transport properties, instead of including ablation products, was also investigated by Esch. Detailed transport properties were found

to be unimportant but inclusion of thermodynamic properties of ablation products proved to alter the radiative wall heating rate significantly. These results are supported by the findings of Rigdon et al. (Ref. 2.15) and Garrett et al. (Ref. 2.21).

Inverse Methods: In solutions which use inverse methods, the shock shape is given and the body shape is to be determined. The shock shape can then be adjusted until the desired body shape is found. The disadvantage of this technique is that, in general, one would like to specify body shape directly (the direct method). Unfortunately, shock shape is somewhat insensitive to body shape. Therefore, different body configurations could have the same shock shape and the inverse method does not always work effectively.

Some success has been achieved using inverse methods. Chou (Ref. 2.29) performed an inverse solution using locally nonsimilar solutions for a radiating shock layer about smooth axisymmetric bodies. In this method, nonsimilar terms in the governing equations were defined as dependent variables; model equations were derived for these variables. Solutions were presented for nonblowing bodies and compared with the Blasius type series solution of Chou and Blake (Ref. 2.30). Excellent agreement was found. Blowing solutions were also obtained (constant blowing around the body), but these cases did not consider radiative transport and no diffusion mechanism was taken into account. Therefore, the injected gas was assumed to be the same as that of the freestream. Results indicated that blowing solutions were more difficult to obtain than nonblowing solutions since the blowing cases were not presented for points very far

downstream of the stagnation region. Another limitation was that only continuum radiation was considered.

In a later report by Chou (Ref. 2.31), the restriction of no blowing for radiative heating cases was removed. In addition the simplifications of constant transport properties and continuum radiation only were also removed. The same locally nonsimilar method was used. Ablation products were allowed to diffuse into the shock layer and the governing equations were modified to allow mass injection rates (carbon injection) to vary with distance around the body. The shock layer was assumed to be in thermo-chemical equilibrium and the FEMP, Free Energy Minimization, computer program (Ref. 2.4) was used to determine the equilibrium composition. Molecular band and atomic line, as well as continuum radiation, was considered. Comparisons were made with the adiabatic inviscid flows of Olstad (Ref. 2.32) and Schneider (Ref. 2.33) and the viscous, nonradiating flow of Davis and Flugge-Lotz (Ref. 2.34).

Chou's results showed good agreement with the work reported in Refs. 2.32, 2.33, and 2.34. A sample calculation was made for a typical Jovian entry which indicated a 40% reduction in total radiative flux at the stagnation point due to carbon gas injection when compared to the nonblowing case. The reduction increased to about 70% of the stagnation line value at about four body nose radii downstream. The radiative flux reduction indicates that most of the radiative energy is absorbed by the injected carbon gas and dumped into the wake. An examination of the spectral distribution of the radiative flux revealed that most of the flux reduction occurred in molecular bands of  $C_2$  and  $C_3$ .

A companion report to Ref. 2.31 was Ref. 2.35 which is a user's manual for the SL4 computer code. The report describes the input and output for the inverse shock layer solution of Chou. The code also provides a basis for obtaining predictions of the surface heating to a body entering any planetary atmosphere at hyperbolic velocities.

In an inverse solution by Olstad (Ref. 2.32) an attempt was made to study the effects of radiation on the flow about smooth symmetric bodies downstream of the stagnation point. Previously Olstad had only presented stagnation region solutions (Refs. 2.22 and 2.36). In the analysis, a simplified flowfield technique developed by Maslen (Ref. 2.37) was modified to account for radiation and large blowing. The radiative transport model included effects of atomic lines and continuum radiation. Blowing was studied for cases of cold air injection, thus only air thermodynamic, transport, and radiative properties were required.

At some point not too large a distance downstream of the stagnation point, the flow direction in the shock layer is nearly parallel to the shock. Maslen (Ref. 2.37) took advantage of this situation to derive a simplified system of equations which adequately describe the flow and thermodynamic properties in the shock layer for an inviscid, nonreacting, nonradiating gas for a given shock shape. The most important step in the derivation was the transformation to the von Mises plane and the subsequent uncoupling of the normal momentum equation from the other equations. Using a modified form of Maslen's method, Olstad studied the effect of cold air blowing on radiative heating. He found that in the stagnation

region blowing was quite effective in reducing the radiative heating to the body surface. However, downstream of the stagnation region, where the radiative heat flux from the hot shock layer decreased, re-emission from the blown layer became increasingly important compared with absorption. Thus the effectiveness of the blown layer decreased and finally became negative, so that the presence of the blown layer actually increased the radiative heating to the body surface downstream of the body station where blowing ceased. Olstad's results must be evaluated in the light of the system under study, i.e., only air blowing was considered and ablation products with different radiative, transport, and thermodynamic properties could very well invalidate the conclusions reached (e.g. Chou, Ref. 2.31).

Falanga and Sullivan (Ref. 2.25) presented an inverse solution for an inviscid shock layer flow about a blunt body. The model included nongray radiative transfer and only considered equilibrium air flow. The technique had the restriction that the shock shape had to be analytic and the solution was only applicable to the subsonic flow region in the shock layer. No blowing from the body was considered. Results compared favorably with the direct method of Rigdon et al. (Ref. 2.15) and the analysis of Callis (Ref. 2.23) who used a time-asymptotic technique.

More recently Sutton and Falanga (Ref. 2.38) published results for the entry of a blunt vehicle into a Venusian atmosphere of 90% carbon dioxide and 10% nitrogen. In this analysis only stagnation region heating results were obtained, but ablation product injection and a boundary layer adjacent to the body were included. Again nongray radiation was accounted for and fully coupled solutions were

presented (i.e. radiative coupling and ablator-flow field coupling). A high density phenolic nylon ablator was considered in the investigation. Results indicated that ablation products reduced radiative heating by as much as 35% for high velocity entries (11 km/sec) and that the boundary layer also significantly reduced the radiative heating (20-30%). For entry velocities below about 10 km/sec, it appeared that the inviscid layer solution gave reasonable estimates of radiative heating.

#### Nonequilibrium Chemistry Solutions

The preceding discussion has been centered on analyses that have sought primarily to describe the important radiative heating phenomena that occur in high temperature shock layers. None of these analyses have considered the effect of nonequilibrium chemistry on the flow field solution. At best equilibrium chemistry was included in some of the previously discussed investigations. This section will attempt to describe some of the important work that has been done in the area of nonequilibrium or finite rate chemistry analyses. In general these studies do not include the effects of radiative heat transport.

Blottner (Refs. 2.39 - 2.43) has done extensive work in the investigation of nonequilibrium flows, especially for boundary layer flows. In an early work by Blottner (Ref. 2.41), the investigation of a viscous nonequilibrium-ionized air boundary layer was performed. Seven air species ( $O_2$ ,  $N_2$ ,  $O$ ,  $N$ ,  $NO$ ,  $NO^+$ ,  $e^-$ ) and 11 chemical reactions were considered. The equations were transformed with the Mangler and Howarth-Dorodnitsyn transformation in order to obtain them in a form more appropriate for numerical solution. The

resulting partial differential equations were then solved using an implicit finite difference scheme. Results were given for slender vehicles entering at 22,000 feet per second at altitudes of 100,000 feet and 150,000 feet. For distances 15 feet downstream from the tip and for conditions considered, none of the chemical reactions were near equilibrium. No mass injection at the wall was included in this analysis.

Blottner published an excellent survey paper on finite difference methods for the solution of the boundary layer equations (Ref. 2.39) in 1970. The paper presented a technique for solving the boundary layer equations for a multicomponent flow with finite chemical reactions. The technique was developed in an effort to attain a method that would solve the governing equations when many chemical species exist in the flow. The solution technique involved the quasilinearization of the equations, uncoupling them, placing them in finite difference form, and solving the resulting tri-diagonal matrices. Blottner found this technique superior in computational efficiency as can be seen from Figure 2.2 to the fully coupled solution. When the equations are fully coupled, the program storage requirements and the computation time increase rapidly as more species are included. Since a matrix inversion must be performed in the fully coupled solution and computation time is proportional to the cube of the matrix size (number of species), large chemical systems become prohibitive. On the other hand, the uncoupled method results in solution times which are approximately proportional to the number of species.



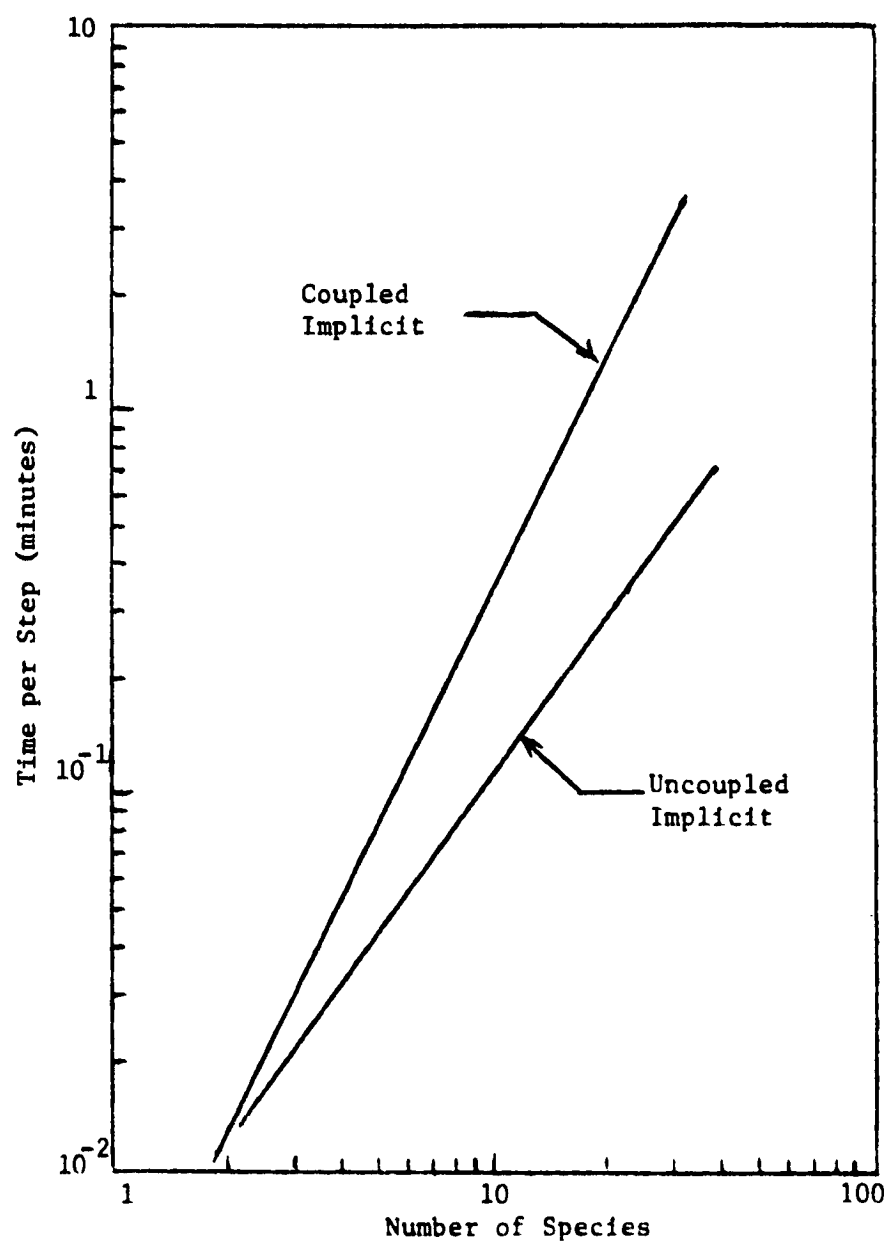


Figure 2.2. Computation Time for Various Numbers of Chemical Species (Binary Diffusion) (Ref. 2.39).

The finite differenced equations were written with a parameter "O" which can vary between 0.0 and 1.0, where O is 0 for explicit methods, 1/2 for Crank-Nicolson method and 1 for the Standard Implicit method. Blottner used  $O = 1/2$  or the Crank-Nicolson difference scheme because the accuracy is of second order compared to first order accuracy for the standard implicit. The explicit scheme was not used since unstable solutions invariably occur when this method is used on boundary layer problems. The boundary conditions for the boundary layer equations were presented by Blottner and placed into the finite difference formulation also. In this analysis, Blottner considered the effect of an ablator injecting mass into the boundary layer. The initial conditions (stagnation line equations) for solution were developed in Ref. 2.42 by Blottner. These ordinary differential equations were treated in the same manner as the partial differential equations which describe the boundary layer.

Results were obtained by Blottner for binary and multicomponent diffusion which showed differences which were almost indistinguishable. The surface boundary conditions employed in the study were such that the wall was either noncatalytic, or fully catalytic. For the case of the non-catalytic wall the mass injection rate and diffusion rate of air species to the surface was zero. A fully catalytic recombination surface for air is defined as a wall where every dissociated and ionized species that strikes the surface is converted to a molecular species due to heterogeneous reactions. Comparison of catalytic to noncatalytic solutions showed negligible differences in temperature and velocity profiles but about a 20%

difference was found in the boundary layer thickness (nong catalytic being higher).

For the solution with a nong catalytic wall, oscillations were obtained with Crank-Nicolson ( $\theta$  equal to 0.5) method. To avoid this problem, the standard implicit method with  $\theta$  equal to 1.0 was employed. The results were compared with  $\theta$  equal to 0.5 and 1.0, and the results were in close agreement downstream. Near the tip (first few inches) there was a difference between the predictions of the two methods for the mass fraction of the species, with the implicit method giving the more accurate results. This type of stability problem has also been observed by Moore (Ref. 2.44) and was corrected by using the standard implicit scheme for the first grid point away from the wall.

In a summary report (Ref. 2.43) Blottner et al. prepared a user's manual for the boundary layer - shock layer computer program that evolved from their work. In the report additional improvements were made in the numerical schemes. Initial profiles can be obtained by the program and variable step sizes across the layer are possible. Results are presented for various boundary layer and shock layer solutions and three example runs are given for input and output of the program which demonstrate its flexibility. One of the major improvements in this version of the program was the capability to handle arbitrary body shapes.

Davis (Refs. 2.45 and 2.46) presented solutions for the viscous flow in the shock layer about reentering vehicles at moderate to high Reynolds number. Shock and wall slip conditions were included in the boundary conditions and were shown to be important for low

Reynolds number cases. The flow field was described by one consistent set of equations which are the thin shock layer equations as Blottner used (Ref. 2.43). The gas was assumed to be a perfect gas so that no chemical reactions occurred and the species continuity equations did not need to be solved.

Davis used a numerical scheme similar to Blottner (Ref. 2.39) and developed the variable step size finite differencing used in Ref. 2.43. Starting with the assumption that the thin shock layer equations applied and that the shock was concentric to the body, the solution proceeded from the stagnation point to points downstream. The thin shock layer assumption is tantamount to omitting normal velocity terms in the normal momentum equations. These assumptions were then removed by making successive iterations around the body using a corrected shock shape and normal components of velocity in the normal momentum equation.

The addition of chemical reactions in the shock layer was later investigated by Davis (Ref. 2.47). In this analysis a chemically reacting binary mixture of oxygen atoms and molecules was considered. Flows past hyperboloids with one inch nose radius asymptotic far downstream to cones of  $60^\circ$ ,  $45^\circ$ , and  $20^\circ$  were investigated at altitudes of 100 to 250 thousand feet and an entry velocity of 20,000 feet per second. The calculations extended 25 nose radii downstream and considered both catalytic and noncatalytic wall boundary conditions. No mass injection at the wall boundary was considered.

Before the calculations were made downstream, Davis performed several stagnation-region solutions to determine what effect rates of chemical reaction given by various authors (Refs. 2.48-2.52) would

have on the solution. Typical results showed that there was about a factor of four between the lowest and highest atom concentration profiles. For the remainder of the work, Davis used the rate data of Bortner (Ref. 2.49).

The wall catalyticity had little effect of skin friction and wall pressure distribution, but had a sizable effect on heat transfer. The wall catalyticity was also shown to have little effect on velocity and temperature profiles in all cases. Davis found by comparing equilibrium concentrations to nonequilibrium calculations that at 100,000 feet the outer inviscid flow was essentially in equilibrium at the stagnation-point, but was not a few nose radii downstream. Caution should be taken, therefore, in using stagnation-point solutions to determine when equilibrium conditions are reached.

In a study that followed Davis, Moss (Ref. 2.53) extended Davis's model (Ref. 2.47) to include multicomponent effects. This investigation studied the effects of chemistry models (frozen, equilibrium, or nonequilibrium) on the flow field solution, the effect of multicomponent diffusion, wall catalyticity, and mass injection. The numerical solution technique was identical to that of Davis (Ref. 2.47) and the body studied was a  $45^\circ$  hyperboloid.

Moss's results indicated that the chemistry models substantially influenced flow parameters and surface transport. Heat transfer rates were higher for both frozen and equilibrium than for nonequilibrium (e.g. 54% higher for equilibrium in the stagnation region). Wall catalyticity also had an effect of increasing heat transfer rates for nonequilibrium cases. The effect of multicomponent diffusion was shown to be negligibly different from the binary diffusion

assumption which agrees with the results of Blottner (Ref. 2.39) and Esch (Ref. 2.28).

Moss considered mass injection of air, water, and ablation products into the shock layer. Only for the case of air injection was nonequilibrium chemistry attempted. Equilibrium solutions were obtained for water and ablation injection. The effect of that mass injection had on heat transfer rates decreased as the injection rate increased. Also, the effect that mass injected in the stagnation region had on the flow field decreased very rapidly downstream. Water was found to be the most effective in reducing heat transfer rates for the injectants considered.

Adams et al. (Ref. 2.54) presented results for chemical nonequilibrium inviscid and laminar viscous flow over spherically blunted cone geometrics. The calculations were made for flight velocities and altitudes where radiative transport was negligible. The chemistry was restricted to air species and injected species of argon, helium, or carbon dioxide. Multicomponent diffusion was discussed; however, no results were presented. For the cases studied, the influence of a noncatalytic wall with mass injection significantly reduced the convective heat transfer.

Using an approximate inverse solution, Grose (Ref. 2.55) presented results for the inviscid nonequilibrium flow in the shock layer about a vehicle in hypersonic flight. Cases were run for Earth, Martian, and Venusian atmospheric entry. The method used a von Mises transformation to place the equations in the form of the stream function similar to Olstad (Ref. 2.32). A large number of reactions and/or species was easily handled. The gas model permitted

consideration of vibrational relaxation, dissociation, recombination, ionization, electronic excitation, and vibration-dissociation coupling. The method had the advantage of being able to solve the subsonic and supersonic portions of the flow field whereas many flow field solutions require the use of two methods for the two regimes. The primary advantage of the method was that computing time was small (79 seconds to about 14 minutes). One disadvantage was that approximations were crude close to the stagnation point where peak heating generally occurs.

#### Summary

The studies reviewed represent the best currently available. Unfortunately the radiating solutions are for equilibrium chemistry flow fields at best. On the other hand the nonequilibrium solutions are presented for cases in which radiation is not a dominant heat transfer mechanism. None of the analyses reviewed incorporate both the effects of radiative transport and nonequilibrium chemistry. With the exception of Chin (Ref. 2.9), Smith et al. (Ref. 2.16), Esch (Ref. 2.28), and Engel (Ref. 2.27), the solutions neglect the coupling of the flow field to the body whereby a surface energy balance is maintained.

As can be seen, no single analysis includes all of the important effects of radiative heating, nonequilibrium chemistry, and ablator coupling. It was the main purpose of this research, therefore, to incorporate all these effects using a computational scheme which is reasonable, accurate and rapid. The model developed includes radiative transport with various line and continuum contributions, nonequilibrium chemistry, and is coupled to a phenolic nylon ablator.

In Chapter 3, the equations that describe this system are developed and placed into a form suitable for numerical solution.



## REFERENCES

- 2.1 Wilson, K. H. and H. Hoshizaki, "Inviscid, Nonadiabatic Flow About Blunt Bodies," AIAA Journal, 3, pp. 67-74, 1965.
- 2.2 Hoshizaki, H. and K. H. Wilson, "Viscous, Radiating Shock Layer About a Blunt Body," AIAA J., Vol. 3, No. 9, Sept. 1965, p. 1614-1622.
- 2.3 Hoshizaki, H. and K. H. Wilson, "Convective and Radiation Heat Transfer During Superorbital Entry," AIAA J., 5 (1), 25 (Jan. 1967).
- 2.4 Smith, A.M.O. and N.A. Jaffe, "General Method for Solving the Laminar Nonequilibrium Boundary-Layer Equations of a Dissociating Gas," AIAA Journal, 4, 611-620 (1966).
- 2.5 Hoshizaki, H. and L. E. Lasher, "Convective and Radiative Heat Transfer to an Ablating Body," AIAA J., 6 (8), pp. 1441-1449, (Aug., 1968).
- 2.6 Wilson, K. H., "The VISC Cose - A User's Manual," NASA CR-2237, March, 1970.
- 2.7 Wilson, K. H., "Evaluation of One-dimensional Approximations for Radiative Transport in Blunt Body Shock Layers," NASA CR-1990, March, 1972.
- 2.8 Howe, J. T. and Y. S. Sheaffer, "Mass Addition in the Stagnation Region for Velocities Up to 50,000 Feet Per Second," TR R-207, Aug. 1964, NASA.
- 2.9 Chin, J. H., "Radiation Transport for Stagnation Flows Including the Effect of Lines and Ablation Layer," AIAA Paper No. 68-664, Fluid and Plasma Dynamics Conference (Los Angeles, Cal.), June 24-26, 1968.
- 2.10 Wilson, K. H. and H. Hoshizaki, "Effect of Ablation Product Absorption and Line Transitions on Shock Layer Radiative Transport," NASA CR-1264, 1969.
- 2.11 Wilson, K. H., "Massive Blowing Effects on Viscous, Radiating, Stagnation-Point Flow," AIAA Paper No. 70-203, Eighth Aerospace Sciences Meeting (New York, N. Y.), January 19-21, 1970.
- 2.12 Wilson, K. H. and R. Greif, "Radiation Transport in Atomic Plasmas," Journal of Quantitative Spectroscopy Research & Technology, Vol. 8, 1968, pp. 1061-1086.

- 2.13 Wilson, K. H., "Stagnation Point Analysis of Coupled Viscous-Radiating Flow with Massive Blowing," NASA CR-1548, June 1970.
- 2.14 Rigdon, W. S., R.B. Dirling, Jr., and M. Thomas, "Radiative and Convective Heating During Atmospheric Entry," NASA CR-1170, September, 1968.
- 2.15 Rigdon, W. S., R. B. Dirling, and M. Thomas, "Stagnation Point Heat Transfer During Hypervelocity Atmospheric Entry," NASA CR-1462, 1970.
- 2.16 Smith, G. L., J. T. Suttles and E. M. Sullivan, "Viscous Radiating Flow Field on an Ablating Blunt Body," AIAA Paper No. 70-218, Eighth Aerospace Sciences Meeting, (New York, New York), January 19-21, 1970.
- 2.17 Garrett, B. L., G. L. Smith, and J. N. Perkins, "An Implicit Finite Difference Solution to the Viscous Shock Layer, Including the Effects of Radiation and Strong Blowing," NASA TR R-388, November 1972.
- 2.18 Sutton, K., "Fully Coupled Nongray Radiating Gas Flows with Ablation Product Effects About Planetary Entry Bodies," AIAA Paper No. 73-672, (Palm Springs, California), July 16-18, 1973.
- 2.19 Suttles, J. T., "A Method of Integral Relations Solution for Radiating, Nonadiabatic, Inviscid Flow Over a Blunt Body," NASA TN D-5480, October, 1969.
- 2.20 Garrett, L. B., "An Implicit Finite Difference Solution to the Viscous Radiating Shock Layer with Strong Blowing," Ph.D. dissertation, North Carolina State University, Raleigh, North Carolina, 1971.
- 2.21 Garrett, L. B., J. T. Suttles and J. N. Perkins, "A Modified Method of Integral Relations Approach to the Blunt-Body Equilibrium Air Flow Field, Including Comparisons with Inverse Solutions," NASA TN D-5434, September, 1969.
- 2.22 Olstad, W. B., "Blunt-Body Stagnation - Region Flow with Nongray Radiation Heat Transfer - A Singular Perturbation Solution," NASA TR R-295, 1968.
- 2.23 Callis, L. B., "Solutions of Blunt-Body Stagnation - Region Flow with Nongray Emission and Absorption of Radiation by a Time-Asymptotic Technique," NASA TR R-299, 1969.
- 2.24 Wilson, K. H. "RATRAP - A Radiation Transport Code," 6-77-67-12, Lockheed Missiles and Space Co., March 14, 1967.
- 2.25 Falanga, R. A. and E. M. Sullivan, "An Inverse-Method Solution for Radiating, Nonadiabatic, Equilibrium Inviscid Flow Over a Blunt Body," NASA TN D-5907, August 1970.

- 2.26 Engel, C. D., R. C. Farmer, and R. W. Pike, "Ablation and Radiation Coupled Viscous Hypersonic Shock Layers," AIAA Journal, 11 (8), pp. 1174-1181, August 1973.
- 2.27 Engel, C. D., "Ablation and Radiation Coupled Viscous Hypersonic Shock Layers," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 2.28 Esch, D. D., "Stagnation Region Heating of a Phenolic-Nylon Ablator During Return from Planetary Missions," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 2.29 Chou, Y. S., "Locally Nonsimilar Solutions for Radiating Shock Layer about Smooth Axisymmetric Bodies," NASA CR 1989, March 1972.
- 2.30 Chou, Y. S. and L. H. Blake, "Thin Radiating Shock Layer About a Blunt Body," NASA CR-1547, March 1970.
- 2.31 Chou, Y. S., "Radiative Coupled Viscous Flow with Massive Blowing," NASA CR-2236, March 1973.
- 2.32 Olstad, W. B., "Nongray Radiating Flow about Smooth Symmetric Bodies," AIAA Journal, 9, (1), pp. 122-130, January 1971.
- 2.33 Schneider, W., "A Uniformly Valid Solution for the Hypersonic Flow Past Blunted Bodies," Journal of Fluid Mechanics, Vol. 31, Part 2, 1968, pp. 379-415.
- 2.34 Davis, R. T. and I. Fluggi-Lotz, "Second Order Boundary-Layer Effects in Hypersonic Flow Past Axisymmetric Blunt Bodies," J. Fluid Mech., Vol. 20, 1964, pp. 593-623.
- 2.35 Chou, Y. S., "SL4 Code - A User's Manual," NASA CR-2235, March 1973.
- 2.36 Olstad, W. B., "Stagnation-Point Solutions for Inviscid Radiating Shock Layers," NASA TN D-5792, June 1970.
- 2.37 Maslen, S. H., "Inviscid Hypersonic Flow Past Smooth Symmetric Bodies," AIAA Journal, 2 (6), pp. 1055-1061, June 1964.
- 2.38 Sutton, K. and R. A. Falanga, "Stagnation Region Radiative Heating with Steady-State Ablation During Venus Entry," JSR, 10 (2), pp. 155-157, February 1973.
- 2.39 Blottner, F. G., "Finite Difference Methods of Solution of Boundary Layer Equations," AIAA J., 8, (2), Feb. 1970.
- 2.40 Blottner, F. G., "Similar and Nonsimilar Solutions of the Non-equilibrium Laminar Boundary Layer," AIAA Journal, 1 (9), pp. 2156-2157, September 1963.

- 2.41 Blottner, F. G., "Nonequilibrium Laminar Boundary-Layer Flow of Ionized Air," AIAA Journal, 2 (11), pp. 1921-1927, November 1964.
- 2.42 Blottner, F. G., "Viscous Shock Layer at the Stagnation Point with Nonequilibrium Air Chemistry," AIAA Journal 7, (12), pp. 2281-2288, December 1969.
- 2.43 Blottner, F. G., M. Johnson and M. Ellis, "Chemically Reacting Viscous Flow Program for Multi-Component Gas Mixtures," SC-RR-70-754, Sandia Laboratories, (Albuquerque, New Mexico), December 1971.
- 2.44 Moore, J. A., "Chemical Nonequilibrium of Viscous Flows," Ph.D. dissertation, State Univ. of New York at Buffalo, May 1967.
- 2.45 Davis, R. T., "The Hypersonic Fully Viscous Shock-Layer Problem," SC-RR-68-840, Sandia Laboratories, December 1968.
- 2.46 Davis, R. T., "Numerical Solution of the Hypersonic Viscous Shock-Layer Equations," AIAA Journal, 8 (5), pp. 843-851, May 1970.
- 2.47 Davis, R. T., "Hypersonic Flow of a Chemically Reacting Binary Mixture Past a Blunt Body," AIAA Paper No. 70-805, Third Fluid and Plasma Dynamics Conference (Los Angeles, Cal.), June 29-July 1, 1970.
- 2.48 Bortner, M. H. and J. A. Golden, "A Critique on Reaction Rate Constants Involved in the Chemical System of High Temperature Air," General Electric TIS R-61SD023, 1961.
- 2.49 Bortner, M. H., "A Review of Rate Constants of Reactions in Re-entry Flow Fields," General Electric TIS R-68SD13, 1968.
- 2.50 Chung, P. M. and A. D. Anderson, "Dissociative Relaxation of Oxygen Over an Adiabatic Flat Plate at Hypersonic Mach Numbers," NASA TN D-140, 1960.
- 2.51 Cresswell, J., B. Kaplan, R. Porter, and C. Sarkos, "Material Effects of Low Temperature Ablators on Hypersonic Wake Properties of Slender Bodies," General Electric TIS 67SD255, 1967.
- 2.52 Hall, G. J., A. W. Eschenroeder, and P. V. Marrone, "Blunt-Nosed Inviscid Airflows with Coupled Nonequilibrium Processes," J. Aerospace Sciences, Vol. 29, pp. 1038-1051, 1962.
- 2.53 Moss, N. N., "Solutions for Reacting and Nonreacting Viscous Shock Layers with Multicomponent Diffusion and Mass Injection," Ph.D. dissertation, Virginia Polytechnic Institute and State University, Blacksburg, Virginia (1971).
- 2.54 Adams, J. G., C. H. Lewis, H. S. Brahensky, et. al., "Nonequilibrium, Mass Transfer, and Viscous Interaction on Spherically

Blunted Cones at Hypersonic Conditions," AEDC-TR-69-237,  
Arnold Engineering Development Center, Jan. 1970.

- 2.55 Grose, W. L., "A Thin-Shock-Layer Solution for Nonequilibrium,  
Inviscid Hypersonic Flows in Earth, Martian, and Venusian  
Atmospheres," NASA TN D-6529, December 1971.

### CHAPTER III

#### DEVELOPMENT OF THE THIN SHOCK LAYER EQUATIONS

##### Introduction

In this chapter the governing equations for the radiating, chemically reacting, viscous flow around a blunt body are developed, transformed, and placed into solution form. A more detailed development of these equations, including an order of magnitude analysis is given by Engel et al. (Ref. 3.1). Engel's development is a complete derivation of the thin shock layer equations from general transport equations which includes intermediate formulations of varying degrees of rigor. The assumptions inherent at each stage of development are clearly stated. This chapter omits some of the tedious mathematical details of the derivation while retaining most of the logical step-by-step procedure used in obtaining the set of equations necessary to describe the flow field around an ablating blunt body.

The thin shock layer equations, which model the flow of the multicomponent gas in the shock heated region of the ablator, are a set of parabolic partial integro-differential equations. By nature, parabolic differential equations require one set of initial conditions and two sets of boundary conditions. In the shock layer, the initial conditions are given by the stagnation line equations, and the two boundary conditions are given by the Rankine-Hugoniot equations at the bow shock and the mass and energy balances at the

ablator surface. The stagnation line equations are simply the thin shock layer equations reduced to ordinary integro-differential equations at the axis of symmetry on the body. The Rankine-Hugoniot equations at the shock and the surface balances at the ablator surface insure conservation of mass, energy, and momentum across these boundaries. Since the initial and boundary conditions are a very integral part of the shock layer solution, this chapter treats the derivation of these equations in some detail.

In order to attain a more feasible coordinate system for solution, the blunt body equations are transformed by the Dorodnitsyn transformation. This transformation is an independent variable transformation which allows for a stretching of the normal body coordinate near the body utilizing the variation in density across the shock. The solution in this form allows for easier numerical solution.

Finally, a general form for numerical solution of the set of parabolic partial differential equations is developed and presented. In Chapter V the equations in this form are finite differenced and arranged in the form for numerical solution.

### The Physical Environment

It becomes necessary in the development of any mathematical model to make simplifying assumptions to the general conservation equations in order to obtain a set of equations that is more feasible to solve. These simplifying assumptions are determined only by understanding the physical system to be modeled and by assessing the magnitude of the terms in the general conservation equations. An assessment of this sort will provide, not only a basis for omitting from the equations terms of negligible consequence, but will also

provide a means of estimating the magnitude of terms impossible or too difficult to evaluate at this stage of technological development. A knowledge of the importance of these various terms indicates which uncertain parameters are most necessary to study.

In order to determine the proper mathematical model to describe the flow field around a blunt body moving at hypersonic velocities, the behavior of the gas the vehicle will encounter must be investigated. In Figure 3.1 based on the work of Reference 3.2 the flight regimes are presented which are encountered by a body during atmospheric entry. The regimes can be grouped into two gasdynamic domains-continuum and noncontinuum. Hayes and Probstein (Ref. 3.2) demonstrates the continuum can be divided into five regimes: (1) classical boundary layer, (2) vorticity interaction, (3) fully viscous, (4) incipient merged layer, and (5) fully merged layer. The general conservation equations can be used to describe the gas flow over a body in these continuum regimes. A brief description of these five continuum regimes follows (Ref. 3.1):

1. Boundary layer regime: The classical boundary layer equations are a valid approximation of the viscous effects for high Reynolds numbers corresponding to lower altitudes. Viscous effects dominate near the wall in a region which is small compared to the shock layer thickness. Vorticity generated by the shock curvature is therefore negligible, having no effect on the boundary layer flow.
2. Vorticity interaction becomes important at lower Reynolds numbers where shock generated vorticity becomes significant in respect to viscous effects near the body. Here the



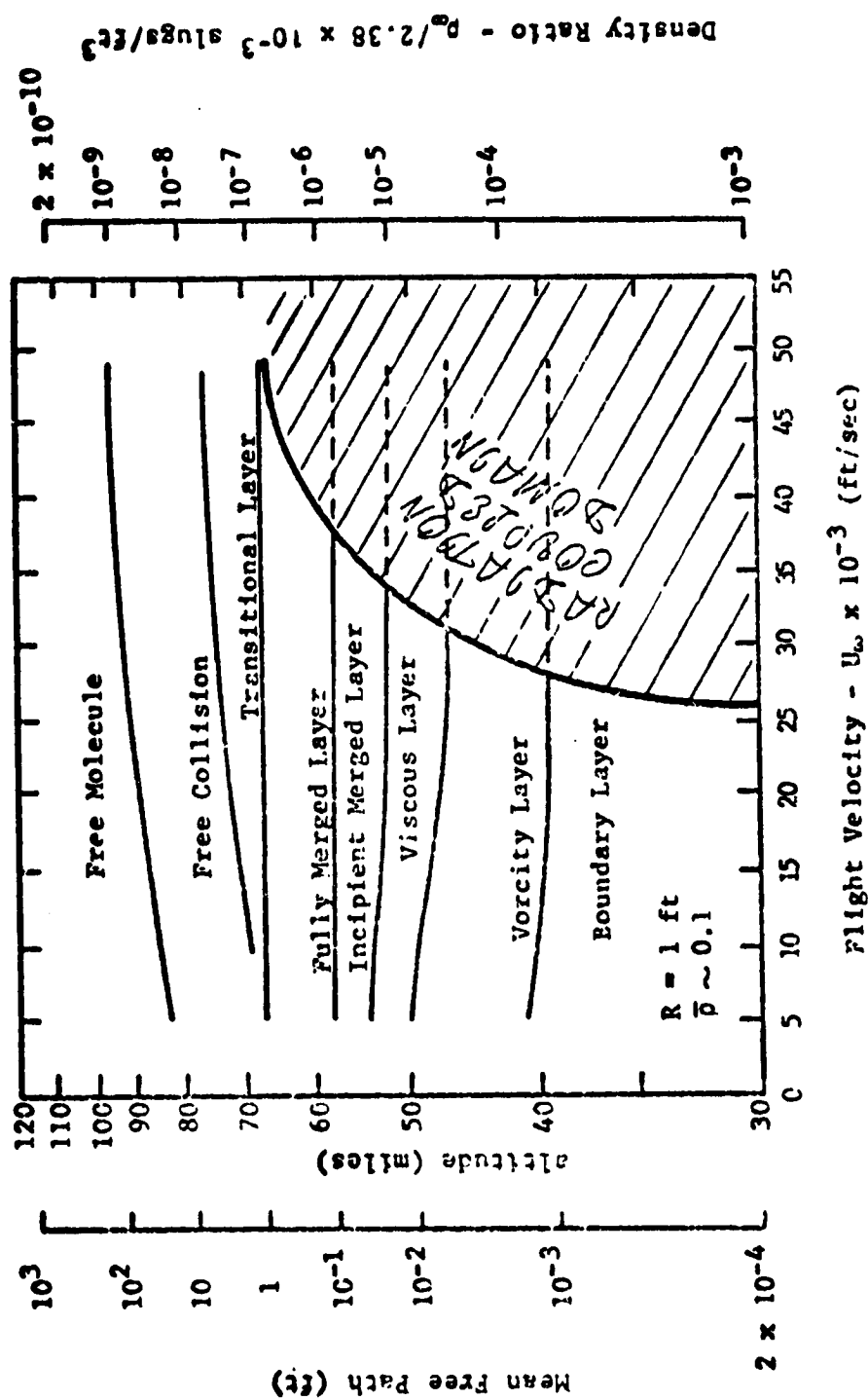


Figure 3.1. Flight Regimes (Modified from Ref. 3.4).

outer region of the shock layer, usually considered the inviscid layer, becomes coupled through momentum transport to the higher shear region near the body, usually thought of as the boundary layer. The high shear region near the body is also larger than experienced at higher Reynolds numbers.

3. Viscous layer regime: Viscous effects from the body interaction are spread through the shock layer. This occurs at lower Reynolds numbers and correspondingly higher altitudes than the vorticity interaction regime. Viscous dissipation at the shock is still small in comparison to dissipation at the body. This condition is true as long as the ratio of the mean free path behind the shock over the shock layer thickness is much smaller than the square root of the density ratio across the shock wave (Ref. 3.2). This implies that the Rankine-Hugoniot shock wave equations are valid for determination of the shock layer boundary conditions.
4. Incipient merged layer regime: The incipient merged layer begins when dissipative effects at the shock are significant. The shock wave is thin relative to the shock layer thickness but the Rankine-Hugoniot relations must be modified to account for viscous effects at the shock boundary.
5. Fully merged layer regime: At higher altitudes and low Reynolds numbers, a distinct shock does not exist. The free stream mean free path over the major body radius is approximately one or less. The flow behaves continuously from the free stream to the body. Above this altitude

range continuum concepts are no longer applicable, and the flow goes through a transition to free molecular flow.

The foregoing discussion of the five continuum flow regimes follows the reasoning of Hayes and Probstein (Ref. 3.2) under the assumption that radiative energy transport and ablative mass injection were negligible. In the present development, these two effects are of primary importance. Figure 3.1 shows the flight regimes where radiative heating to a one-foot body becomes significant. For the most part, significant ablation rates are also encountered in these regimes when using present day charring ablators such as phenolic carbon and phenolic nylon. Engel et al. (Ref. 3.1) concludes that these same mechanisms for momentum transport hold true when radiation and ablation are present, therefore the basic characteristics of the shock layer are unchanged in these regimes.

The basic conservation equations are appropriate to describe the flow of a continuum reacting and radiating gas mixture over a blunt surface when thermal equilibrium exists. For the present work, a reduced set of equations will be determined which describes the flow in the shock layer over a blunt body when the outer boundary of the shock layer is a shock wave described by the Rankine-Hugoniot equations. Thus, the equations governing the flow in the shock layer will be applicable to the three higher Reynolds number regimes. It is in these three regimes where the heating rates to a vehicle's surface are the most significant.

#### The Conservation Equations

The conservation equations can now be written in general vector-tensor notation. The statement of these equations is given by

Engel et al. (Ref. 3.1) in time independent vector form as follows:

Global Continuity:

$$\bar{\nabla} \cdot \rho \bar{V} = 0 \quad [3.1]$$

Species Continuity:

$$\bar{\nabla} \cdot (\rho_i \bar{V}) + \bar{\nabla} \cdot \bar{J}_i = \omega_i \quad [3.2]$$

Momentum:

$$\rho(\bar{V} \cdot \bar{V}) \bar{V} + \bar{\nabla} \cdot (\bar{\tau} - \bar{I}P) = 0 \quad [3.3]$$

Energy:

$$\rho(\bar{V} \cdot \bar{V}) H + \bar{\nabla} \cdot (\bar{q}_D + \bar{q}_R) - \bar{\nabla} \cdot (\bar{\tau}) \cdot \bar{V} = 0 \quad [3.4]$$

where the radiation pressure tensor and external force field terms have been omitted from the energy and momentum equations.

In order to describe the flow over blunt bodies, moving at hypersonic velocities, it is found convenient to state the conservation equations in orthogonal body oriented coordinate systems. The type of body under consideration, i.e. three dimensional, axisymmetric or two-dimensional, determines the stretching functions which are used in transforming the equations from general vector-tensor notation into a usable coordinate system. The classes of bodies considered in this development are axisymmetric and two dimensional which have the same stretching functions therefore the same form of the equations. The statement of the body oriented shock layer equations is given below (after Engel, Ref. 3.1):

Global Continuity:

$$\frac{\partial}{\partial x} (\rho u r^A) + \frac{\partial}{\partial y} (\tilde{\kappa} r^A \rho v) = 0 \quad [3.5]$$

Species Continuity:

$$\begin{aligned} \frac{\partial}{\partial x} (r^A \rho C_1 u) + \frac{\partial}{\partial y} (\tilde{\kappa} r^A \rho C_1 v) = & - \frac{\partial}{\partial x} (r^A J_{1,x}) \\ & - \frac{\partial}{\partial y} (\tilde{\kappa} r^A J_{1,y}) + \tilde{\kappa} r^A \omega_1 \end{aligned} \quad [3.6]$$

x-momentum:

$$\begin{aligned} \rho r^A u \frac{\partial u}{\partial x} + \rho \tilde{\kappa} r^A v \frac{\partial v}{\partial y} - \rho \kappa r^A u v + r^A \frac{\partial P}{\partial x} \\ - \frac{\partial}{\partial x} (r^A \tau_{xx}) - \frac{\partial}{\partial y} (\tilde{\kappa} r^A \tau_{xy}) - r^A \kappa \tau_{xy} + \tau_{zz} \frac{\partial r^A}{\partial x} = 0 \end{aligned} \quad [3.7]$$

y-momentum:

$$\begin{aligned} \rho r^A u \frac{\partial v}{\partial x} + \rho \tilde{\kappa} r^A v \frac{\partial v}{\partial y} - \rho \kappa r^A u^2 + \tilde{\kappa} r^A \frac{\partial P}{\partial y} \\ - \frac{\partial}{\partial x} (r^A \tau_{xy}) - \frac{\partial (\tilde{\kappa} r^A \tau_{yy})}{\partial y} + \kappa r^A \tau_{xx} + \tilde{\kappa} \tau_{zz} \frac{\partial r^A}{\partial y} \end{aligned} \quad [3.8]$$

Energy:

$$\begin{aligned} \rho r^A u \frac{\partial H}{\partial x} + \rho \tilde{\kappa} r^A v \frac{\partial H}{\partial y} = & - \frac{\partial}{\partial x} (r^A q_{D,x}) - \frac{\partial}{\partial y} (\tilde{\kappa} r^A q_{D,y}) \\ & - \frac{\partial}{\partial x} (r^A q_{R,x}) - \frac{\partial}{\partial y} (\tilde{\kappa} r^A q_{R,y}) + \frac{\partial}{\partial x} [r^A u \tau_{xx} + r^A v \tau_{xy}] \\ & + \frac{\partial}{\partial y} [\tilde{\kappa} r^A u \tau_{xy} + \tilde{\kappa} r^A v \tau_{yy}] = 0 \end{aligned} \quad [3.9]$$

where  $\kappa$  is the local body curvature and  $r$  is defined in Figure 3.2.

Using Figure 3.2, the following relationships may be found

$$\delta = \int_0^x (1 + k\delta) \tan \epsilon dx + \delta_0$$

$$\theta = \int_0^x k(x) dx$$

**Figure 3.2. Body-Oriented Coordinate System.**

$$r = r_b + y \sin \theta \quad [3.10]$$

$$dr = \sin \theta dy + \tilde{\kappa} \cos \theta dx \quad [3.11]$$

where

$$\tilde{\kappa} = 1 + \kappa y \quad [3.12]$$

$$A = 0, \quad \text{two dimensional}$$

$$A = 1, \quad \text{axisymmetric} \quad [3.13]$$

Defining the diffusional mass flux vectors,  $\bar{J}_i$ , and the heat flux vectors,  $\bar{q}_D$ , yields

$$\bar{J}_i = \bar{J}_i^{(D)} + \bar{J}_i^{(T)}; \quad \bar{q} = \bar{q}_D + \bar{q}_R \quad [3.14]$$

The components (neglecting force and pressure mass diffusion) are  
Concentration diffusion:

$$J_{i,x}^{(D)} = \frac{n_T^2}{\rho RT} \sum_{j=1}^n M_i M_j D_{ij} \left( Y_i \sum_{\substack{k=1 \\ k \neq j}}^n \left( \frac{\partial G_i}{\partial Y_k} \right)_{T, \varphi, Y_1} \frac{1}{\tilde{\kappa}} \frac{\partial Y_k}{\partial x} \right) \quad [3.15]$$

$$J_{i,y}^{(D)} = \frac{n_T^2}{\rho RT} \sum_{j=1}^n M_i M_j D_{ij} \left( Y_i \sum_{\substack{k=1 \\ k \neq j}}^n \left( \frac{\partial G_i}{\partial Y_k} \right)_{T, \varphi, Y_1} \frac{\partial Y_k}{\partial y} \right) \quad [3.16]$$

Thermal diffusion:

$$J_{i,x}^{(T)} = - \frac{D_i^T}{\tilde{\kappa}} \frac{\partial \ln T}{\partial x} \quad [3.17]$$

$$J_{i,y}^{(T)} = - D_i^T \frac{\partial \ln T}{\partial y} \quad [3.18]$$

Diffusional heat flux:

$$q_{D,x} = - \frac{k}{\tilde{\kappa}} \frac{\partial T}{\partial x} + \sum h_i J_{i,x} - \frac{P}{N^2} \sum_i \sum_{j \neq i} \frac{N_i}{M_i} \frac{D_i^T}{D_{ij}} \left( \frac{J_{j,x}}{\rho_j} - \frac{J_{i,x}}{\rho_i} \right) \quad [3.19]$$

$$q_{D,y} = -k \frac{\partial T}{\partial y} + \sum h_1 J_{1y} - \frac{P}{N^2} \sum_1 \sum_{j \neq 1} \frac{N_1}{m_1} \frac{D_1^T}{D_{1j}} \left( \frac{J_{1,y}}{\rho_j} - \frac{J_{1,y}}{\rho_1} \right) \quad [3.20]$$

Radiative heat flux:

$$q_{R,x} = \int_{x(\bar{r}_0)}^{x(\bar{r}_1)} \int_0^\infty \alpha_v \left( 4\pi B_v - \int_0^{4\pi} I_v(\bar{r}) d\Omega \right) dv \kappa dx \quad [3.21]$$

$$q_{R,y} = \int_{y(\bar{r}_0)}^{y(\bar{r}_1)} \int_0^\infty \alpha_v \left( 4\pi B_v - \int_0^{4\pi} I_v(\bar{r}) d\Omega \right) dv dy \quad [3.22]$$

In the momentum and the energy equations, the components of the stress tensor are:

$$\begin{aligned} \tau_{xx} = & \frac{\lambda}{\kappa r^A} \left( \frac{\partial}{\partial x} (r^A u) + \frac{\partial}{\partial y} (\kappa r^A v) \right) \\ & + \frac{2\mu}{\kappa} \left( \frac{\partial u}{\partial x} + \kappa v \right) \end{aligned} \quad [3.23]$$

$$\tau_{yy} = \frac{\lambda}{\kappa r^A} \left( \frac{\partial}{\partial x} (r^A u) + \frac{\partial}{\partial y} (\kappa r^A v) \right) + 2\mu \frac{\partial v}{\partial y} \quad [3.24]$$

$$\begin{aligned} \tau_{zz} = & \frac{\lambda}{\kappa r^A} \left( \frac{\partial}{\partial x} (r^A u) + \frac{\partial}{\partial y} (\kappa r^A v) \right) \\ & + 2\mu \left[ \frac{u}{\kappa r^A} \frac{\partial r^A}{\partial x} + \frac{v}{r^A} \frac{\partial r^A}{\partial y} \right] \end{aligned} \quad [3.25]$$

$$\tau_{xy} = \tau_{yx} = \mu \left( \frac{1}{\kappa} \frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} - \frac{\kappa}{\kappa} u \right) \quad [3.26]$$

The statement of these vector and tensor components completes the set of conservation equations expressed in body-oriented, orthogonal coordinates. These equations are subsequently simplified by an order



of magnitude analysis, yielding a much more manageable set of equations called the thin shock layer equations.

### The Thin Shock Layer Equations

The statement was made earlier that the basic conservation equations were to be used for the flow of a continuum reacting and radiating gas mixture over a blunted surface when thermodynamic equilibrium exists. It was also stated that a reduced set of equations could be used which would describe the flow when the outer boundary of the shock layer is a shock wave described by the Rankine-Hugoniot relations. This set of equations is called the thin shock layer equations and is arrived at by an order of magnitude analysis in which all terms of order one [1] are retained (Ref. 3.1). Further simplifications can be made by neglecting the shock layer thickness in comparison to the local body radius. This implies

$$\kappa \rightarrow 0, \quad \bar{\kappa} \rightarrow 1, \quad \text{and} \quad r^A \rightarrow r_b^A \quad [3.27]$$

Applying these limits and dropping terms using an order of magnitude analysis gives

### Global Continuity:

$$\frac{\partial}{\partial x} (r_b^A \rho u) + r_b^A \frac{\partial}{\partial y} (\rho v) = 0 \quad [3.28]$$

### Species Continuity:

$$\frac{\partial}{\partial x} (r_b^A \rho C_i u) + r_b^A \frac{\partial}{\partial y} (\rho C_i v) = -r_b^A \frac{\partial}{\partial y} (J_{i,y}) + r_b^A \omega_i \quad [3.29]$$

x-momentum:

$$\rho u \frac{\partial u}{\partial x} + \rho v \frac{\partial u}{\partial y} = - \frac{\partial P}{\partial x} + \frac{\partial}{\partial y} \left( \mu \frac{\partial u}{\partial y} \right) \quad [3.30]$$

y-momentum:

$$\frac{\partial P}{\partial y} = \frac{\kappa}{\bar{\kappa}} \rho u^2 \quad [3.31]$$

Energy:

$$\begin{aligned} \rho u \frac{\partial H}{\partial x} + \rho v \frac{\partial H}{\partial y} = & \frac{\partial}{\partial y} \left( \frac{k \partial T}{\partial y} \right) - \frac{\partial}{\partial y} \left\{ \sum_i h_i J_{i,y} \right. \\ & \left. - \frac{P}{N^2} \sum_i \sum_{j \neq i} \frac{N_j}{m_j} \frac{D_{ij}^T}{D_{ij}} \left( \frac{J_{i,y}}{\rho_j} - \frac{J_{j,y}}{\rho_i} \right) \right\} - \frac{\partial q_{R,y}}{\partial y} \\ & + \frac{\partial}{\partial y} \left( \mu u \frac{\partial u}{\partial y} \right) \end{aligned} \quad [3.32]$$

The shock layer equations 3.28 to 3.32 are now rewritten with some changes noted. The global continuity and the two momentum equations are unchanged. The global continuity is removed from the species equation, and the energy equation is written with enthalpy and temperature as the principle dependent variable. Furthermore, the resulting temperature form of the energy equation was obtained by subtracting the momentum equation, substituting the right hand side of the species equation from the left hand side and assuming binary diffusion. Thermal diffusion is also neglected. The resulting set of dimensional axisymmetric (A=1) shock layer equations is

Global Continuity:

$$\frac{\partial}{\partial x} (\rho u r_b) + r_b \frac{\partial}{\partial y} (\rho v) = 0 \quad [3.33]$$

Species Continuity:

$$\rho u \frac{\partial C_1}{\partial x} + \rho v \frac{\partial C_1}{\partial y} = \frac{\partial}{\partial y} (\rho D_{12} \frac{\partial C_1}{\partial y}) + \omega_1 \quad [3.34]$$

x-momentum:

$$\rho u \frac{\partial u}{\partial x} + \rho v \frac{\partial u}{\partial y} = - \frac{\partial P}{\partial x} + \frac{\partial}{\partial y} (\mu \frac{\partial u}{\partial y}) \quad [3.30]$$

y-momentum

$$\frac{\partial P}{\partial y} = \frac{\kappa}{\kappa} \rho u^2 \quad [3.31]$$

Energy (Enthalpy):

$$\begin{aligned} \rho u \frac{\partial H}{\partial x} + \rho v \frac{\partial H}{\partial y} &= \frac{\partial}{\partial y} (k \frac{\partial T}{\partial y}) - \frac{\partial}{\partial y} \left[ h_1 J_{1,y} - \frac{\partial q_{R,y}}{\partial y} \right] \\ &+ \frac{\partial}{\partial y} (\mu u \frac{\partial u}{\partial y}) \end{aligned} \quad [3.35]$$

Energy (Temperature):

$$\begin{aligned} \rho u C_p \frac{\partial T}{\partial x} + \rho v C_p \frac{\partial T}{\partial y} &= - \sum h_1 \omega_1 + u \frac{\partial P}{\partial x} \\ &+ \frac{\partial}{\partial y} (k \frac{\partial T}{\partial y}) - \frac{\partial q_R}{\partial y} + \mu \left( \frac{\partial u}{\partial y} \right)^2 + \sum C_{p1} \rho D_{12} \frac{\partial C_1}{\partial y} \frac{\partial T}{\partial y} \end{aligned} \quad [3.36]$$

These equations are in dimensional form and body oriented coordinates. In the following section these equations are nondimensionalized and an independent variable transformation is performed.

Nondimensionalized and Transformed Equations

The thin shock layer equations 3.30, 3.31, 3.33, 3.34, 3.35 and 3.36 can be nondimensionalized by using the following

dimensionless variables which are appropriate to the problem under consideration.

$$\begin{aligned}
 x &= \frac{x^*}{R^*} & y &= \frac{y^*}{R^*} & u &= \frac{u^*}{U_\infty^*} & v &= \frac{v^*}{U_\infty^*} \\
 \rho &= \frac{\rho^*}{\rho_{s,o}^*} & \mu &= \frac{\mu^*}{\mu_{s,o}^*} & P &= \frac{P^*}{\rho_{s,o}^* U_\infty^{*2}} & \delta &= \frac{\delta^*}{R^*} \\
 r_b &= \frac{r_b^*}{R^*} & \kappa &= \kappa^* R^* & k &= \frac{k^* T_{s,o}^*}{R^* \rho_{s,o}^* U_\infty^{*3}} & T &= \frac{T^*}{T_{s,o}^*} \\
 h &= \frac{h^*}{1/2 U_\infty^{*2}} & \omega_1 &= \frac{R^* \omega_1^*}{\rho_{s,o}^* U_\infty^*} & C_p &= \frac{C_p^* T_{s,o}^*}{1/2 U_\infty^{*2}} \\
 \tilde{\kappa} &= 1 + \kappa y & Re_s &= \frac{\rho_{s,o}^* U_\infty^* R^*}{\mu_{s,o}^*} & Sc &= \frac{\mu_{s,o}^*}{\rho_{s,o}^* D_{12}^*} \\
 E &= \frac{E^* p^*}{\rho_{s,o}^* U_\infty^{*3}} & \text{where } E^* &= \frac{\partial q_R^*}{\partial y^*}
 \end{aligned}$$

The dimensional variables used in the preceding section for the thin shock layer equations are denoted by a superscript asterisk (\*) in this discussion. The nondimensional variables have no superscript, and the nondimensional quantities for Reynolds number ( $Re_s$ ) and Schmit ( $Sc$ ) number have been incorporated. The resulting form of the shock layer equations in nondimensional form is shown in Table 3.1.

These equations can now undergo an independent variable transformation of the Dorodnitsyn type. The Dorodnitsyn transformation reduces the effect of a large variation in density in the body normal coordinate by incorporating the density in the transformed normal coordinate. The new independent variables now become

Table 3.1

## Nondimensional Thin Shock Layer Equations

Global Continuity:

$$\frac{\partial}{\partial x} (\rho u r_b) + r_b \frac{\partial}{\partial y} (\rho v) = 0 \quad [3.37]$$

Species Continuity:

$$\rho u \frac{\partial C_1}{\partial x} + \rho v \frac{\partial C_1}{\partial y} = \frac{1}{Re_s} \frac{\partial}{\partial y} \left( \frac{\mu}{Sc} \frac{\partial C_1}{\partial y} \right) + \omega_1 \quad [3.38]$$

x-momentum:

$$\rho u \frac{\partial u}{\partial x} + \rho v \frac{\partial u}{\partial y} = -\frac{\partial P}{\partial x} + \frac{1}{Re_s} \frac{\partial}{\partial y} \left( \mu \frac{\partial u}{\partial y} \right) \quad [3.39]$$

y-momentum:

$$\frac{\partial P}{\partial y} = \frac{\kappa}{\bar{\kappa}} \rho u^2 \quad [3.40]$$

Energy (Enthalpy):

$$\begin{aligned} \rho u \frac{\partial h}{\partial x} + \rho v \frac{\partial h}{\partial y} + 2\rho \left( u^2 \frac{\partial u}{\partial x} + uv \frac{\partial v}{\partial x} + v^2 \frac{\partial v}{\partial y} + vu \frac{\partial u}{\partial y} \right) \\ = 2 \frac{\partial}{\partial y} \left( \frac{k}{C_p} \frac{\partial h}{\partial y} \right) - 2 \frac{\partial}{\partial y} \left( \frac{k}{C_p} \sum h_1 \frac{\partial C_1}{\partial y} \right) + \frac{1}{Re_s} \sum \frac{\mu}{Sc} \frac{\partial C_1}{\partial y} \frac{\partial h_1}{\partial y} \\ + \frac{1}{Re_s} \sum h_1 \frac{\partial}{\partial y} \left( \frac{\mu}{Sc} \frac{\partial C_1}{\partial y} \right) - 2 \frac{\partial}{\partial y} \left( \frac{2}{Re_s} \mu \frac{\partial u}{\partial y} \right) \end{aligned} \quad [3.41]$$

Energy (Temperature):

$$\begin{aligned} \rho u C_p \frac{\partial T}{\partial x} + \rho v C_p \frac{\partial T}{\partial y} = \sum h_1 \omega_1 + 2u \frac{\partial P}{\partial x} \\ + 2 \frac{\partial}{\partial y} \left( k \frac{\partial T}{\partial y} \right) - 2 \frac{\partial}{\partial y} \left( \frac{2}{Re_s} \mu \left( \frac{\partial u}{\partial y} \right)^2 \right) + \frac{\mu}{Re_s} \left( \sum \frac{C_{p1}}{Sc} \frac{\partial C_1}{\partial y} \frac{\partial T}{\partial y} \right) \end{aligned} \quad [3.42]$$

$$\xi = x \quad [3.43]$$

$$\eta = \frac{\int_0^y \rho \, dy}{\int_0^{\tilde{\delta}} \rho \, dy} \quad [3.44]$$

where  $\delta$  is the shock standoff distance, and  $\tilde{\delta}$  is the transformed shock standoff distance which is used to normalize the variable  $\eta$  from 0 to 1. The derivatives become

$$\frac{\partial}{\partial x} = \frac{\partial}{\partial \xi} + \frac{\partial \eta}{\partial x} \frac{\partial}{\partial \eta} \quad [3.45]$$

$$\frac{\partial}{\partial y} = \frac{\rho}{\tilde{\delta}} \frac{\partial}{\partial \eta} \quad [3.46]$$

When the new independent variables are introduced and the transformations are applied, the nondimensional shock layer equations become the following in the transformed plane:

Global Continuity:

$$\frac{1}{r_b} \frac{\partial}{\partial \xi} (\rho u r_b) + \frac{\partial \eta}{\partial x} \frac{\partial}{\partial \eta} (\rho u) + \frac{\rho}{\tilde{\delta}} \frac{\partial}{\partial \eta} (\rho v) = 0 \quad [3.47]$$

Species Continuity:

$$\rho u \frac{\partial C_1}{\partial \xi} + \rho u \frac{\partial \eta}{\partial x} \frac{\partial C_1}{\partial \eta} + \frac{\rho^2 v}{\tilde{\delta}} \frac{\partial C_1}{\partial \eta} =$$

$$\frac{\rho}{Re_\delta \delta^2} \frac{\partial}{\partial \eta} \left( \frac{\rho \mu}{Sc} \frac{\partial C_1}{\partial \eta} \right) + \omega_1 \quad [3.48]$$

x-momentum:

$$\begin{aligned} \rho u \frac{\partial u}{\partial \xi} + \rho u \frac{\partial \eta}{\partial x} \frac{\partial u}{\partial \eta} + \frac{\rho v^2}{\delta} \frac{\partial u}{\partial \eta} = - \frac{\partial P}{\partial \xi} - \frac{\partial \eta}{\partial x} \frac{\partial P}{\partial \eta} \\ + \frac{\rho}{Re_s \delta^2} \frac{\partial}{\partial \eta} \left( \rho \mu \frac{\partial u}{\partial \eta} \right) \end{aligned} \quad [3.49]$$

y-momentum:

$$\frac{\partial P}{\partial \eta} = \frac{\kappa}{\delta} u^2 \quad [3.50]$$

Energy (Enthalpy):

$$\begin{aligned} \rho u \frac{\partial h}{\partial \xi} + \rho u \frac{\partial \eta}{\partial x} \frac{\partial h}{\partial \eta} + \rho v \frac{\rho}{\delta} \frac{\partial h}{\partial \eta} + 2\rho \left( u^2 \frac{\partial u}{\partial \xi} + u^2 \frac{\partial \eta}{\partial x} \frac{\partial u}{\partial \eta} \right. \\ \left. + v u \frac{\partial v}{\partial \xi} + v u \frac{\partial \eta}{\partial x} \frac{\partial v}{\partial \eta} + \frac{\rho v^2}{\delta} \frac{\partial v}{\partial \eta} + \frac{\rho v u}{\delta} \frac{\partial u}{\partial \eta} \right) = \\ \frac{2\rho}{\delta^2} \frac{\partial}{\partial \eta} \left( \frac{\rho k}{C_p} \frac{\partial h}{\partial \eta} \right) - \frac{2\rho}{\delta^2} \frac{\partial}{\partial \eta} \left( \frac{\rho k}{C_p} \sum h_i \frac{\partial C_i}{\partial \eta} \right) + \\ \frac{\rho}{Re_s \delta^2} \sum \left( \frac{\rho \mu}{Sc} \right) \frac{\partial C_i}{\partial \eta} \frac{\partial h_i}{\partial \eta} + \frac{\rho}{Re_s \delta^2} \sum h_i \frac{\partial}{\partial \eta} \left( \frac{\rho \mu}{Sc} \frac{\partial C_i}{\partial \eta} \right) \\ - 2 IE + \frac{2\rho}{Re_s \delta^2} \frac{\partial}{\partial \eta} \left( \rho \mu u \frac{\partial u}{\partial \eta} \right) \end{aligned} \quad [3.51]$$

Energy (Temperature):

$$\begin{aligned} \rho u C_p \frac{\partial T}{\partial \xi} + \rho u C_p \frac{\partial \eta}{\partial x} \frac{\partial T}{\partial \eta} + \rho^2 \frac{v}{\delta} C_p \frac{\partial T}{\partial \eta} = \sum h_i \omega_i \\ + 2u \frac{\partial P}{\partial \xi} + 2u \frac{\partial \eta}{\partial x} \frac{\partial P}{\partial \eta} + \frac{2\rho}{\delta^2} \frac{\partial}{\partial \eta} \left( \rho k \frac{\partial T}{\partial \eta} \right) - 2 IE \\ + \frac{2\rho^2 \mu}{Re_s \delta^2} \left( \frac{\partial u}{\partial \eta} \right)^2 + \frac{\rho^2 \mu}{Re_s \delta^2} \left( \sum \frac{C_{pi}}{Sc} \frac{\partial C_i}{\partial \eta} \right) \frac{\partial T}{\partial \eta} \end{aligned} \quad [3.52]$$

This nondimensional, transformed set of equations is a set of parabolic partial differential equations which needs two sets of

boundary conditions and one set of initial conditions. The formulation of the boundary and initial conditions will now be discussed.

### Boundary and Initial Conditions

The parabolic nature of the thin shock layer equations mathematically requires initial conditions as well as boundary conditions in order to obtain a solution. The entry vehicle axis of symmetry is the appropriate location of the starting line for zero angle of attack problems and provides the initial conditions for the shock equations. The determination of conditions along this line, called the stagnation line, is a major and important problem in itself. The stagnation line solution also requires the same two sets of boundary conditions as the shock layer solution. Consequently, development of the method to obtain these initial conditions (i.e. stagnation line solution) is delayed until after the boundary conditions are established. The boundary conditions are specified at the bow shock and at the ablator surface.

### Boundary Conditions at the Bow Shock

If the shock geometry is known, the Rankine-Hugoniot equations can be used to obtain the shock boundary conditions. The development of these equations in curvilinear coordinates follows directly from Ref. 3.5. The dimensional Rankine-Hugoniot equations written in body oriented coordinates (See Figures 3.2 and 3.3) are:

$$v_g = U_\infty^* \sin \phi \sin \epsilon - \bar{\rho} U_\infty^* \cos \phi \cos \epsilon \quad [3.53]$$

$$u_g = U_\infty^* \sin \phi \cos \epsilon + \bar{\rho} U_\infty^* \cos \phi \sin \epsilon \quad [3.54]$$

$$P_g^* + \rho_g^* (\bar{\rho} U_\infty^* \cos \phi)^2 = P_\infty^* + \rho_\infty^* (U_\infty^* \cos \phi)^2 \quad [3.55]$$



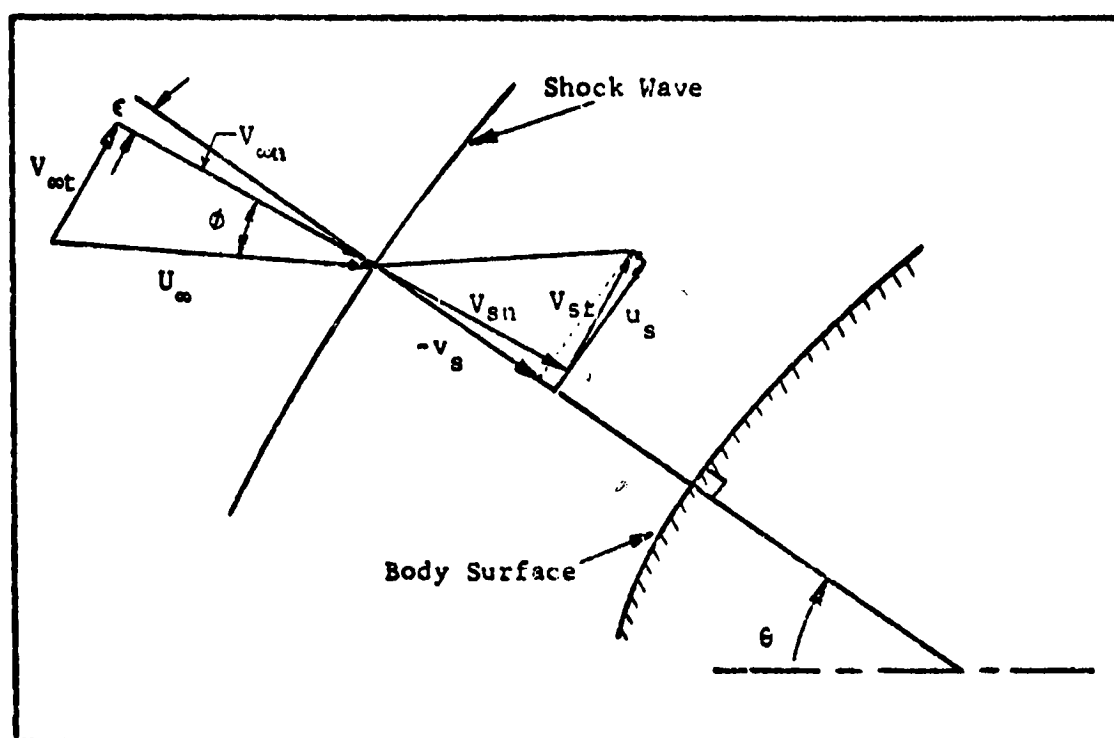


Fig. 3.3. Resolution of Velocity Components in a Body Oriented Coordinate System

$$h_s^* = \frac{U_\infty^{*2}}{2} - \frac{1}{2} (u_s^{*2} + v_s^{*2}) + h_\infty^* \quad [3.56]$$

The tangential velocity is negligible near the stagnation line thus reducing Equation 3.56 to

$$h_s^* = \frac{1}{2} U_\infty^{*2} (1 - \bar{\rho}^2) \cos^2 \phi + h_\infty^* \quad [3.57]$$

Nondimensionalizing Equations 3.53 through 3.57 and dropping  $P_\infty^*$  and  $h_\infty^*$  which are of order  $(\bar{\rho}^2)$  yields the following shock boundary conditions:

$$v_s = \sin \phi \sin \epsilon - \bar{\rho} \cos \phi \cos \epsilon \quad [3.58]$$

$$u_s = \sin \phi \cos \epsilon + \bar{\rho} \cos \phi \sin \epsilon \quad [3.59]$$

$$P_s = (1 - \bar{\rho}) \cos^2 \phi \quad [3.60]$$

$$h_s = (1 - \bar{\rho}^2) \cos^2 \phi \quad [3.61]$$

or

$$h_s = 1 - (u_s^2 + v_s^2) \quad [3.62]$$

where the following nondimensionalization convention has been used.

$$v = \frac{v^*}{U_\infty^*}, \quad u = \frac{u^*}{U_\infty^*}, \quad h = \frac{h^*}{\frac{1}{2} U_\infty^{*2}}, \quad P = \frac{P^*}{\rho_\infty^* U_\infty^{*2}}, \quad \bar{\rho} = \frac{\rho_\infty^*}{\rho_s^*}$$

The preceding Rankine-Hugoniot relations were obtained by performing mass, energy, and momentum balances across a thin shock wave. These relations now enable us to write the shock boundary conditions at  $y = \delta$ .

$$u = u_s$$

$$v = v_s$$

$$P = P_s$$

$$h = h_s$$

$$C_1 = C_{1,s} (P_s, h_s) \text{ (Assuming chemical equilibrium)}$$

$$I_v^-(\tau_{v,s}) = 0 \quad [3.63]$$

The Rankine-Hugoniot equations provide expressions for  $u_s$ ,  $v_s$ ,  $P_s$ , and  $h_s$ . The equation of state and free stream elemental mass fraction provides the post shock species mass fractions assuming chemical equilibrium. The specific intensity coming through the shock towards the body is specified as zero. It is noted that, in total, four boundary conditions are required for the energy equation because of its integro-differential nature. Thus, two boundary conditions, enthalpy and specific intensity, have been specified at the shock.

#### Boundary Conditions at the Ablator Surface

In order to specify the conditions at the ablator boundary, surface balances of mass, momentum, and energy are required. In connection with these balances, an evaluation is required of the surface removal mechanisms: sublimation, chemical reactions, and erosion. The purpose of this section is to integrate the various surface phenomena through which flow field calculations and the ablator response calculations are coupled. In Appendix A, the quasi-steady state equations that describe the ablator response are presented along with results from numerical solutions. The arrangement of the surface species and energy balances into a form suitable for coupling of the flow field to the ablator response is

given in Appendix B. This development in Appendix B is based on the assumption that sublimation is the major surface removal mechanism.

The boundary conditions for this investigation can be derived in either of two ways. The first technique consists of simply formulating a physical balance across the boundaries of the system. The second technique involves the use of the flow field equations themselves, which are integrated across the system boundaries and then contracted by taking the limit as the spacial increment approaches zero. Both of these methods should yield the same results. In this development both of these methods will be employed. The first method has the advantage that the physical significance of each term in the resulting equation is more readily evident. The integration technique will assure that all of the necessary terms have been included.

Species Boundary Conditions: At the char surface the following general surface balance is known to exist:

$$\begin{aligned}
 & \left( \begin{array}{l} \text{convective flux} \\ \text{of species 1 on} \\ \text{the char side} \\ \text{of the interface} \end{array} \right) + \left( \begin{array}{l} \text{diffusive flux of} \\ \text{species 1 on the} \\ \text{char side of the} \\ \text{interface} \end{array} \right) + \left( \begin{array}{l} \text{all contributions} \\ \text{to the net flux} \\ \text{of species 1 due} \\ \text{to the surface} \\ \text{phenomena} \end{array} \right) \\
 & = \left( \begin{array}{l} \text{convective flux of} \\ \text{species 1 on flow} \\ \text{field side of the} \\ \text{interface} \end{array} \right) + \left( \begin{array}{l} \text{diffusive flux of} \\ \text{species 1 on flow} \\ \text{field side of the} \\ \text{interface} \end{array} \right) \quad [3.64]
 \end{aligned}$$

or

$$\rho^* v^* C_1^- + J_1^{*-} + S_1^* = \rho^* v^* C_1^+ + J_1^{*+} \quad [3.65]$$

Nondimensionalizing by  $\rho_{s,0}^*$  and  $U_\infty^*$  gives

$$\rho v \bar{C}_1 + J_1^- + S_1 = \rho v C_1^+ + J_1^+ \quad [3.66]$$

The surface generation,  $S_1$ , can now be quantitatively defined and the remaining terms of the equation verified by examining the integral derivation of the species surface balance from the nondimensional shock layer species continuity equation. Equation 3.38 is now rewritten as

$$\rho u \frac{\partial C_1}{\partial x} + \rho v \frac{\partial C_1}{\partial y} = - \frac{\partial J_1}{\partial y} + \omega_1 \quad [3.67]$$

where  $J_1$  has been substituted back into the equation. Equation 3.67 can be further simplified when we realize that a no slip stipulation at the surface implies that  $u = 0$  which gives

$$\rho v \frac{dC_1}{dy} = - \frac{dJ_1}{dy} + \omega_1 \quad [3.68]$$

The generation term in this equation can be expressed as the sum of contributing effects

$$\omega_1 = \omega_{1\text{homo}} + \omega_{1\text{het}} + \omega_{1\text{subl}} \quad [3.69]$$

where  $\omega_{1\text{homo}}$  is the net generation of species 1 by means of homogeneous chemical reactions,  $\omega_{1\text{het}}$  is the net generation by heterogeneous reactions, and  $\omega_{1\text{subl}}$  is the rate of formation of gaseous species by sublimation.

Each of the previous terms represents a generation of mass of species 1 by reaction per unit time per unit volume. For the ablator this unit volume contains both gas and solid. Thus if  $R_1^*$  is the rate of formation of species 1 by heterogeneous reaction per

unit of solid surface,  $A_r^*$ , then  $\omega_{het}^*$  is given by the following equation

$$\omega_{het}^* = \frac{R_1^* A_r^*}{A^* \Delta y} \quad [3.70]$$

The ratio of  $A_r^*/A^* \Delta y$  represents the concentration of surface area, i.e., the available surface area per unit volume of the reacting system.

In the subsequent steps of this analysis Equation 3.68 will be integrated across the char interface a distance of  $\Delta y^*$  which will define the thickness of a control volume of cross-sectional area  $A^*$  which contains a total reactive surface area  $A_r^*$ . Equation 3.70 can be written in non-dimensional form by the following manipulations.

$$\omega_{het} = \left( \frac{R_1^* A_r^*}{A^* \Delta y^*} \right) \left( \frac{R^*}{\rho_{s,o}^* U_{\infty}^*} \right) = \frac{R_1 A_r^*}{A^* \Delta y} \quad [3.71]$$

This equation can also be conveniently expressed in terms of char porosity,  $\epsilon_p$  (volume of voids per unit volume), which for an isotropic material is equal to  $(A^* - A_r^*)/A^*$ . Thus,

$$\omega_{het} = R_1 \frac{(1 - \epsilon_p)}{\Delta y} \quad [3.72]$$

In a similar manner the sublimation term can be derived as,

$$\omega_{subl} = \left( \frac{f_1^* A_r^*}{A^* \Delta y^*} \right) \left( \frac{R^*}{\rho_{s,o}^* U_{\infty}^*} \right) = f_1 \frac{(1 - \epsilon_p)}{\Delta y} \quad [3.73]$$

where  $f_1^*$  is the mass rate of sublimation per unit area of solid surface.

Following the above procedure, the homogeneous reaction rate term can be written as

$$\omega_{i,homo} = (\omega_{i,homo}^* \epsilon_p) \left( \frac{R^*}{\rho_{s,o} u_\infty^*} \right) \quad [3.74]$$

where  $\omega_{i,homo}^*$  is the net rate of formation of species  $i$  in the gas phase.

Having now defined, quantitatively, the total generation term in the species continuity equation, the integration technique will be used to define  $f_i$  in terms of  $\omega_i$ . Equation 3.68 can be integrated as follows:

$$\int_{y^-}^{y^+} \rho v \frac{dC_i}{dy} dy = - \int_{y^-}^{y^+} \frac{dJ_i}{dy} dy + \int_{y^-}^{y^+} (\omega_{i,homo} + \omega_{i,hetero} + \omega_{i,subl}) dy \quad [3.75]$$

Substituting Equations 3.72 and 3.74 into the above gives:

$$\rho v C_i^+ - \rho v C_i^- = -J_i^+ + J_i^- + \omega_{i,homo} \Delta y + \int_{y^-}^{y^+} \frac{(R_i + f_i)(1-\epsilon_p)}{\Delta y} dy \quad [3.76]$$

Integrating the above equation noting that  $\rho v$  is a constant gives

$$\rho v C_i^+ - \rho v C_i^- = J_i^- + J_i^+ + \omega_{i,homo} \Delta y + (R_i + f_i)(1-\epsilon_p) \Delta y \quad [3.77]$$

Taking the limit as  $\Delta y \rightarrow 0$  and rearranging gives:

$$\rho v C_i^- + J_i^- + (R_i + f_i)(1-\epsilon_p) = \rho v C_i^+ + J_i^+ \quad [3.78]$$

Comparing the above equation with Equation 3.66 confirms our previous surface balance and defines  $S_i$ , the dimensionless surface generation term, as:

$$S_1 = (R_1 + J_1)(1 - \epsilon_p) \quad [3.79]$$

Although Equation 3.78 is completely rigorous for all species which may exist either in the flow field or the char, a more specific interpretation of this equation can be arrived at for various types of species which can be grouped into distinct categories. For example: pyrolysis gases (excluding carbon), carbon gas, all remaining gases, and finally the solid carbon. A more complete treatment of how these various categories are handled is given in Appendix B.

Energy Boundary Conditions: Starting with Equation 3.32 the energy balance at the surface will be developed in a fashion analogous to the preceding one for the species equations. Equation 3.32 is the energy equation in a dimensionless form. This equation becomes (omitting thermal diffusion)

$$\begin{aligned} \rho u \frac{\partial H}{\partial x} + \rho v \frac{\partial H}{\partial y} = \frac{\partial}{\partial y} \left[ k \frac{\partial T}{\partial y} - \sum h_i J_i \right] - \frac{\partial q_{R,y}}{\partial y} \\ + \frac{\partial}{\partial y} (\mu u \frac{\partial u}{\partial y}) \end{aligned} \quad [3.80]$$

Again noting that the tangential velocity,  $u$ , equals zero at the surface, Equation 3.80 reduces to

$$\rho v \frac{dH}{dy} = \frac{d}{dy} \left( k \frac{dT}{dy} \right) - \frac{d}{dy} \left[ \sum h_i J_i \right] - \frac{\partial q_{R,y}}{\partial y} \quad [3.81]$$

Integrating the above equation across the char surface gives



$$\begin{aligned} \rho v H^+ - \rho v H^- &= [k \frac{dT^+}{dy} - \sum h_i J_i^+ + q_R^-] \\ &- [k \frac{dT^-}{dy} - \sum h_i J_i^- + q_R^+] \end{aligned} \quad [3.82]$$

The above equation can be placed in a more convenient form using the following equations. The equation for the total enthalpy is

$$H = h + \frac{\bar{v}^2}{2} \quad [3.83]$$

and, using the definition of the static enthalpy it can be written in terms of the mass fraction,  $C_i$ , and enthalpy per unit mass of  $i$ ,

$$h_i \cdot H = \sum h_i C_i + \frac{\bar{v}^2}{2} \quad [3.84]$$

Using the above relationships and noting that kinetic energy terms are small compared to the enthalpy terms and can be deleted, the left hand side of Equation 3.82 can be written as

$$\rho v \sum h_i C_i^+ - \rho v \sum h_i C_i^- \quad [3.85]$$

The terms on the right hand side of Equation 3.82 will now be evaluated. Considering the first bracketed term on the right hand side of Equation 3.82, all three terms in this bracket are to be evaluated in the flow field and will be retained. The first term is the conductive heat flux, the second is the diffusive energy flux, and the third,  $q_R^-$ , is the energy radiated to the surface from the flow field. In the second bracketed term the only terms of importance are the conduction of energy into the char ( $k \frac{dT^-}{dy}$ )

and the energy radiated by the surface ( $q_R^+$ ). The resulting equation after substituting in these relations becomes

$$\rho v \sum h_i (C_i^+ - C_i^-) + k \frac{dT}{dy} - q_{RR} = k \frac{dT}{dy} + \sum h_i J_i^+ - q_R \quad [3.86]$$

where  $q_R$  is the energy radiated by the flow field to the surface and  $q_{RR}$  is the energy radiated by the surface toward the flow field. The terms with the superscript  $(-)$  are evaluated with an ablator response analysis described in Appendices A and B. The terms with the superscript  $(+)$  are evaluated in the flow field analyses.

The above form of the surface energy balance shows the coupling between the compositions, mass fluxes and surface temperature. Thus a simultaneous solution of the species surface balances and the energy surface balance is required.

Momentum Boundary Conditions: If a no slip condition is imposed at the ablator surface, the tangential velocity,  $u$ , is zero for all  $x$  at  $y = 0$ . This boundary condition has been used in the previous species and energy surface balance derivations. This condition can also be used in the determination of the surface pressure from the  $y$ -momentum equations, Equation 3.31, yielding

$$\frac{dP}{dy} = 0 \quad [3.87]$$

Integrating Equation 3.87 gives the simple result

$$P^+ = P^- \quad [3.88]$$

which states that the pressures on both sides of the surface are equal when considering terms in the y-momentum equation that are first order accurate.

Global Continuity Boundary Condition: The global continuity equation is a statement of the conservation of mass and is a first order partial differential equation. The solution of this equation requires the specification of the normal velocity at the surface. More normally the mass injection rate is specified  $(\rho v)$ . Figure 3.4 shows that the flow of gases  $(\rho_g v_g)$  to the surface along with the rate of solid addition  $(\rho_c v_r)$  is equal to the total mass injection rate,  $(\rho v)_{\text{wall}}$ . This same conclusion can be arrived at by solving the global continuity equation, Equation 3.37, at the surface

$$\frac{d}{dy} (\rho v) = 0 \quad [3.89]$$

which upon integrating becomes

$$\rho v|^- = \rho v|^+ \quad [3.90]$$

or

$$\rho_g v_g + \rho_c v_r = \rho v_{\text{wall}} \quad [3.91]$$

The terms  $\rho_g v_g$  and  $\rho_c v_r$  are determined from the ablator response characteristics.

In summary, the derivation of the wall boundary conditions has been completed by applying surface mass, energy, and momentum balances. These balances describe the interaction of the ablator with the flow field. Appendix A develops the equations which describe the ablator response and Appendix B shows the coupling

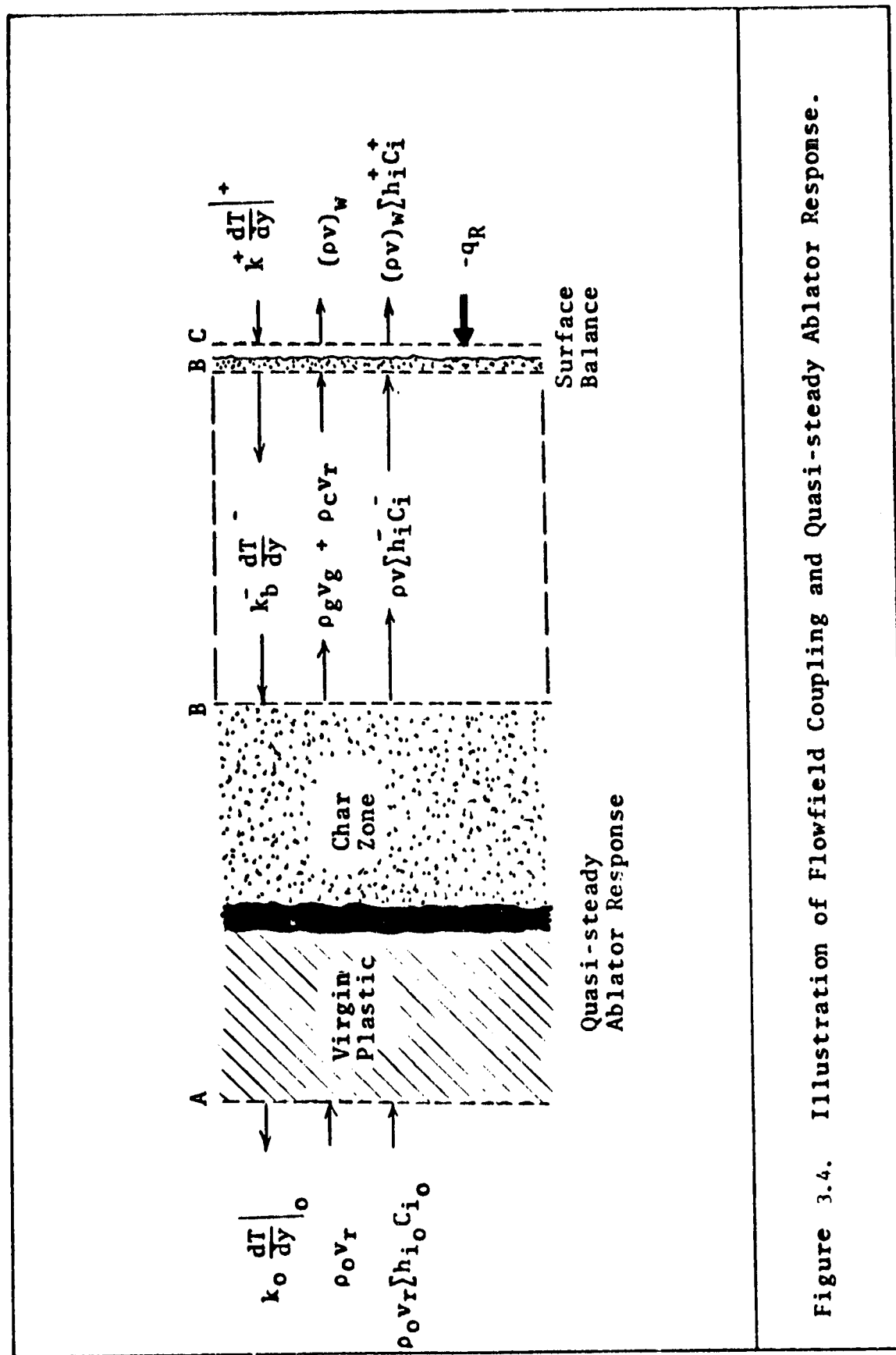


Figure 3.4. Illustration of Flowfield Coupling and Quasi-steady Ablator Response.

that exists between the species and energy surface balances. The behavior of the flow field is governed by the shock layer equations along with its boundary conditions and initial conditions. To complete the flow field specification, the initial conditions will now be discussed.

### Initial Conditions

In order to obtain initial values for the shock layer solution, a reduced set of thin shock-layer equations must be solved at  $x = 0$  along  $y$ , the stagnation line (see Figure 3.2). The solution of this set of equations is important because any distributional shock layer solution, due to its parabolic nature, is only as valid as its initial values.

The solution of the stagnation line equations by direct methods has been approached in two ways. The work of Ho and Probstein (Ref. 3.6) typifies the stagnation region solutions which use expansions of the dependent variables in  $x$  to obtain the stagnation and near stagnation line equations. The work of Hoshizaki and Wilson (Ref. 3.7) typifies the stagnation line solutions which determine the stagnation line equations by formally taking the limit of the terms in the shock layer equations at  $x = 0$  using symmetry conditions. The latter method is used in this analysis.

Engel (Ref. 3.8) develops the stagnation line equations by a detailed derivation from the thin shock layer equations. The global continuity, species continuity, and energy equations are reduced by taking limits and noting that at  $x = 0$ ,  $u = 0$ , the species and energy equations become ordinary differential equations. The

x-momentum equation yields a trivial solution in its unaltered form, therefore the x-momentum equation is differentiated with respect to x and its limiting form along the stagnation line is determined.

The normal curvature effects are dropped and only first order terms are kept. The resulting equations are

Global Continuity:

$$2 \frac{\partial u}{\partial x} = \frac{-1}{\rho} \frac{\partial}{\partial y} (\rho v) \quad [3.92]$$

Species Continuity:

$$\rho v \frac{\partial C_1}{\partial y} = \frac{-\partial}{\partial y} (J_{1,y}) + \omega_1 \quad [3.93]$$

x-momentum:

$$\frac{\partial}{\partial y} \left[ \mu \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial y} \right) \right] - \rho v \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial x} \right) + \rho \left( \frac{\partial u}{\partial x} \right)^2 + \left( \frac{\partial^2 p}{\partial x^2} \right)_{x=0} = 0 \quad [3.94]$$

y-momentum:

$$\frac{\partial p}{\partial y} = 0 \quad [3.95]$$

Energy (Enthalpy):

$$\begin{aligned} \rho v \frac{\rho}{\delta} \frac{dh}{dn} + \frac{2(\rho v)}{\delta} \frac{dv}{dn} &= \frac{2\rho}{\delta} \frac{d}{dn} \left( \rho \frac{k}{C_p} \frac{dh}{dn} \right) \\ &- \frac{2\rho}{\delta^2} \frac{d}{dn} \left( \rho \frac{k}{C_p} \int h_1 \frac{dC_1}{dn} \right) + \frac{\rho}{Re_s \delta^2} \int \left( \frac{\rho \mu}{Sc} \right) \frac{dC_1}{dn} \frac{dh_1}{dn} \\ &+ \frac{\rho}{Re_s \delta^2} \int h_1 \frac{d}{dn} \left( \frac{\rho \mu}{Sc} \frac{dC_1}{dn} \right) - 2 E \end{aligned} \quad [3.96]$$

Energy (Temperature):

$$\begin{aligned} \rho v C_p \frac{dT}{dy} &= - \int h_1 \omega_1 + \frac{d}{dy} \left( k \frac{dT}{dy} \right) - \frac{dq_R}{dy} \\ &+ \sum C_{p1} J_1 \frac{dT}{dy} \end{aligned} \quad [3.97]$$

Using the nondimensional quantities given previously, the global continuity and momentum equations can be placed in dimensionless form as follows:

Global Continuity:

$$2 \frac{\partial u}{\partial x} = - \frac{1}{\rho} \frac{\partial}{\partial y} (\rho v) \quad [3.98]$$

x-momentum:

$$\begin{aligned} \frac{1}{Re_s} \frac{\partial}{\partial y} \left( \mu \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial x} \right) \right) - \rho v \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial x} \right) - \rho \left( \frac{\partial u}{\partial x} \right)^2 \\ - \frac{\partial^2 p}{\partial x^2} = 0 \end{aligned} \quad [3.99]$$

y-momentum:

$$\frac{\partial p}{\partial y} = 0 \quad [3.100]$$

Species Continuity:

$$\rho v \frac{dC_1}{dy} = - \frac{1}{Re_s} \frac{dJ_1}{dy} + \omega_1 \quad [3.101]$$

Energy (Enthalpy):

$$\begin{aligned} \rho v \frac{\partial h}{\partial y} + 2\rho v^2 \frac{\partial v}{\partial y} = 2 \frac{\partial}{\partial y} \left( \frac{k}{C_p} \frac{\partial h}{\partial y} \right) - 2 \frac{\partial}{\partial y} \left( \frac{k}{C_p} \right) \sum h_i \frac{\partial C_i}{\partial y} \\ + \frac{1}{Re_s} \sum \frac{\mu}{Sc} \frac{\partial C_i}{\partial y} \frac{\partial h_i}{\partial y} + \frac{1}{Re_s} \sum h_i \frac{\partial}{\partial y} \left( \frac{\mu}{Sc} \frac{\partial C_i}{\partial y} \right) - 2 E \end{aligned} \quad [3.102]$$

Energy (Temperature):

$$\begin{aligned} \rho v C_p \frac{dT}{dy} = \sum h_i \omega_i + 2 \frac{d}{dy} \left( k \frac{dT}{dy} \right) - 2 E \\ + \frac{\mu}{Pe_s} \sum C_{pi} J_i \frac{dT}{dy} \end{aligned} \quad [3.103]$$

The second derivative of pressure term in the x-momentum equation can be evaluated at the shock using the Rankine-Hugoniot equations if  $\frac{\partial P}{\partial y}$  is assumed zero. This pressure term is

$$\frac{\partial^2 P}{\partial x^2} = -2 \bar{\rho} (1 - \bar{\rho}) \left( \frac{\partial \phi}{\partial x} \right)^2 \quad [3.104]$$

For a concentric shock  $\frac{\partial \phi}{\partial x} = 1$ . The concentric shock assumption is usually rather good for hypersonic flow, however in this derivation  $\frac{\partial \phi}{\partial x}$  will be treated as a parameter and thus will be left general. Substituting this term into the momentum equation yields

$$\begin{aligned} \frac{1}{Re_s} \frac{\partial}{\partial y} \left( \mu \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial x} \right) \right) - \rho v \frac{\partial}{\partial y} \left( \frac{\partial u}{\partial x} \right) - \rho \left( \frac{\partial u}{\partial x} \right)^2 \\ + 2 \bar{\rho} (1 - \bar{\rho}) \left( \frac{\partial \phi}{\partial x} \right)^2 = 0 \end{aligned} \quad [3.105]$$

To obtain a more classical form of the momentum equation, a velocity function is defined

$$f' = \frac{\partial u}{\partial x} / \frac{\partial u_{s,o}}{\partial x} = \lim_{x \rightarrow 0} \frac{u}{u_s} = \text{a function of } y \quad [3.106]$$

where

$$\frac{\partial u_{s,o}}{\partial x} = \frac{\partial \phi}{\partial x} + \bar{\rho} \frac{\partial \epsilon}{\partial x} \quad [3.107]$$

from the Rankine-Hugoniot equations. Substituting into the momentum equation yields

$$\begin{aligned} \frac{1}{Re_s} \frac{\partial}{\partial y} \left( \mu \frac{\partial f'}{\partial y} \right) - \rho v \frac{\partial f'}{\partial y} - \rho \frac{\partial u_{s,o}}{\partial x} (f')^2 \\ + \frac{2 \bar{\rho} (1 - \bar{\rho})}{\partial u_{s,o} / \partial x} \left( \frac{\partial \phi}{\partial x} \right)^2 = 0 \end{aligned} \quad [3.108]$$

The stagnation line equations can now be transformed using the



Dorodnitsin transformation as was used for the thin shock layer equations. For this transformation, Equations 3.43 through 3.46 are applied

$$\xi = x \quad [3.43]$$

$$\eta = \frac{\int_0^y \rho dy}{\delta} \quad [3.44]$$

$$\frac{\partial}{\partial x} = \frac{\partial}{\partial \xi} + \frac{\partial \eta}{\partial x} \frac{\partial}{\partial \eta} \quad [3.45]$$

$$\frac{\partial}{\partial y} = \frac{\rho}{\delta} \frac{\partial}{\partial \eta} \quad [3.46]$$

where  $\partial \eta / \partial x$  equals zero at the stagnation line. The resulting transformed equations are

Global Continuity:

$$2 \frac{\partial u}{\partial \xi} = - \frac{1}{\delta} \frac{\partial}{\partial \eta} (\rho v) \quad [3.109]$$

Species Continuity:

$$\frac{\rho v}{\delta} \frac{dC_1}{d\eta} = \frac{1}{Re_s \delta^2} \frac{d}{d\eta} \left( \frac{\rho \mu}{S_c} \frac{dC_1}{d\eta} \right) + \frac{\omega_1}{\rho} \quad [3.110]$$

Velocity function:

$$f' = \frac{\partial u / \partial \xi}{\partial u_{s,0} / \partial \xi} = \frac{-\frac{1}{2\delta} \frac{\partial}{\partial \eta} (\rho v)}{\frac{\partial u_{s,0}}{\partial \xi}} = \frac{\partial f}{\partial \eta} \quad [3.111]$$

x-momentum:

$$\frac{d}{d\eta} \left( \rho \mu \frac{df'}{d\eta} \right) - \rho v Re_s \delta \frac{df'}{d\eta} + \frac{Re_s \delta}{\mu \left( \frac{\partial u_{s,0}}{\partial \xi} \right)} \left[ 2 \bar{u} (1 - \bar{u}) \left( \frac{\partial f'}{\partial \xi} \right)^2 - \mu \left( \frac{\partial u_{s,0}}{\partial \xi} f' \right)^2 \right] = 0 \quad [3.112]$$

y-momentum:

$$\frac{dP}{d\eta} = 0 \quad [3.113]$$

Energy (Enthalpy):

$$\begin{aligned} \rho v \frac{\rho}{\delta} \frac{\partial h}{\partial \eta} + 2 \frac{\rho v^2}{\delta} \frac{\partial v}{\partial \eta} &= \frac{\rho_0}{\delta^2} \frac{\partial}{\partial \eta} \left( \rho \frac{k}{C_p} \frac{\partial h}{\partial \eta} \right) - \frac{2\rho}{\delta^2} \frac{\partial}{\partial \eta} \\ &\left( \frac{\rho k}{C_p} \int h_1 \frac{\partial C_1}{\partial \eta} \right) + \frac{\rho}{Re_s \delta^2} \int \left( \frac{\rho \mu}{Sc} \right) \frac{\partial C_1}{\partial \eta} \frac{\partial h_1}{\partial \eta} + \\ &\frac{\rho}{Re_s \delta^2} \int h_1 \frac{\partial}{\partial \eta} \left( \frac{\rho \mu}{Sc} \frac{\partial C_1}{\partial \eta} \right) - 2 E \end{aligned} \quad [3.114]$$

Energy (Temperature):

$$\begin{aligned} \frac{\rho v C_p}{\delta} \frac{dT}{d\eta} &= \int h_1 \left( \frac{\omega_1}{\rho} \right) + \frac{2}{\delta^2} \frac{d}{d\eta} \left( \rho k \frac{dT}{d\eta} \right) \\ &- \frac{2 E}{\rho} + \frac{\rho \mu}{Re_s \delta^2} \int \frac{C_{p1}}{Sc} \frac{dC_1}{d\eta} \frac{dT}{d\eta} \end{aligned} \quad [3.115]$$

Integrating Equation 3.111 yields

$$\rho v = - 2 \left( \frac{\partial u_{s,0}}{\partial \xi} \right) \delta f \quad [3.116]$$

Equation 3.116 can be used to eliminate  $\rho v$  from the x-momentum equation. This procedure results in a third order nonlinear ordinary differential equation.

$$\begin{aligned} (\rho \mu f'')' + [2 Re_s \delta^2 \left( \frac{\partial u_{s,0}}{\partial \xi} \right)] f f'' + \\ \frac{Re_s \delta^2}{\left( \frac{\partial u_{s,0}}{\partial \xi} \right)} \left[ \frac{2 \bar{\rho} (1-\bar{\rho})}{\rho} \left( \frac{\partial \phi}{\partial \xi} \right)^2 - \left( \frac{\partial u_{s,0}}{\partial \xi} \right)^2 (f')^2 \right] = 0 \end{aligned} \quad [3.117]$$

The boundary conditions for this equation are

$$\begin{aligned} f' &= 1 & \text{at} & \eta = 1 \\ f' &= 0 & \text{at} & \eta = 0 \\ f &= f_w = \frac{-(\rho v)_w}{2\tilde{\delta} (\partial u_{s,0}/\partial \xi)} & \text{at} & \eta = 0 \end{aligned} \quad [3.118]$$

In addition, Equation 3.116 has a boundary condition imposed on it to determine the transformed standoff distance,  $\tilde{\delta}$ .

$$f = f_s = \frac{-(\rho v)_s}{2\tilde{\delta} (\frac{\partial u_{s,0}}{\partial \xi})} \quad \text{at} \quad \eta = 1 \quad [3.119]$$

The momentum equation can be reduced to a first and a second order equation by defining

$$\zeta = \frac{f'}{\tilde{\delta}} \quad [3.120]$$

and substituting into Equation 3.117

$$\begin{aligned} (\rho\mu)\zeta'' + [2 \operatorname{Re}_s \tilde{\delta}^2 (\frac{\partial u_{s,0}}{\partial \xi}) f + (\rho\mu)'] \zeta' \\ - \operatorname{Re}_s \tilde{\delta}^3 (\frac{\partial u_{s,0}}{\partial \xi}) \zeta^2 = -(\frac{2 \operatorname{Re}_s \tilde{\delta} \bar{\rho} (1-\bar{\rho})}{f (\frac{\partial u_{s,0}}{\partial \xi})}) (\frac{\partial \phi}{\partial \xi})^2 \end{aligned} \quad [3.121]$$

The resulting boundary conditions for Equation 3.121 are:

$$\begin{aligned} \zeta &= 0 & \text{at} & \eta = 0 \\ \zeta &= 1/\tilde{\delta} & \text{at} & \eta = 1 \end{aligned} \quad [3.122]$$

Equations 3.120 and 3.121 can now be solved simultaneously to give  $\zeta$  and  $f$ . Equation 3.119 can then be solved for the transformed standoff distance,  $\tilde{\delta}$ .

The foregoing discussion of the stagnation line equations completes the formulation for the initial conditions needed in the solution of the thin shock layer equations. The boundary conditions for the energy and species stagnation line equations are the very same as those discussed for the shock layer equations, i.e. the Rankine-Hugoniot relations and the surface balances. The shock layer solution requires one more piece of information to be completely specified - the shock geometry. The following section describes the geometrical relation necessary in determining shock shape.

#### Shock Geometry

There are three methods which may be used to determine the shock shape and pressure distribution around the body. First, the technique used in Reference 3.9 and 3.10 and others can be used. The shock shape is specified a priori from which the wall pressure distribution is calculated as the solution proceeds around the body. An output shock shape is calculated from the geometrical relation (Ref. 3.11).

$$\epsilon = \tan^{-1} [(d\delta/d\xi)/(1 + \kappa\delta)] \quad [3.123]$$

where  $\delta$  is calculated as a result of the x-momentum solution. This output angle is compared with the input angle and if the input and output meet a specified tolerance, the solution is said to be converged.

The second technique of predicting shock shape involves specifying a wall pressure distribution a priori. Preferably this distribution is known from experimental data for hypersonic Mach

numbers. The change in pressure due to radiation coupling is justifiably neglected (Ref. 3.9), and a shock shape is also assumed. The shock layer equations are solved around the body and the calculated and input pressure distributions are compared. The shock shape is then numerically adjusted according to the pressure differences. This calculation is repeated until satisfactory pressure convergence is obtained.

The third technique involves a simultaneous solution of the geometric relation

$$\delta = \int_0^{\xi} (1 + \kappa \xi) \tan \epsilon \, d\xi + \delta_0 \quad [3.124]$$

with the shock layer equations. The pressure behind the shock and the pressure at the surface are calculated as part of the thin shock layer solution. Only one around the body iteration is required for this technique.

The first two techniques have been implemented in a modified version of the computer program described in Reference 3.9 and 3.10. Unfortunately both of these techniques have severe limitations. The first method consumes a large amount of computer time because many around the body iterations are necessary to insure convergence. In addition, each estimate of shock shape must be made by hand. The second method also consumes a ~~great deal~~ of computer time for the same reason - too many iterations. The second technique does have the advantage over the first that the input and output pressures are smooth allowing an automatic update of the shock shape by the computer. However, this technique does not necessarily satisfy the geometrical differential equation, Equation 3.124.

Engel (Ref. 3.8) concludes, therefore, that since the first two techniques are subject to undesirable limitation, the third technique should be implemented.

This section concludes the development of the mathematical equations which describe the shock layer surrounding a reentering blunt vehicle. The solution of these mathematical equations is taken up in Chapter V where a numerical procedure is developed which requires that the thin shock layer and stagnation line equations first be linearized. The linearization of these equations, therefore, is carried out in the next section.

#### Quasilinearization of Shock Equations

The differential equations can be made linear in the unknown variables by performing a Taylor series expansion and truncating after the first derivative term of the Taylor series. Coupling between the equations is eliminated in this analysis because computer storage becomes excessive with many chemical species. Coupling between the equations is required for the quasilinearization technique as defined by Bellman and Kulaba (Ref. 3.12), but the uncoupled procedure is shown to be computationally superior by Blottner (Ref. 3.13). Therefore, the present scheme uncouples the conservation equations.

The thin shock layer equations and the stagnation line equations, with the exception of global continuity and y-momentum, can be written in the following general form

$$\frac{\partial^2 W}{\partial \eta^2} + \alpha_1 \frac{\partial W}{\partial \eta} + \alpha_2 W + \alpha_3 + \alpha_4 \frac{\partial W}{\partial \xi} = 0 \quad [3.125]$$

where  $\alpha_1$ ,  $\alpha_2$ ,  $\alpha_3$ , and  $\alpha_4$  are the linearized coefficients for  $W = \zeta$ ,  $W = u$ ,  $W = C_1$ ,  $W = H$ , and  $W = T$ . These coefficients are as follows

x momentum (stagnation line,  $W = \zeta$ )

$$\begin{aligned}\alpha_1 &= \frac{2 \operatorname{Re}_s \tilde{\delta}^2}{\rho \mu} \left( \frac{\partial u_{s,0}}{\partial \xi} \right) f + \frac{(\rho \mu)'}{\rho \mu} \\ \alpha_2 &= - \frac{2 \operatorname{Re}_s \tilde{\delta}^3}{\rho \mu} \left( \frac{\partial u_{s,0}}{\partial \xi} \right) \zeta \\ \alpha_3 &= 2 \operatorname{Re}_s \tilde{\delta} \left[ \frac{\bar{\rho} (1-\bar{\rho})}{\rho^2 \mu} \left( \frac{\partial \phi^2}{\partial \xi} \right) + \frac{\tilde{\delta}^2}{2 \rho \mu} \left( \frac{\partial u_{s,0}}{\partial \xi} \right) \zeta^2 \right] \\ \alpha_4 &= 0\end{aligned}\quad [3.126]$$

x-momentum (shock layer,  $W = u$ )

$$\begin{aligned}\alpha_1 &= \frac{1}{\rho \mu} \left[ (\rho \mu)' - \operatorname{Re}_s \tilde{\delta} \rho v - \operatorname{Re}_s \tilde{\delta}^2 u \frac{\partial \eta}{\partial x} \right] \\ \alpha_2 &= \frac{-\operatorname{Re}_s \tilde{\delta}^2}{\rho \mu} \left[ \frac{\partial u}{\partial \xi} + \frac{\partial \eta}{\partial x} \frac{\partial u}{\partial \eta} + \frac{2 \tilde{\delta} u}{\rho} \frac{\partial \eta}{\partial x} \right] \\ \alpha_3 &= \frac{\operatorname{Re}_s \tilde{\delta}^2}{\rho \mu} \left[ u \frac{\partial u}{\partial \xi} - \frac{1}{\rho} \frac{\partial P}{\partial \xi} + u \frac{\partial \eta}{\partial x} \frac{\partial u}{\partial \eta} + \frac{u^2 \tilde{\delta}}{\rho} \frac{\partial \eta}{\partial x} \right] \\ \alpha_4 &= - \frac{\operatorname{Re}_s \tilde{\delta}^2 u}{\rho \mu}\end{aligned}\quad [3.127]$$

Species Continuity,  $W = C_1$

$$\begin{aligned}\alpha_1 &= \frac{1}{\beta} \left[ \beta' - \operatorname{Re}_s \tilde{\delta} \rho v - \operatorname{Re}_s \tilde{\delta}^2 u \frac{\partial \eta}{\partial x} \right] \\ \alpha_2 &= \frac{\operatorname{Re}_s \tilde{\delta}^2}{\beta} \frac{\partial}{\partial C_1} \left( \frac{\omega_1}{\rho} \right)\end{aligned}$$

$$\alpha_3 = \frac{Re_s \tilde{\delta}^2}{\beta} \left[ \frac{\omega_1}{\rho} - \left( \frac{\partial}{\partial C_1} \left( \frac{\omega_1}{\rho} \right) \right) C_1 \right] \quad [3.128]$$

$$\alpha_4 = \frac{-Re_s \tilde{\delta}^2 u}{\beta}$$

where

$$\beta = \rho u / Sc \quad [3.129]$$

Energy (Enthalpy):  $W = h$

$$\alpha_1 = \frac{C_p}{\rho k} \left[ \frac{\partial}{\partial \eta} \left( \frac{C_p}{\rho k} \right) - \frac{u \tilde{\delta}^2}{2} \frac{\partial \eta}{\partial x} - \frac{\rho v \tilde{\delta}}{2} \right]$$

$$\alpha_2 = 0$$

$$\alpha_3 = \frac{C_p}{\rho k} \left\{ \left( \frac{\rho \mu}{2 Sc Re_s} - \frac{\rho k}{C_p} \right) \left[ \frac{\partial C_1}{\partial \eta} \frac{\partial h_1}{\partial \eta} + \frac{1}{2 Re_s} \frac{\partial}{\partial \eta} \left( \frac{\rho \mu}{Sc} \right) - \frac{\partial}{\partial \eta} \left( \frac{\rho k}{C_p} \right) \right] \right.$$

$$\left. \left[ h_1 \frac{\partial C_1}{\partial \eta} + \left( \frac{\rho \mu}{2 Re_s Sc} - \frac{\rho k}{C_p} \right) \left[ h_1 \frac{\partial^2 C_1}{\partial \eta^2} - \frac{\tilde{\delta}^2}{\rho} \right] \right] \right.$$

$$\left. + \frac{1}{Re_s} \left( \frac{\partial}{\partial \eta} \left( \rho \mu u \right) \frac{\partial u}{\partial \eta} + \rho \mu u \frac{\partial^2 u}{\partial \eta^2} \right) - \right.$$

$$\left. \tilde{\delta}^2 \left[ u^2 \frac{\partial u}{\partial \xi} + \left( u^2 \frac{\partial \eta}{\partial x} + \frac{\rho v u}{\tilde{\delta}} \right) \frac{\partial u}{\partial \eta} + uv \frac{\partial v}{\partial \xi} + \right. \right.$$

$$\left. \left( uv \frac{\partial \eta}{\partial x} + \frac{\rho v^2}{\tilde{\delta}} \right) \frac{\partial v}{\partial \eta} \right]$$

$$\alpha_4 = \frac{-u C_p \tilde{\delta}^2}{2 \rho k}$$

[3.130]

Energy (Temperature):  $W = T$

$$\alpha_1 = \frac{1}{\rho k} \left[ (\rho k)' - \frac{\rho v \tilde{\delta} C_p}{2} + \frac{\rho \mu}{2 Re_s} \left[ C_p \frac{\partial C_1}{\partial \eta} \right. \right.$$

$$\left. - \frac{u C_p \tilde{\delta}^2}{2} \frac{\partial \eta}{\partial x} \right]$$



$$\alpha_2 = \frac{\tilde{\delta}^2}{2\rho k} \left[ \sum (C_{p1} \left(\frac{\omega_1}{\rho}\right) + h_1 \frac{\partial}{\partial T} \left(\frac{\omega_1}{\rho}\right)) \right] T_{s,o}$$

$$\alpha_3 = \frac{\tilde{\delta}^2}{2\rho k} \left[ \sum h_1 \left(\frac{\omega_1}{\rho}\right) + \frac{2u}{\rho} \frac{\partial P}{\partial \xi} + \frac{\tilde{\delta} u^3}{\rho} \frac{\partial \eta}{\partial x} - 2 \frac{E}{\rho} + \frac{2}{Re_s} \left(\frac{\partial u}{\partial \eta}\right)^2 \right] - \alpha_2 T$$

$$\alpha_4 = -u C_p \tilde{\delta}^2 / 2\rho k \quad [3.131]$$

Equations 3.125 through 3.131 can be used to numerically solve the shock layer problem. The species and energy equations can be applied at the stagnation line simply by noting that the tangential velocity,  $u$ , equals zero, thus making  $\alpha_4$  zero along with terms in the energy equation that contain the normal derivative of  $u$ . Chapter V treats these equations further by placing them in finite difference form and describing the numerical solution procedure.

### Summary

The thin shock layer equations have been derived from the general conservation equations in vector-tensor notation. The resulting partial integro-differential equations require two sets of boundary conditions and one set of initial conditions due to their parabolic nature. The boundary conditions are given at the bow shock by the Rankine-Hugoniot relations and at the ablator surface by the surface balances. The initial conditions were derived from the thin shock layer equations by taking the limit as  $x$  approached zero which yields the stagnation line equations (along the line of symmetry). The shock layer equations were then

nondimensionalized, transformed, and quasilinearized for ease of numerical solution.

In order to solve the final set of equations, thermodynamic, transport, kinetic, and radiation properties must be known. The next chapter describes how these properties are obtained and gives the values used in the solution of the equations. In Chapter V the quasilinearized shock layer equations are finite differenced and the method of solution, stability, and results of numerical experimentation are demonstrated.

## REFERENCES

- 3.1 Engel, C. D., D. D. Esch, R. C. Farmer and R. W. Pike, "Formulation, Derivation and Development of the Analysis of Interaction of Ablating Protection Systems and Stagnation Region Heatings", NASA CR-66931, 1970.
- 3.2 Hayes, W. D. and R. F. Probstein, Hypersonic Flow Theory 2nd Ed., Ch. 5, Academic Press, New York (1966).
- 3.3 Probstein, R. F., "Shock Wave and Flow Field Development in Hypersonic Reentry," Am. Rocket Soc. Journal, 31, pp. 184-194 (Feb. 1961).
- 3.4 Edelman, R. K. and D. Fortune, "Mixing and Combustion in the Exhaust Plumes of Rocket Engines Burning RP1 and Liquid Oxygen," TR-631 General Applied Science Laboratories, Inc., Nov. 1966.
- 3.5 Spradley, L. W. and C. D. Engel, "Formulation of a Method for Predicting Coupled Convective and Radiative Heat Transfer About A Blunt Body", NASA CR-61200 (April, 1968).
- 3.6 Ho, H. T. and R. F. Probstein, "The Compressible Viscous Shock Layer in Rarefied Hypersonic Flow", in Rarefied Gas Dynamics, ed. L. Talbot, Academic Press, New York (1961).
- 3.7 Hoshizaki, H. and K. H. Wilson, "Convective and Radiation Heat Transfer During Superorbital Entry," AIAA Jour., 5 (1), 25 (January, 1967).
- 3.8 Engel, C. D., "Ablation and Radiation Coupled Viscous Hypersonic Shock Layers," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 3.9 Spradley, L. W. and C. D. Engel, "The Effects of Shock Layer Radiation and Viscous Coupling in the Total Heating Rate to a Reentering Blunt Body," NASA CR-61224, April, 1968.
- 3.10 Spradley, L. W. and C. D. Engel, "A Computer Program for Predicting Coupled Convective and Radiative Heat Transfer to a Body During Superorbital Reentry," LMSC/HREC A791350, Lockheed Missiles & Space Company, Huntsville, Alabama, (May, 1968).
- 3.11 Hoshizaki, H. and K. H. Wilson, "Viscous, Radiating Shock Layer About a Blunt Body," AIAA J. 3, No. 9, Sept. 1965, pp. 1614-1622.

- 3.12 Bellman, R. E. and R. E. Kulaba, Quasilinearization and Non-linear Boundary-Value Problems, American Elsevier, New York, 1965.
- 3.13 Blottner, F. G., "Finite Difference Methods of Solution of Boundary Layer Equations," AIAA Jour., 8, (2), (February, 1970), pp. 193-205.

CHAPTER IV  
TRANSPORT, THERMODYNAMIC, RADIATION, AND  
KINETIC PROPERTIES

Introduction

In the previous chapter the conservation equations were developed into a form that will be finite differenced and placed into numerical solution form in the following chapter. However, first the coefficients in these equations contain not only the dependent variables, but also transport and thermodynamic properties along with generation terms of mass and energy which must be evaluated. The production or generation of mass is described by the chemical kinetics of the multicomponent reacting system. The generation terms for energy include that due to chemical reaction and that due to radiative heat flux.

It is important to have an accurate description of these terms in order to have a valid mathematical model. In general, the transport and thermodynamic properties and the chemical and radiation generation terms are functions of temperature, pressure, and chemical composition. Theoretical values of the transport and thermodynamic properties for pure components are curve fit and used in correlation formulas to predict mixture values. These curve fits allow a reasonably fast method of computation in contrast to the more rigorous theoretical computations which are extremely time consuming. This chapter is devoted to presenting the form to be

used in the numerical computation of the transport, thermodynamic, kinetic, and radiation properties of a multicomponent system. The numerical values for the curve fits and kinetic equations are given in tabular form for the ablating-flow field system under consideration.

### Chemical Kinetics

The generation of mass in the species equation and the generation of heat in the temperature form of the energy equation are given by the reaction kinetics of the system. The reaction rate is a function of density, temperature, and mixture composition.

$$\omega_i = \omega_i(\rho, T, C_i) \quad [4.1]$$

The rate is given by the Law of Mass Action as

$$\omega_i = \sum_{j=1}^M \nu_{ij} \frac{k_{rj}}{\rho M_i} \prod_{k=1}^n \left( \frac{C_k}{\rho M_k} \right)^{\nu_{kj}} - \sum_{j=1}^M \nu'_{ij} \frac{k_{fj}}{\rho M_i} \prod_{k=1}^n \left( \frac{C_k}{\rho M_k} \right)^{\nu'_{kj}} \quad [4.2]$$

where  $n$  is the number of species present and  $M$  is the total number of chemical reactions. The coefficients  $k_{fj}$  and  $k_{rj}$  are the forward and reverse rate constants respectively for the  $j$ th reaction and are taken as a function of temperature only. The forward rate constant has the form

$$k_{fj} = A_{fj} T^{S_{fj}} \exp \left[ -E_{fj} / RT \right] \quad [4.3]$$

where  $A_{fj}$  is referred to as the frequency factor and  $E_{fj}$  is known as the activation energy.  $A_{fj}$ ,  $S_{fj}$ , and  $E_{fj}$  are constants for each reaction. The reverse rate constants can be determined from the equilibrium constant,  $K_{eq}$ , for elementary reactions using the

following relation

$$k_{rj} = k_{fj}/K_{eqj} \quad [4.4]$$

The stoichiometric coefficients  $v''_{ij}$  and  $v'_{ij}$  are defined by the following expression for the  $j$ th chemical reaction equation

$$\sum_{i=1}^n v'_{ij} R_i \rightarrow \sum_{i=1}^n v''_{ij} P_i \quad j = 1, \dots, M \quad [4.5]$$

where  $R_i$  and  $P_i$  are the chemical symbols of the reactants and products respectively.

Also appearing in the quasilinearized species equation, Equation 3.128 is the reaction rate linearized with respect to species "i". The quasilinearized form of the energy equation, Equation 3.131, requires the reaction rate linearized with respect to temperature. These quantities can be derived from Equation 4.2 and are expressed as follows:

$$\begin{aligned} \frac{\partial}{\partial C_i} \left( \frac{\omega_i}{\rho} \right) = & \sum_{j=1}^M (v''_{ij})^2 k_{rj} C_i^{(v''_{ij}-1)} \prod_{\substack{k=1 \\ k \neq i}}^n \left( \rho \frac{C_k}{M_k} \right)^{v''_{kj}} \\ & - \sum_{j=1}^M (v'_{ij})^2 k_{fj} C_i^{(v'_{ij}-1)} \prod_{\substack{k=1 \\ k \neq i}}^n \left( \rho \frac{C_k}{M_k} \right)^{v'_{kj}} \end{aligned} \quad [4.6]$$

and

$$\begin{aligned} \frac{\partial}{\partial T} \left( \frac{\omega_i}{\rho} \right) = & \sum_{j=1}^M v''_{ij} \left( \frac{s_{rj}}{T} + \frac{E_{rj}}{RT^2} \right) k_{rj} \prod_{k=1}^n \left( \rho \frac{C_k}{M_k} \right)^{v''_{kj}} \\ & - \sum_{j=1}^M v'_{ij} \left( \frac{s_{fj}}{T} + \frac{E_{fj}}{RT^2} \right) k_{fj} \prod_{k=1}^n \left( \rho \frac{C_k}{M_k} \right)^{v'_{kj}} \end{aligned} \quad [4.7]$$

These partial derivatives can be analytically evaluated if the chemical kinetics of the system are known and used in the numerical solution of the species and energy equations for the shock layer.

The problem now rests in the specification of the chemical kinetics that describe the shock layer. There are innumerable possible chemical reactions that can take place among the air and ablation product species that are present in the shock layer adjacent to an ablating blunt body. However, to be able to solve the transport equations, a set of important chemical reactions must be selected that accurately describe the chemistry of the reactions of ablation products and air species. This information is stored in the form of a computer implemented data management file which presently contains several thousand reactions.

It was necessary to examine this extensive set of reactions to arrive at a listing of most probable reactions. These probable reactions were selected from the kinetic data management system. This set represented a total of 44 reactions involving 28 species. Currently, it is not possible to consider a solution of the transport equations with this many species and reactions. Therefore, this preliminary list was then re-examined for key reactions to represent the chemical system. The selection was based on the species anticipated to have the largest compositions and presumably, therefore, dominate the energy absorption of the system. These 16 key reactions among 19 species are given in Table 4.1 along with the recommended values of the rate constants. The forward rate constant listed was the best experimental value or was computed theoretically. The reverse rate constant was computed using Equation 4.4



Table 4.1  
Shock Layer Reactions with Kinetic Coefficients

$$k_f = \exp [A + B \ln T - C \times 1000/T]$$

$$K_{eq}^{(a)} = \exp [D - E \times 1000/T]$$

Reaction	A	$\gamma$	C	D	E	Source
1. $N_2 + M \xrightarrow{+} 2N + M^*$	48.35	-1.5	112.5	3.07	114.3	5.7
2. $O + M \xrightarrow{+} O^+ + e^- + M$	28.66	0.5	156.5	-5.33	165.9	5.8
3. $N + M \xrightarrow{+} N^+ + e^- + M$	28.70	0.5	166.5	-3.69	175.7	5.8
4. $H_2 + M \xrightarrow{+} 2H + M$	42.73	-0.8	51.5	0.94	51.2	5.8
5. $CN + M \xrightarrow{+} C + N + M$	46.84	-1.0	65.9	2.06	93.8	5.9
6. $C_2H + H \xrightarrow{+} C_2 + H_2$	26.83	0.5	17.5	0.33	14.6	5.10
7. $CO + M \xrightarrow{+} C + O + M$	45.89	-1.0	129.0	3.58	127.0	5.9
8. $CO + N \xrightarrow{+} CN + O$	30.40	0.5	35.0	1.55	33.4	5.10
9. $C_2 + M \xrightarrow{+} 2C + M$	31.84	0.5	70.5	1.28	70.7	5.11

\* M is a third body

Table 4.1 (Cont'd)

Reaction	A	B	C	D	E	Source
10. $C + M \xrightarrow{\pm} C^+ + e^- + M$	32.97	0.5	130.7	-5.11	136.3	5.11
11. $HCN + M \xrightarrow{\pm} CN + H + M$	31.84	0.5	60.0	3.97	72.4	(b)
12. $C_2H + M \xrightarrow{\pm} C_2 + H + M$	31.84	0.5	70.0	1.00	64.0	(b)
13. $C_2H_2 + M \xrightarrow{\pm} C_2H + H + M$	31.84	0.5	58.5	2.28	54.8	(b)
14. $C_3 + M \xrightarrow{\pm} C_2 + C + M$	31.84	0.5	95.0	4.14	78.7	(b)
15. $C_3H + M \xrightarrow{\pm} C_2H + C + M$	31.84	0.5	41.5	5.08	80.9	(b)
16. $C_4H + M \xrightarrow{\pm} C_3H + C + M$	31.84	0.5	36.5	1.02	67.0	(b)

(a)  $K_{eq}$ , equilibrium constant, computed using the method of free energy minimization

(b) Kinetic data estimated employing Ref. 5.12.

assuming the equilibrium constant could be expressed in the form,  $K_{eq} = \exp [A+B/T]$  for the temperature range from 3000°K to 15000°K. The system shown in Table 4.1 includes 16 chemical reactions with 19 species.

In Table 4.2 the previously omitted ablation product and combustion reactions are given. Shown in Table 4.3 are the additional air and hydrogen combustion reactions. Some kinetic data is available for these; but, as a matter of convenience, it has not been included.

#### Transport Properties

Generally, investigators in the area of shock layer solutions have resorted to the classical Chapman-Enskog kinetic theory relations for estimation of the required transport properties. The modification of these relationships to account for polyatomic, reacting mixtures results in very cumbersome equations. In some cases there are simplifications which can be applied without substantial loss in accuracy. At this point, it becomes desirable to optimize between accuracy and computation time. A wide variety of methods for estimating these properties has been developed in just this manner. In this section, a formulation is provided from which an optimum method to accurately compute high temperature transport properties with reasonable computational convenience is presented. These properties, along with comparisons with other investigators, are developed in more detail by Esch et al. (Ref. 4.1). The following discussion is based on Esch's work.

Viscosity: The relationship employed for the prediction of mixture viscosity is the commonly used Buddenberg-Wilke correlation (Ref. 4.2).

Table 4.2. Additional Ablation Product and Combustion Reactions

CH<sub>4</sub>, CH<sub>3</sub>, CH<sub>2</sub>, CH Reactions

1.  $\text{CH}_4 \rightleftharpoons \text{CH}_3 + \text{H}$
2.  $\text{CH}_4 \rightleftharpoons \text{CH}_2 + \text{H}_2$
3.  $\text{CH}_3 \rightleftharpoons \text{CH}_2 + \text{H}$
4.  $\text{CH}_2 \rightleftharpoons \text{CH} + \text{H}$
5.  $\text{CH}_2 + \text{H} \rightleftharpoons \text{CH} + \text{H}_2$
6.  $\text{CH} \rightleftharpoons \text{C} + \text{H}$
7.  $\text{C} + \text{H}_2 \rightleftharpoons \text{CH} + \text{H}$
8.  $\text{CH}_2 + \text{O} \rightleftharpoons \text{CO} + \text{H}_2$

CN Reactions

1.  $2\text{CN} \rightleftharpoons \text{C}_2 + \text{N}_2$
2.  $\text{CN} + \text{O} \rightleftharpoons \text{N} + \text{CO}$

C<sub>2</sub>H<sub>2</sub>, C<sub>2</sub>H Reactions

1.  $\text{C}_2\text{H}_2 + \text{H} \rightleftharpoons \text{C}_2\text{H} + \text{H}_2$
2.  $\text{C}_2\text{H}_2 + \text{O} \rightleftharpoons \text{CH}_2 + \text{CO}$
3.  $\text{C}_2\text{H}_2 + \text{OH} \rightleftharpoons \text{C}_2\text{H} + \text{H}_2\text{O}$
4.  $\text{C}_2\text{H} + \text{H} \rightleftharpoons \text{C}_2 + \text{H}_2$
5.  $\text{C}_2\text{H} + \text{O} \rightleftharpoons \text{CH} + \text{CO}$

Other Reactions

1.  $\text{C}_4\text{H} \rightleftharpoons \text{C}_3\text{H} + \text{C}$
2.  $\text{C}_3\text{H} \rightleftharpoons \text{C}_2\text{H} + \text{C}$
3.  $\text{C}_2 + \text{H} \rightleftharpoons \text{C} + \text{CH}$

Table 4.3. Additional Air and Hydrogen Combustion Reactions

Air Reactions

1.  $\text{O}_2 + \text{M} \rightleftharpoons 2\text{O} + \text{M}$
2.  $\text{N}_2 + \text{M} \rightleftharpoons \text{N}^+ + \text{e}^- + \text{M}$
3.  $\text{NO} + \text{M} \rightleftharpoons \text{N} + \text{O} + \text{M}$
4.  $\text{NO} + \text{O} \rightleftharpoons \text{O}_2 + \text{N}$
5.  $\text{N}_2 + \text{O} \rightleftharpoons \text{NO} + \text{N}$
6.  $\text{N} + \text{O} \rightleftharpoons \text{NO}^+ + \text{e}^-$

Hydrogen Combustion Reactions

1.  $\text{H} + \text{O}_2 \rightleftharpoons \text{OH} + \text{O}$
2.  $\text{O} + \text{H}_2 \rightleftharpoons \text{OH} + \text{H}$
3.  $\text{OH} + \text{H}_2 \rightleftharpoons \text{H}_2\text{O} + \text{H}$
4.  $2\text{OH} \rightleftharpoons \text{H}_2\text{O} + \text{O}$
5.  $\text{H} + \text{OH} + \text{M} \rightleftharpoons \text{H}_2\text{O} + \text{M}$
6.  $\text{H} + \text{O} + \text{M} \rightleftharpoons \text{OH} + \text{M}$

The mixture viscosity is calculated from the pure component viscosities with the relation

$$\mu = \frac{\sum_{i=1}^n Y_i \mu_i}{\sum_{j=1}^n Y_j \phi_{1j}} \quad [4.8]$$

The coefficients  $\phi_{1j}$  are a function of the pure component viscosities and molecular weight ratios.

$$\phi_{1j} = \frac{1}{\sqrt{8}} \left( 1 + \frac{M_i}{M_j} \right)^{-1/2} \left[ 1 + \left( \frac{\mu_i}{\mu_j} \right)^{1/2} \left( \frac{M_i}{M_j} \right)^{1/4} \right]^2 \quad [4.9]$$

where the pure component viscosity in this analysis is calculated by the second order polynomial curve fit

$$\mu_i = a_i + b_i T + c_i T^2 \quad [4.10]$$

The above curve fit is obtained using the data from theoretical predictions (Refs. 4.3-4.7). This curve fit can then be used for computer implementation. A summary for the empirical constants required for Equation 4.10 is given in Table 4.4.

Thermal Conductivity: Mixture thermal conductivity is calculated in the same manner as mixture viscosity

$$k = \frac{\sum_{i=1}^n Y_i k_i}{\sum_{j=1}^n Y_j \phi_{1j}} \quad [4.11]$$

where  $k$  is the frozen mixture thermal conductivity and  $\phi_{1j}$  is defined by Equation 4.9. As with the viscosity, the pure component thermal conductivities were obtained from the theoretical predictions of Refs. 4.3-4.7. Examination of the data revealed that a

Table 4.4  
Empirical Constants for Viscosity Correlation

$$\mu_1 = a_1 + b_1 T + c_1 T^2 \frac{\text{lbm}}{\text{ft-sec}} \quad (\text{Ref. 4.1})$$

Species	$a \times 10^5$	$b \times 10^7$	$c \times 10^{12}$	Temperature Range ( $^{\circ}\text{K}$ )
$\text{O}_2$	1.693	0.1496	-0.2276	2,000-10,000
$\text{N}_2$	0.970	0.1613	-0.1916	2,000-10,000
O	1.519	0.1875	-0.2228	2,000-10,000
N	0.253	0.2206	-0.3737	2,000-10,000
$\text{O}^+$	0.0	0.0500	-0.1000	8,000-15,000
$\text{N}^+$	0.0	0.0500	-0.1000	8,000-15,000
$\text{e}^-$	0.0	0.0500	-0.1000	8,000-15,000
C	1.997	0.1772	-0.3378	5,000-10,000
H	0.294	0.0889	-0.0811	4,000-10,000
$\text{H}_2$	-0.079	0.0791	-0.0886	4,000-10,000
CO	2.404	0.1363	-0.2184	4,000- 9,000
$\text{C}_3$	2.019	0.1179	-0.1655	1,000- 5,000
CN	2.404	0.1363	-0.2184	4,000- 9,000
$\text{C}_2\text{H}$	2.404	0.1363	-0.2184	4,000- 9,000
$\text{C}_2\text{H}_2$	1.396	0.0842	-0.6939	1,000- 5,000
$\text{C}_3\text{H}$	2.019	0.1179	-0.1655	1,000- 5,000
$\text{C}_4\text{H}$	2.019	0.1179	-0.1655	1,000- 5,000
HCN	1.378	0.0965	-0.0948	1,000- 5,000
$\text{C}_2$	1.931	0.1393	-0.2575	4,000- 9,000
$\text{C}^+$	0.0	0.0500	-0.1000	8,000-15,000

linear fit would be satisfactory for accurate correlation.

$$k_1 = a_1 + b_1 T \quad [4.12]$$

A summary of the coefficients required for Equation 4.12 is presented in Table 4.5.

Binary Diffusion Coefficients: In view of the multiplicity of binary interactions required, it was decided that the following Chapman-Enskog equation for the prediction of binary diffusion coefficients would be used (Ref. 4.8).

$$D_{1j} = \frac{28.28 \times 10^{-7} T^{3/2}}{P \sigma_{1j}^2 \Omega_{1j}^{(1,1)}} \left( \frac{M_1 + M_j}{2 M_1 M_j} \right)^{1/2} \frac{\text{ft}^2}{\text{sec}} \quad [4.13]$$

where  $\sigma_{1j} = (\sigma_1 + \sigma_j)/2$ , the quantities  $\sigma_1$  and  $\sigma_j$  being the collision diameters of the interacting species. The quantity,  $\Omega_{1j}^{(1,1)}$ , is the Leonard-Jones collision integral for diffusion as determined by the following empirical equation

$$\Omega_{1j}^{(1,1)} = 1.061 (T_{1j}^*)^{-1.56} \quad [4.14]$$

These constants were obtained from a curve fit of the Leonard-Jones potential as reported by Hirschfelder in Ref. 4.8 for  $10 \leq T_{1j}^* \leq 1000$  which includes all species and temperatures considered in the current study. The quantity  $T_{1j}^*$  is computed as

$$T_{1j}^* = \frac{T}{\epsilon_{1j}/k} \quad [4.15]$$

where  $\epsilon_{1j} = \sqrt{\epsilon_1 \epsilon_j}$ ,  $\epsilon_1$  and  $\epsilon_j$  being the characteristic interaction energies of species 1 and j.

Table 4.5

Empirical Constants for Thermal Conductivity Correlation

$$k_1 = a + bT \text{ (BTU/ft-sec-}^\circ\text{R)} \quad (\text{Ref. 4.1})$$

Species	$a \times 10^5$	$b \times 10^8$	Temperature Range ( $^\circ\text{K}$ )
$\text{O}_2$	1.019	0.4901	2,000-10,000
$\text{N}_2$	0.654	0.6457	2,000-10,000
O	1.250	0.7092	2,000-10,000
N	1.281	0.8593	2,000-10,000
$\text{O}^+$	26.0	0.0	8,000-15,000
$\text{N}^+$	26.0	0.0	8,000-15,000
$\text{e}^-$	26.0	0.0	8,000-15,000
C	2.506	0.7479	—5,000-10,000
H	2.496	5.129	4,000-10,000
$\text{H}_2$	3.211	5.344	4,000-10,000
CO	0.859	0.6233	1,000- 5,000
$\text{C}_3$	0.630	0.5804	1,000- 5,000
CN	0.859	0.6233	2,000-10,000
$\text{C}_2\text{H}$	1.126	0.7439	1,000- 5,000
$\text{C}_2\text{H}_2$	1.126	0.7439	1,000- 5,000
$\text{C}_3\text{H}$	0.630	0.5804	1,000- 5,000
$\text{C}_4\text{H}$	0.630	0.5804	1,000- 5,000
HCN	0.486	0.8714	1,000- 5,000
$\text{C}_2$	0.859	0.6233	1,000- 5,000
$\text{C}^+$	26.0	0.0	8,000-15,000



Values for the collision parameters,  $\sigma_1$  and  $\epsilon_{1/k}$  for  $O_2$ ,  $N_2$ , C, H,  $H_2$ , CO, CN,  $C_2H_2$ , HCN, and  $C_2$ , were obtained from Svehla (Ref. 4.6). Since no data was available for  $C_3$ ,  $C_2H$ ,  $C_3H$ , and  $C_4H$ , it was necessary to develop correlations based upon similar species. Esch (Ref. 4.1) developed the following correlation structures for the prediction of the collision parameters of the above light hydrocarbon species.

Collision Diameter:

$$\sigma_1 = 2.69 + 0.0514 M_1 \quad [4.16]$$

Interaction Energy ( $C_n$  molecules):

$$\epsilon_{1/k} = -17.0 + 4.02 M_1 \quad [4.17]$$

Interaction Energy ( $C_n H_n$  molecules):

$$\epsilon_{1/k} = -105. + 12.4 M_1 \quad [4.18]$$

The collision parameters for O, N,  $O^+$ ,  $N^+$ , and  $e^-$  were estimated from the theoretically determined binary diffusion coefficients reported by Yun, Weissman, and Mason (Ref. 4.3) and by Yos (Ref. 4.7). The behavior of  $C^+$  was then assumed to be similar to that of  $N^+$ . A summary of the collision parameters employed in the present study is given in Table 4.6.

#### Thermodynamic Properties

The thermodynamic properties of heat capacity and enthalpy appear in the energy equation of the thin shock layer equations. Data for thermodynamic properties was obtained from References 4.9-4.13. In Reference 4.13 the appropriate species for the shock layer are presented with polynomial fits for a range of temperature of 300°K

Table 4.6.  
Collision Parameters Employed in the Current Study

Species	$M_1$	$\sigma_1$	$\epsilon_1/k$
O <sub>2</sub>	32.000	3.467	106.7
N <sub>2</sub>	28.016	3.798	71.4
O	16.000	7.990	106.7
N	14.008	7.940	71.4
O <sup>+</sup>	16.000	14.220	106.7
N <sup>+</sup>	14.008	14.930	71.4
e <sup>-</sup>	$5.486 \times 10^{-4}$	14.930	71.4
C	12.001	3.385	30.6
H	1.008	2.708	37.0
H <sub>2</sub>	2.016	2.827	59.7
C <sub>3</sub>	36.033	4.450	128.0
CN	26.019	3.856	75.0
C <sub>2</sub> H	25.030	3.880	205.0
C <sub>2</sub> H <sub>2</sub>	26.038	4.033	231.8
C <sub>3</sub> H	37.041	4.600	356.0
C <sub>4</sub> H	49.052	5.210	504.0
HCN	27.027	3.630	469.1
C <sub>2</sub>	24.022	3.913	78.8
C <sup>+</sup>	12.011	15.000	30.6

to 15000°K. The heat capacity of species "i" is expressed in terms of temperature in the following form

$$\frac{C_{pi}}{R} = a_1 + a_2 T + a_3 T^2 + a_4 T^3 + a_5 T^4 \quad [4.19]$$

The coefficients  $a_1$  through  $a_5$  are given in Appendix C for the shock layer chemistry constituents. In addition to heat capacity data, enthalpy is required. The enthalpy of species "i" is obtained from the relation

$$h_i = \int_{298}^T C_{pi} dT + h_{i,298}^{\circ} \quad [4.20]$$

where  $h_{i,298}^{\circ}$  is the enthalpy of formation of species "i". The polynomial form for enthalpy follows directly from the integration of Equation 4.20 by substituting in the heat capacity from Equation 4.19.

$$\frac{h_i}{RT} = a_1 + \frac{a_2 T}{2} + \frac{a_3 T^2}{3} + \frac{a_4 T^3}{4} + \frac{a_5 T^4}{5} + \frac{a_6}{T} \quad [4.21]$$

The mixture values of heat capacity and enthalpy have the simple relationship

$$C_p = \sum_{i=1}^n Y_i C_{pi} \quad [4.22]$$

and

$$h = \sum_{i=1}^n Y_i h_i \quad [4.23]$$

where  $C_p$  is the frozen total heat capacity and  $h$  is the total static enthalpy.

### Radiation Properties

Radiating shock layer analyses can be categorized by means of the radiation transport model employed, i.e. whether the gas is treated as transparent, gray, or nongray. In a transparent gas, fluid elements are assumed to emit locally but not absorb radiative energy. Both gray and nongray gas fluid elements locally emit and absorb some of the incident radiative intensity that originates from the other fluid elements in the shock layer. However, the effect of this self-absorption differs markedly between a gray gas and a nongray gas. The gas absorption coefficient is independent of wavelength for the gray gas assumption, whereas the nongray gas model includes the wavelength variation of the gas absorption coefficient, either approximately or in detail. The nongray model is the most realistic treatment for atmospheric reentry although it is also by far the most time-consuming computationally. In addition, each of the three categories can be subdivided according to whether or not the gas dynamic flow field is assumed to be coupled with the radiative energy transport. If no coupling is assumed, the computation of local flow field variables neglects local emission or absorption of radiation, and the radiative energy flux is subsequently calculated by integration over the resulting (nonradiating) flow field density and temperature profiles. The more realistic case is that of radiative gas dynamic coupling, where the fact that the flow field variables depend on radiative emission is accounted for.

It is not intended to present a complete mathematical formulation of the radiative gas dynamics in this section since a formulation of this type is given in detail by Engel (Ref. 4.14).

The following discussion will be more qualitative and will attempt to justify the radiation model used in this analysis by comparing results with other investigators. The present analysis is a nongray radiation model which is coupled to the flow field variables.

The radiative flux divergence ( $\partial q_R / \partial y$ ) appearing in Equation 3.35 is defined as follows (Ref. 4.15):

$$\frac{\partial q_R}{\partial y} = \int_0^{\infty} \int_0^{4\pi} \alpha(y, \nu) [B(y, \nu) - I(y, \nu, \Omega)] d\Omega d\nu \quad [4.24]$$

where  $q_R$  = the radiative flux in the normal direction from the body,

$\alpha(y, \nu)$  = volumetric absorption coefficient,

$B(y, \nu)$  = Plankian radiation intensity,

$I(y, \nu, \Omega)$  = radiation intensity,

$d\Omega$  = solid angle about the unit vector  $\bar{\Omega}$ ,

$\nu$  = frequency.

With given temperature and species profiles, Equation 4.24 can be evaluated theoretically. However, due to the discontinuous nature of the absorption coefficients, numerical integration is a formidable task. Typical absorption coefficient distributions are shown in Figure 4.1 to illustrate the extent of this problem (Ref. 4.16).

To overcome the numerical difficulties associated with the integration of such distributions, the frequency range is sub-divided into regions (bands) in which the discontinuous variations are averaged. Continuum radiation bands are used to represent wide regions of continuous radiation while line bands are used to model the effects of the various discontinuous (line) contributions. The use of more bands obviously leads to a more accurate representation of

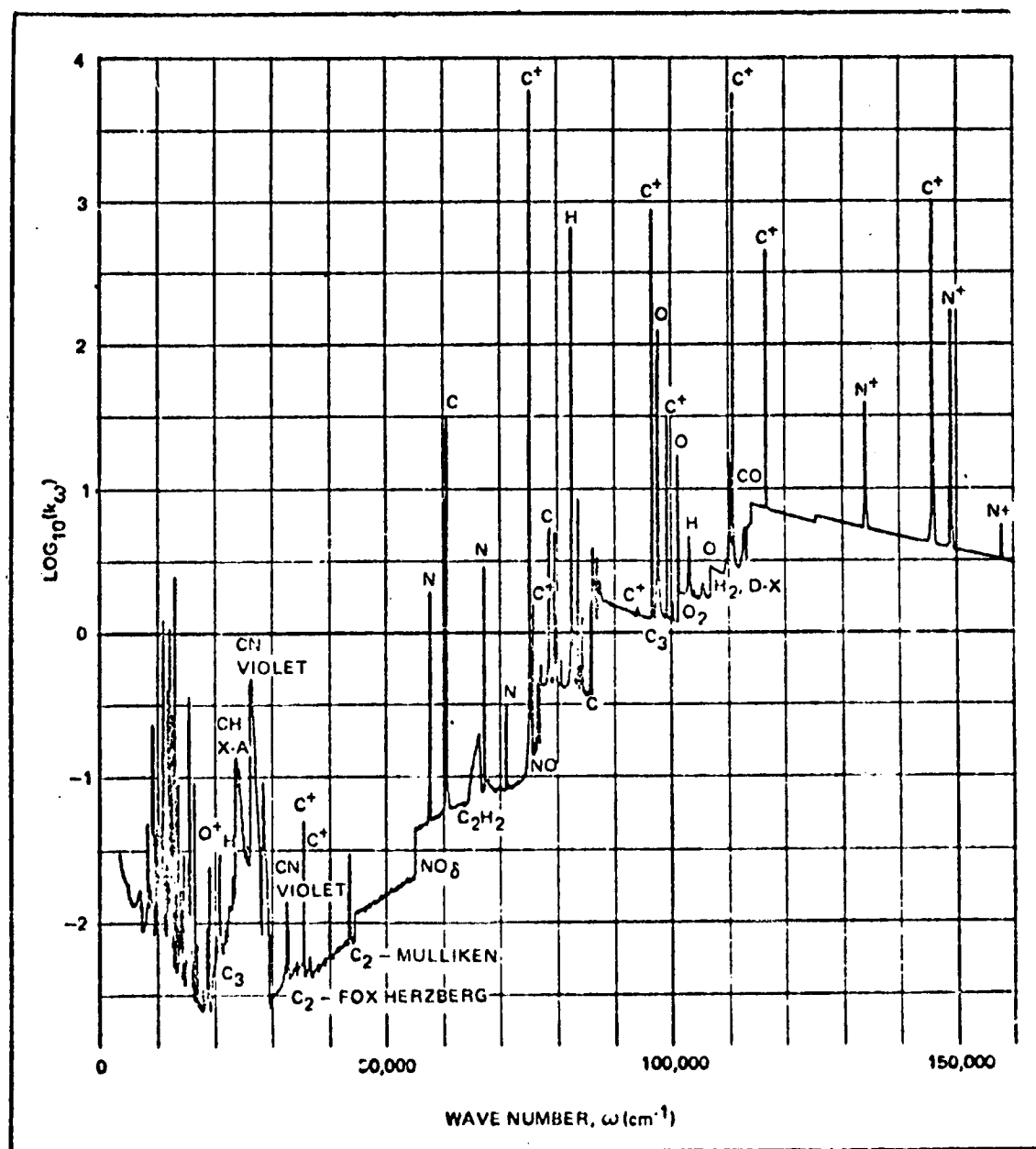


Figure 4.1. Spectral Absorption Coefficient of Air-Carbon Phenolic Ablation Products Mixture for  $P = 1$  and  $T = 10,000^\circ\text{K}$  (Ref. 4.19).

the radiative process. However, in developing radiation models for computer implementation, some compromise must be reached between computer time and the number of bands.

The radiation model used in the present work is a revision (Engel, Ref. 4.14) of a coupled line and continuum model originally developed by Wilson (Ref. 4.15). This model provides a useful tool for evaluating the radiation flux and the radiative flux divergence across a slab of gas containing both air and ablation species.

The existing program (LRAD) contains twelve continuum frequency bands and nine line bands. A comparison with a more detailed model (RATRAP) in Reference 4.17, demonstrated that the existing analysis predicted total heat flux values to within 5%. The species considered in the model used in the current study are shown in Table 4.7.

The radiation model was included in reduced form (air species only) in a program called VISRADI which was developed by Spradley and Engel (Ref. 4.18). Figure 4.2 presents a comparison of dimensionless radiative heating rates as a function of free stream velocity by several investigators. For this no mass injection case the comparison is quite reasonable. All investigators reported on this figure have line and continuum radiation calculations for air at the stagnation line. Although this case does not contain the effects of ablation species, it is a standard for comparison of computational techniques.

#### Summary

The terms in the thin shock layer equations have now been completely specified. This chapter has presented the values to be

Table 4.7  
Species Considered in LRAD Radiation Model

Continuum and Lines	Continuum
C	CO
H	C <sub>2</sub>
N	C <sub>3</sub>
O	O <sub>2</sub>
	H <sub>2</sub>
	N <sub>2</sub>
	C <sub>2</sub> H
	e <sup>-</sup>



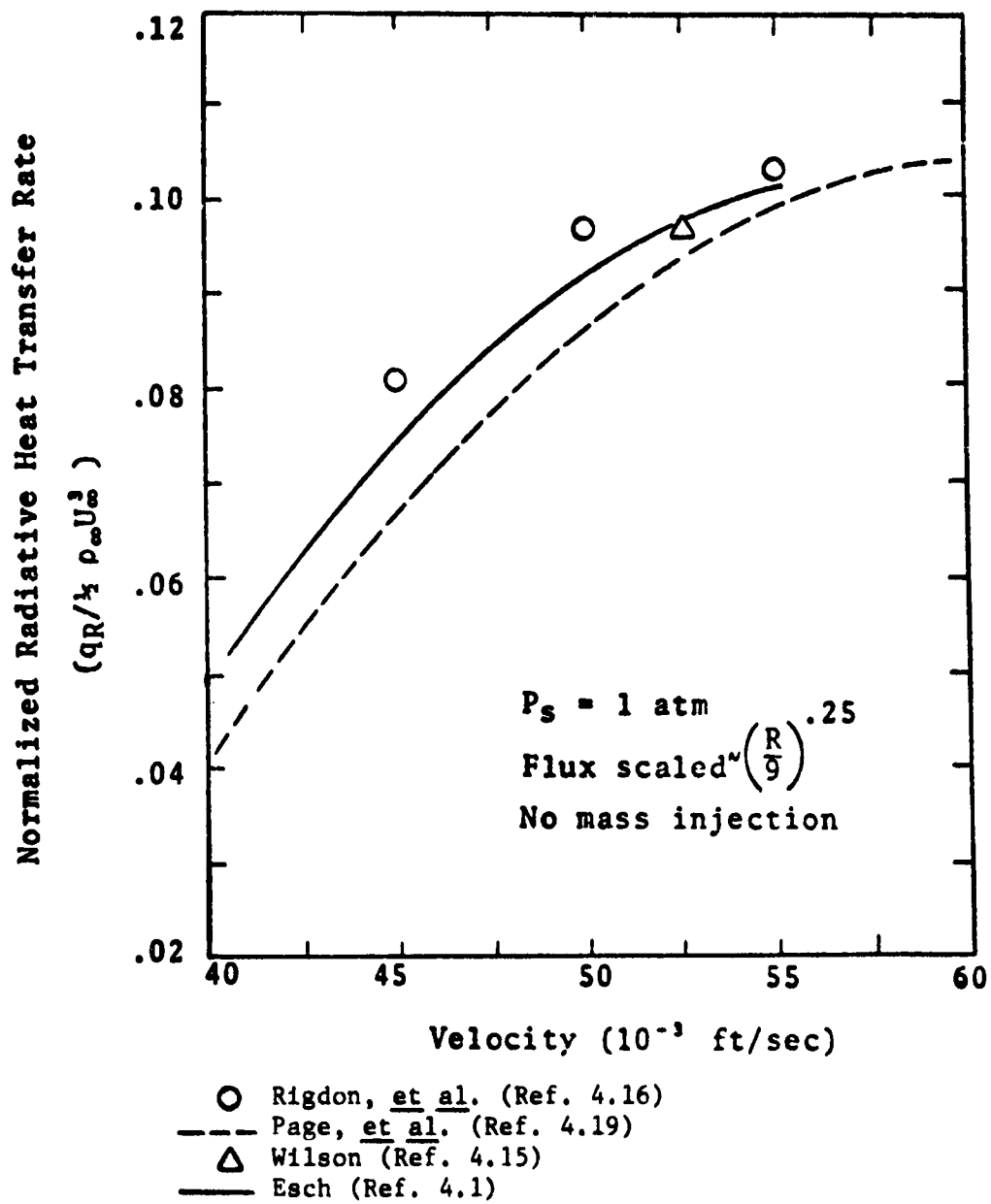


Figure 4.2. Comparison of Dimensionless Radiative Heating Rates.

used for the transport properties, viscosity, thermal conductivity, and binary diffusion coefficients; the thermodynamic properties, heat capacity and enthalpy; and the production terms, chemical production and radiative energy production. With these properties known, it is now possible to solve the complete simultaneous set of shock layer equations along with their initial and boundary conditions. In Chapter V these equations, as given in linearized form in Chapter III, are finite differenced, and they are placed in numerical solution form.

## REFERENCES

- 4.1 Esch, D. D., "Stagnation Region Heating of a Phenolic-Nylon Ablator During Return from Planetary Missions," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 4.2 Buddenberg, J. W. and C. R. Wilke, Ind. Eng. Chem., 41, pp. 1345-1347 (1949).
- 4.3 Yun, Kwank-Sik, S. Weissman, and E. A. Mason, "High Temperature Transport Properties of Dissociating Nitrogen and Dissociating Oxygen," Phys. Fluids 5, 672 (1962).
- 4.4 Lasher, L. E., "The Effect of Carbon Gas Injection (Graphite Ablation) on the Stagnation Point Compressible Boundary Layer," Masters thesis, Cornell University, Ithaca, New York (1965).
- 4.5 Valderslice, J. F., et al., "High-Temperature Transport Properties of Dissociating Hydrogen," Phys. Fluids 5, 155 (1962).
- 4.6 Svahla, R. S., "Estimated Viscosities and Thermal Conductivities of Gases at High Temperatures," NASA TR R-132 (1962).
- 4.7 Yos, J. M., "Transport Properties of Nitrogen, Hydrogen, Oxygen, and Air to 30,000°K," AVCO RAD Technical Memorandum 63-7, (1963) as cited by Hoshizaki, op cit.
- 4.8 Hirschfelder, J. O., C. F. Curtiss, and R. B. Bird, Molecular Theory of Gases and Liquids, 2nd printing, corrected with notes added, John Wiley and Sons, Inc., New York (1964).
- 4.9 Browne, W. G., "Thermodynamic Properties of Some Atoms and Atomic Ions," General Electric Company, MSVD Engr. Physics Tech. Memo. No. 2 (May 14, 1962).
- 4.10 Browne, W. G., "Thermodynamic Properties of Some Diatoms and Diatomic Ions," General Electric Co. MSVD Engr. Physics Tech. Memo. 8, (May 14, 1962).
- 4.11 Browne, W. G., "Thermodynamic Properties of the Species CN, C<sub>2</sub>, C<sub>3</sub>, C<sub>2</sub>N<sub>2</sub>, and C<sup>-</sup>," General Electric MSVD Engr. Physics Tech. Memo., No. 9, (May 14, 1962).
- 4.12 Browne, W. G., "Thermodynamic Properties of Some Ablation Products from Plastic Heat Shields in Air," General Electric Co. MSVD Engr. Physics Tech. Memo. No. 11, (March 15, 1964).

- 4.13 Esch, D. D., A. Siripong, and R. W. Pike, "Thermodynamic Properties in Polynomial Form for Carbon, Hydrogen, Nitrogen, and Oxygen Systems from 300°K to 15,000°K," NASA RFL-TR-70-3, 1970.
- 4.14 Engel, C. D., "Ablation and Radiation Coupled Viscous Hypersonic Shock Layers," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 4.15 Wilson, K. H., "Stagnation Point Analysis of Coupled Viscous-Radiating Flow with Massive Blowing," LMSC-687209, April 1969.
- 4.16 Rigdon, W. S., R. B. Dirling, and M. Thomas, "Stagnation Point Heat Transfer During Hypervelocity Atmospheric Entry," NASA CR-1462, 1970.
- 4.17 Wilson, K. H. and H. Hoshizaki, "Effect of Ablation Product Absorption and Line Transitions on Shock Layer Radiative Transport," NASA CR-1264, 1969.
- 4.18 Spradley, L. W. and C. D. Engel, "A Computer Program for Predicting Coupled Convective and Radiative Heat Transfer to a Body During Superorbital Reentry," LMSC/HREC A791350, Lockheed, Missiles & Space Company, Huntsville, Alaba, May 1968.
- 4.19 Page, W. A., et al., "Radiative Transport in Inviscid Non-adiabatic Stagnation-Region Shock Layers," AIAA Paper No. 68-784, AIAA 3rd Thermophysics Conference, June 1968.

CHAPTER V  
NUMERICAL IMPLEMENTATION OF THE  
SHOCK LAYER EQUATIONS

Introduction

In Chapter III the thin shock layer equations were developed into a general quasilinearized form suitable for numerical solution. The coefficients ( $\alpha$ 's) in these equations are functions of the local temperature, composition, and pressure. The transport, thermodynamic, kinetic, and radiative properties of the system required for complete description of these coefficients were discussed in the previous chapter. The general quasilinear equations are a set of parabolic partial differential equations which can be solved by an appropriate finite difference technique. The numerical finite differenced form is presented in this chapter along with a discussion of stability, convergence, and general computational experience. The numerical technique uncouples the finite differenced equations and results in a tridiagonal matrix which can be solved efficiently using the Thomas Algorithm.

Detailed results are presented in the following chapter using the numerical scheme described in this chapter for heating rates to an ablation protection system typical of reentry from planetary missions.

### Coupled Ablation Flow Field Analysis

A simplified flow diagram of the overall flow field-ablation analysis, as developed in this study, is given in Figure 5.1. The analysis is initiated at the stagnation line or line of symmetry on an axisymmetric body of arbitrary shape. The major input data consist of free-stream velocity ( $U_\infty$ ), free-stream density ( $\rho_\infty$ ) which is indicative of altitude, body radius ( $R$ ), and mass injection rate (RVW). Initial estimates of all other variables and profiles are automatically developed by the program unless the user overrides them. The parameter, mass injection rate (RVW), is varied until the heating rate from flow field to ablator surface agrees with the energy absorbed and reradiated by the ablator - in essence, satisfying a surface energy balance for a fully coupled solution. The results are punched out on cards for the stagnation line at the free-stream conditions specified.

The analysis proceeds around the body by using the results from the stagnation line as initial conditions. The ablator mass loss rate profile is specified as a function of distance around the body which also determines surface heating rate and surface temperature from the surface energy balance. When the correct mass loss rate profile has been selected, this fact will be reflected by a matching heating rate profile from the flow field.

The program first determines boundary conditions at the bow shock and the ablator surface. Thermodynamic, transport, and radiation properties are then calculated using assumed temperature, pressure, and composition profiles. Subsequently the shock layer equations are solved yielding profiles of pressure from the

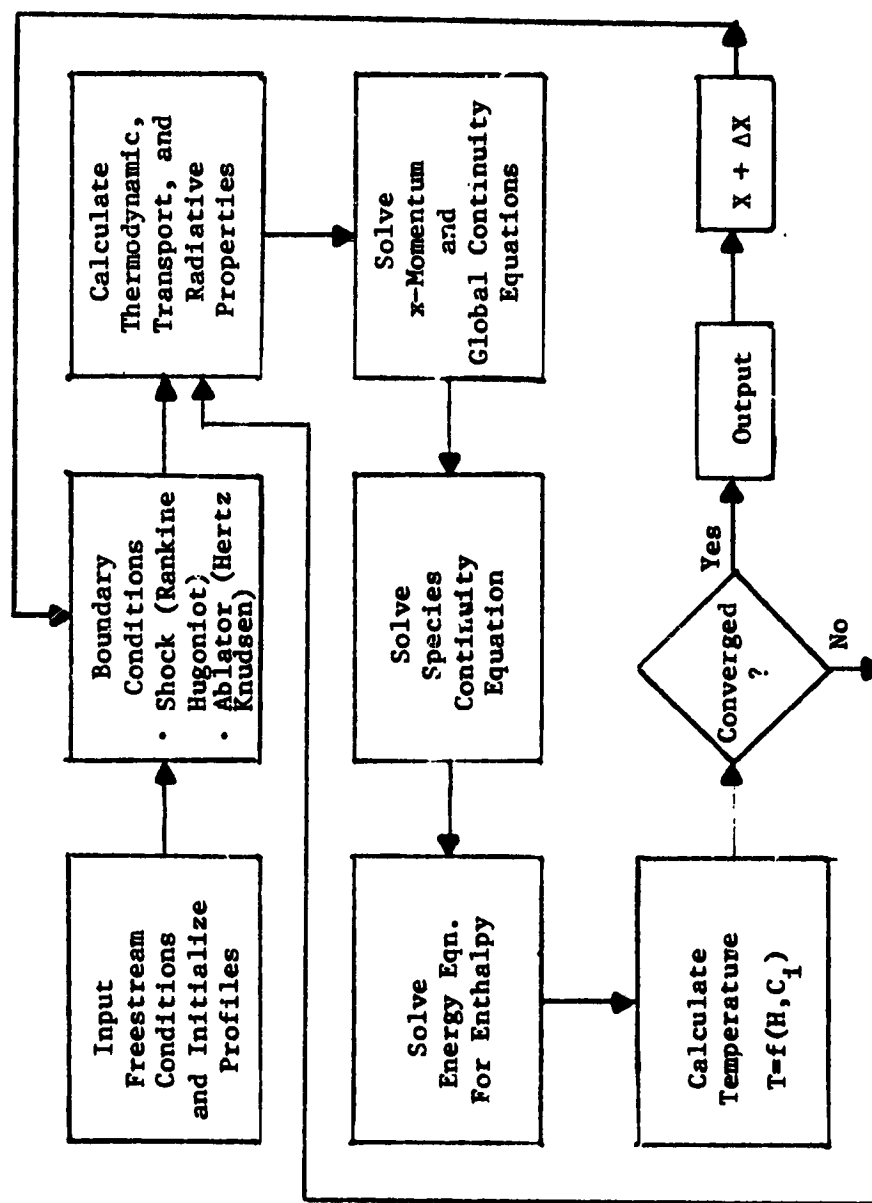


Figure 5.1. Simplified Flow Diagram of Coupled Shock Layer Analysis.

y-momentum equation, tangential velocity (or velocity function,  $f'$ ) from the x-momentum equation, normal velocity from global continuity, composition from species continuity, and temperature (or enthalpy) from the energy equation. The shock layer equations have been uncoupled during the finite differencing procedure and coupling is then achieved by iteration until all profiles are converged simultaneously. The finite difference procedures used to solve the flow field equations are presented in the following sections.

#### Numerical Solutions of the Conservation Equations

The general quasilinearized form for the x-momentum, species continuity, and energy equations was given in Chapter III as

$$-\frac{\partial^2 W}{\partial \eta^2} + \alpha_1 \frac{\partial W}{\partial \eta} + \alpha_2 W + \alpha_3 + \alpha_4 \frac{\partial W}{\partial \xi} = 0 \quad [3.125]$$

where  $\alpha_4$  equals zero along the stagnation line and  $W$  is the dependent variable. Using a three point, variable step size, finite difference approximation at  $(s-2)$  points (Fig. 5.2) across the flow field, the numerical form of the equations becomes

$$\bar{A} \cdot \bar{W} = \bar{B} \quad [5.1]$$

where  $\bar{A}$  is a tridiagonal matrix of order  $(s-2)$ . This form readily lends itself to numerical solution by a tridiagonal matrix inversion algorithm (Ref. 5.1). The matrix has the form



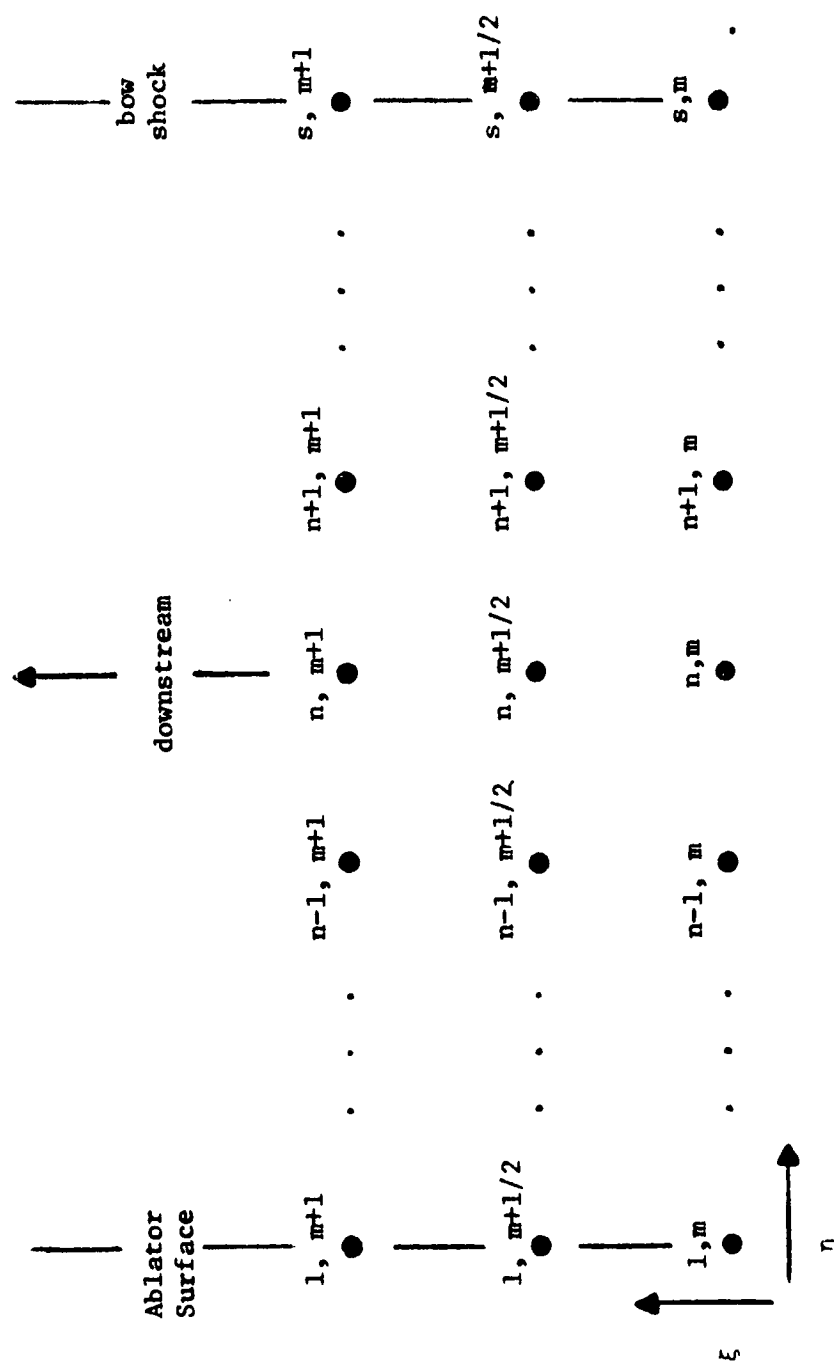


Figure 5.2. Crank Nicolson Finite Difference Grid.

$$\begin{bmatrix}
 B_1 & C_1 & & & \\
 A_2 & B_2 & C_2 & & \\
 & A_3 & B_3 & C_3 & \\
 & & \cdot & \cdot & \cdot \\
 & & & \cdot & \cdot \\
 & & & & A_{s-2} & B_{s-2} & C_{s-2} \\
 & & & & & A_{s-1} & B_{s-1}
 \end{bmatrix}
 \begin{bmatrix}
 W_2 \\
 W_3 \\
 W_4 \\
 \cdot \\
 \cdot \\
 W_{s-2} \\
 W_{s-1}
 \end{bmatrix}
 =
 \begin{bmatrix}
 D_1 - A_1 W_1 \\
 D_2 \\
 D_3 \\
 \cdot \\
 \cdot \\
 D_{s-2} \\
 D_{s-1} - C_{s-1} W_s
 \end{bmatrix}
 \quad [5.2]$$

The coefficients  $A_n, B_n, C_n$  and  $D_n$  have the following values.

$$\begin{aligned}
 A_n &= \Theta (H_n^2 + \alpha_1 H_n') \\
 B_n &= \Theta (G_n^2 + \alpha_1 G_n' + \alpha_2) + \alpha_4 / \Delta \xi \\
 C_n &= \Theta (F_n^2 + \alpha_1 F_n') \\
 D_n &= \frac{(\Theta - 1)}{\Theta} [A_n W_{n-1}^m + B_n W_n^m + C_n W_{n+1}^m] - \alpha_3 + \frac{\alpha_4}{\Theta \Delta \xi}
 \end{aligned}
 \quad [5.3]$$

where  $\Theta$  is a parameter which will yield the various finite difference schemes as given below:

$$\Theta = \begin{cases} 0 & \text{Explicit} \\ 1/2 & \text{Crank Nicolson} \\ 1 & \text{Implicit} \end{cases}$$

and the  $F$ ,  $G$ , and  $H$  coefficients are obtained from the three point finite difference formulae (Ref. 5.2).

$$\frac{\partial W}{\partial \eta} = F'_n W_r + G'_n W_n + H'_n W_{n-1} \quad [5.4]$$

$$\frac{\partial^2 W}{\partial \eta^2} = F_n^2 W_{n+1} + G_n^2 W_n + H_n^2 W_{n-1} \quad [5.5]$$

The shock layer is divided with a variable grid size  $\Delta \eta$  and  $\Delta \xi$  in which the subscript  $n$  refers to the  $\eta$  or normal coordinate and the superscript  $m$  refers to the  $\xi$  or tangential coordinate. The tridiagonal matrix is solved for the dependent  $W$  variables at the  $(m+1)^{th}$  body station.

To complete the system of Eqs. (5.2), the boundary conditions are included as follows:

$$\begin{aligned} B_1 W_2^{m+1} + C_1 W_3^{m+1} &= D_1 - A_1 W_1^{m+1} \\ A_{s-1} W_{s-2}^{m+1} + B_{s-1} W_{s-1}^{m+1} &= D_{s-1} - C_{s-1} W_s^{m+1} \end{aligned} \quad [5.6]$$

where  $W_1^{m+1}$  is the boundary condition at the ablator surface and  $W_s^{m+1}$  is the boundary condition of the shock. These boundary conditions were discussed in some detail in Chapter III.

#### Global Continuity

The global continuity equation is a first order partial differential equation which has one form downstream of the stagnation line and a limiting form at the stagnation line. The downstream form is

$$\frac{1}{r_b} \frac{\partial}{\partial \xi} (\rho u r_b) + \frac{\partial \eta}{\partial X} \frac{\partial}{\partial \eta} (\rho u) + \frac{\rho}{\delta} \frac{\partial}{\partial \eta} (\rho v) = 0 \quad [5.8]$$

and the limiting form at the stagnation line is

$$2 \frac{\partial u}{\partial \xi} = - \frac{1}{\tilde{\delta}} \frac{\partial}{\partial \eta} (\rho v) \quad [5.9]$$

Both of these forms can be solved readily for  $\rho v$  by integrating yielding

$$\rho v = (\rho v)_w - \int_0^\eta \frac{\tilde{\delta}}{\rho} \left[ \frac{1}{r_b} \frac{\partial}{\partial \xi} (\rho u r_b) + \frac{\partial \eta}{\partial x} \frac{\partial}{\partial \eta} (\rho u) \right] d\eta \quad [5.10]$$

and

$$\rho v = (\rho v)_w - 2 \int_0^\eta \tilde{\delta} \left( \frac{\partial u}{\partial \xi} \right) d\eta \quad [5.11]$$

Numerically this integration is carried out quite satisfactorily by a simple trapezoidal scheme. The transformed standoff distance,  $\tilde{\delta}$ , is computed at the stagnation line and around the body using the respective global continuity equations in the following manner

$$\tilde{\delta} = - \frac{(\rho v)_s - (\rho v)_w}{\int_0^1 f' d\eta} \text{ (stagnation)} \quad [5.12]$$

and

$$\tilde{\delta} = - \frac{(\rho v)_s - (\rho v)_w}{\int_0^1 \frac{1}{\rho} \left[ \frac{1}{r_b} \frac{\partial}{\partial \xi} (\rho u r_b) + \frac{\partial \eta}{\partial x} \frac{\partial}{\partial \eta} (\rho u) \right] d\eta} \quad [5.13]$$

The integrals for Eqs. 5.12 and 5.13 are identical to those in 5.10 and 5.11 and are obtained by using the trapezoidal rule integration. This computed value of  $\tilde{\delta}$  is then compared to the assumed value and a new value is estimated. The calculation is repeated until convergence of  $\tilde{\delta}$  is achieved (<0.1%). The actual standoff distance is then computed from the flow field as

$$\delta = \tilde{\delta} \int_0^1 \frac{d\eta}{\rho} \quad [5.14]$$

Again, the integral is evaluated using the simple trapezoidal scheme.

The standoff distance around the body can be calculated by the following geometrical relationship (trapezoidal integration)

$$\delta^{m+1} = \int_0^{\xi^{m+1}} (1 + \kappa\delta) \tan \epsilon \, d\xi \quad [5.15]$$

The value of  $\delta$  from the flow field calculation and the geometrical relation can then be compared to indicate if numerical convergence has been obtained on shock standoff distance. In practice, an  $\epsilon$  profile is assumed and the geometrical relation is integrated using this assumed profile. The  $\epsilon$  profile is then updated until input shock shape (geometrical relation) and the output shock shape (flow field calculation) agree within about 2%.

#### Calculation of $\partial\eta/\partial x$

The value,  $\partial\eta/\partial x$ , appears in the conservation equations and must therefore be evaluated. The term can be expanded as follows using the chain rule

$$\frac{\partial\eta}{\partial x} = \frac{\partial\eta}{\partial y} \frac{\partial y}{\partial x} \quad [5.16]$$

where

$$\frac{\partial\eta}{\partial y} = \frac{\rho}{\delta} \quad [5.17]$$

since by definition

$$d\eta = \frac{\rho}{\delta} dy \quad [5.18]$$

The term  $\partial y / \partial x$  can be evaluated numerically as

$$\frac{\partial y}{\partial x} = \frac{y^{m+1} - y^m}{\Delta x} \quad [5.19]$$

$y^{m+1}$  is approximated using

$$y^{m+1} \approx \frac{y^m}{\delta^m} \delta^{m+1} \quad [5.20]$$

where  $\delta^m$  and  $\delta^{m+1}$  are evaluated from the geometrical relation, Eq. 5.15.

#### y-Momentum Equation

The final partial differential equation to be solved is the y-momentum equation for pressure along the normal coordinate. This equation is actually an ordinary differential equation when the thin shock layer approximation is made as in this study. The transformed y-momentum equation becomes

$$\frac{dP}{d\eta} = \tilde{\delta} u^2 \quad [5.21]$$

Upon integrating

$$P(\eta) = P_s - \int_1^\eta \tilde{\delta} u^2 d\eta \quad [5.22]$$

which can be solved by a simple trapezoidal method as was the global continuity.

The conservation equations have thus been specified as to their numerical solution form. Now a discussion of how the boundary conditions are incorporated follows.

#### Boundary Conditions

The boundary conditions must be specified at two points since the conservation equations are second order parabolic equations. The conditions for this system are given at the shock and the ablator surface. First, the method of determining shock boundary conditions will be discussed followed by the boundary specification at the ablator surface.

Shock boundary conditions: The conditions at the shock are determined by using the Rankine-Hugoniot equations (Eqns. 3.58-3.62). Chemical equilibrium is assumed at the shock and an initial guess of density ratio across the shock wave is made. From the density ratio, enthalpy can be calculated from the Rankine-Hugoniot relations. Temperature is determined from the enthalpy, and the density ratio can then be calculated using equilibrium air compositions (from the equilibrium air subroutine, GAS). This density ratio is then compared with the assumed density ratio. This value is then updated until convergence is obtained ( $<0.5\%$ ). When density ratio is known, the shock pressure, the normal velocity, and the tangential velocity can be determined from the Rankine-Hugoniot relations. The temperature and species compositions are products of the chemical equilibrium calculation.

Ablator surface conditions: The boundary conditions at the ablator surface are a little more complicated to evaluate than the

shock boundary conditions. The material response is actually determined from the solution to a set of nonlinear ordinary differential equations which describes the ablation of the plastic composite heat shield. In this research a Padé approximation numerical technique (CHAR) was used to describe the flow of decomposition gases through the charred heat shield. The technique is described in Appendix A along with a discussion of results and their eventual incorporation in the overall analysis. Essentially, the analysis yields heat conducted at the surface of the ablator as a function of mass loss rate and surface temperature as shown in Figure 5.3.

The results are input into the flow field analysis as

$$q_{\text{cond}} = a_1 T^3 \rho v + a_2 T^2 \sqrt{\rho v} + a_3 T^3 \sqrt{\rho v} + a_4 \rho v \quad [5.23]$$

Coefficients for  $a_1$ ,  $a_2$ ,  $a_3$ , and  $a_4$  are given in Table 5.1 where  $T$  and  $\rho v$  are evaluated at the surface.

The overall analysis is initiated at the stagnation line with an assumed value of mass loss rate ( $\rho v$ ). From this assumed value of mass loss rate, the surface temperature ( $T$ ) can be calculated using the Hertz-Knudsen equation as described in Appendix B.

$$T = f(\rho v, P) \quad [5.24]$$

The pressure is determined at the shock from the Rankine-Hugoniot equations. The wall pressure is equivalent to the shock pressure at the stagnation line since the thin shock layer assumption has been assumed valid.

Boundary conditions have been established for energy ( $T=T_w$ ),



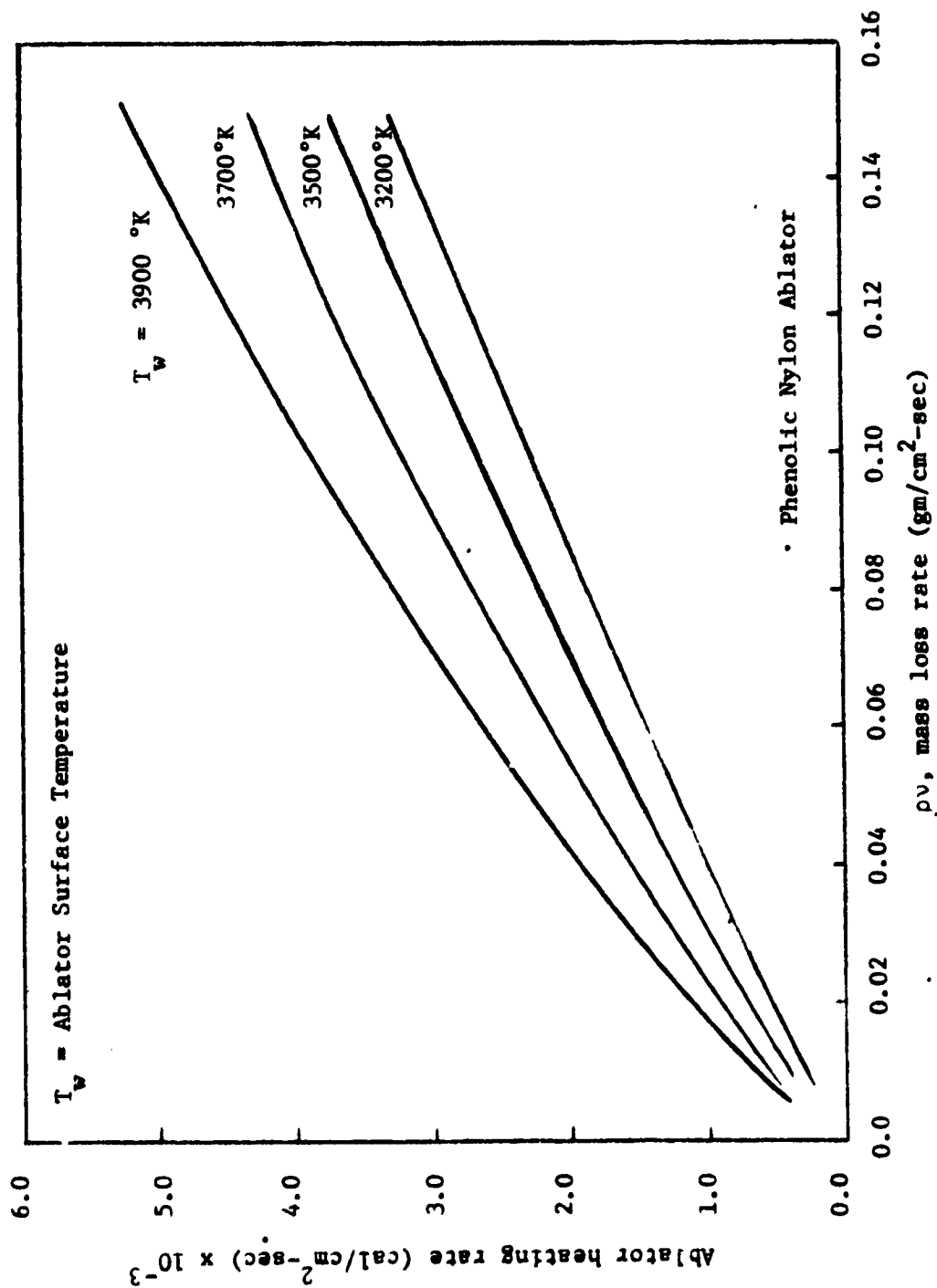


Figure 5.3. Ablator Heating Response vs. Mass Loss Rate.

Table 5.1

Curve Fit Coefficients for Heat Conducted  
Through the Surface of a Phenolic Nylon Ablator

$$(q = a_1 T^3 \rho v + a_2 T^2 \sqrt{\rho v} + a_3 T^3 \sqrt{\rho v} + a_4 \rho v)$$

	q in * <u>cal/sq cm-sec</u>	q in † <u>BTU/sq ft-sec</u>
a <sub>1</sub>	0.9292 x 10 <sup>-8</sup>	0.1674 x 10 <sup>-7</sup>
a <sub>2</sub>	-0.1292 x 10 <sup>-3</sup>	-0.3330 x 10 <sup>-3</sup>
a <sub>3</sub>	0.4268 x 10 <sup>-7</sup>	0.1099 x 10 <sup>-6</sup>
a <sub>4</sub>	0.1516 x 10 <sup>4</sup>	0.273x x 10 <sup>4</sup>

\* T is temperature in °K

ρv is mass loss rate in gm/sq cm-sec

† T is temperature in °K

ρv is mass loss rate in lbm/sq ft-sec

x-momentum ( $f^* = 0$ ), and global continuity ( $\rho v = \rho v_w$ ). The wall values of species composition remain to be specified. This specification is made using constant values of wall composition. The values chosen were determined from the CHAR program at conditions considered typical of the ablator surface during planetary reentry (3600°K and 0.04 lbs/ sq. ft.-sec. mass loss rate) and are given in Table 5.2.

The stagnation line solution can now be determined since the problem has been completely specified. The solution of the shock layer equations yields a value of heating rate (radiative + conductive + diffusive) from flow field to ablator surface. This value must be equivalent to the energy absorbed by the ablator surface by conduction, sublimation phase change and energy reradiated by the surface,  $q_w$ .

$$q_w = q_{\text{cond}} + q_{\text{subl}} + q_{\text{RR}} \quad [5.25]$$

The first term,  $q_{\text{cond}}$ , is attained by evaluating Eq. 5.23 for the conduction through the surface. The heat absorbed by sublimation phase change,  $q_{\text{subl}}$ , is simply the heat of sublimation,  $\Delta H_{\text{vap}}$  ( $\sim 4.4 \times 10^3$  BTU/lb), multiplied by the mass loss rate of solid carbon,  $(\rho v)_c$

$$q_{\text{subl}} = 4.4 \times 10^3 (\rho v)_c \quad [5.26]$$

and the energy reradiated by the char surface is found by applying the Stephan-Boltzmann equation with an emissivity,  $\epsilon$ , of 0.66 (Ref. 5.3).

Table 5.2  
Compositions at the Surface of a Phenolic Nylon Ablator\*

(Surface Temperature = 3600°K,  
Pressure = 1 atm, Mass Loss Rate = .04 lbm sq ft-sec)

	<u>Component</u>	<u>Mass Fraction</u>	<u>Mole Fraction</u>
1	H <sub>2</sub>	0.04135	0.2745
2	H	0.0221	0.2934
3	C <sub>2</sub> H <sub>2</sub>	0.0589	0.0301
4	C	0.3695	0.1363
5	C <sub>2</sub>	0.0422	0.0233
6	C <sub>3</sub>	0.0211	0.0233
7	C <sub>2</sub> H	0.1388	0.0737
8	C <sub>3</sub> H	1 x 10 <sup>-10</sup>	-
9	C <sub>4</sub> H	1 x 10 <sup>-10</sup>	-
10	CO	0.257	0.1219
11	O	0.292 x 10 <sup>-8</sup>	-
12	N <sub>2</sub>	0.04951	0.0235
13	HCN	1 x 10 <sup>-10</sup>	-
14	CN	1 x 10 <sup>-10</sup>	-
15	N	1 x 10 <sup>-10</sup>	-
16	N <sup>+</sup>	1 x 10 <sup>-10</sup>	-
17	O <sup>+</sup>	1 x 10 <sup>-10</sup>	-
18	C <sup>+</sup>	1 x 10 <sup>-10</sup>	-
19	O <sub>2</sub>	1 x 10 <sup>-10</sup>	-
20	e <sup>-</sup>	1 x 10 <sup>-10</sup>	-

---

\* Calculated using CHAR program described in Appendix A.

$$q_{RR} = \epsilon \sigma T^4 = 3.3 \times 10^{-12} T^4 \quad [5.27]$$

where  $q_{RR}$  is the reradiated energy expressed in BTU/sq ft-sec and  $T$  is the surface temperature expressed in °K.

The surface heating rates from the flow field solution and Eq. 5.25 can then be compared. If they differ, then a new value of mass loss rate must be assumed and the procedure repeated. A typical solution progression is shown in Figure 5.4. Normally a reasonably (within 2%) coupled flow field-surface solution can be obtained in three attempts in this manner at the stagnation line.

The solution downstream of the stagnation region is coupled by assuming an initial mass loss rate profile and comparing heating rates as at the stagnation line. In Figure 5.5 a typical mass loss profile is shown and in Figure 5.6, the corresponding heating rates are given. The mass loss rate profile is updated until coupling is achieved at all points downstream of the stagnation region.

#### Stability of the Energy Equation

The energy equation was found to present the most severe stability problems of all the conservation equations. Two forms of the energy equation were presented in Chapter III - the temperature formulation and the enthalpy formulation. Unsurmountable problems with stability were encountered when attempting to use the temperature form whereas the enthalpy form gave excellent results when properly "tuned". A discussion of both methods follows.

Temperature form: The temperature form of the energy equation proved to be quite unsatisfactory for reasons of extreme instability. Initially it was believed that the temperature formulation

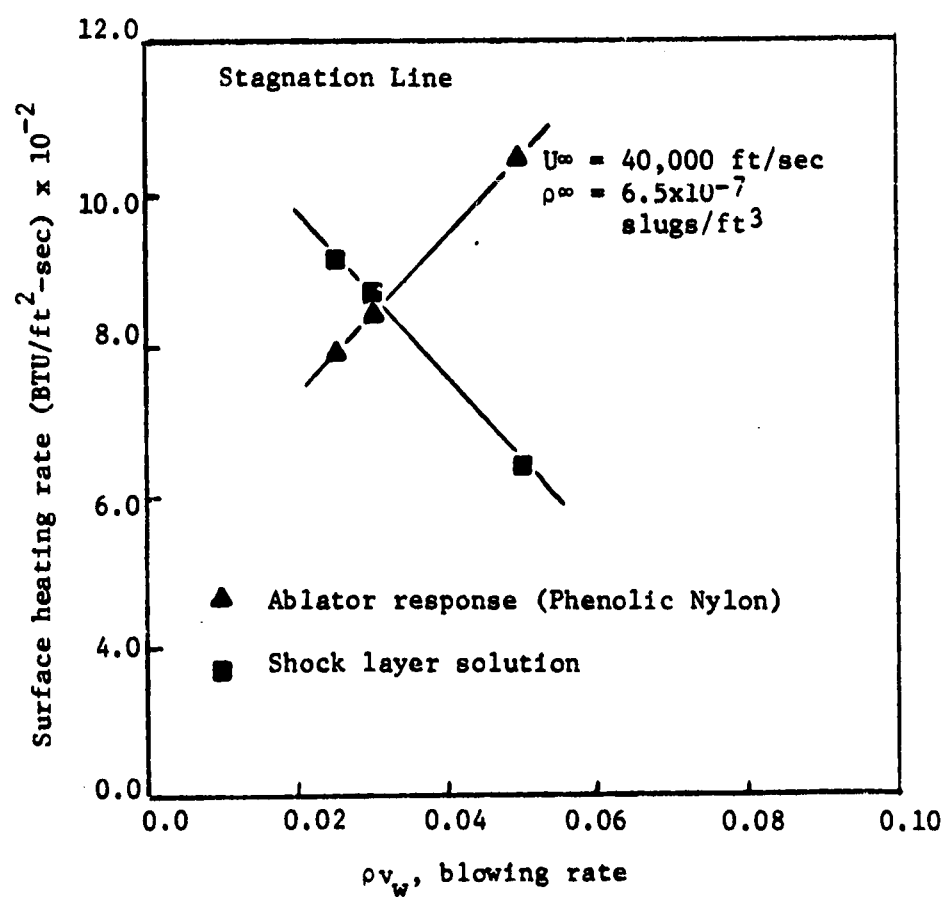


Figure 5.4. Coupling of the Shock Layer to the Ablator by a Surface Energy Balance.

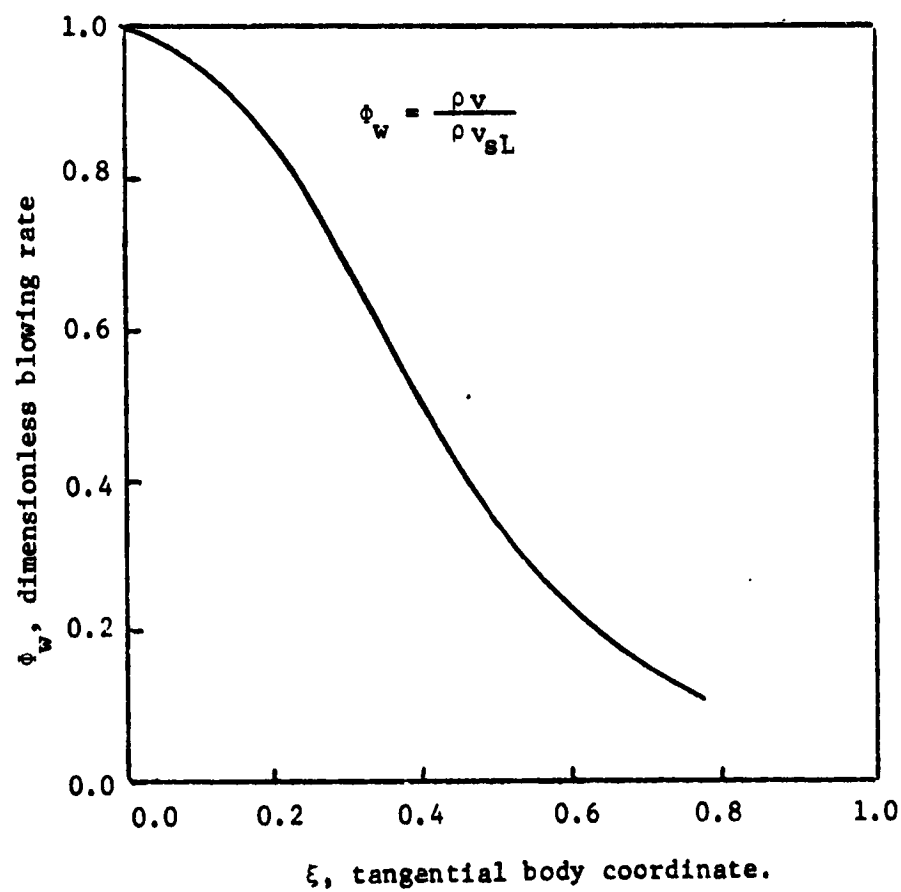


Figure 5.5. Dimensionless Blowing Rate Profile Around the Body.

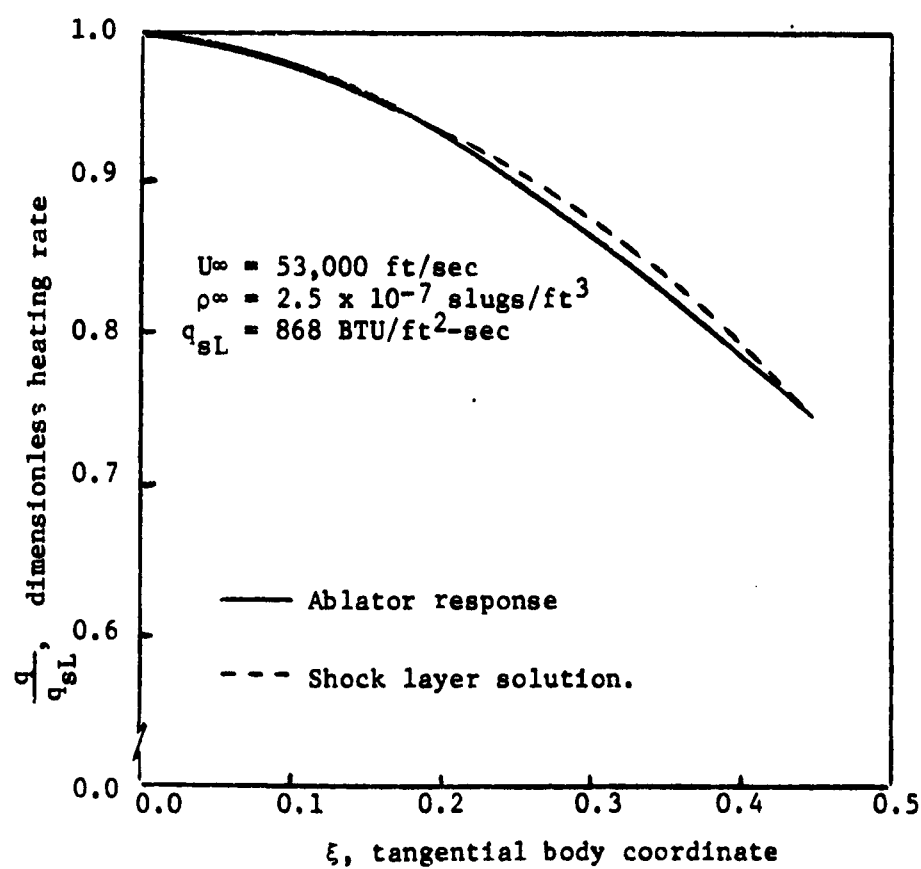


Figure 5.6. Dimensionless Heating Rate Profile Around the Body - Coupled.



would be more desirable than the enthalpy form since temperature is a direct result of the solution of the tridiagonal matrix. In the enthalpy formulation, the tridiagonal matrix solution yields enthalpy which must then be converted to temperature using an iteration scheme with species compositions at each nodal point. The temperature formulation, however, rapidly goes unstable with disastrous oscillations which create havoc on thermodynamic, transport, kinetic, and radiative properties. Within a few iterations the temperature goes negative at some nodal points and at other points, the shock temperature is greatly exceeded. The first few iterations with the energy equation in the temperature form are shown in Figure 5.7. All attempts to dampen these oscillations were unsuccessful.

The probable reason for this instability in the temperature formulation might be in the quasilinearized chemical production term

$$\alpha_2 = \frac{\partial}{\partial T} \left( \sum h_i \frac{\omega_i}{\rho} \right) \quad [5.28]$$

This quasilinearized term should technically include the radiative flux divergence contribution along with the chemical production term. Unfortunately, there is no convenient analytical expression for the partial of radiative flux with respect to temperature. In an attempt to include the quasilinearized radiative flux divergence into  $\alpha_2$ , the flux divergence term was numerically differentiated. The radiative flux divergence is highly nonlinear as depicted in Figure 5.8. In fact with a slight variation in temperature profile,

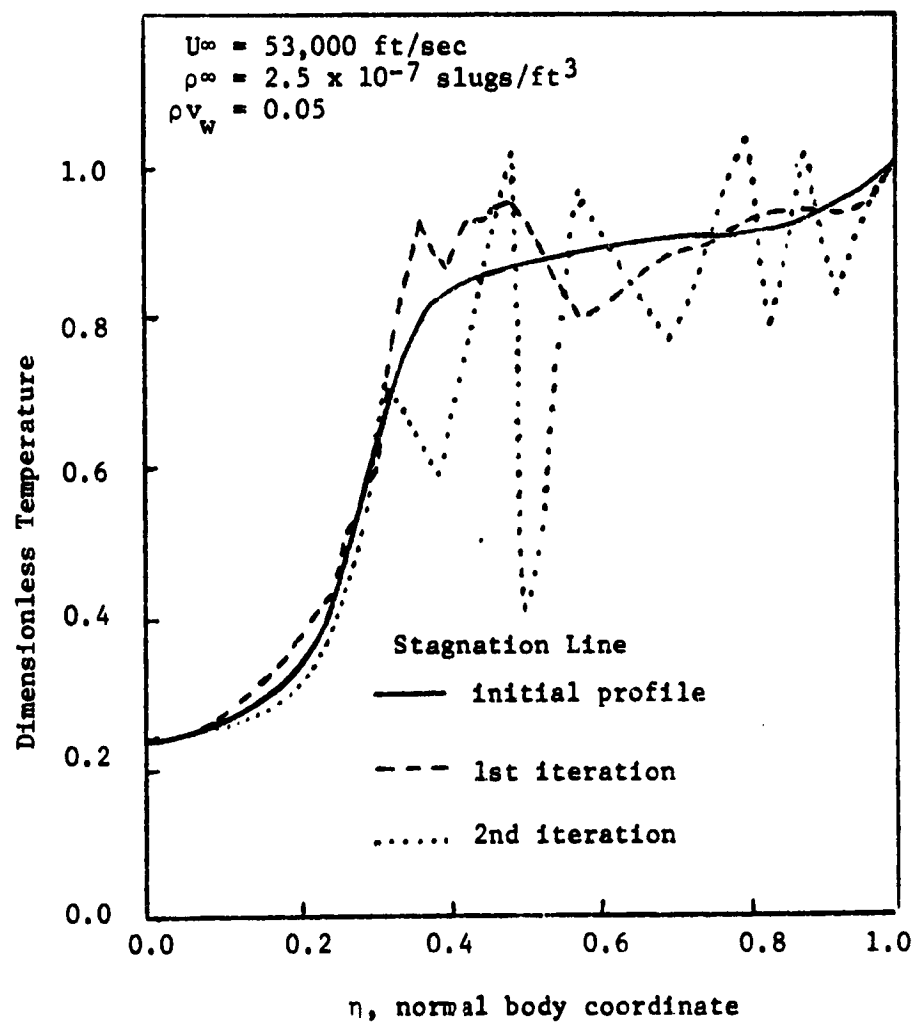


Figure 5.7. Instability of the Temperature Form of the Energy Equation.

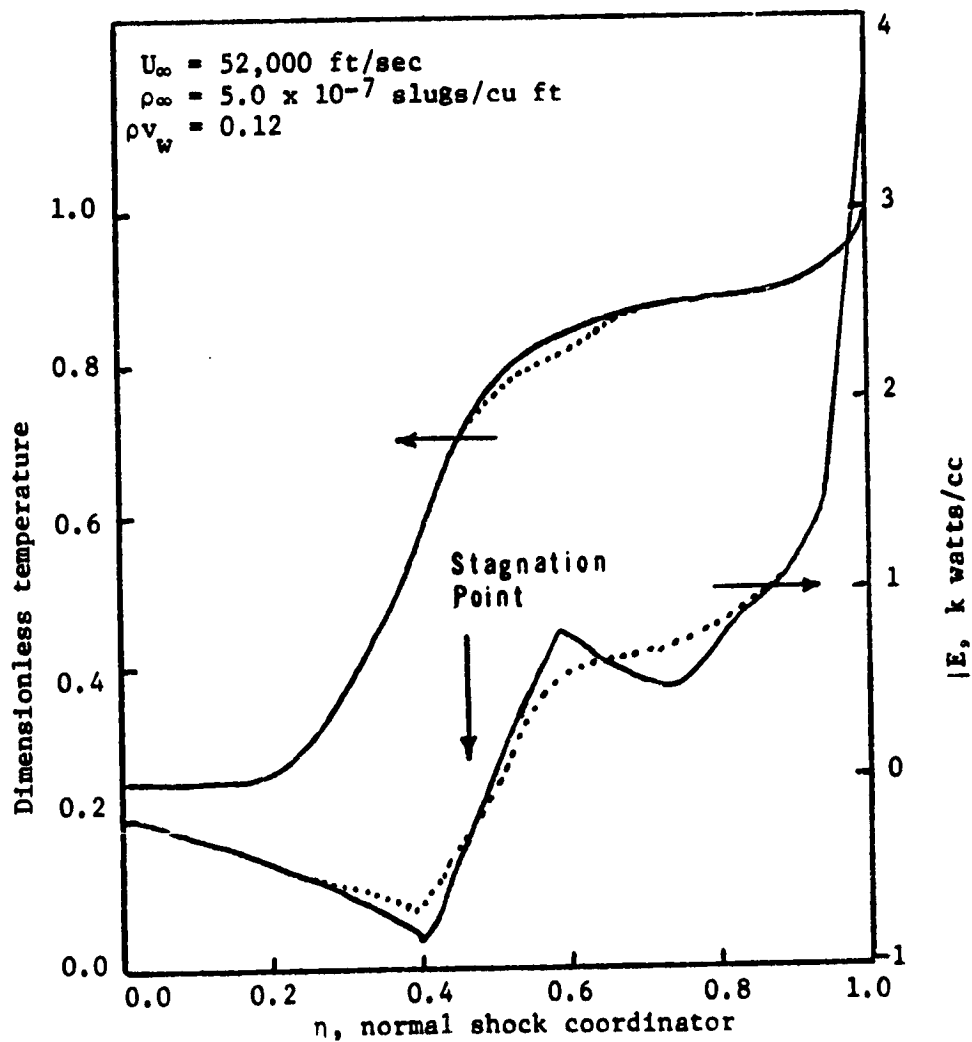


Figure 5.8. Illustration of the Effect of Temperature Variations Upon Radiative Flux Divergence.

the radiative flux divergence changes dramatically inside the flow field. This extreme nonlinearity led to a very unstable numerical differentiation; and, therefore, a very unstable numerical solution of the energy equation.

With the quasilinearized part of the radiative flux term being excluded,  $\alpha_2$  is left with only the chemical production term. The problem with this arrangement is that the primary energy production in the flow field is not due to chemical reactions but to radiation. The resulting situation appeared to place undue emphasis on chemical production which provided for an unstable energy equation formulation. The energy equation in enthalpy form circumvents these problems as discussed below.

Enthalpy form: When the energy equation is solved in its enthalpy form (Eq. 3.130), the result is a quite stable solution. In Figure 5.9 the consecutive enthalpy profiles obtained from using the enthalpy formulations are presented. In Figure 5.10 the temperature profiles which correspond to the enthalpy profiles are presented. The results very clearly show a smooth, stable behavior of the solution.

Comparing the enthalpy formulation of this work with the temperature formulation (equilibrium chemistry model) of Esch (Ref. 5.4) demonstrates the superior stability of the enthalpy formulation (Fig. 5.11). Esch reported that when several iterations were performed the solution sometimes became unstable enough that it could not recover. The profile was then manually adjusted and the solution continued. In the present enthalpy formulation, manual adjustment

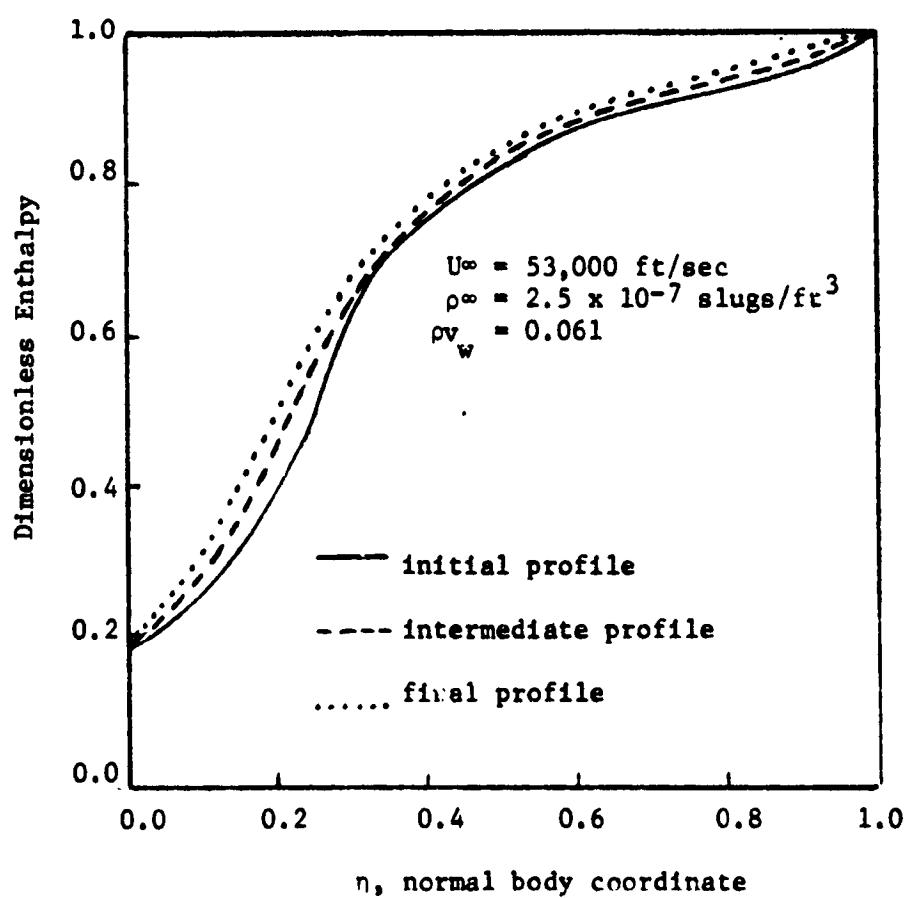


Figure 5.9. Stable Enthalpy Profiles Using Enthalpy Form of the Energy Equation.

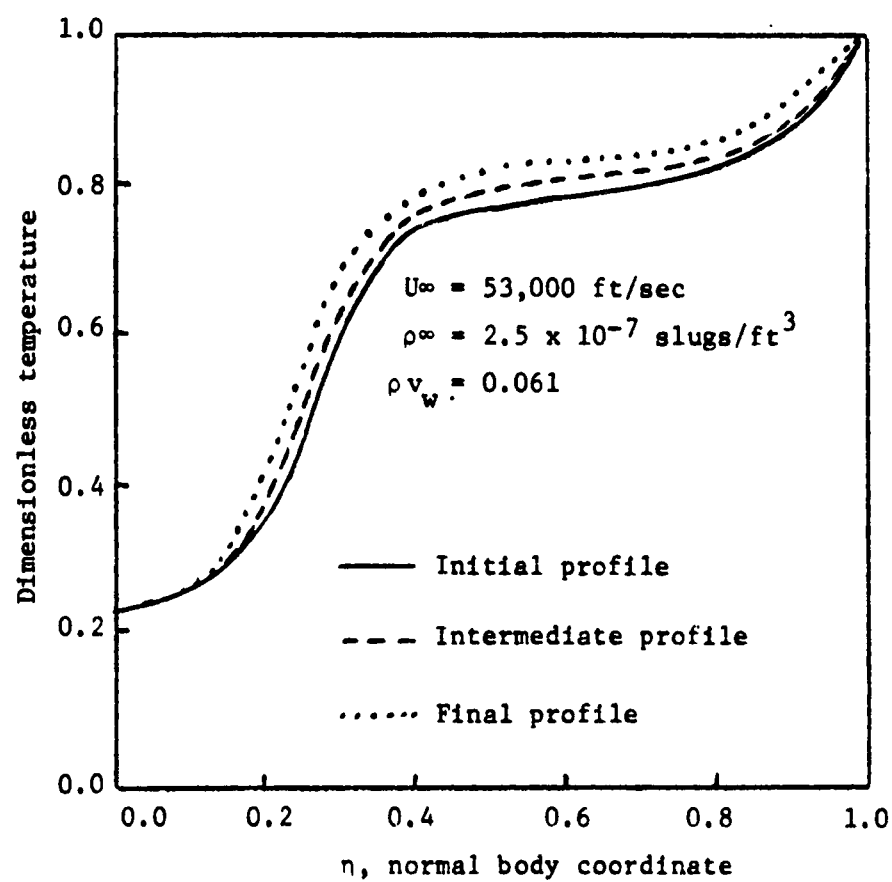


Figure 5.10. Stable Temperature Profiles Using Enthalpy Form of the Energy Equation.

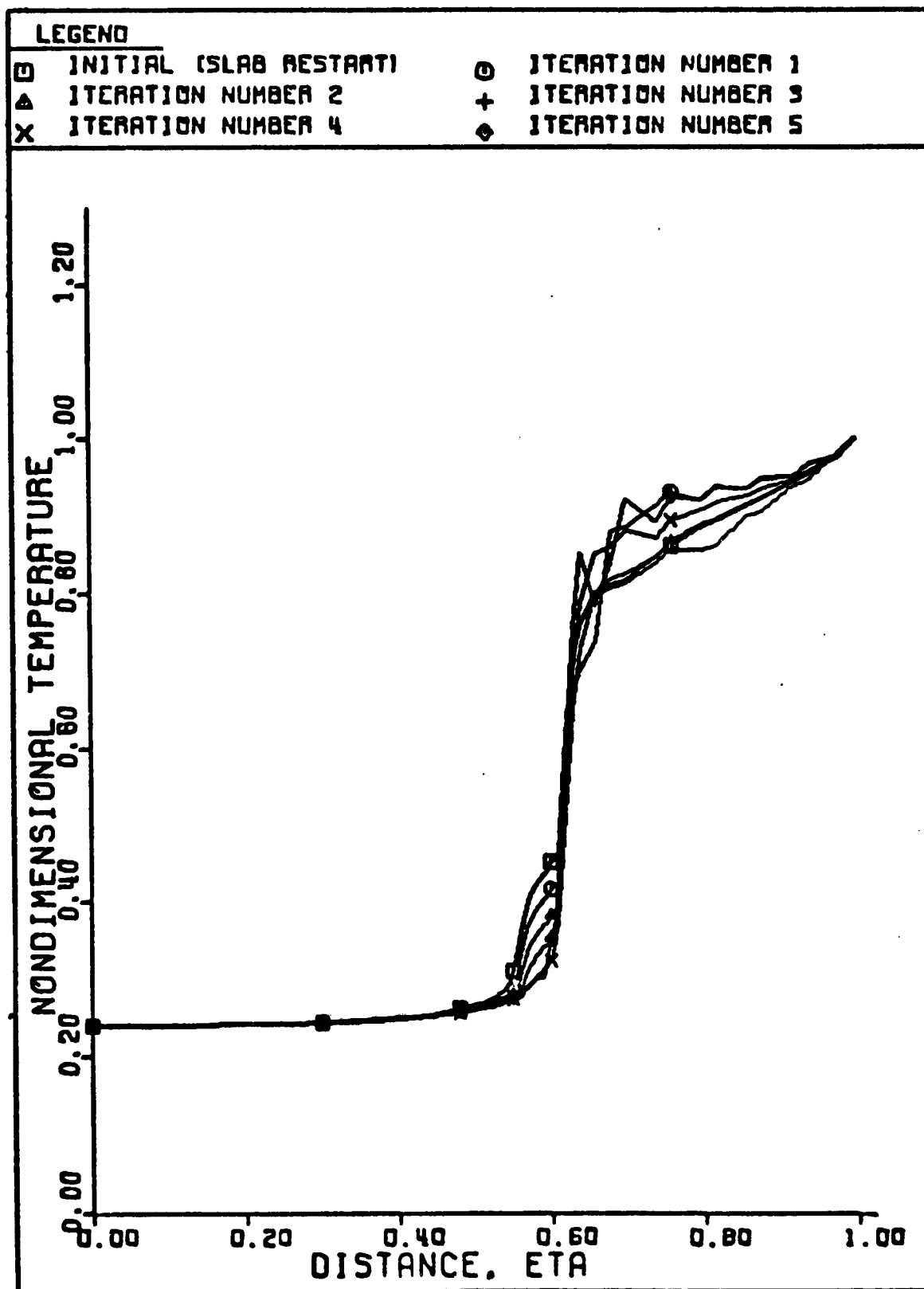


Figure 5.11. Variation in Equilibrium Chemistry Temperature Profiles from Iteration to Iteration Using the Temperature Form of the Energy Equation. (Esch, Ref. 5.4)

was not found necessary. Damping of successive iterations was, however, required as described in the next section.

#### Damping Coefficients

The conservation equations, especially the energy equation, can be made more stable by the use of damping coefficients. The solution is thus permitted to vary only slowly from the initial guessed profile. Damping was achieved by using  $i$ th iteration values and calculated values of the profiles in the following manner

$$W^{i+1} = d W_{\text{calc}} + (1-d) W^i \quad [5.29]$$

where the damping coefficient,  $d$ , is between 0.0 and 1.0 and can be different for each conservation equation. For fast response (rapid convergence), the damping coefficient should approach 1.0. This cannot always be done in practice since stability sometimes becomes a severe problem. In case of instability, the lower value of damping coefficient must be utilized. The best value of damping coefficient can only be determined by numerical experimentation with the specific case and the specific conservation equation in question. The most sensitive equation by far in this study was the energy equation due to its extreme nonlinear production terms of radiation and chemical reaction. For cases in which radiative heating was not severe, damping values of 0.5 on enthalpy yielded reasonably stable solutions, but when radiative heating became predominant, the damping coefficient had to be reduced to 0.25.

The species continuity equations are heavily dependent upon temperature since temperature appears in the Arrhenius form of the



chemical production terms. Due to the sensitivity that the species exhibited on temperature, convergence was often difficult to attain (although stability was no problem) unless the temperature was damped with the very low coefficient of 0.1. The best results were obtained when a damping coefficient of 0.25 was used to bring the solution close to convergence. The damping coefficient was then reduced to 0.1 to eliminate small oscillations in the temperature profile which caused corresponding but slightly more pronounced oscillations in the species profiles. Solutions could then be obtained in which all profiles were converged to within 0.5%.

The "best" damping coefficients for the x-momentum and species continuity equations were found to be 1.0. The damping coefficient used for the enthalpy profile was 0.5 (this value directly affects the temperature profile since the temperature is calculated from enthalpy). Besides damping coefficients on these profiles, a damping coefficient was applied to the calculation of the transformed shock standoff distance,  $\tilde{\delta}$  (from global continuity), as follows

$$\tilde{\delta}^{i+1} = d \tilde{\delta}_{calc} + (1-d) \tilde{\delta}^i \quad [5.30]$$

The value of the damping coefficient used for  $\tilde{\delta}$  for 0.5. The use of this damping coefficient on transformed standoff distance increased stability and made convergence more rapid.

#### Shock Standoff Distance

The calculation of shock standoff is not as simple as it would appear at first glance. The calculation of standoff distance from

the geometrical relation of Equation 5.15 requires the specification of the angle,  $\epsilon$ , which is the difference between the body angle and the shock angle (refer to Fig. 3.2). This angle is also used in the calculation of the boundary conditions at the shock with the Rankine-Hugoniot relations. These shock boundary conditions in turn alter the flow field solution significantly thus altering the value of shock standoff distance calculated from Equation 5.14.

In Figure 5.12 two  $\epsilon$  angle profiles are shown as a function of distance around the body. The corresponding shock standoff distances as computed from Equations 5.14 and 5.15 are shown in Figure 5.13. The different  $\epsilon$  angle profiles create only a minor change in the solution of the geometrical relation (Eq. 5.15) but the same difference in profiles causes a drastic change in standoff distance as computed from the flow field solution, Equation 5.14. This sensitivity of the flow field shock standoff distance calculation to input  $\epsilon$  profile causes difficult convergence problems. Because of the difficulty associated with achieving a converged shock standoff profile, the convergence criteria was not as stringent as for the rest of the flow field parameters. A difference of about 2 to 3% was considered acceptable.

In order to determine the sensitivity of the overall solution on the correct shock standoff distance, the heating rate profiles were compared for the two profiles. Results in Table 5.3 using the two profiles in Figure 5.12 yielded heating rate profiles that differed by only 0.5% at a body station of  $x = 0.7$ . This result

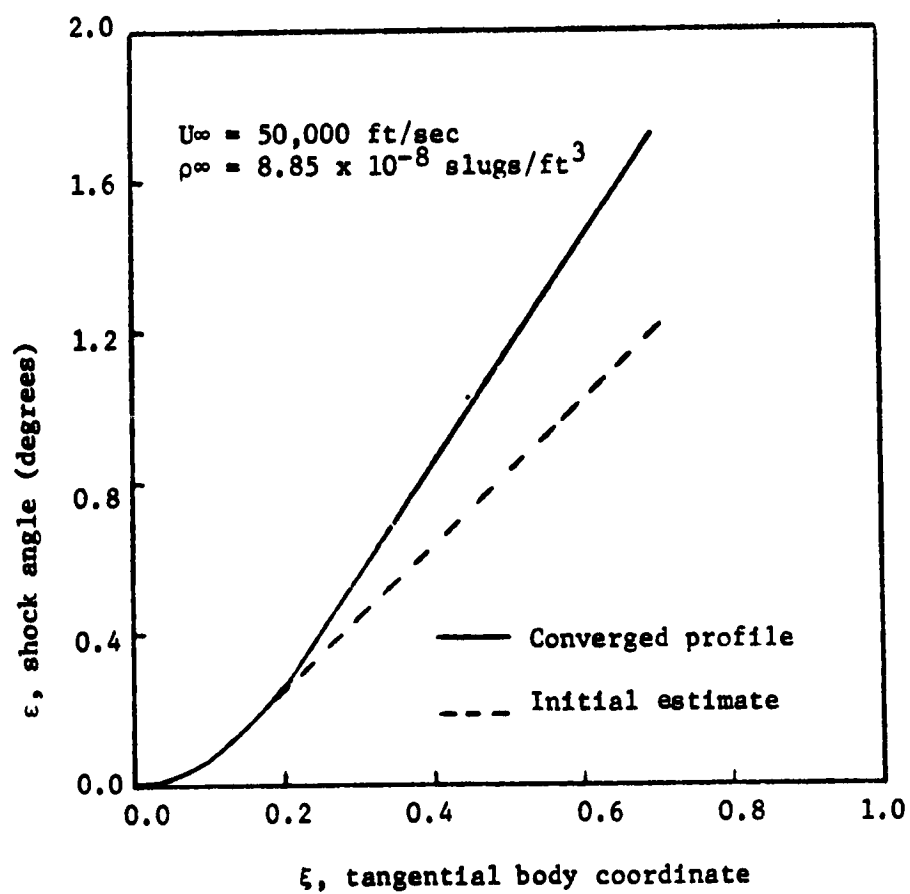


Figure 5.12. Convergence of shock angle,  $\epsilon$ , Around the Body.

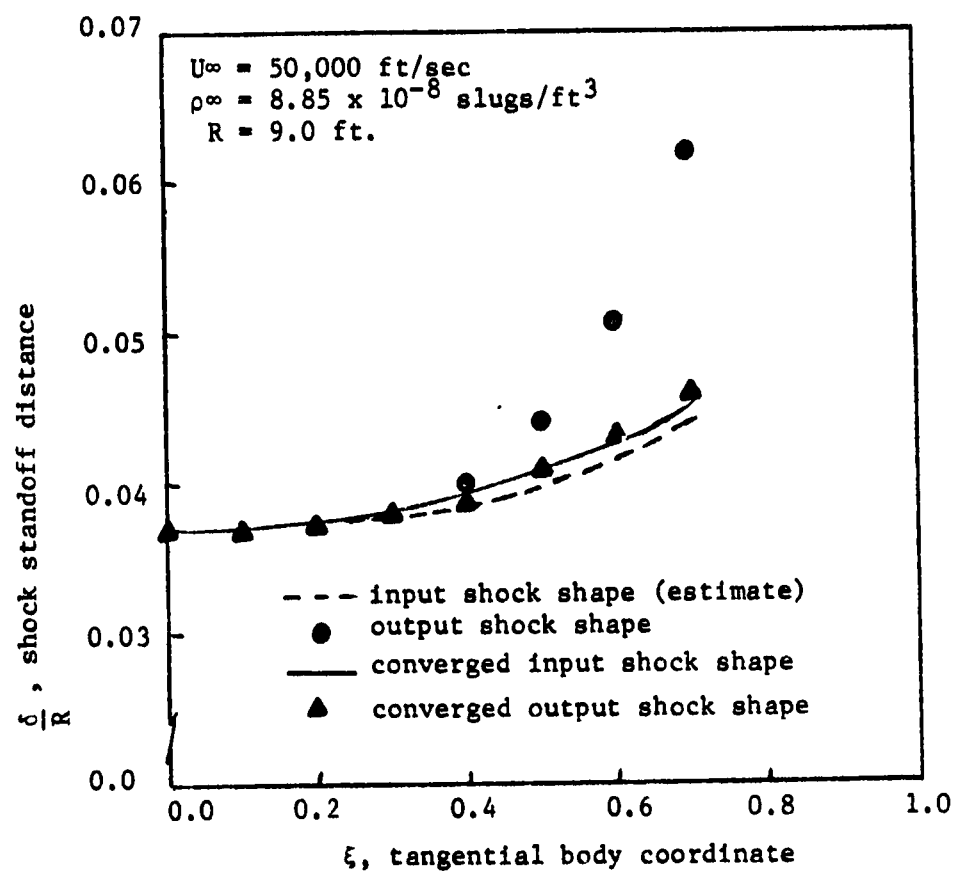


Figure 5.13. Convergence of shock shape profile around the body.

Table 5.3

Insensitivity of Heating Rate to Convergence  
of Shock Standoff Distance,  $\delta$

$U_{\infty} = 50,000 \text{ ft/sec}$ $\rho_{\infty} = 8.85 \times 10^{-8} \text{ slugs/cu ft}$		Heating Rate (BTU/sq ft - sec)	
$\xi$	$\delta$ converged	$\delta$ not converged	
0.0	182.2	182.2	
0.2	163.1	163.1	
0.4	114.5	114.3	
0.6	58.1	58.0	
0.8	24.2	24.0	

indicates that the solution is essentially insensitive to shock standoff distance.

The specification of the shock stand-off distance does appear to affect the speed of convergence. When the incorrect angle profile was used, the solution took 150 minutes of computer time (IBM 360/65) to reach a body station of  $\xi = 0.6$ ; and when the correct angle profile was used, the solution time took only 120 minutes of computer time to reach a body station of  $\xi = 0.9$ .

The results from this study indicated that a highly accurate profile of shock angle was not completely necessary for an accurate solution of the heat load on the ablation heat shield. Comparable results were obtained by Moss (Ref. 5.5). However, the closer one can approximate the correct profile, the seemingly more efficient is the utilization of computer time. Using this result as a basis, therefore, relatively little emphasis was placed on convergence of shock standoff distance.

### Initial Profiles

The specification of initial profiles was made along the stagnation line using the following

$$\rho v = \eta^2 [(\rho v)_s - (\rho v)_w] + (\rho v)_w \quad [5.31]$$

$$f' = \eta \quad [5.32]$$

$$C_1 = \begin{cases} C_{1w} & \rho v < 0 \\ C_{1\delta} & \rho v > 0 \end{cases} \quad [5.33]$$

$$\frac{T - T_w}{T_\delta - T_w} = \frac{T' - T_w'}{T_\delta' - T_w'} \quad [5.34]$$

where  $T'$  is an internally guessed profile for nondimensional temperature as shown in Figure 5.14. This profile can be input by the user if a better idea of its shape is known. In all cases studied, it was found that the solution converged satisfactorily using the temperature profile in Figure 5.14.

It was found from experience that the species profile estimates in Equation 5.33 yielded negative values which when used to calculate temperature from enthalpy, yielded an unstable temperature profile. This problem was solved by iterating a number of times with the species continuity equation alone (no energy or momentum) and using absolute values of species composition. The species profiles eventually lined-out and the solution could proceed normally. A satisfactory number of iterations was found to be 50.

The initial guesses of profiles downstream of the stagnation line are determined as follows

$$\left. \frac{v - v_w}{v_s - v_w} \right|_{\xi_{m+1}} = \left. \frac{v - v_w}{v_s - v_w} \right|_{\xi_m} \quad [5.35]$$

$$f'(\xi_{m+1}) = f'(\xi_m) \quad [5.36]$$

$$C_i(\xi_{m+1}) = C_i(\xi_m) \quad [5.37]$$

$$\left. \frac{T - T_w}{T_\delta - T_w} \right|_{\xi_{m+1}} = \left. \frac{T - T_w}{T_\delta - T_w} \right|_{\xi_m} \quad [5.38]$$

The solution iterates at each body station until convergence is attained and then proceeds to the  $m+1^{\text{th}}$  body station with initial guesses as given in Equations 5.35 - 5.38.

#### Iteration Sequence

The iteration scheme used in this study is shown in Figure 5.15. With initial estimates of the flow field variables the radiative flux divergence was computed in LRAD. The normal and tangential velocity profiles along with the transformed shock standoff distance were then computed by converging simultaneously on the global continuity, x-momentum, and y-momentum equations. Convergence was assumed when the tangential velocity profile change by less than 0.5% and the transformed shock standoff distance changed by less than 0.1%.

The species continuity equations (SPECIE) were then solved with five consecutive iterations. These multiple iterations were found necessary to stabilize the solution since it appears that the species equations have a much slower response than the energy equation. If the species solution is not allowed to keep pace with the energy solution, stability problems are encountered when the temperature profile is computed from enthalpy and species profiles. The problem soon compounds itself and the solution wanders



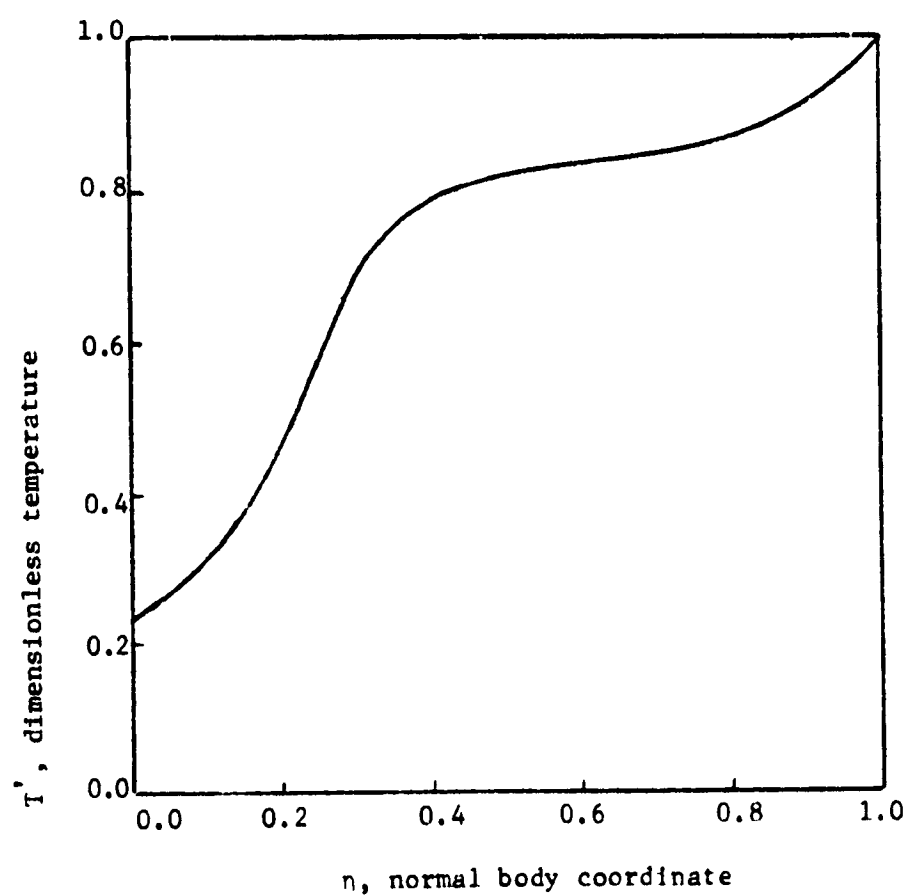


Figure 5.14. Dimensionless Temperature Profile for Initial Estimate at Stagnation Line.

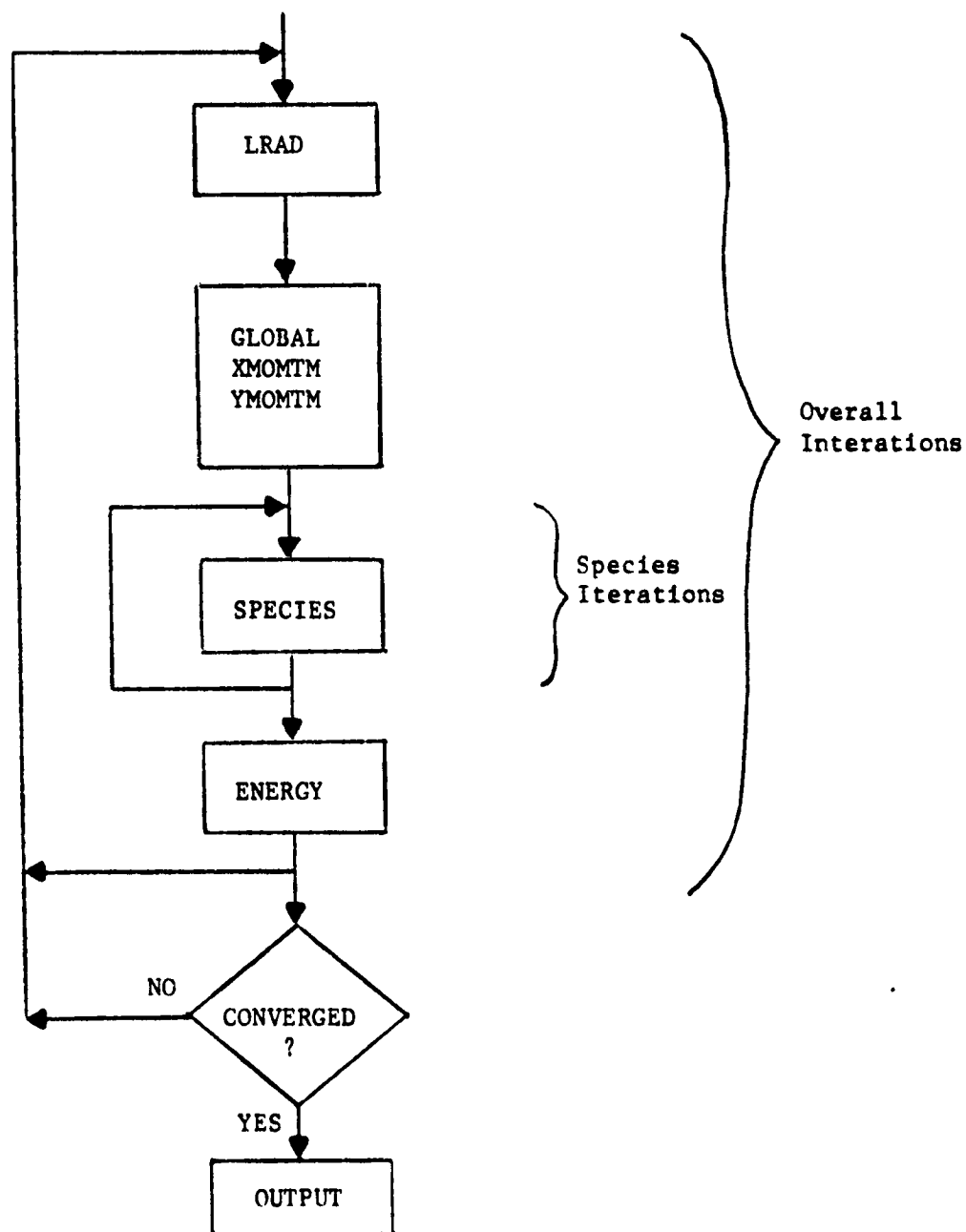


Fig. 5.15. Iteration Sequence of Shock Layer Analysis

into oblivion. Five iterations on species continuity for every iteration on energy proved quite satisfactory.

The energy equation (ENERGY) was then solved in enthalpy form. The enthalpy profile thus obtained was used in conjunction with the species profiles to determine a temperature profile. This cycle of x-momentum, y-momentum, global, species, and energy was repeated five times using the same radiative flux divergence profile in the energy equation. The flux divergence profile was then updated and the sequence repeated until convergence of the overall analysis was achieved.

#### Convergence

Convergence must be satisfied on two levels -- first, the solution must agree from one iteration to the next, and second, the number of nodal points used must be sufficient to give the true solution. Both of these convergence criteria are discussed below.

Iteration convergence: Initially, an attempt was made to require all the profiles to be within 0.5% of each other on consecutive iterations before converged was assumed. Computer time using this convergence criteria, however, became astronomical. An alternate convergence criteria was then used which stipulated that heat flux from flow field to ablator agreed within 0.5% on consecutive iterations. The heating rate, after all, is the factor that determines the rate at which the ablator loses mass - the direct purpose of this research. Reasonable computation time was experienced using this convergence criteria.

Nodal convergence: The shock layer equations are a set of parabolic partial differential equations in two dimensions. The finite difference grid, therefore, has two spacings - one the normal coordinate and the other the tangential coordinate. Convergence must be demonstrated for both coordinates. The finite difference grid system was shown schematically in Figure 5.2 where  $n$  is the normal coordinate nodal point and  $m$  is the tangential coordinate nodal point.

Convergence was verified in the tangential direction by halving the stepsize from  $\Delta\xi = 0.2$  to  $\Delta\xi = 0.1$  and the heating rate profile was examined (since this is the most important variable in the analysis, and it determines mass loss rate for the ablator). This profile was used for the convergence criteria, and the heating rate profiles differed by less than 1% for the two stepsizes investigated at a body station of  $\xi = 0.8$  as indicated by the comparison of results in Table 5.4. This result implied that convergence had been attained so a step size of  $\Delta\xi = 0.05$  was used in all subsequent analyses.

Esch (Ref. 5.4) showed convergence along the stagnation line by approximately halving the step size (126 points versus 59 points). He demonstrated that the profiles of dependent variables remained essentially the same. The flow field analysis by Esch was very similar to the present flow field analysis in that it was the solution of the reacting, radiating, viscous flow equations. The present analysis also used the same numerical technique as used by Esch. The major differences are that Esch's analysis was only for the stagnation line, it used equilibrium chemistry, and it

Table 5.4  
Nodal Convergence in Tangential Direction

$U_{\infty} = 50,000 \text{ ft/sec}$		Heating Rate Results	
$\rho_{\infty} = 8.85 \times 10^{-8} \text{ slugs/cu ft}$		BTU/sq ft.-sec	
$\xi$	$\Delta\xi = 0.1$	$\Delta\xi = 0.2$	
0.0	182.2	182.2	
0.2	163.1	163.0	
0.4	114.5	113.8	
0.6	58.1	57.8	
0.8	24.2	24.0	

solved the energy equation in temperature form. The present analysis applied to the entire flow field in which nonequilibrium chemistry was employed, and the energy equation was solved in its enthalpy form. These differences did not significantly alter the validity of the numerical solution, but they did significantly affect computational time and core requirements. Convergence was therefore established by calling upon the Lax equivalence theorem (Ref. 5.6). This theorem states: "Given a properly posed initial boundary value problem and a finite difference approximation to it that satisfies the consistency condition, then stability is the necessary and sufficient condition for convergence." The finite difference equation is said to be consistent (compatible) with the differential equation if the local truncation errors tend to zero as  $h, k \rightarrow 0$ . The explicit and Crank Nicolson formulae are compatible (Ref. 5.6). Since the Crank Nicolson method was used in this analysis, and stability had been achieved, it can be argued that convergence has been attained based on the Lax equivalence theorem.

#### Summary

The numerical formulation of shock layer equations has been specified in this chapter. Using the  $\alpha$  coefficients developed in Chapter III and the flow field properties (thermodynamic, transport, kinetic, and radiative) developed in Chapter IV, results can be achieved for a wide variety of cases.

In this chapter numerical incorporation of the boundary conditions was discussed along with stability, convergence, and numerical experience with the numerical techniques. It was

concluded that the energy equation could be most easily solved in its enthalpy formulation with proper use of damping coefficients. The temperature formulation of the energy equation leads to a highly unstable solution which appears impossible to control. Convergence to a true solution was shown by calling upon the Lax Equivalence Theorem. The shock standoff distance demonstrated severe sensitivity to shock angle,  $\epsilon$ , but this was shown to have an insignificant effect on the overall analysis.

The next chapter will use the numerical procedures presented in this chapter and investigate heating rates and mass loss rates for the flow field around an ablating blunt body (phenolic-nylon ablator).

## References

- 5.1 Conte, S. O., Elementary Numerical Analysis, McGraw-Hill Book Co., New York (1965).
- 5.2 Davis, R. T., "Numerical Solution of the Hypersonic Viscous Shock-Layer Equations," AIAA Journal, 8 (5), pp. 843-851, May 1970.
- 5.3 Pope, R. B., "Measurement of the Total Surface Emittance of Charring Ablators," AIAA Journal, Vol. 5, No. 12, (1967), pp. 2285-2287.
- 5.4 Esch, D. D., "Stagnation Region Heating of a Phenolic-Nylon Ablator During Return from Planetary Missions," PhD dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 5.5 Moss, J. N., "Solutions for Reacting and Nonreacting Viscous Shock Layers with Multicomponent Diffusion and Mass Injection," PhD dissertation, Virginia Polytechnic Institute and State University, Blacksburg, Virginia (1971).
- 5.6 Ames, Williams F., Numerical Methods for Partial Differential Equations, Barnes & Noble, Inc., New York, N.Y., pp. 61-62, (1969).
- 5.7 Bahn, G. S., "Reaction Rate Compilation for the H-O-N System", p. 23, Gordon and Breach Science Publication, New York (1968).
- 5.8 Pike, R. W., "Evaluation of the Literature for Chemical Reaction Rates for the Decomposition Products from Charring Ablators," Langley Working Paper, LWP-181, NASA (Jan. 21, 1966).
- 5.9 Menees, G. P. and R. L. McKenzie, "A Simplified Chemical Model for Estimating Non-Equilibrium Radiant Emission of CN (Violet) in Shock-Heated Mixtures of CO<sub>2</sub> and N<sub>2</sub>," AIAA J., 6 (3), 554 (March 1968).
- 5.10 Davies, William O., "Radiative Energy Transfer on Entry Into Mars and Venus," Rept. No. IITRI-T200-13, IIT Research Institute, Chicago, Illinois, Dec. 1965.
- 5.11 Allen, C. W., Astrophysical Quantities, Second edition, University of London, Athlone Press, 1963.
- 5.12 Semenov, N. N., Some Problems of Chemical Kinetics and Reactivity, Vol. I, Pergamon Press, New York, p. 2, 1958.



## CHAPTER VI

### RESULTS OF THE SHOCK LAYER ANALYSIS

#### Introduction

The ground work has been laid in the foregoing chapters for the results presented in this chapter. All the assumptions pertinent to obtaining the results presented here were discussed in Chapter III where the conservation equations which describe the flow-field were developed. The species and momentum conservation equations have the form of a set of parabolic partial differential equations and the energy equation takes the form of an integro-partial differential equation. In Chapter IV, the thermodynamic, transport, and radiative properties of the flow field were discussed along with the chemical reactions taking place in the flow field. This work was based on the development of a useful analytical tool to be applied in the determination of the heating load on a manned reentry space vehicle. This reentry vehicle is to be equipped with a heat shield made of a phenolic nylon composite capable of sustaining velocities of earth entry on return from planetary missions such as Mars. Typically, a reentering spacecraft will make contact with the earth's thin upper atmosphere at about 55,000 feet per second. When the craft encounters the denser air at low altitudes, it will be slowed by the frictional drag of the atmosphere thus reducing its kinetic energy.

The results presented in this chapter will be for six key points along the reentry trajectory. The reentry body was taken to have a spherical shape with a nine (9.0) foot radius which is typical of the Apollo spacecraft. Heating rates were determined at the stagnation line for these six cases and downstream heating was determined for four of these six cases - all of which were coupled to the ablator by a surface energy balance. These six cases required more than 100 hours of IBM 360/65 CPU time. A study of chemical reaction rates in the flow field was also performed which demonstrated a marked effect on the heating load on the ablator.

After the stagnation line heating rate results are presented in this chapter, a comparison of the stagnation line is made with several other investigators. Around the body results are then presented followed by a discussion of chemical reactions in the shock layer. The chapter closes with a discussion of computer limitations and other difficulties encountered in the analysis.

#### Stagnation Region Heating

The stagnation line solution is important, not only because this is normally the region of peak heating, but also because this solution is required as initial conditions for the solution of the flow field around-the-body. Conditions for the six coupled stagnation line cases run for this investigation are given in Table 6.1. Heating rate results are shown in Figure 6.1 and Table 6.2 for these cases. The trajectory along which these conditions exist is shown in Figure 6.2. All six cases were coupled to the ablator by a surface energy balance as described in Chapter III.

Table 6.1  
 Freestream and Ablator Conditions for Cases Studied  
 in the Present Investigation

Case	Altitude (ft)	$U_{\infty}$ (ft/sec)	$\rho_{\infty}$ (slugs/ft <sup>3</sup> )	Surface * Temperature (°K)	Mass Loss Rate * (lbm/sq ft-sec)
1	247,000	55,000	$7.5 \times 10^{-8}$	3390	0.005
2	218,000	53,000	$2.5 \times 10^{-7}$	3576	0.026
3	199,500	52,000	$5.0 \times 10^{-7}$	3741	0.102
4	193,000	50,000	$6.5 \times 10^{-7}$	3781	0.135
5	193,000	45,000	$6.5 \times 10^{-7}$	3685	0.059
6	193,000	40,000	$6.5 \times 10^{-7}$	3605	0.026

\* Stagnation line quantities.

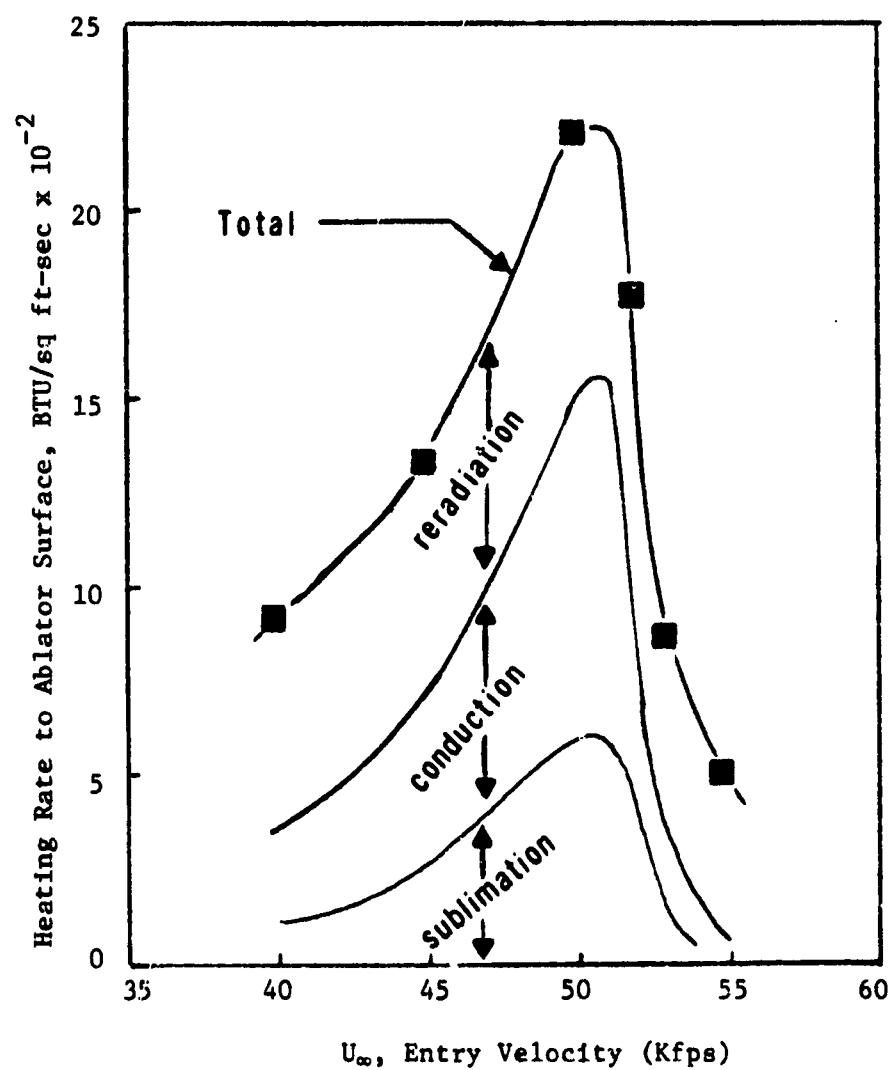


Figure 6.1. Stagnation Heating Rate at a Phenolic Nylon Ablator Surface Along a Reentry Trajectory.

Table 6.2  
Stagnation Line Heating Rate Results for Cases Investigated

Case	Heating Rate, BTU/sq ft-sec						
	Ablator Surface			Shock Layer			
	$q_{RR}^1$	$q_c^2$	$q_s^3$	$q_T^4$	$q_R^5$	$q_c^2$	$q_D^6$
1	436	48	22	506	153	255	93
2	540	214	115	868	671	156	49
3	646	716	451	1813	1793	-36	25
4	674	926	600	2200	2232	-72	33
5	609	450	262	1320	1266	39	45
6	558	225	116	899	649	199	75

1.  $q_{RR}$  is energy reradiated from char surface
2.  $q_c$  is energy transport by conduction
3.  $q_s$  is energy absorbed by sublimation phase change
4.  $q_T$  is total energy
5.  $q_R$  is radiative energy transmitted from the shock layer to the surface
6.  $q_D$  is energy transport by diffusion

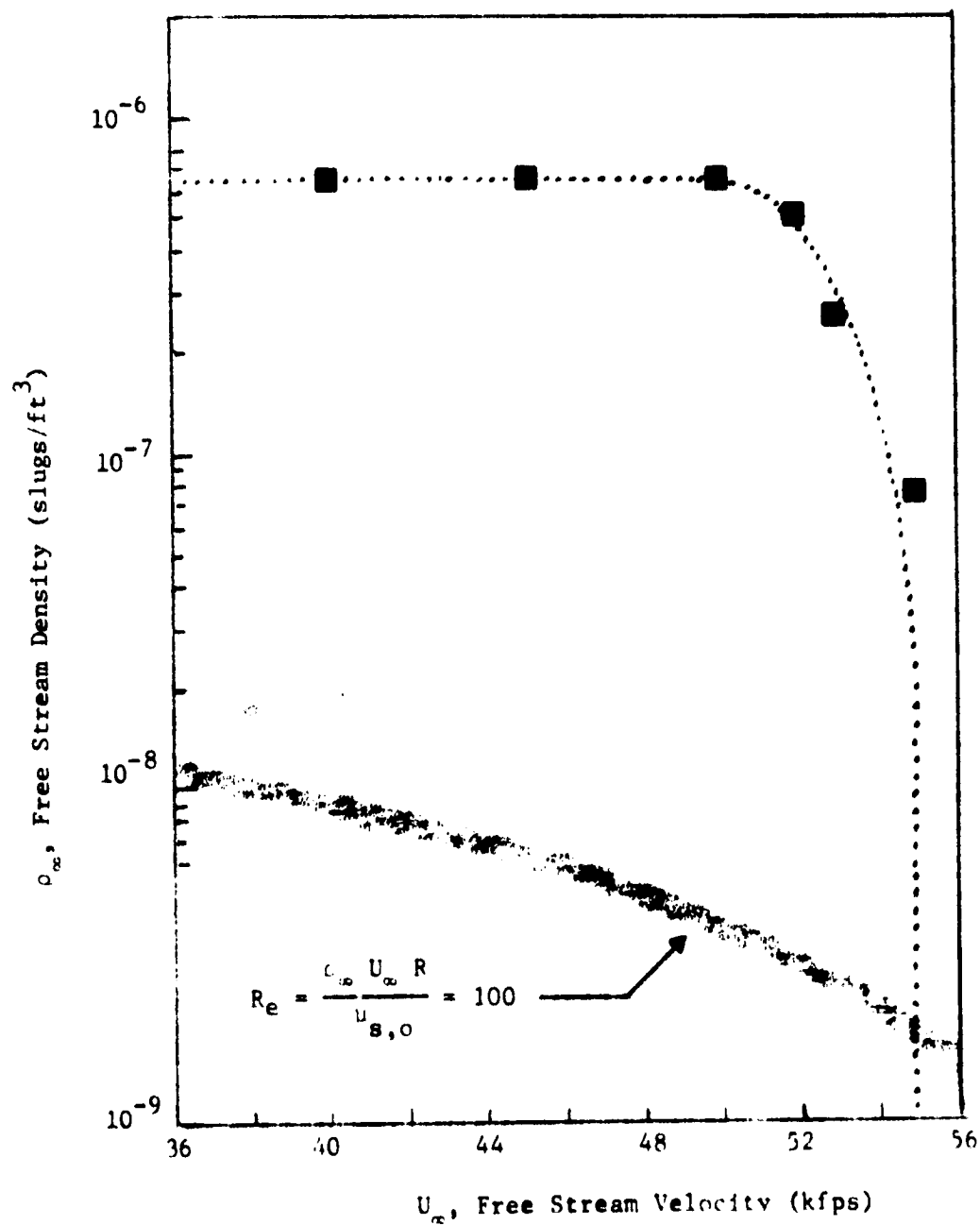


Figure 6.2. Reentry Trajectory for Present Investigation.

Note in Figure 6.1 that, as the vehicle first encounters the earth's atmosphere at a high rate of speed (55,000 feet per second), the heating load on the protective heat shield is low. As reentry proceeds, the capsule slows down but the load on the heat shield attains a maximum at around 50,000 to 51,000 feet per second. The heating rate then decreases as the capsule is slowed further. The reason for this phenomenon is that the upper atmosphere first encountered at around 247,000 feet is an extremely low density layer ( $7.5 \times 10^{-8}$  slugs/c.ft.). The amount of frictional heat developed in this region is therefore only minor and the capsule is slowed only slightly. However, as the vehicle encounters the denser earth atmosphere at lower altitudes, the frictional heating becomes more intense. On this trajectory the capsule is slowed from 53,000 feet per second to 50,000 feet per second while traversing densities of  $2.5 \times 10^{-7}$  to  $6.5 \times 10^{-7}$  slugs per cubic foot. At an altitude of 193,000 feet where the density is  $6.5 \times 10^{-7}$  slugs per cubic foot, the capsule is then slowed rapidly to a point where the heating load becomes much less severe. Results show that the stagnation heating rate is only 900 BTU/sq. ft.-sec. at a velocity of 40,000 feet per second and an atmospheric density of  $6.5 \times 10^{-7}$  slugs per cubic foot -- 41% of the heating rate at 50,000 feet per second. The heating rate will fall off even more as the capsule slows further at this same density (see trajectory curve of Figure 6.2).

In Figure 6.1 the primary heat absorption mechanisms of conduction and sublimation of carbon along with reradiated energy are shown for the energy reaching the ablator surface. The magnitude

of all the heat absorption mechanisms increases as total heat absorption increases but the fraction of energy accommodated by each mechanism changes through the trajectory.

At the low heat load conditions experienced at 55,000 ft/sec., for example, sublimation only accounts for 4% of the heat absorbed, (Figure 6.3), conduction accounts for about 11% and reradiation is the dominant heat transfer mechanism in reradiating 85% of the heat reaching the ablator surface. Conversely, as the heat load increases, the fraction of energy reradiated decreases and the mechanisms of conduction and sublimation account for a much more substantial portion of the heat absorption. For instance, at the maximum heating rate conditions experienced during the trajectory (about 50,000 ft/sec. and 193,000 feet), the energy reradiated is only 30% of the total energy reaching the capsule surface. Conduction and sublimation account for 42% and 28% respectively at these peak heating conditions. This fact means that the heat shield is losing mass at a much faster rate. The mass loss rate is shown in Figure 6.4. It is seen to be characteristically similar to the heat load profile.

These stagnation line results are the first attempt made which incorporate both nonequilibrium chemistry and detailed line and continuum radiative transport in the calculation at severe heat loads. Until now all analyses performed have been for equilibrium or non-radiating flow. The results depict the solution along a typical reentry trajectory for manned interplanetary missions. Points were chosen on the trajectory which represent typical heat loads on an ablative heat shield at the stagnation point. Previously



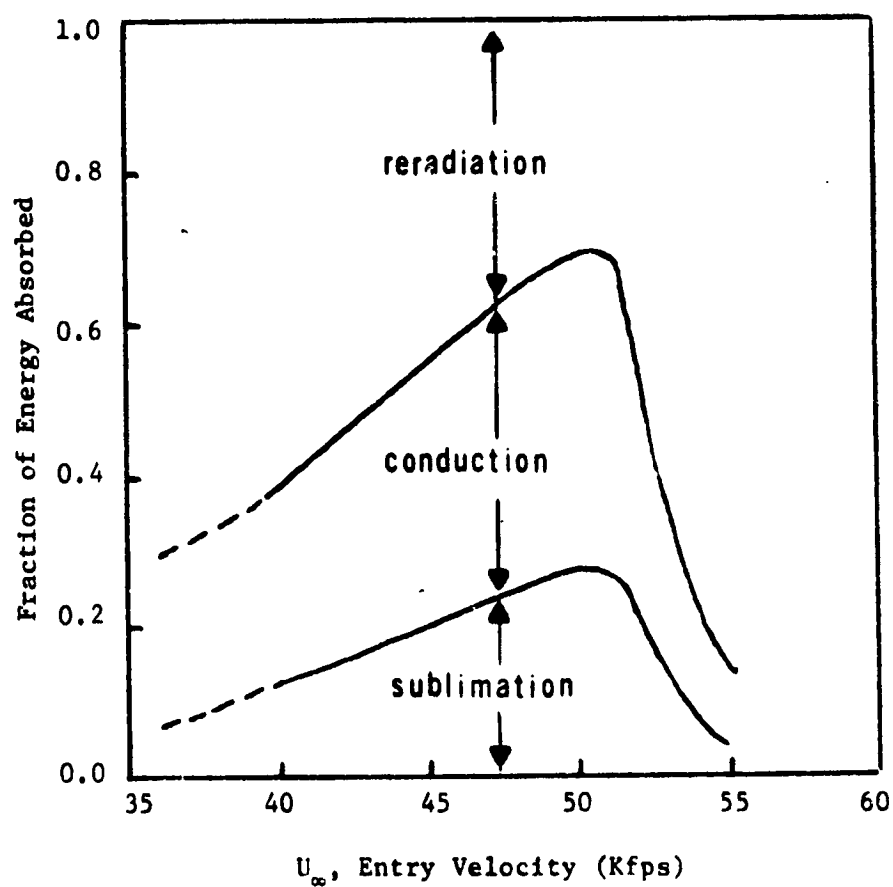


Figure 6.3. Effectiveness of Heat Absorption Mechanisms for a Phenolic Nylon Ablator Along the Reentry Trajectory.

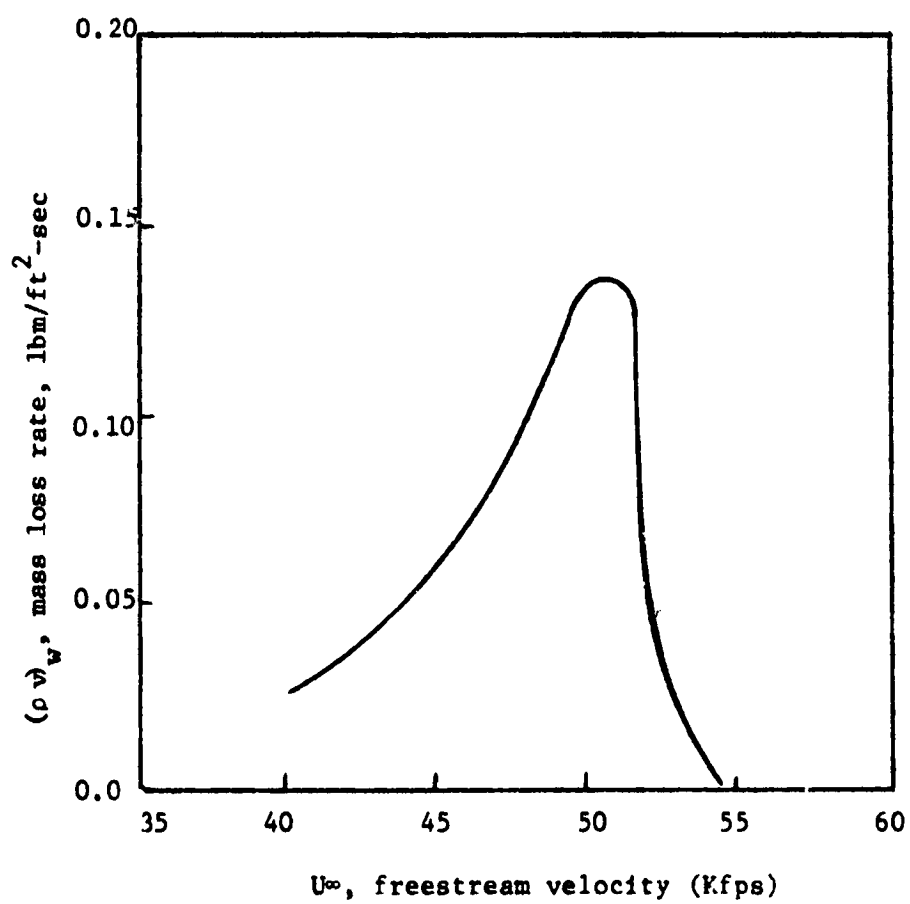


Figure 6.4. Mass Loss Rate at the Stagnation Line of a Phenolic Nylon Ablator Along the Reentry Trajectory.

analyses were performed along similar trajectories using the simpler equilibrium chemistry approximations. Nonequilibrium analyses were generally not attempted in this regime since the results are not accurate when radiation heating is not included. This analysis is also the first to couple a nonequilibrium shock layer to a nonequilibrium charring ablator.

The following section shows a comparison of the heating results from the present investigation with equilibrium results from previous investigators. A comparison is also made with a low heating load case of Perez (Ref. 6.1) which included nonequilibrium chemistry and radiative heating.

#### Comparison with Other Stagnation Line Investigators

In Table 6.3 is a list of results for the present investigation along with a comparison with previous investigators for conditions of low radiative heating to the ablator. These results are shown for the stagnation line with a phenolic-nylon ablator.

The present investigation used finite rate chemistry whereas, with the exception of one case run by Perez (Ref. 6.1), the rest of the analyses assumed equilibrium chemistry. The reason for this comparison with equilibrium chemistry is that there are no other finite rate analyses including radiative heat transfer. The only conditions for which a direct comparison can be made with Esch (Ref. 6.2), Engel (Ref. 6.3), and Perez (Ref. 6.1) is a freestream velocity of 50,000 feet per second, a freestream density of  $8.85 \times 10^{-8}$  slugs per cubic foot, and a blowing rate of 5% of freestream with a phenolic nylon ablator.

Table 6.3

Comparison of Moderate Heating Rate Results at the Stagnation Line  
(Phenolic Nylon Ablator)

$U_{\infty} = 50,000 \text{ ft/sec}, \quad \rho_{\infty} = 8.85 \times 10^{-8} \text{ slugs/ft}^3, \quad \rho v_w = 0.05$		
Reference	Chemistry	$q_R$ (BTU/ft <sup>2</sup> -sec)
Present investigation	Nonequilibrium	121
Perez (Ref. 6.1)	Nonequilibrium	121
Perez (Ref. 6.1)	Equilibrium	340
Esch (Ref. 6.2)	Equilibrium	349
Engel (Ref. 6.3)	Equilibrium	305

To verify the validity of the present analysis a comparison was made with the nonequilibrium results of Perez (Ref. 6.1). Both analyses predict 121 BTU/sq ft-sec for the case shown in Table 6.3 which is excellent agreement. The analysis of Perez can be validated by comparing his equilibrium results at these same conditions to the work of Esch (Ref. 6.2) and Engel (Ref. 6.3). His radiative heating results ( $q_R = 340$  BTU/sq ft-sec) compare very favorably with the analyses of Esch ( $q_R = 349$  BTU/sq ft-sec), and Engel ( $q_R = 304$  BTU/sq ft-sec). Perez used identical equilibrium and nonequilibrium numerical solutions of the conservation equation to those of Esch and Engel except for the solution of the species equations. Actually, a comparison with Esch is more meaningful since in both these analyses the species equations were actually solved whereas Engel used a step function model of elemental composition across the diffusion zone (discussed in Chapter II).

Perez's equilibrium results differ with Esch by less than 3% for a very favorable agreement. However Perez's equilibrium results are about 190% higher than the nonequilibrium results. This was the first time nonequilibrium chemistry and radiative heating effects were included simultaneously in a shock layer solution. The results show that nonequilibrium chemistry effects can significantly effect the heating rate to a phenolic nylon ablator during reentry.

Direct comparisons between investigators are difficult if not impossible at high heat loads for several reasons. Analyses have been performed for widely different radiation and chemistry models, and different ablator composites. The result is that lack of agreement is not only probable but expected.

For higher heating loads at denser atmospheric conditions, the results from Case 3 of the present investigation were used for comparison with equilibrium analyses of other investigators as shown in Table 6.4. The blowing rate of  $\rho v_w = 0.1$  shown in Table 6.4 was a case run in the process of attempting to couple Case 3 to the ablator. It should be noted that the results reported in Table 6.4 are for solutions which are not coupled to the ablator by a surface energy balance with the exception of Chin (Ref. 6.4). At 52,000 feet per second, this case has a higher freestream velocity than the other cases shown at 50,000 feet per second but because of the lower freestream density ( $5.0 \times 10^{-7}$  slugs per cubic foot), the kinetic energy is within 7% of the cases run at  $5.3 \times 10^{-7}$  slugs per cubic foot which means that the energy absorbed in both cases should be the same order of magnitude. This is the case, and Chin obtained  $1445 \text{ BTU/ft}^2 \text{ sec}$  as compared to  $1820 \text{ BTU/ft}^2 \text{ sec}$  in this work.

The nonequilibrium results at these high heating rates could be postulated to compare quite favorably with equilibrium analyses since it is theorized that equilibrium is approached at these severe (high temperature) conditions. The radiative heating results of Garrett (Ref. 6.5) at a blowing rate of 0.1,  $q_R = 2455 \text{ BTU/sq ft-sec}$ , are about 35% higher than the heating results of the present investigation at the same blowing rate,  $q_R = 1820 \text{ BTU/sq ft-sec}$ . Although the nonequilibrium heating rate is appreciably lower than equilibrium for this case, the difference is much less than the previous low heating rate case (35% vs 190% difference). This fact lends

Table 6.4  
Comparison of Severe Heating Rate Results at the Stagnation Line

$U_\infty$ (kfps)	$\rho_\infty$ (slugs/ft <sup>3</sup> ) $\times 10^{-7}$	$\rho V_\infty$	$q_R$ (BTU/ft <sup>2</sup> -sec)	Ablator	References
52	5.0	0.1	1820	PN*	Present investigation
50	5.3	0.1	2455	CP	Garrett et al. (Ref. 6.5)
50	5.3	0.2	2253	CP	Garrett et al. (Ref. 6.5)
50	5.3	0.2	2350	CP	Rigdon et al. (Ref. 6.6)
50	5.3	0.2	2856	CP	Smith et al. (Ref. 6.7)
50	4.5	0.2	1816	PN	Esci (Ref. 6.2)
50	4.5	0.2	1661	PN	Engel (Ref. 6.3)
50	3.45	0.076	1445	CP	Chin (Ref. 6.4)
50	3.45	0.076	1925	CP	Smith et al. (Ref. 6.7)
50	3.45	0.076	1423	CP	Garrett et al. (Ref. 6.5)

PN = Phenolic Nylon, CP = Carbon Phenolic

\* Nonequilibrium

credibility to the assumption that equilibrium chemistry can be used as an accurate approximation at the higher heating rates.

The equilibrium analyses of Garrett (Ref. 6.5) predicts a radiative heating rate of 2253 BTU/sq ft-sec at a reentry velocity of 50,000 feet per second, an atmospheric density of  $5.3 \times 10^{-7}$  slugs per cubic foot, and a wall blowing rate of 0.2. Rigdon, et al. (Ref. 6.6) attained similar results in predicting a radiative heating rate of 2350 BTU/sq ft-sec at these same conditions. The difference in these two analyses is only about 4%. In contrast to these results, Smith, et al. (Ref. 6.7) studied these identical conditions and determined a radiative heating rate of 2856 BTU/sq ft-sec which is 22% greater than either Garrett or Rigdon, et al. The major reason for this discrepancy appears to be the radiation model used by these investigators. Smith, et al. have neglected to include atomic line absorption for the highly absorbing carbon species. This fact accounts for the over-prediction of heating rate by Smith, et al.

A comparison of the present nonequilibrium results with Garrett and Rigdon, et al. shows that the nonequilibrium prediction is about 24% lower than these equilibrium results. A greater difference is seen between the nonequilibrium case and Smith, et al. but the reason for this is that carbon atomic lines were included in the nonequilibrium radiation analyses whereas Smith, et al. omitted this absorption mechanism in his analysis.

A comparison of the present results with the equilibrium analyses of Esch (Ref. 6.2) and Engel (Ref. 6.3) are also shown in Table 6.4. Flight conditions for the equilibrium runs were 50,000



feet per second,  $4.5 \times 10^{-7}$  slugs per cubic foot, and 0.2 blowing rate. Again these flight conditions differ from the nonequilibrium case but are close enough for comparison purposes. The radiative heating rate predicted by Engel was 1661 BTU/sq ft-sec which is 8.7% lower than the 1820 BTU/sq ft-sec from the present analysis. Esch determined the radiative heating rate to be 1816 BTU/sq ft-sec yielding only a 0.2% difference. The comparison with Esch is more realistic since he actually solved the equilibrium species continuity equations whereas Engel assumed a step function of elemental compositions across the stagnation point. The fact that blowing rates are different (0.2 for equilibrium and 0.1 for present nonequilibrium) does not affect the results very significantly (probably less than 10%) since Engel demonstrated in a parametric study that the effect of increased blowing rate above 0.1 is not very great on radiative heating. The agreement between the present analysis and Esch is then seen to be excellent despite the differences in conditions.

The results presented at 50,000 feet per second,  $3.45 \times 10^{-7}$  slugs per cubic foot, and 0.076 blowing rate were included for completeness. These conditions are much less stringent than the 52,000 feet per second,  $5.0 \times 10^{-7}$  slugs per cubic foot, and 0.1 blowing rate for the nonequilibrium case. The radiative heating rates of Chin (Ref. 6.4), 1445 BTU/sq ft-sec, and Garrett, et al., 1423 BTU/sq ft-sec, are very close to each other in contrast to the 1925 BTU/sq ft-sec predicted by Smith, et al. The kinetic energy for the equilibrium cases is seen to be 36% lower than the nonequilibrium case (1820 BTU/sq ft-sec) but the heating rates are only about 25% lower for Chin and Garrett, et al., and 5.8% higher

for Smith, et al. Again the reason for the large discrepancy between Smith, et al. and other investigators is believed to be due to differences in radiation data.

In summarizing, radiative heating predictions from the present nonequilibrium investigation are identical to those of Perez (Ref. 6.1) who did a nonequilibrium study at 50,000 feet per second and  $8.85 \times 10^{-8}$  slugs per cubic foot. Perez's numerical solution was verified by comparing his equilibrium results at these same conditions to the equilibrium results of Esch (Ref. 6.2) and Engel (Ref. 6.3) yielding less than 11% difference with Engel and less than 3% difference with Esch.

At higher heat load conditions, comparisons were made with equilibrium solutions at slightly different conditions. In general the results from the present investigation agree quite well considering the fact that equilibrium and nonequilibrium chemistry were used and flight conditions were not identical. Heating rate predictions for nonequilibrium chemistry in the present investigation is consistently lower than equilibrium predictions by other investigations. This phenomenon is due to the differences in species profiles across the shock layer as will be shown later in a discussion on kinetics.

#### Around the Body Results

Six coupled solutions were obtained at the stagnation line using nonequilibrium chemistry and were presented in Figure 6.1 and Tables 6.1 and 6.2. Of these six cases, four cases were continued downstream using the stagnation line results as initial

conditions. The shock layer, was coupled to the ablator surface at all points around-the-body using the blowing rate profile. Heating rate profiles could then be compared with ablator response heating rates to insure a coupled solution. Only four of the six stagnation line cases were run because of the immense computer time required. For instance, Case 4 took four hours of CPU time to reach a downstream point of  $\xi = 0.30$ . A more complete discussion of computer time is provided at the end of the chapter.

In Figure 6.5 a plot of nondimensional blowing rate versus around the body distance is presented. The blowing rate is nondimensionalized with the stagnation line blowing rate,  $(\rho v)_{\xi = 0}$ . Dimensionless heating rates are presented in Figures 6.6 through 6.9 for cases 1 through 4 using the blowing rate profile in Figure 6.5. The heating rate is nondimensionalized with the stagnation heating rate,  $q_{SL}$ , for convenience. The ablator-flow field coupling was found to be excellent, differing by less than 3% along the wall using the one generalized blowing rate profile with the exception of Case 1 where the flow field and ablator solutions diverge downstream of the stagnation line where a difference of 16.3% was observed. The blowing rate profile could have been altered until the ablator-flow field solutions matched for Case 1 but this would have required more computer time (approximately four to six hours of IBM/360/65 CPU time). The expense of the computer time was deemed too great in view of the fact that Case 1 is not a high heating rate case since the stagnation line heating rate is only 22% of the maximum heating rate investigated (Case 4). In addition, the difference in heating rate at the downstream

point of  $\xi = 0.45$  between ablator and shock layer solutions is only 16.3% which is not that great a deviation.

The shock layer results for the various cases run were obtained farther downstream for some than others. The reason for this is that some cases are more difficult computationally than others and therefore computer time becomes a major factor. The high heating rate cases of Case 3 and Case 4 were the most time consuming. For Case 3, a downstream value of  $\xi = 0.35$  was reached, where for Case 4, a downstream value of only 0.30 was reached. Over four hours of computer time was used for each of these cases to reach these downstreams points. The heating rate severity appears to be a measure of the computational difficulty.

An interesting result of the around the body analysis is that one blowing rate profile can be used to describe all of the points on the reentry trajectory reasonably satisfactorily. What this means is that given the mass injection rate at the stagnation line, an estimate of the downstream mass loss rate can be made using Figure 6.5. The total mass injection rate,  $\dot{m}$ , can be estimated using the following relationship

$$\dot{m} = (\rho v)_{SL} 2\pi \int_0^{\xi} \phi_w \xi d\xi$$

which is the mass flux from the ablator integrated over the ablator surface area. The mass flux at the stagnation line  $(\rho v)_{SL}$ , can be determined from Figure 6.4 and the nondimensionalized mass flux around the body,  $\phi_w$ , is the function in Figure 6.5. The value of the integral is a constant. By knowing the time elapsed at each point along the trajectory, the total mass loss for reentry,  $m$ , can

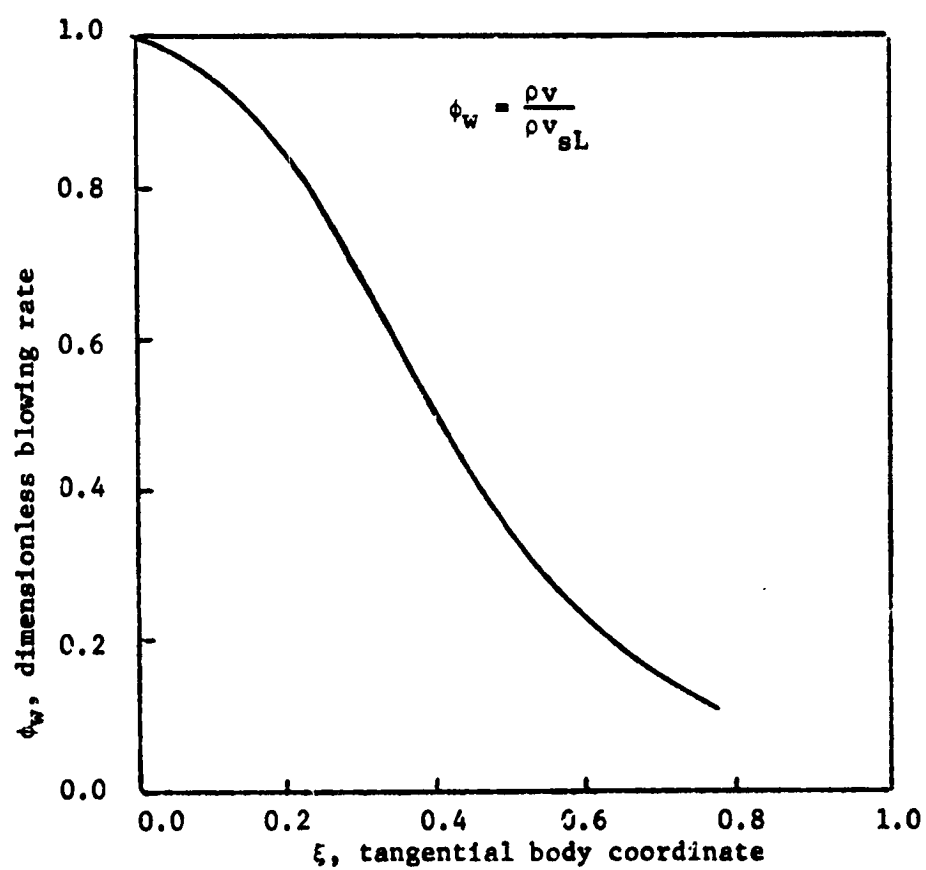


Figure 6.5. Dimensionless Blowing Rate Profile Around the Body.

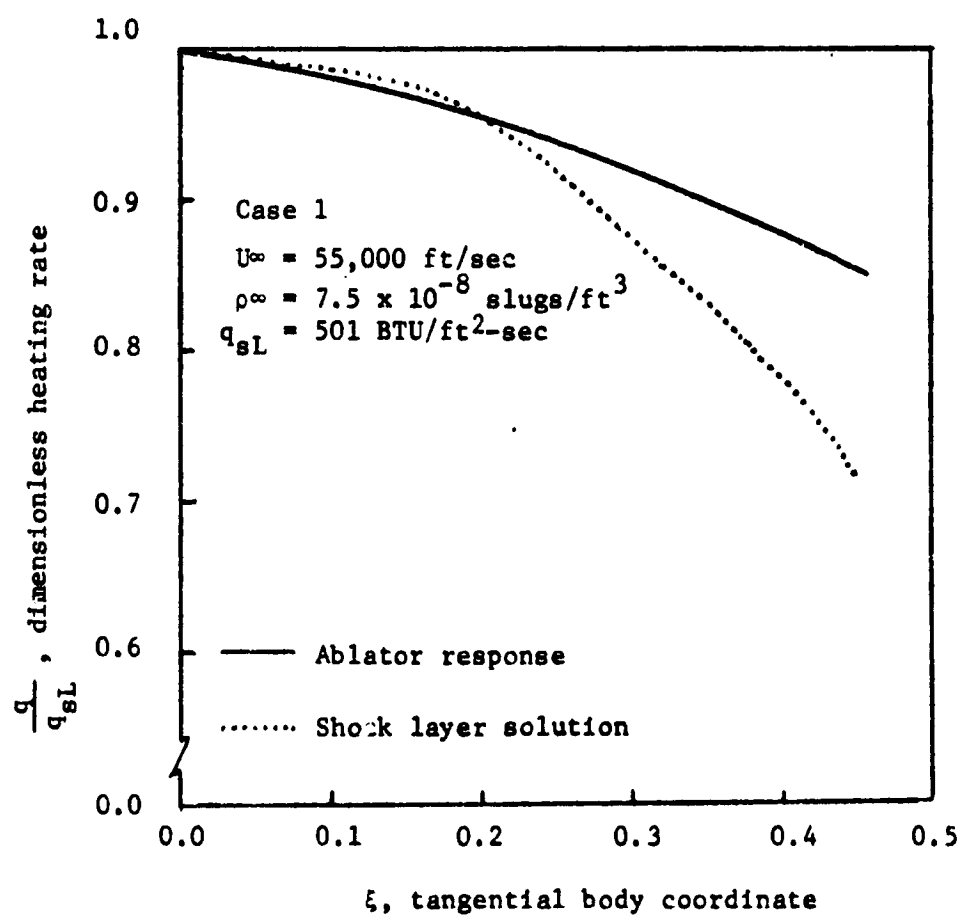


Figure 6.6. Dimensionless Heating Rate Profile  
For Case 1 - Coupled

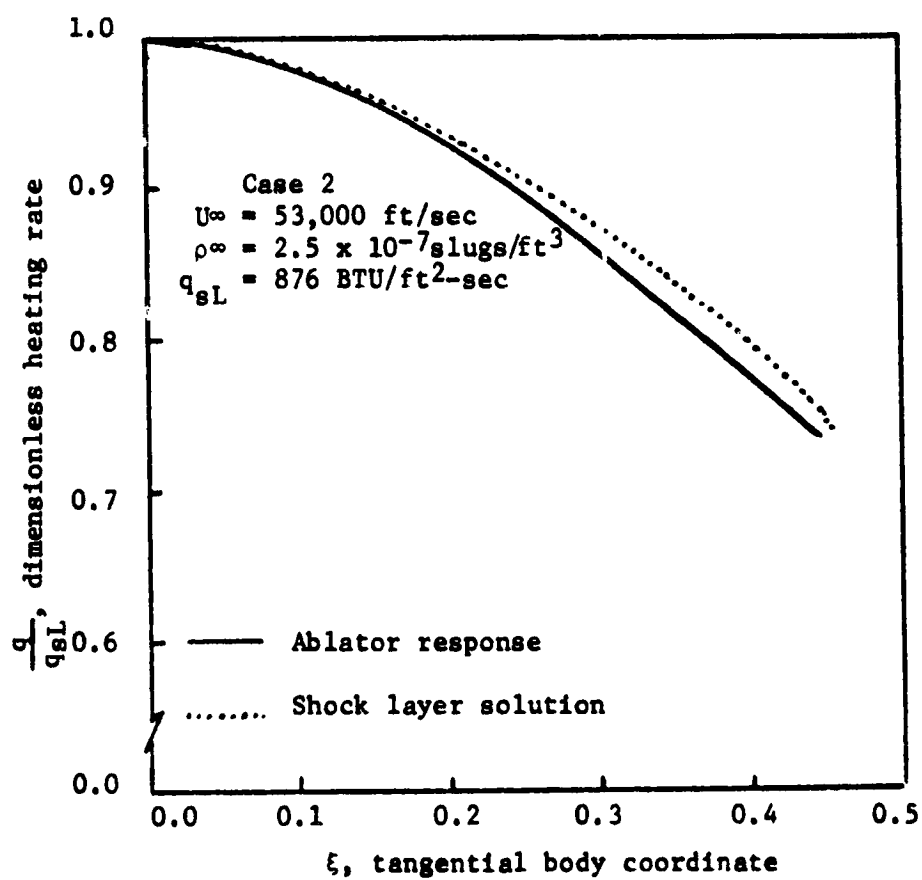


Figure 6.7. Dimensionless Heating Rate Profile for Case 2 - Coupled.

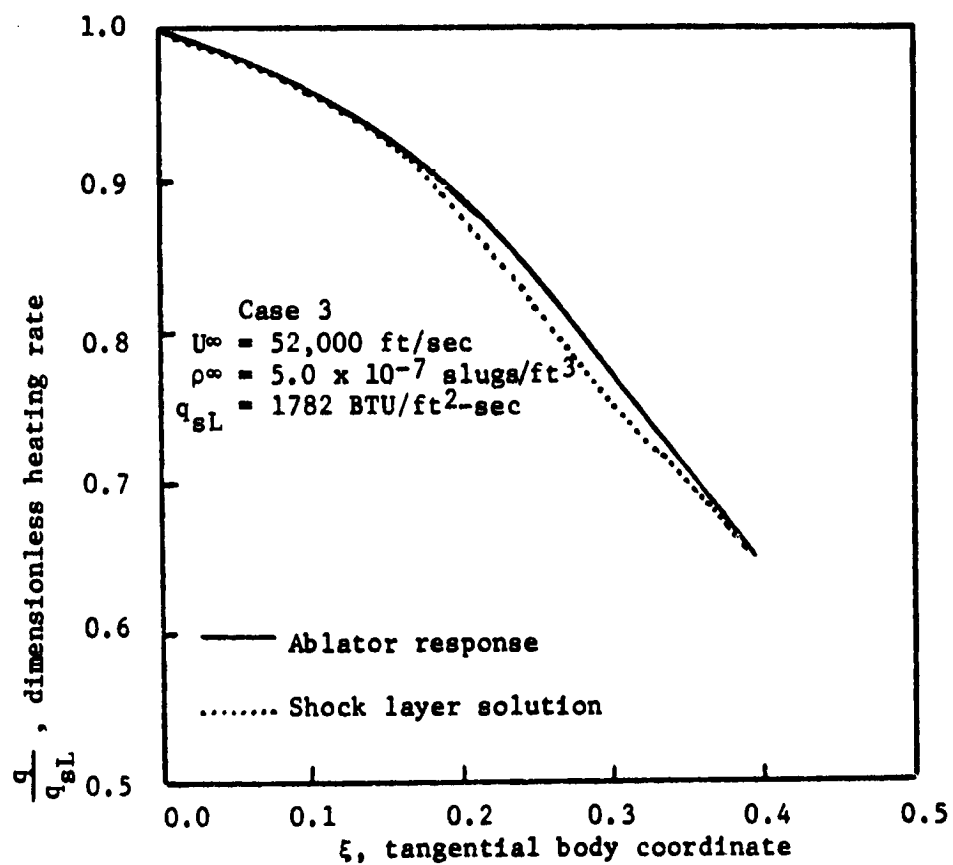


Figure 6.8. Dimensionless Heating Rate Profile for Case 3 - Coupled



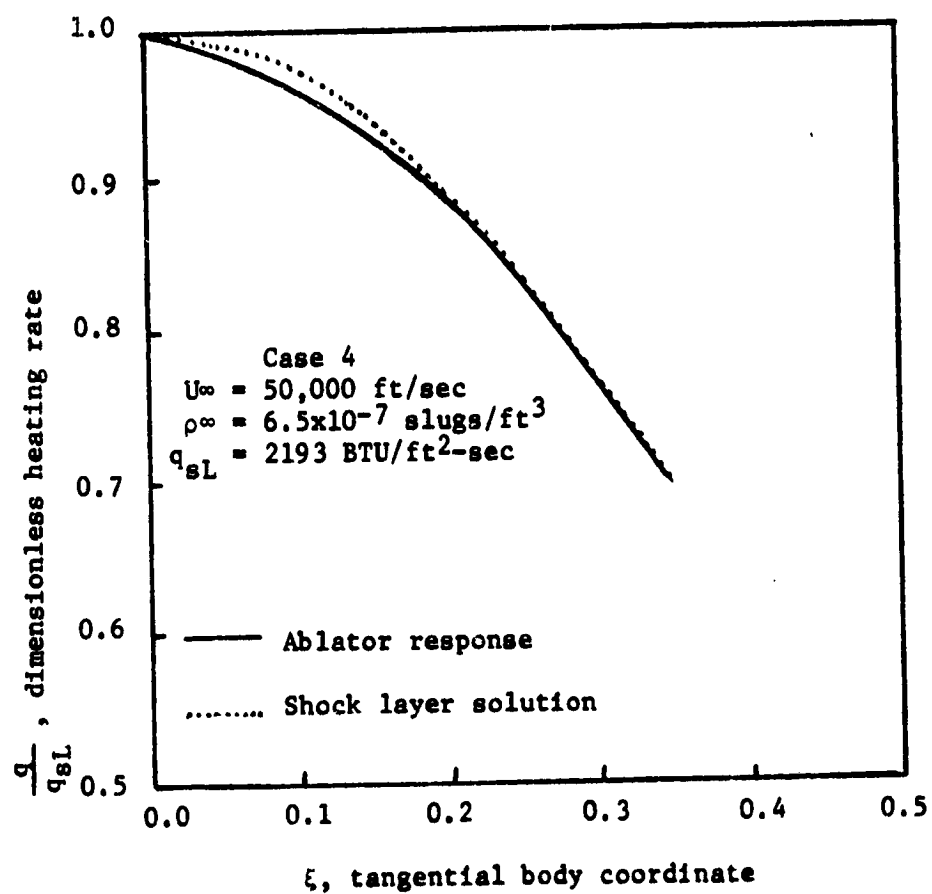


Figure 6.9. Dimensionless Heating Rate Profile for Case 4 - Coupled.

be estimated as

$$m = \int_0^t \dot{m} dt$$

or

$$m = 2\pi \int_0^{\xi} \phi_w \xi d\xi \int_0^t (\rho v)_{SL} dt$$

The result of the solution of the integral equation is the total mass loss from the ablative heat shield from the stagnation point to a radius  $\xi$  between initial contact with earth's upper atmosphere and some elapsed time "t". All that is required now is a time plot for the trajectory and the solution could be completed.

In Figure 6.10 the wall surface temperature profiles are shown for the four around the body cases. These surface temperatures were calculated using the Hertz-Knudsen analysis described in Appendix B. The wall temperature is seen to decrease gradually as the solution progresses downstream. This decrease in temperature is due to the two factors of decreased surface pressure and decreased heating rate.

Temperature profiles for stagnation line and downstream points  $\xi = 0.30$  and  $\xi = 0.45$  are presented in Figure 6.11 for Case 3. The profiles are very similar although the temperature gradient at the wall increases slightly as the solution proceeds downstream. The same kind of results were obtained for the other cases.

A comparison of around the body heating results is given in Figure 6.12 between Case 4 of the present analysis and those of Chou and Blake (Ref. 6.8) and Smith et al. (Ref. 6.7). The comparison is made at different conditions shown in the figure but the heating rate has been nondimensionalized with the stagnation line

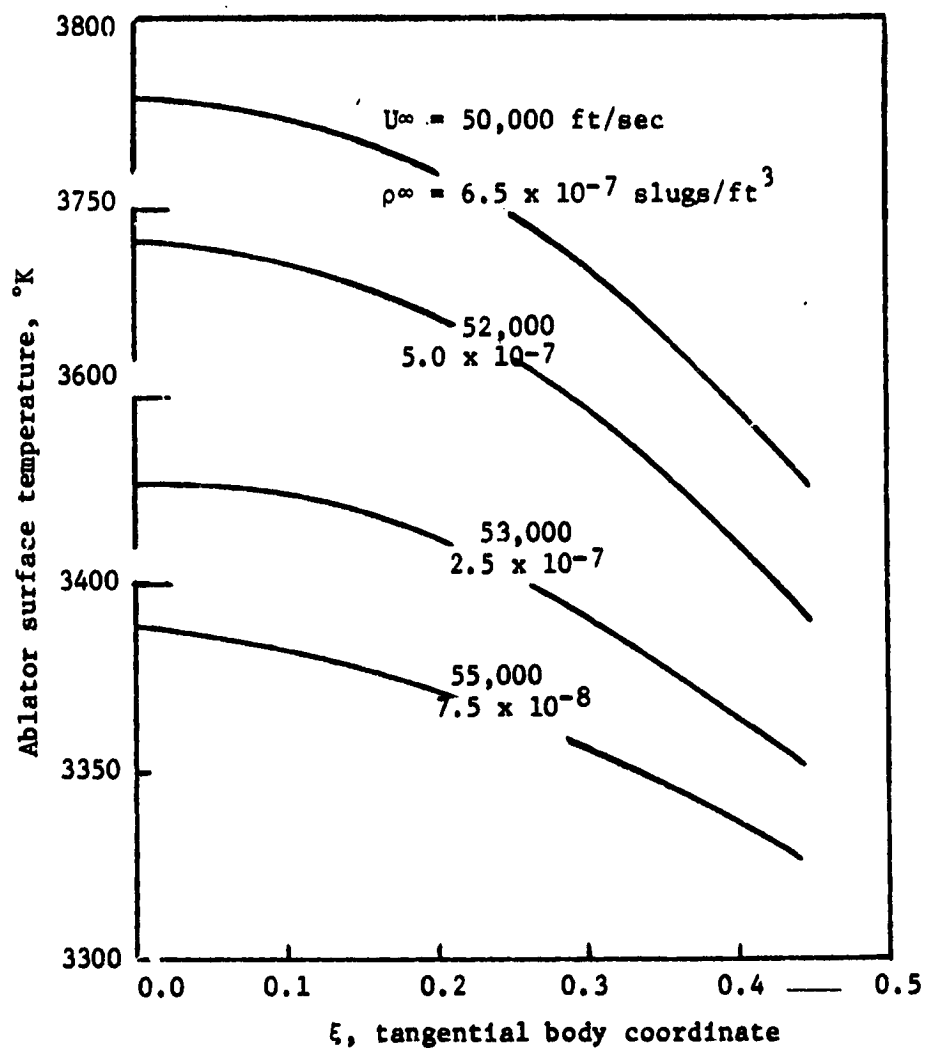


Figure 6.10. Ablator surface temperature profiles for cases studied.

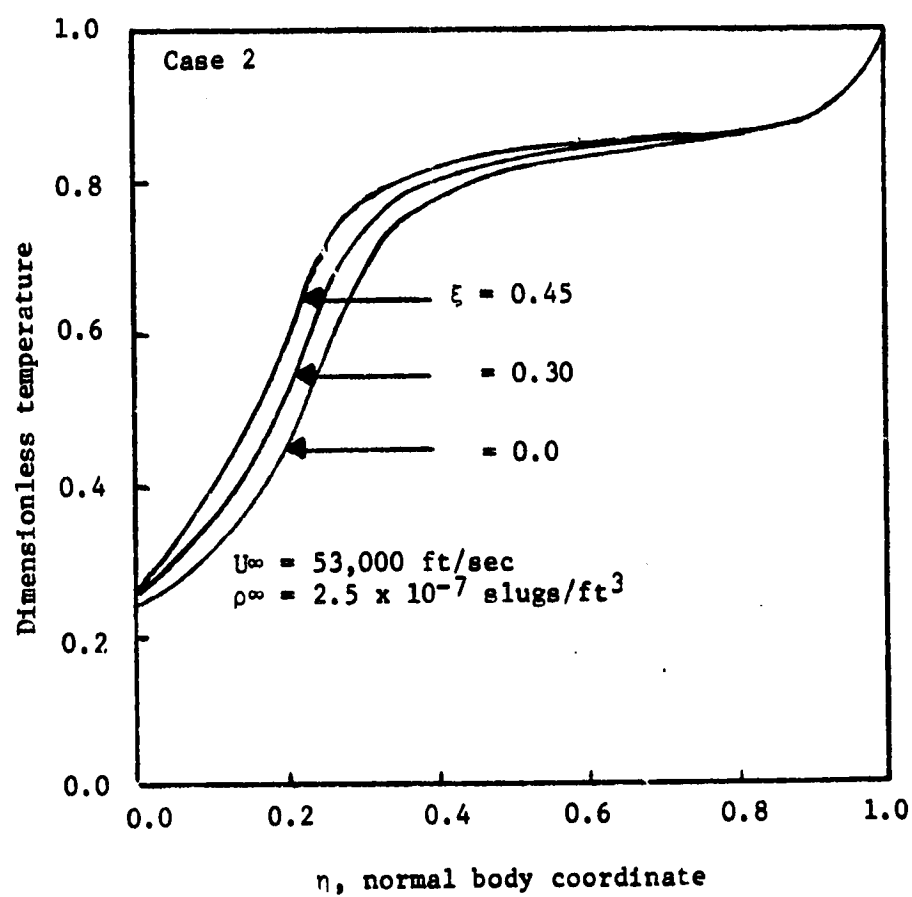


Figure 6.11. Dimensionless Temperature Profiles for the Stagnation Line and Downstream.

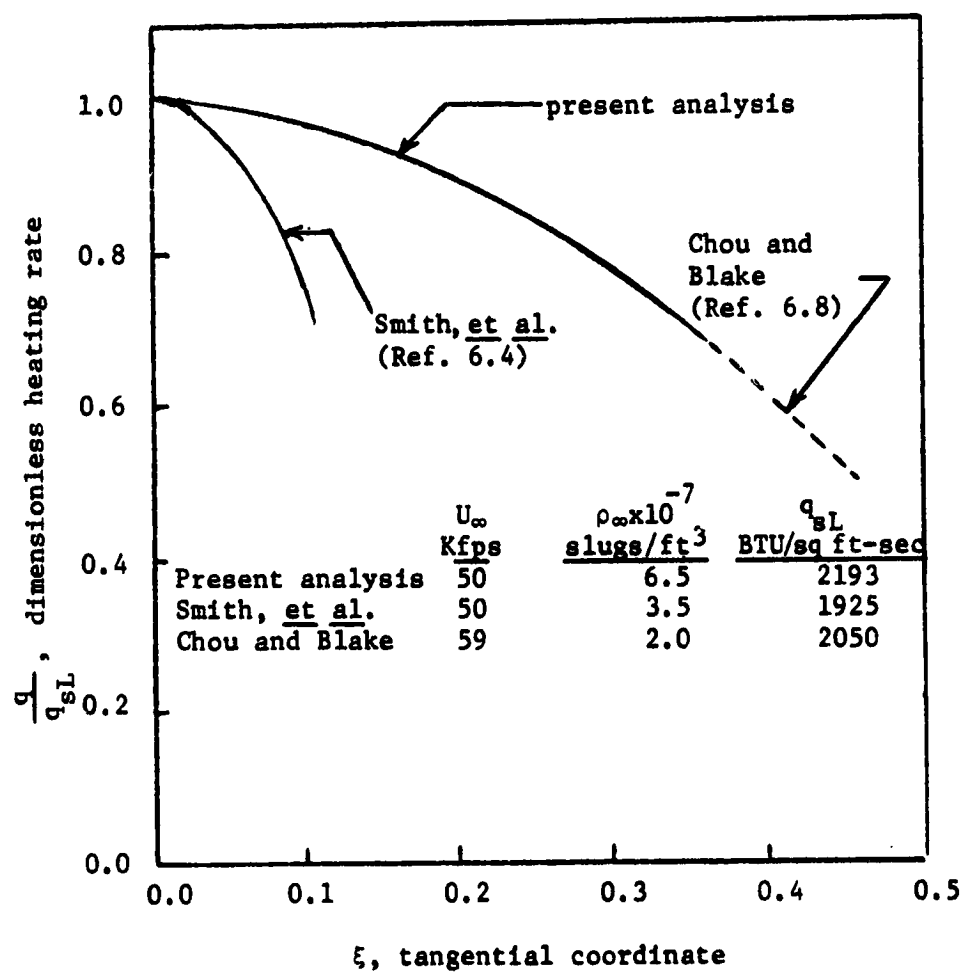


Figure 6.12. Comparison of Dimensionless Heating Rates Around the Body

heating rate. The results of Chou and Blake and the present analysis are identical as the solution proceeds downstream (with the results of Chou and Blake extending farther downstream). In contrast to these results, Smith et al. predicts a much more rapid decrease in heating rate away from the stagnation line. The heating rate has already been reduced to 70% of the stagnation line value at  $\xi = 0.1$  whereas the present analysis along with Chou and Blake do not decrease to 70% until a body station of  $\xi = 0.35$  is reached. Differences occur in radiation properties included for the three analyses, along with differences in numerical solutions, and chemistry (nonequilibrium for the present analysis vs. equilibrium for the others). It is not clear which of these effects or what combination of these effects causes the similarities and differences in results presented. It is somewhat surprising that Chou and Blake and the present analysis agree so well actually since the analysis of Chou and Blake did not include wall blowing.

These around the body results are the first to be presented for a reentering space vehicle with the inclusion of finite rate chemistry and line and continuum radiation. Although some previous investigators have coupled their flow fields to the ablator, this is the first attempt to couple a nonequilibrium flow field to a nonequilibrium ablator. Previous investigators have used equilibrium sublimation models to yield temperature and species concentration at the surface. The present investigation employed the nonequilibrium Hertz-Knudsen sublimation model described in Appendix B and the nonequilibrium char analysis described in Appendix A to determine surface boundary conditions and further couple them with

the flow field by an energy balance. The results presented have been for the stagnation point and around the body for points along a typical reentry trajectory from planetary missions. These results make feasible a realistic estimate of the total mass loss during reentry and therefore provide the necessary information for calculation of the ablator heat shield thickness.

#### Radiation - Chemical Reactions Interaction

The results presented in the previous sections for stagnation line and around-the-body solutions were based on a set of chemical reactions and chemical species for the flow field given in Table 4.1. The choice of species and chemical reactions to include was based upon equilibrium and nonequilibrium considerations along with computer limitations. The choice of these reactions and species was discussed briefly in Chapter IV but the choice of kinetic data is an altogether different problem.

Although the reactions for the flow field had been chosen, the task of deciding what kinetic data to use remained. For some reactions data was reported by several investigators - sometimes in disagreement with one another. Other reactions had data from only one source, and for still others there was no data. When one source of data was found, this was used; but when many were found, the one that appeared to be most reputable was used. The most difficult case was when no data was found. In this instance the data had to be estimated from theory.

For estimated data, two constants had to be evaluated in the rate constant shown below - the frequency factor,  $k_0$ , and the

activation energy, E.

$$k = k_0 T^{1/2} \exp [-E/RT]$$

Initially no rate data was found for the reactions listed in Table 6.5. These are nine out of a total of 16 reactions and these nine are ablation species reactions. For these nine reactions the frequency factor was calculated using collision theory (Ref. 6.10) and the activation energy was estimated using the method of Semenov (Ref. 6.9). The activation energy estimation appeared quite satisfactory since the estimated values agreed with data when it was available.

The collision theory frequency factor, however, was often orders of magnitude different from published data. In Table 6.5 three sets of frequency factors are labeled best, fast perturbation, and slow perturbation. Chronologically, the first values used were the fast perturbation frequency factors which were calculated when collision theory was applied in error. When the frequency factor calculation was corrected, the slow perturbation values were achieved. Data was subsequently found in the literature for the first three reactions in Table 6.5 (Refs 6.10, 6.11) which was not in agreement with the estimated frequency factors ( $\sim 10^{14}$  vs  $10^{10}$ ). Based on the data for these three reactions the decision was made to alter all of the nine collision theory frequency factors from  $10^{10}$  to  $10^{14}$ . The values thus obtained were decided to be the best frequency factor estimates available and were used for all coupled stagnation line and around the body results.



Table 6.5  
Frequency Factors for Three Finite Rate Cases  
for Selected Reactions of Ablation Species in the Shock Layer

Reaction	Frequency Factor, log k <sub>0</sub>		
	Best Kinetics	Slow Perturbation	Fast Perturbation
$\text{CO} + \text{N} \xrightarrow{+} \text{CN} + \text{O}$	13.2	10.0	18.9
$\text{C}_2 + \text{M} \xrightarrow{+} 2\text{C} + \text{M}$	13.8	10.1	19.0
$\text{C} + \text{M} \xrightarrow{+} \text{C}^+ + \text{e}^- + \text{M}$	14.3	10.1	19.0
$\text{HCN} + \text{M} \xrightarrow{+} \text{CN} + \text{H} + \text{M}$	13.8	10.0	18.9
$\text{C}_2\text{H} + \text{M} \xrightarrow{+} \text{C}_2 + \text{H} + \text{M}$	13.8	10.0	18.9
$\text{C}_2\text{H}_2 + \text{M} \xrightarrow{+} \text{C}_2\text{H} + \text{H} + \text{M}$	13.8	10.0	18.9
$\text{C}_3 + \text{M} \xrightarrow{+} \text{C}_2 + \text{C} + \text{M}$	13.8	10.0	18.9
$\text{C}_3\text{H} + \text{M} \xrightarrow{+} \text{C}_2\text{H} + \text{C} + \text{M}$	13.8	10.0	18.9
$\text{C}_4\text{H} + \text{M} \xrightarrow{+} \text{C}_3\text{H} + \text{C} + \text{M}$	13.8	10.0	18.9

As a consequence of these changes a collection of results were in hand with different values of the frequency factor for nine of the 16 reactions. This information was used to give an estimate of the effects of uncertainty of kinetic data on radiative heating. Three sets of data were available at moderate heating conditions of 50,000 feet per second,  $8.85 \times 10^{-8}$  slugs per cubic foot, a wall blowing rate of 5% of freestream, and a wall surface temperature of 3450°K at the stagnation line. The ablator was a phenolic-nylon composite. These conditions and ablator were the same as those studied by Esch (Ref. 6.2) who used an equilibrium chemistry model. The study was extended to include a comparison of the best kinetics with equilibrium at more severe heating conditions.

It would have been more desirable to conduct a parameter study on all frequency factors and activation energies. However, since one case requires approximately 60 minutes of computer time it was felt that the available results would serve this purpose. In all, these runs required three hours of computer time.

Moderate heating conditions in ablation layer: For the case of moderate heating to a phenolic nylon ablator, the freestream conditions of 50,000 feet per second and  $8.85 \times 10^{-8}$  slugs per cubic foot were used. Wall blowing rate was 5% of freestream blowing and wall surface temperature was 3450°K. All three non-equilibrium chemistry cases were investigated and compared to chemical equilibrium of Esch (Ref. 6.2).

The basis of comparison is the radiative heating rate arriving at the ablator surface since this is the most important parameter in design of thermal protection systems. It will be shown that as a

result of variations in species concentration profiles in the shock layer for the different frequency factors of nine reactions the radiative heating can be appreciably altered. In the ablation layer the important species appears to be the highly absorbing carbon atom.

In Figure 6.13 the radiative flux divergence profiles are shown for the three reaction rate sets investigated and the equilibrium results of Esch (Ref. 6.2). The radiative wall heating rates for the cases are also shown in Figure 6.13 - best (121 BTU/sq ft-sec), slow (160), fast (90), and equilibrium (349). In Figures 6.14, 6.15, and 6.16 are the species profiles for the finite rate cases and equilibrium species profiles are presented in Figure 6.17.

The interaction between species and radiative heating in the ablation layer is quite pronounced. Note that there is a trend in radiative flux divergence profiles near the stagnation point ( $\eta=0.1$  to  $\eta=0.4$ ) for the four cases presented. The equilibrium model (the infinitely fast reaction rate) demonstrates a large dip in this region, the fast kinetics perturbation dips slightly less, the best kinetics shows no dip at all, and the slow kinetics perturbation actually reaches a local maximum in the stagnation region. The results indicate that, at increasingly faster reaction rates, the radiative flux divergence drops off due to some absorption mechanism in the stagnation region. For an explanation of this phenomenon, inspection of the species profiles in Figures 6.14 through 6.17 is helpful. Notice in particular the carbon atom profiles for the four cases. Both equilibrium and the fast kinetics perturbation show high peak mass fractions of carbon atom ( $\sim 0.5$  to  $\sim 0.6$ ). The best kinetics case shows a moderate peak for carbon atom (mass fraction

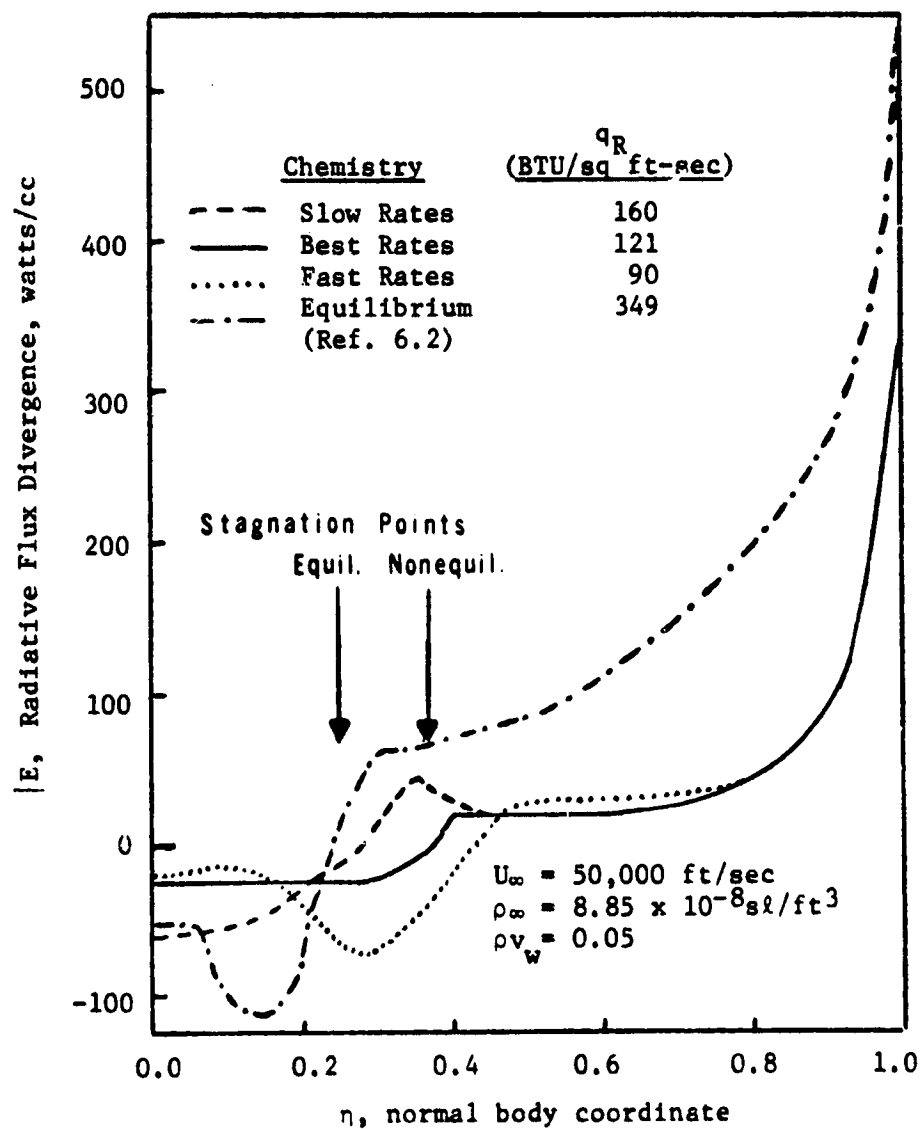


Figure 6.13. Comparison of Flux Divergence Profiles for Various Chemistry Models at the Stagnation Line.

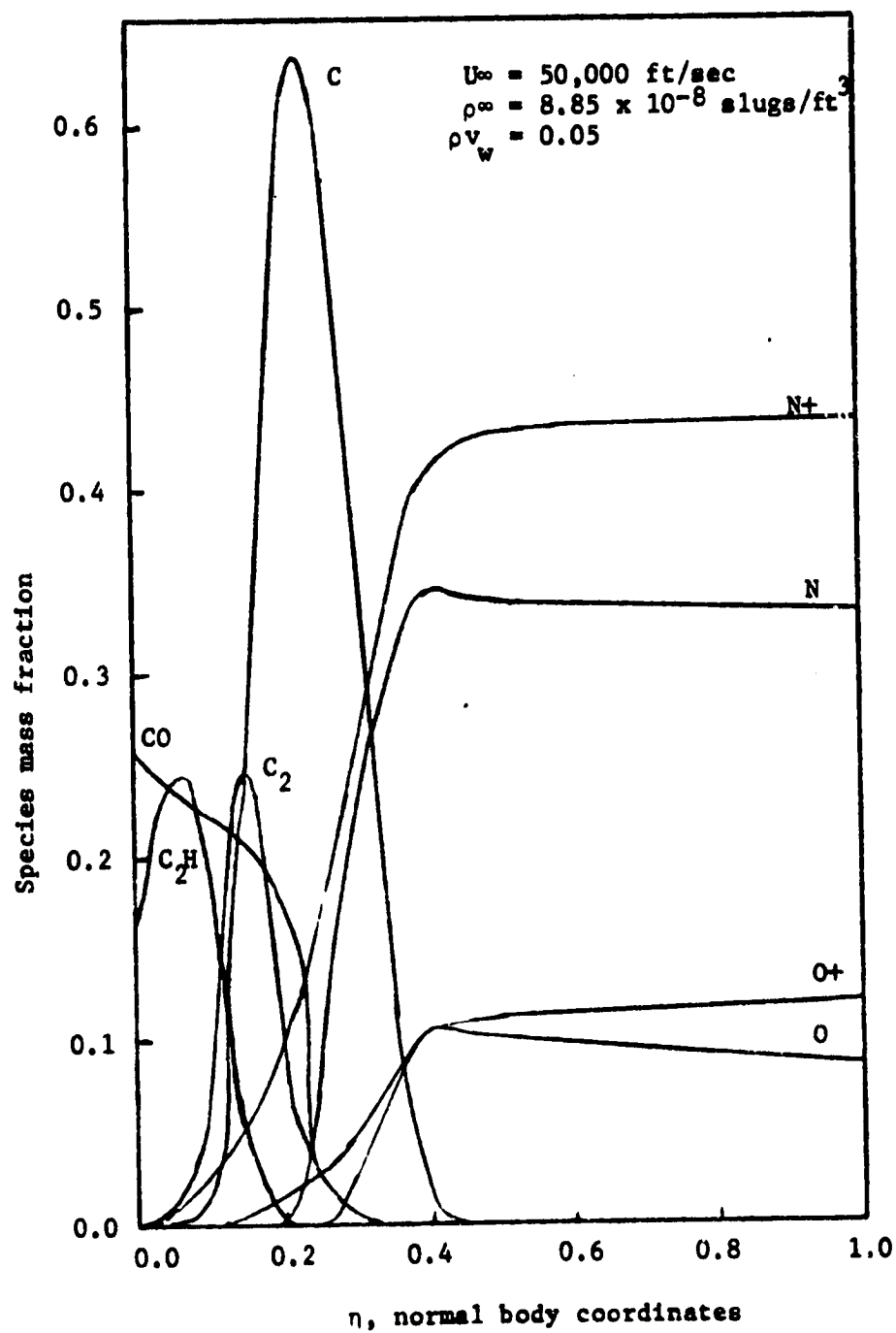


Figure 6.14. Species Profiles for "Fast Perturbation" Kinetics Case at the Stagnation Line.

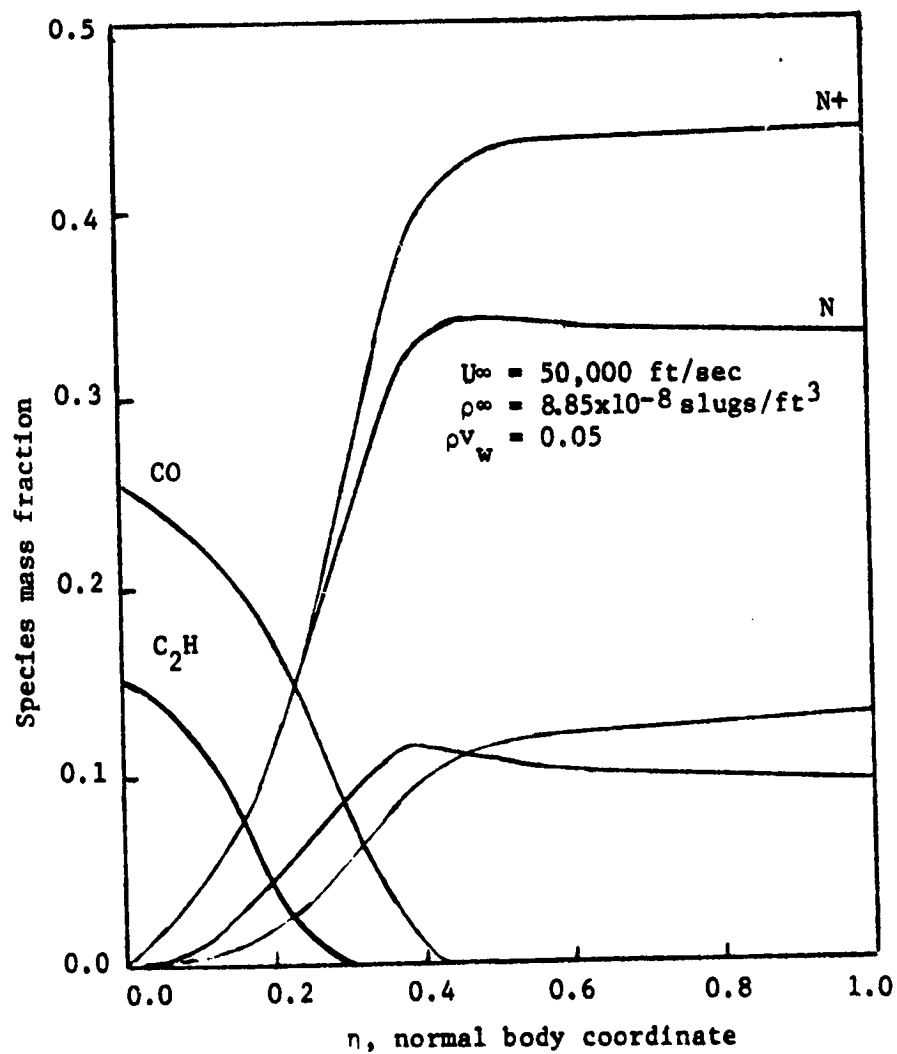


Figure 6.15. Species Profiles for "Slow Perturbation" Kinetics Case at the Stagnation Line.

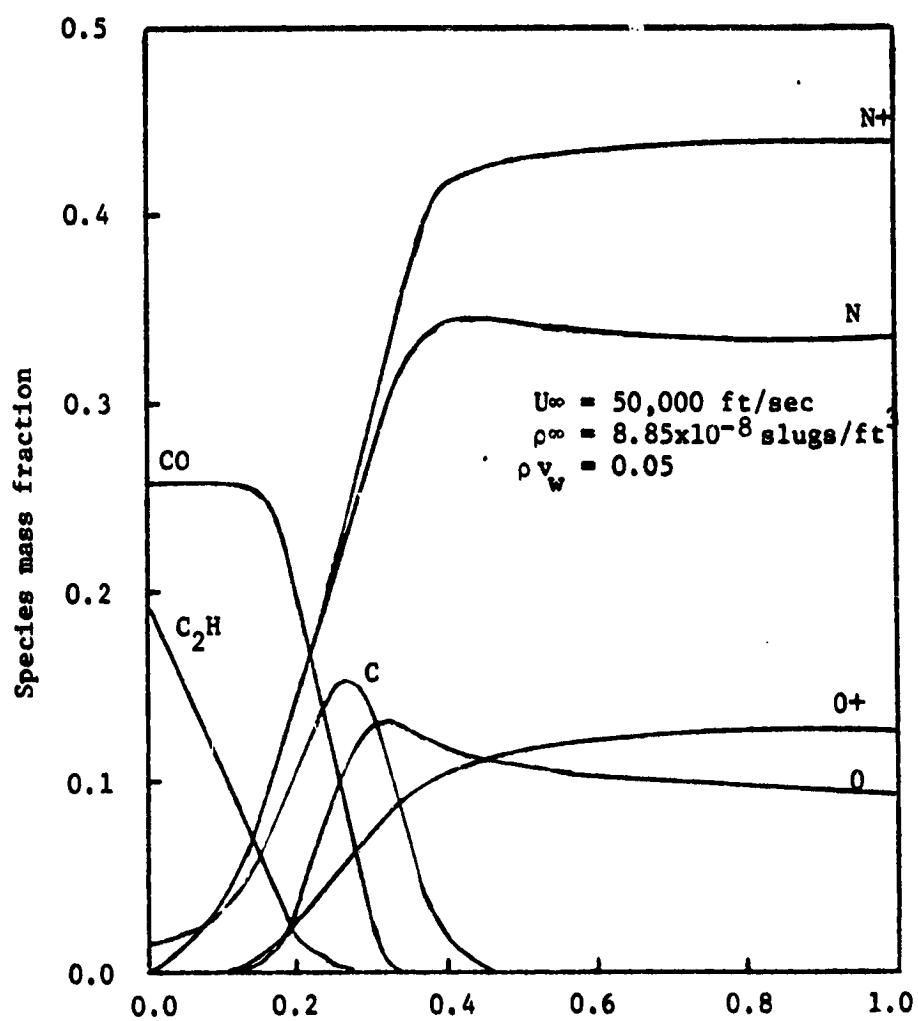


Figure 6.16. Species Profiles for "Bast" Kinetics Case at the Stagnation Line.

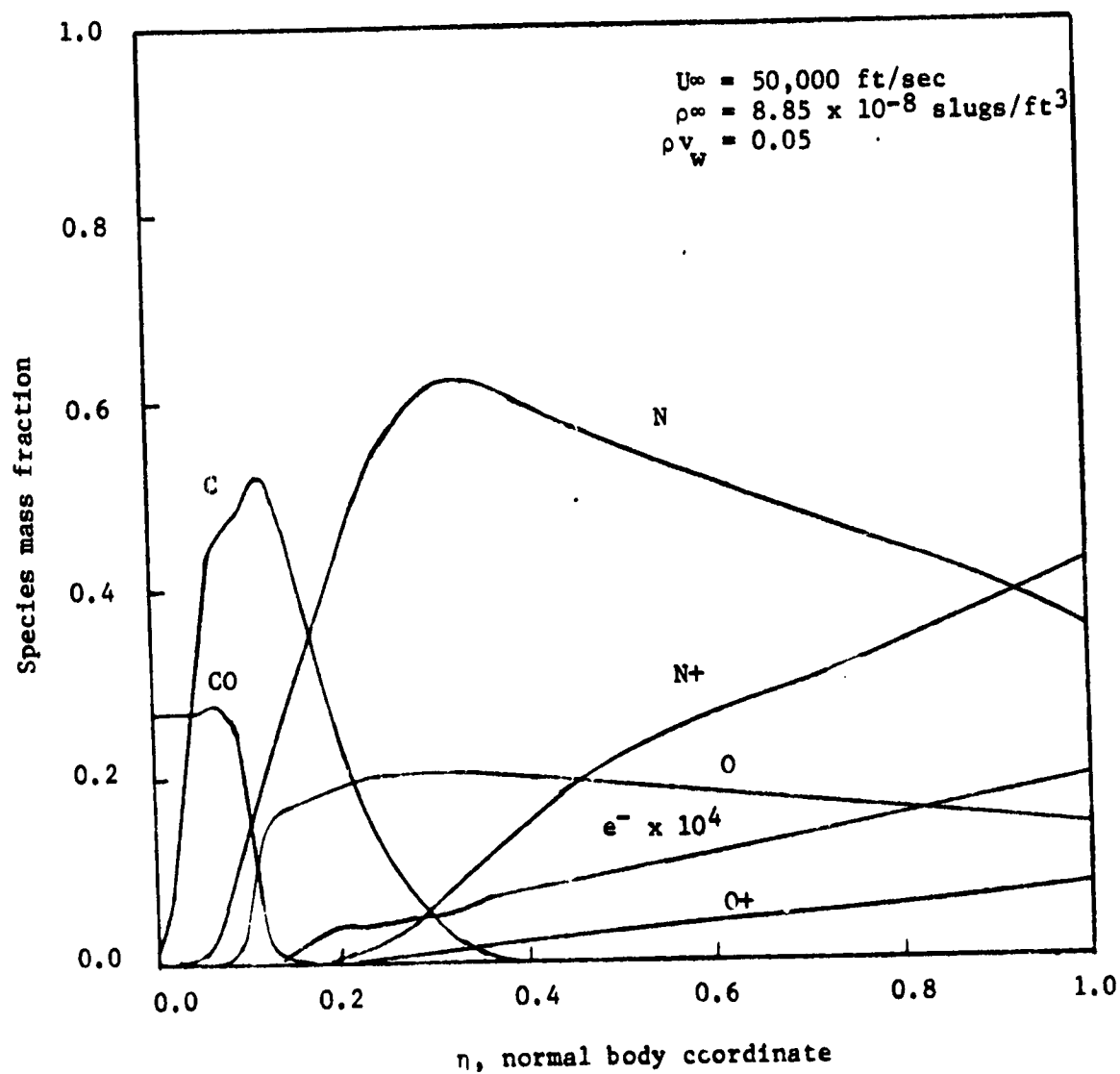


Figure 6.17. Species Profiles for Equilibrium Chemistry at the Stagnation Line (Esch, Ref. 6.2)



of 0.15), and in the slow perturbation kinetics case there was not a significant concentration of carbon atom anywhere within the shock layer. The key to explain the nature of the radiative flux divergence profiles near the stagnation point is these carbon atom profiles. Carbon is a highly absorbing species in this temperature range, therefore, high concentrations result in significant reductions in radiative heating as evidenced by the results presented in Figure 6.13. For the three kinetics models, the heating rate increases (90 BTU/sq ft-sec for fast, 121 for best, 160 for slow) as the reaction rate is decreased. The exception to the rule is the infinitely fast rate of equilibrium which has the highest heating rate, 349 BTU/sq ft-sec. This discrepancy will be discussed in the following section.

The importance of carbon atom radiative absorption can be corroborated with the results of Engel (Ref. 6.3). In Figure 6.18 are the temperature profiles and radiative heating rates obtained by Engel for three cases at the same conditions. The first case included all species in the radiative flux divergence calculation (1661 BTU/sq ft-sec), the second case excluded carbon atom (2530 BTU/sq ft-sec), and the third case omitted molecules (2605 BTU/sq ft-sec). The effect of which species are included in the radiation calculation are seen to have a dramatic effect on radiative heating at the ablator surface. Of special significance is the increased heating effect when carbon atom is omitted allowing a 52% increase. The same type of effect is realized with respect to slow and fast perturbation kinetics of the present investigation. By slowing down the reaction rate to a point where no atomic carbon is produced

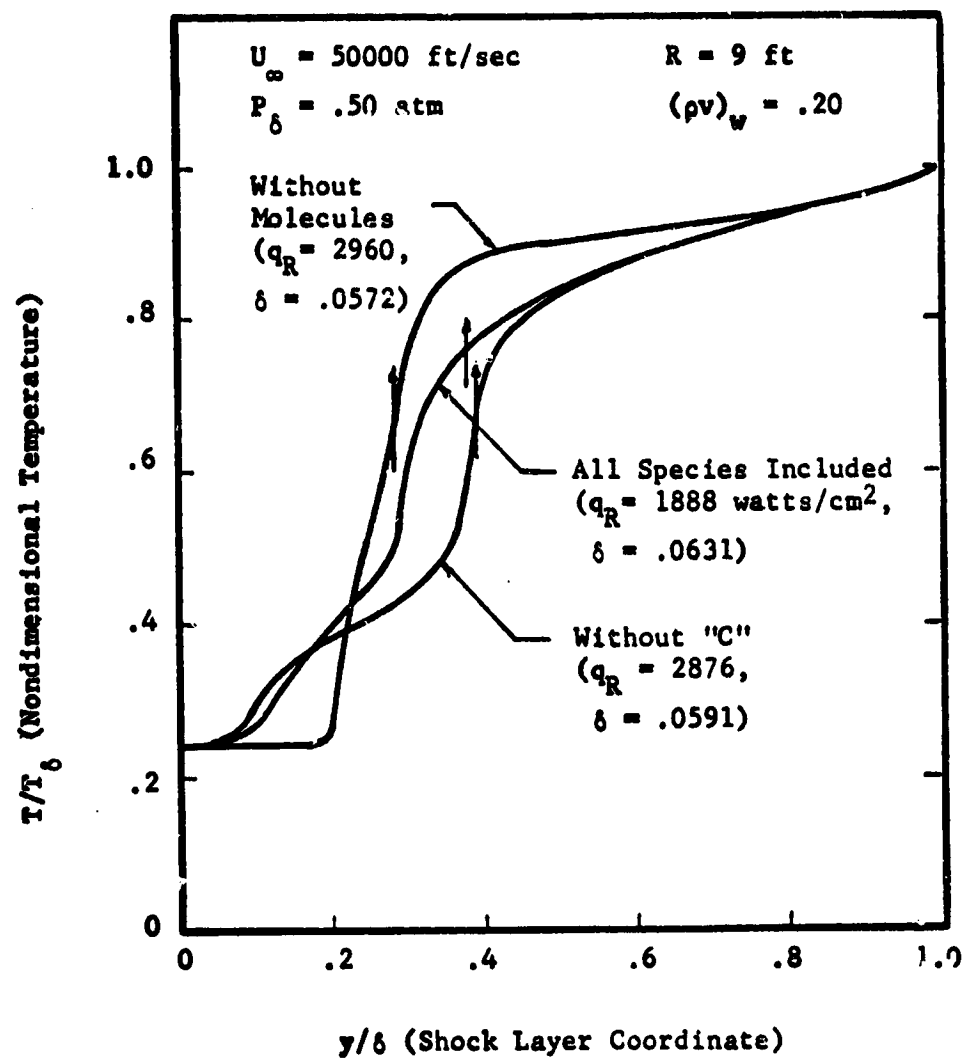


Figure 6.18. Effects of Carbon Atom and Molecular Radiative Coupling on the Temperature Profile and Radiative Heating (Ref. 6.3).

(slow perturbation kinetics, 160 BTU/sq ft-sec), a 78% higher heating rate is observed than the fast perturbation kinetics, 90 BTU/sq ft-sec. The best kinetics case, with an intermediate buildup of atomic carbon, results in an intermediate heating rate, 121 BTU/sq ft-sec.

In summary, the three kinetics cases of the present investigation and the equilibrium analysis of Esch follow a pattern in their ablation layer species profiles with atomic carbon increasing as reaction rate increases. The radiative flux divergence profiles of these cases reflect this trend near the stagnation point by demonstrating a significant reduction for cases where the highly absorbing carbon atom is present in high concentrations. The radiative heating rates of the three kinetics cases of the present investigation also reflect this trend by increasing as atomic carbon decreases corresponding to a decrease in reaction rate. The exception to this trend in heating rate is the equilibrium results which show the highest radiative heating rate although the atomic carbon concentration is high for this infinitely fast reaction rate. The reasons for the apparent contradiction are discussed in the following section.

Moderate heating conditions in air layer: In Figure 6.17 the species profiles from Esch's (Ref. 6.2) equilibrium analysis were presented. As discussed above, these results compare well with the results of the fast finite rate analysis in the ablation layer. Note, however, that the species mass fraction profiles in the air layer for chemical equilibrium are very much different than any of the finite rate analyses (Compare Figure 6.16 with Figures 6.13, 6.14

and 6.15). There appears to be little reaction occurring in the air layer for the nonequilibrium analysis (Figures 6.14, 6.15, and 6.16) if one inspects the flat nature of the  $N$ ,  $N^+$ ,  $O$  and  $O^+$  profiles. The  $N^+$  and  $O^+$  ion concentration levels remain high until they reach the diffusion zone near the stagnation point indicating no reactions are proceeding in the air layer. The equilibrium analysis of Esch produced species profiles as shown in Figure 6.17. The mass fractions of ionic species  $N^+$  and  $O^+$  decrease from shock to stagnation point much faster for equilibrium than for nonequilibrium. The equilibrium mass fractions of atomic species  $N$  and  $O$  compensate by increasing readily. In other words; at equilibrium reactions take place which alter the species profiles in the air layer. If the shock layer were truly near chemical equilibrium, the finite rate reactions would reflect this by occurring fast enough to experience similar profiles to the equilibrium profiles. They do not; therefore, one can conclude from these results that equilibrium is not a valid assumption even for this hot outer air layer -- at least at these conditions (50,000 feet per second and  $8.85 \times 10^{-8}$  slugs per cubic foot) and a nonequilibrium shock calculation should be performed.

The postulate that the shock wave is not at equilibrium, as assumed, is substantiated by the work of Carlson (Ref. 6.12). Carlson did a nonequilibrium calculation across the shock wave using an air chemistry model. He states that a high degree of chemical and thermal nonequilibrium can exist for low pressure conditions as in this study. The conclusion is made that these effects cannot be neglected if accurate solutions are desired. Zeldovich et al.

(Ref. 6.13) explains that it is not until 20,000°K that rapid atomic dissociation is achieved, and the rate of single ionization is still considered a slow mechanism at this time. Slow refers to a process which requires a large number of gas kinetic collisions. When the pressure is reduced, the number of collisions is reduced. The result is that the shock wave is only in partial equilibrium and probably has only achieved atomic dissociation at best.

The radiative heating implications are evident when an inspection is made of these species mass fractions in light of the radiative properties. The air layer is a predominantly emitting zone. The species that exist in the air layer are N, N<sup>+</sup>, O, O<sup>+</sup>, and e<sup>-</sup> of which radiative properties exist for N, O, and e<sup>-</sup>. In the equilibrium case ( $q_R = 349$  BTU/sq ft-sec) the emitting species of N and O atoms are present in higher concentrations than in the nonequilibrium cases ( $q_R = 160$  BTU/sq ft-sec). The concentrations of the nonemitting N<sup>+</sup> and O<sup>+</sup> ions and the emitting electron are present in high concentrations for the nonequilibrium cases. This shift from nitrogen and oxygen atom radiative properties to electron radiative properties for equilibrium and nonequilibrium respectively appears to have a significant effect on the radiative heating. Equilibrium predicts 190% higher radiative heating than nonequilibrium best kinetics. The result is a lower net radiative heating (emission) when the air layer is highly ionized.

In Figure 6.13 the radiative heating differences of the various chemistry models were dramatically demonstrated. The three nonequilibrium models show similarity in the vicinity of the air layer for both species profiles and radiative flux divergence. The

equilibrium model, on the other hand, depicts a much greater positive flux divergence in the air layer thus demonstrating its widely different species composition in this air layer (atomic rather than ionic).

The apparent paradox that equilibrium has the highest heating rate (349 BTU/sq ft-sec) and fast kinetics has the lowest heating rate (90 BTU/sq ft-sec) can now be explained. The reason for this is that "fast" applies only to some select reactions. These reactions all pertain to species in the ablation layer. Therefore, the air layer is practically unaffected by changes made for the three kinetics cases. This is the reason the radiative flux divergence and species profiles are almost identical in the air layer for the kinetics cases of the present investigation. The equilibrium analysis, in contrast, is seen to be dominated by the radiation from the air layer (Figure 6.13).

Severe heating rate conditions: In Figure 6.19 is a plot of species profiles at the stagnation line for the coupled solution of a severe heating rate case (Case 3) run using the frequency factors from the best kinetics case in Table 6.5. In this case a freestream velocity of 52,000 feet per second and a freestream density of  $5.0 \times 10^{-7}$  slugs per cubic foot were used. The radiative heating rate at the ablator surface for the coupled solution was 1782 BTU/sq ft-sec with a blowing rate of 12% of freestream. A comparison was made with an equilibrium case of Esch (Ref. 6.2) run at 50,000 feet per second,  $4.5 \times 10^{-7}$  slugs per cubic foot, and a blowing rate of 20% of freestream. The radiative heating rate for this case was 1816 BTU/sq ft-sec. These conditions are not exactly the

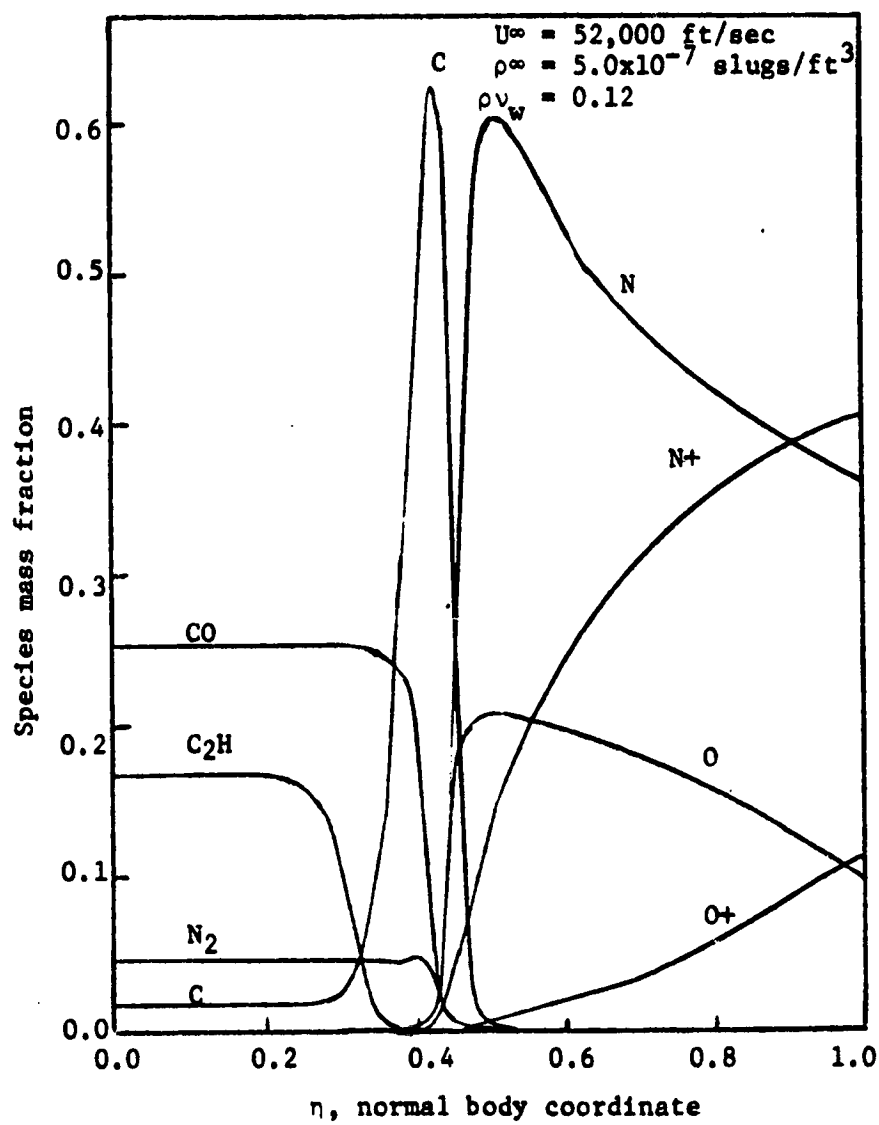


Figure 6.19. Species Profiles for Nonequilibrium Chemistry at a Severe Heating Load at the Stagnation Line.

same as the finite rate conditions, but they are close enough for comparative purposes.

A plot of species mass fraction profiles for Esch's equilibrium case is shown in Figure 6.20. The stagnation point of the non-equilibrium analysis (Case 3) in Figure 6.18 is about  $\eta = 0.4$  whereas Esch's equilibrium analysis has a stagnation point displaced further toward the shock ( $\eta = 0.6$ ). The reason for the difference in stagnation points is that Case 3 has a blowing rate of 12% and Esch assumed a blowing rate of 20%. Other than this difference the species profiles are remarkably similar -- even in the air layer where the moderate heating rate analysis showed a great deal of difference (Figures 6.16 and 6.17). For both the severe heating rate cases, atomic nitrogen and oxygen build up in concentration from shock to stagnation point and ionic species  $N^+$  and  $O^+$  decrease in this region. At the stagnation point itself, the concentration of all air species decreases due to diffusion. In the ablation layer the major species again have very similar profiles. Flat profiles are observed for  $CO$ ,  $C_2H$ , and  $N_2$ . Carbon atom attains a peak of about 0.6 mass fraction at the stagnation point and drops off rapidly on either side.

For years investigators have postulated that at high heating load conditions, the shock layer approaches chemical equilibrium and could therefore be modeled as such. The results presented here for the severe heating case only support this contention. The conditions in the hot outer air layer ( $14,500^\circ K$ ) are such that reactions occur fast enough to compare with equilibrium. The reason for this result is a combination of both high temperature



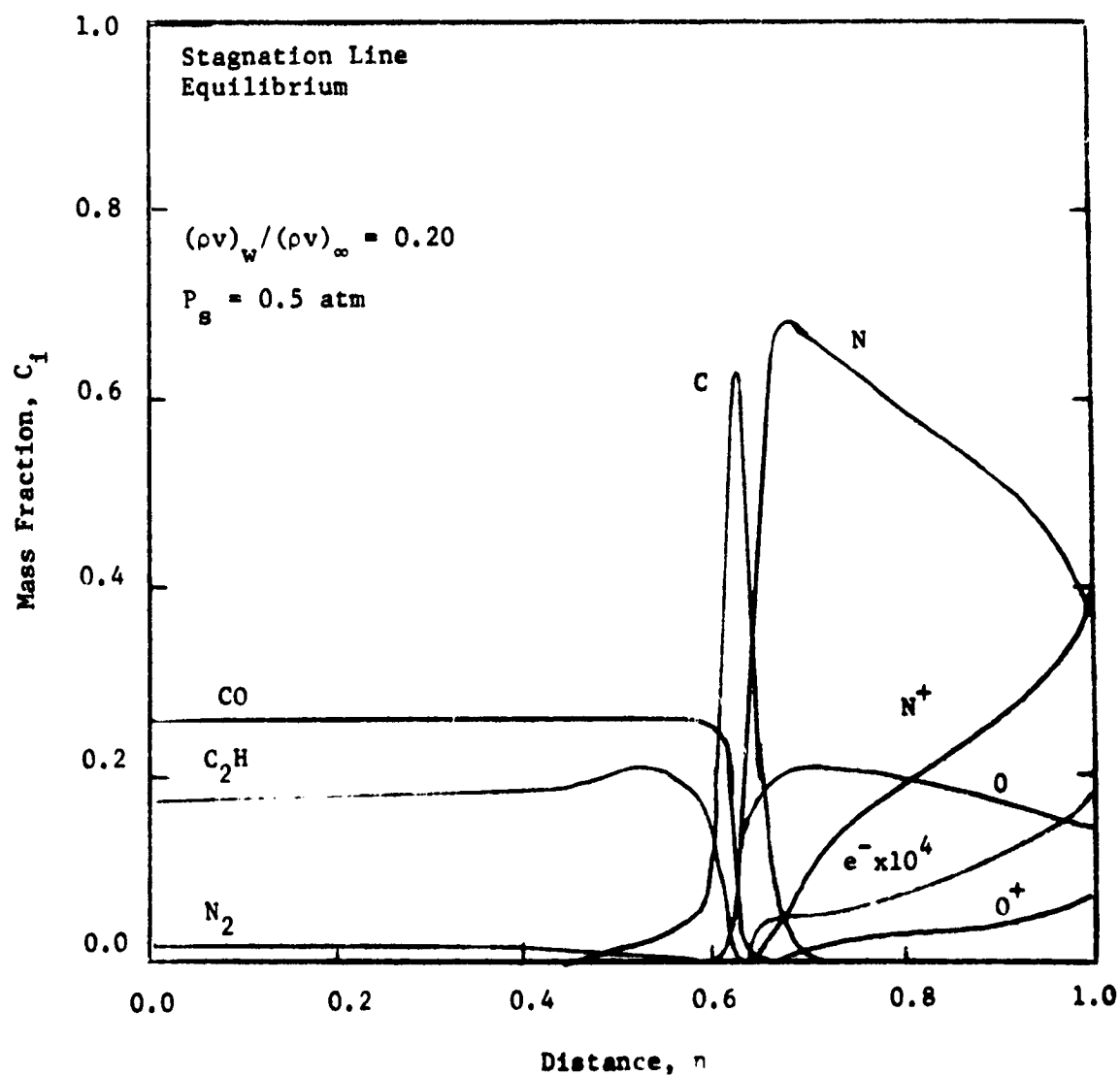


Figure 6.20. Species Profiles for Equilibrium Chemistry at a Severe Heating Load at the Stagnation Line. (Esch, Ref. 6.2)

and post shock pressure (0.5 atm). For the moderate heating cases discussed earlier, the post shock pressure was only 0.1 atm and the shock temperature was 13,000°K. Since the phenomenological reaction rate expression has a collision theory temperature dependence and is proportional to the product of the partial pressures, the thermodynamic variables of temperature and pressure are very important to the finite rate chemistry. The combined effect of high temperature and high pressure provides sufficiently fast chemical reaction rates that the shock layer approaches equilibrium at severe heat loads.

The radiative flux divergence profiles of these two cases are presented in Figure 6.21. The profiles are very similar showing a characteristic dip in the stagnation region and increasing rapidly in the air layer. The dip in flux divergence for Case 3 is closer to the ablator since the stagnation point is closer to the ablator (due to a 12% vs 20% blowing rate). The radiative heating rates for the equilibrium (1816 BTU/sq ft-sec) and the nonequilibrium (1782 BTU/sq ft-sec) differ by only 2% demonstrating excellent agreement.

#### Species Boundary Conditions

At the shock wave chemical equilibrium air is normally assumed to provide boundary conditions at the shock. This assumption is often a good one especially at high heat loads (as discussed in the previous section). However, the boundary conditions at the ablator are most likely different from equilibrium. In Appendices A and B a nonequilibrium char layer analysis and the nonequilibrium Hertz-Knudsen

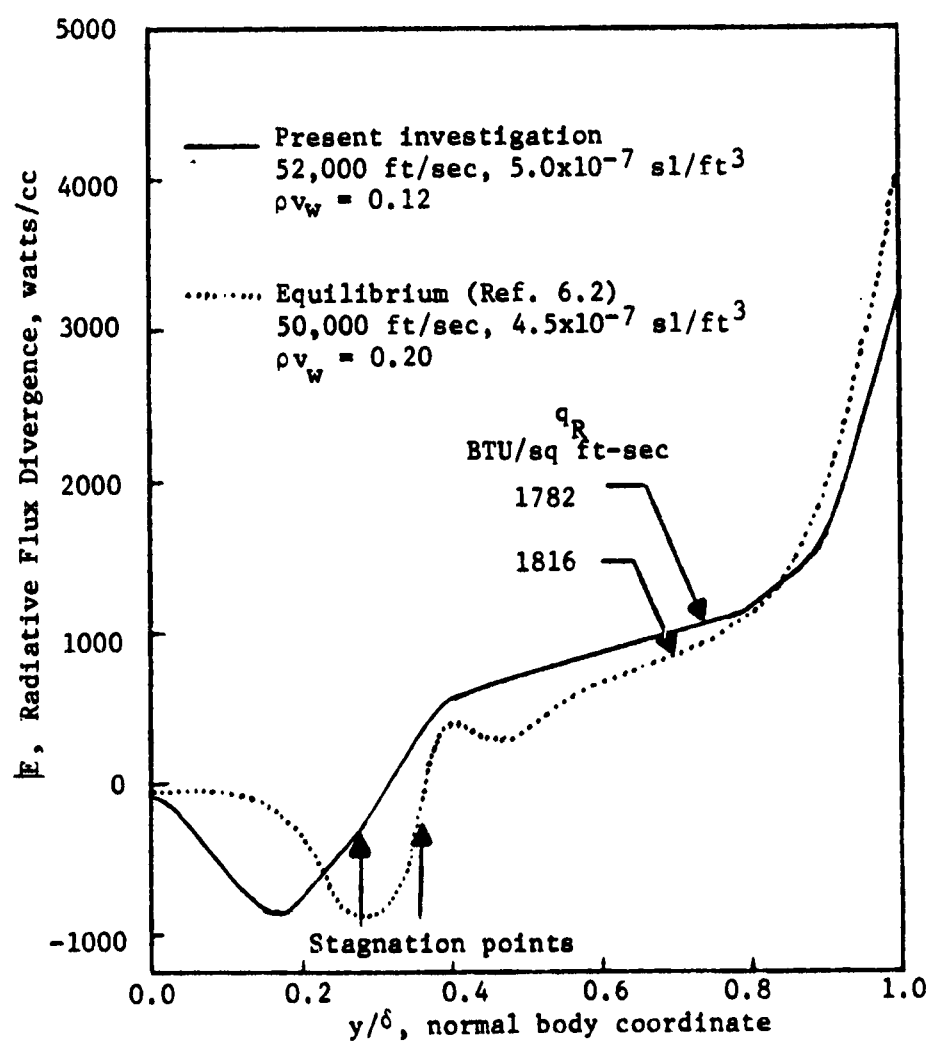


Figure 6.21. Comparison of Radiative Flux Divergence Profiles for Severe Heating Rate Conditions of Equilibrium and Finite Rate Analyses.

equation are used to determine the conditions at the ablator surface. Ablator surface compositions of twenty species vary as functions of surface temperature, pressure, and blowing rate. To couple species compositions from the nonequilibrium ablator analyses to the shock layer or even their results would be a formidable task in itself. Based on the results presented in this section and the tremendous computational simplification involved, one nonequilibrium set of ablator surface mass fractions was used for all coupled stagnation line and around the body results. The conditions that these species compositions were obtained from were a surface temperature of 3600°K, a surface pressure of 1.0 atmosphere, and a mass loss rate of 0.04 lbs/sq ft-sec.

To evaluate the effect of this simplification on radiative heating rates two cases were investigated for the following conditions -- freestream velocity of 50,000 feet per second, freestream density of  $8.85 \times 10^{-8}$  slugs per cubic foot, and a wall blowing rate of 5% of freestream. The frequency factors of the best kinetics in Table 6.5 were used and coupling of shock layer to ablator surface was not attempted. The compositions chosen for the study are shown in Table 6.6. One case is equilibrium compositions at 3450°K for a phenolic-nylon ablator with an elemental composition shown in Table 6.7. The other case used the nonequilibrium values at 3600°K, 1.0 atmosphere, and 0.4 mass loss rate achieved from the nonequilibrium analyses described in Appendices A and B.

The radiative heating rates were 125 and 121 BTU/sq ft-sec for nonequilibrium and equilibrium ablator boundary conditions respectively. The radiative flux divergence profiles for the two cases are

Table 6.6  
Species Composition at the Ablator Surface

Component	Composition (mass fraction)	
	Equilibrium	Nonequilibrium
1. O <sub>2</sub>	$4.81 \times 10^{-15}$	$1.00 \times 10^{-10}$
2. N <sub>2</sub>	$5.37 \times 10^{-2}$	$4.95 \times 10^{-2}$
3. O	$9.74 \times 10^{-8}$	$1.00 \times 10^{-10}$
4. N	$9.44 \times 10^{-5}$	$1.00 \times 10^{-10}$
5. O <sup>+</sup>	$1.66 \times 10^{-20}$	$1.00 \times 10^{-10}$
6. N <sup>+</sup>	$5.72 \times 10^{-18}$	$1.00 \times 10^{-10}$
7. e <sup>-</sup>	$1.13 \times 10^{-9}$	$1.00 \times 10^{-10}$
8. C	$1.34 \times 10^{-2}$	$2.11 \times 10^{-2}$
9. H	$3.86 \times 10^{-2}$	$2.21 \times 10^{-2}$
10. H <sub>2</sub>	$1.36 \times 10^{-2}$	$4.13 \times 10^{-2}$
11. CO	$2.69 \times 10^{-1}$	$2.57 \times 10^{-1}$
12. C <sub>3</sub>	$8.17 \times 10^{-2}$	$3.69 \times 10^{-1}$
13. CN	$7.79 \times 10^{-2}$	$1.00 \times 10^{-10}$
14. C <sub>2</sub> H	$1.65 \times 10^{-1}$	$1.39 \times 10^{-1}$
15. C <sub>2</sub> H <sub>2</sub>	$1.67 \times 10^{-2}$	$5.89 \times 10^{-2}$
16. C <sub>3</sub> H	$1.30 \times 10^{-1}$	$1.00 \times 10^{-10}$
17. C <sub>4</sub> H	$8.42 \times 10^{-2}$	$1.00 \times 10^{-10}$
18. HCN	$3.10 \times 10^{-2}$	$1.00 \times 10^{-10}$
19. C <sub>2</sub>	$2.47 \times 10^{-2}$	$4.22 \times 10^{-2}$
20. C <sup>+</sup>	$1.43 \times 10^{-11}$	$1.00 \times 10^{-10}$

Table 6.7

## Elemental Composition of a Phenolic Nylon Ablator

Element	Elemental Composition (mass fraction)
Carbon	0.7303
Hydrogen	0.0729
Nitrogen	0.0496
Oxygen	0.1472

shown in Figure 6.22. They compare almost identically along the shock layer with the exception of a slight difference near the surface which is expected. The difference in radiative heating rates is only 3.2% which is definitely within the accuracy of the calculation. If the extreme case of equilibrium at the surface varies so little from the nonequilibrium conditions used, then the true nonequilibrium values should differ even less. Based on the insensitivity of the analysis to the wall boundary conditions, it was decided to use the one simplified set of compositions in Table 6.6.

#### Computation Time and Core Usage

The stagnation line and around the body results that have been presented in this chapter represent a sizeable amount of computer time and core usage. The computer utilized in this study was an IBM 360/65 and the program was compiled on the FORTRAN H-compiler for maximum computational optimization. The total size of the program was about 252K bytes which was overlaid down to 186K bytes. The radiation package, LRAD, developed by Engel (Ref. 6.3) was itself 96K (overlaid from 104K bytes) of the 186K. The computer time usage was in the order of 100 hours of actual CPU time with about 60 hours being used on stagnation line solutions alone.

The major time consumption calculations were radiation (LRAD) which took 1.5 minutes per pass and the species solutions which took 20 seconds per pass. The LRAD calculation included line and continuum radiative properties for twelve ionic, atomic, and molecular species. This same calculation took 2.5 minutes by Esch

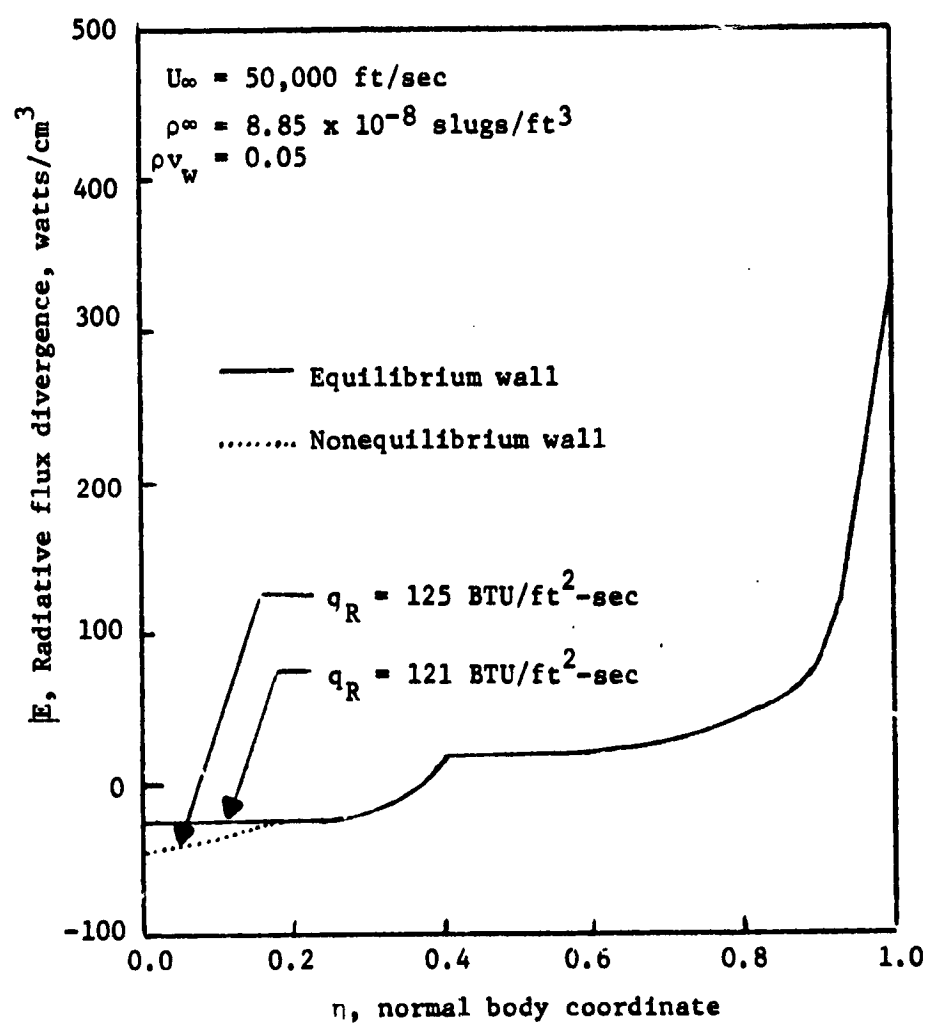


Figure 6.22. Comparison of Radiative Flux Divergence Profiles for Equilibrium and Nonequilibrium Walls at the Stagnation Line.



(Ref. 6.2) when using the same computer but the FORTRAN G compiler instead of the FORTRAN H compiler. Since there were 20 species, it took 20 seconds of total time to solve the species conservation equations taking approximately one second per species. The iteration scheme employed in this study called for many more solutions to the species equations than the radiation flux divergence in LRAD, therefore the total time utilized in the species solution was about three times the LRAD calculation.

The total time for a major iteration loop was about 8 minutes. The total time for stagnation line uncoupled solutions took anywhere from 0.5 hours to about 6 hours. In comparison the same case that took Perez (Ref. 6.1) about 30 hours of CPU time only took 40 minutes for the present investigation. The savings in computer time is believed to come from different initialization, iteration scheme, and efficiency in programming. The initialization of species profiles, for example, was achieved by using flat or frozen flow compositions from shock and ablator to the stagnation point. Since chemical reactions were slow for this case especially in the air layer, the initially assumed compositions were close from the start. Perez, on the other hand, assumed equilibrium profiles and iterated until the profiles became flat. For other cases Perez's initialization scheme may prove more advantageous. For a coupled stagnation line solution, at least three stagnation line solutions had to be run so typical coupled solutions required in the order of 10 hours for completion.

The around the body solutions required differing amounts of time depending on the severity of heating rate. Usually the more

severe heating rates took longer to reach the same distance around the body as a less severe case. For instance Case 1 ( $q_R = 153$  BTU/sq ft-sec) took about 2.5 hours to reach a point of  $\xi = 0.45$  whereas Case 4 ( $q_R = 2232$  BTU/sq ft-sec) only reached a point of  $\xi = 0.30$  after 4.0 hours (see Figures 6.5 and 6.6).

One reason for long computation time in the around the body solution was the shock standoff distance,  $\delta$ . If the calculated value was very different than the input value, convergence on all the conservation equations was very difficult (especially momentum). When the input shock standoff distance was altered to yield a more realistic value, the heating rates were essentially unchanged but the computation time decreased significantly -- sometimes by a factor of two. This result was discussed in Chapter V and it was decided that since heating rates remained unchanged, the input shock standoff distance would only be altered to improve the convergence rate of the conservation equations. No real attempt was made to obtain absolute convergence of the shock standoff distance.

The important point to be made from these results is that shock layer solutions do require large amounts of computer time and core. Although the analysis was diligently trimmed to reduce both time and core usage, the expense is still seen to be phenomenal. Such small inputs as shock standoff distance or iteration scheme can greatly alter the CPU time required. Further improvements in such things to reduce this computational expense should be studied in the future.

Summary

An investigation of the performance of a phenolic-nylon ablator has been made for flight conditions characteristic of return from planetary flight. The conclusions and recommendations resulting from this study are presented in the following chapter.

## REFERENCES

- 6.1 Perez, C., "Finite-Rate Chemistry Effects Upon Convective and Radiative Heating of an Atmospheric Entry Vehicle," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1974).
- 6.2 Esch, D. D., "Stagnation Region Heating of a Phenolic-Nylon Ablator During Return from Planetary Missions," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 6.3 Engel, C. D., "Ablation and Radiation Coupled Viscous Hypersonic Shock Layers," Ph.D. dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).
- 6.4 Chin, J. H., "Radiation Transport for Stagnation Flows Including the Effect of Lines and Ablation Layer," AIAA Paper No. 68-664, Fluid and Plasma Dynamics Conference (Los Angeles, Calif.), June 24-26, 1968.
- 6.5 Garrett, L. B., G. L. Smith, and J. N. Perkins, "An Implicit Finite Difference Solution to the Viscous Shock Layer, Including the Effects of Radiation and Strong Blowing," NASA TR R-388, November 1972.
- 6.6 Rigdon, W. S., R. B. Dirling, and M. Thomas, "Stagnation Point Heat Transfer During Hypervelocity Atmospheric Entry," NASA CR-1462, 1970.
- 6.7 Smith, G. L., J. T. Suttles, and E. M. Sullivan, "Viscous Radiating Flow Field on an Ablating Blunt Body," AIAA Paper No. 70-218, Eighth Aerospace Sciences Meeting (New York, New York), January 19-21, 1970.
- 6.8 Chou, Y. S. and L. H. Blake, "Thin Radiating Shock Layer About a Blunt Body," NASA CR-1547, March 1970.
- 6.9 Semenov, N. N., Some Problems of Chemical Kinetics and Reactivity, Vol. I, Pergamon Press, New York, p. 2, 1958.
- 6.10 Davies, William O., "Radiative Energy Transfer on Entry Into Mars and Venus," Report No. IITRI-T200-13, IIT Research Institute, Chicago, Illinois, December 1965.
- 6.11 Allen, C. W., Astrophysical Quantities, Second edition, University of London, Athlone Press, 1963.
- 6.12 Carlson, L. A., "Radiative-Gasdynamics Coupling and Nonequilibrium Effects behind Reflected Shock Waves," AIAA J., 9 (5), May 1971, 858-864.

- 6.13 Zeldovich, Y. B., Y. P. Raizer, Physics of Shock Waves and High Temperature Hydrodynamic Phenomena, Vol. II, ed. W. D. Hayes and R. F. Probstein, Academic Press, New York and London, p. 489-90 (1967).

## CHAPTER VII

### CONCLUSIONS AND RECOMMENDATIONS

#### Conclusions

Based upon the results of the present investigation, the following conclusions are drawn:

1. The numerical solution of the conservation equations were found to yield stable and accurate solutions to the nonequilibrium, radiating shock layer surrounding a blunt reentry body.
2. Nonequilibrium results, when compared to equilibrium results at comparable conditions, yield consistently lower wall heating rates.
  - (a) The largest differences in equilibrium and nonequilibrium heating rates occurred at low head load conditions. This fact was due to the inability of the finite rate reactions in air to proceed as fast as these conditions at equilibrium. The air layer species profiles in this layer were very different for equilibrium and nonequilibrium thus drastically changing the radiant emission back to the ablator wall.
  - (b) The nonequilibrium heating rate, although lower than equilibrium, did approach the equilibrium prediction for high heat load conditions. The reactions did proceed fast enough in these cases to yield near equilibrium species profiles resulting in similar radiative emission.

3. One normalized around-the-body blowing rate profile was sufficient to adequately couple all the cases studied on the trajectory (coupling required a surface energy balance on the ablator).

4. The specification of equilibrium boundary conditions at the shock wave may be an invalid assumption; especially for low heating rate conditions since the air layer is itself far from chemical equilibrium at these conditions.

5. An incorrect estimate of input shock stand-off distance in the around-the-body solutions does not appreciably affect the coupled surface heating rate around-the-body but can increase computer time quite significantly.

6. The choice of species wall boundary conditions does not greatly affect the shock layer solution. For heating rate at the surface there was a 4% difference in heating rate between non-equilibrium and equilibrium wall cases.

7. Computer time for the stagnation line and around-the-body analyses is large (over 100 hours) for the radiating, nonequilibrium chemistry model used. The radiation calculation (LRAD) and the convergence of the species conservation equations were the major time consuming steps.

8. The temperature form of the energy conservation equation is extremely unstable for the shock layer system with finite rate chemistry and radiation. The enthalpy form of the energy equation is quite stable and converges satisfactorily when the proper damping coefficients are chosen.

### Recommendations

Considering the conclusions presented above, it is recommended that:

1. Nonequilibrium predictions should be employed when low heating rate conditions are involved ( $<1500$  BTU/sq. ft.-sec). Equilibrium models are sufficient for high heat load conditions ( $>1500$  BTU/sq. ft.-sec).
2. For a conservative estimate of the mass loss rate for the phenolic nylon ablator, equilibrium predictions should be used throughout since this will overpredict the mass loss.
3. One representative set of species compositions should be used at the ablator surface to reduce the computational difficulty and computer time required. The difference in radiative heating was shown to be small (4%) for the cases investigated.
4. An investigation should be performed to determine the effect of doing a finite rate analysis across the bow shock. The equilibrium assumption was shown to be in error at low heat load conditions.
5. The numerical analysis in its present form is extremely computer time sensitive to the specification of this profile around-the-body. A more satisfactory method of solution should be found for shock standoff distance calculation.
6. The enthalpy form of the energy equation should be preferred to the temperature form since the enthalpy form is much more stable.



NOMENCLATUREEnglish

$B_v$	Planckian radiation intensity ( $m/t^2 \times \text{no. of particles}$ )
$C_i$	Mass fraction (mass of $i$ /unit mass of fluid)
	$C_i = \frac{\rho_i}{\rho} \frac{Y_i}{M} = \frac{n_i}{\rho} \frac{M}{M}, \quad \sum_i C_i = 1$
$C_p$	Specific heat at constant pressure ( $L^2/t^2 \times T$ )
$D_{ij}$	Multicomponent diffusion coefficient ( $L^2/t$ )
$D_{1T}$	Effective multicomponent diffusion coefficient ( $L^2/t$ )
$D_1$	Thermal diffusion coefficient ( $m/L \times t$ )
$D_{ij}$	Binary diffusion coefficient ( $L^2/t$ )
$E$	Stagnation internal energy ( $mL^2/t^2$ )
$\dot{E}$	Radiative flux divergence ( $m/L \times t^2$ )
$\bar{e}$	Strain tensor ( $1/t$ )
$f'$	Velocity function, $u/u_g$
$G_j$	Gibbs free energy ( $mL^2/t^2 \times \text{mole of } j$ )
$H$	Total enthalpy, $H = h + \bar{V}^2/2$
$h$	Static enthalpy, $h = Q + P/\rho$ ( $L^2/t^2$ ), also Planck's constant
$\bar{J}_i$	Mass flux vector of species $i$ ( $m/L^2 \times t$ )
$k$	Ordinary coefficient of thermal conductivity ( $mL/t^3 \times T$ )
$k_c$	Boltzmann's constant ( $mL^2/t^2 T$ )
$k_b$	Bulk thermal conductivity ( $mL/t^3 \times T$ )
$M_i$	Molecular weight of species $i$ (mass of $i$ /mole of $i$ )
$m_i$	Mass of $i$ ( $m$ )

$n$	Number density (particles/L <sup>3</sup> )
$n_1$	Molal volume (moles of 1/L <sup>3</sup> )
$n_t$	Molal density (total moles/L <sup>3</sup> )
$P$	Static pressure (m/L x t <sup>2</sup> ) or (F/L <sup>2</sup> )
$P_r$	Prandtl number, $C_p \mu / k$
$\bar{P}$	Radiative stress tensor (m/L x t <sup>2</sup> )
$Q$	Internal energy per unit mass, including chemical energy (L <sup>2</sup> /t <sup>2</sup> )
$\bar{q}_R$	Radiative heat flux vector defined by (m/t <sup>3</sup> ) or (E/L <sup>2</sup> x t)
$q_c$	Convective energy flux to a surface (m/t <sup>3</sup> )
$q_R$	Radiative energy flux to a surface (m/t <sup>3</sup> )
$\bar{q}_D$	Diffusional energy flux vector (m/t <sup>3</sup> )
$R$	Body nose radius (L)
$Re_s$	Reynolds number, $\rho_{\delta,0} U_{\infty} R / \mu_{s,0}$
$R_1$	Mass rate of formation of species 1 by heterogeneous reactions (m/tL <sup>2</sup> )
$R$	Universal gas constant (mL <sup>3</sup> /t <sup>2</sup> x T x no. of moles)
$r$	Cylindrical body radius
$S_1$	Total surface generation of species 1 (m/t L <sup>2</sup> )
$T$	Thermodynamic temperature (T)
$t$	Time (t)
$U_{\infty}$	Freestream velocity (L/t)
$x$	Body oriented coordinate
$y$	Body oriented coordinate
$Y_1$	Mole fraction of species 1, $Y = n_1/n_t$ , $\sum_1 Y_1 = 1$

Greek

$\alpha$	Sublimation accommodation coefficient
$\alpha_v$	Volumetric absorption coefficient, effective ( $L^2 \times \text{no. of particles}/L^3$ )
$\nabla$	Del operator (1/L)
$\delta$	Shock detachment distance (L)
$\tilde{\delta}$	Transformed shock detachment distance
$\epsilon$	Difference between the body and shock angle $\epsilon = \theta - \phi$ (radians)
$\epsilon_p$	Volume of voids per unit volume
$\eta$	Dorodnitsyn variable
$\theta$	Body angle (radians)
$I_v$	Spectral radiation intensity ( $m/t^2 \times \text{no. of particles} \times$ no. of steradians)
$\bar{I}$	Unit tensor
$\kappa$	Absorption coefficient ( $L^2 \times \text{no. of particles}/L^3$ )
$\kappa$	Local body curvature (1/L)
$\tilde{\kappa}$	$1 + \kappa y$
$\lambda$	$(\tilde{\mu} - 2/3 \mu) (m/L \times t)$
$M$	$\sum_i Y_i M_i$ mean molecular weight of the mixture (m of mixture/ mole of mixture)
$\mu$	Ordinary viscosity (m/L x t)
$\tilde{\mu}$	Bulk viscosity (m/L x t)
$\nu$	Frequency (1/t)
$\rho$	Density (m/L <sup>3</sup> ), $\rho = n_t M$
$\rho_i$	Partial density of species i, $\rho_i = n_i M_i$ (m of i/L <sup>3</sup> )
$\bar{\rho}$	Density ratio across shock $\rho_\infty / \rho_s$

$\sigma$	Radiative absorption cross section ( $L^2$ )
$\tau_\nu$	Optical depth at frequency $\nu$
$\tau$	Viscous stress tensor ( $m/L \times t^2$ )
$\phi$	Shock angle (radians)
$\bar{\Omega}_1$	Unit vector in the direction of photon propagation
$\Omega$	Solid angle (stearadians)
$\Omega_{ij}$	Collision integral of colliding species $i$ and $j$
$\omega_1$	Generation of species $i$ ( $m/L^3 \times t$ )
$u$	Component of $\bar{V}$ in the $\xi_1$ direction (parallel to body surface) ( $L/t$ )
$\bar{V}$	Velocity vector, $u\bar{i} + v\bar{j} + w\bar{k}$ ( $L/t$ )
$v$	Component of $\bar{V}$ in the direction normal to the body surface ( $L/t$ )

#### Subscripts

$i$	Species $i$
$w$	Wall quantities
$o$	Stagnation line quantities
$\infty$	Freestream conditions
$s, \delta$	Quantities immediately behind the shock

#### Superscripts

$D$	Diffusion
$A$	0 or 1 denoting two-dimensional or axisymmetric respectively (an exponent)
$T$	Thermal
$*$	Denotes dimensional variables
$o$	Standard state quantity

- .. Evaluated on char side of ablator interface
- + Evaluated on flow field side of ablator interface

APPENDIX A  
ENERGY TRANSFER IN THE CHAR ZONE

Introduction

An important step in the development of a coupled solution for the heat transfer in the flow field of a reentry body, is the description of the ablator itself. For the phenolic nylon ablator studied in this investigation, there are two zones of interest in the ablator - the decomposition zone and the char zone. This section will concentrate on the solution of the conservation equations in the char zone since this is the zone of most significant heat and mass transfer. For results of the phenomena occurring in the decomposition zone, the work of del Valle (Ref. A.1) was used.

The solution of the char zone conservation equations requires a specification of the thermodynamic and transport properties for the species involved along with a set of chemical reactions which occur at the conditions present. For this information the work of del Valle (Ref. A.1) was again used as a basis.

The problem of using del Valle's results directly as input to the flow field was a severe shortcoming in the numerical technique used. The conditions studied in del Valle's work allowed the use of fourth order Runge-Kutta for numerical solution of the energy and species equations. Unfortunately, these conditions were for the case where chemical reactions were not very rapid. The conditions

required for a coupled solution to the flow field surrounding the ablator during planetary return have much faster chemical reaction rates. The step size for the Runge-Kutta integration became so small (in order to insure stability) that computation time became excessive and accuracy became questionable (due to round off errors). A new, absolutely stable numerical technique was then used which was many times faster than Runge-Kutta. This implicit technique is called the Padé integration scheme.

In this section a brief development of the one-dimensional conservation equations that describe the char zone will be presented. This development will be followed by a discussion of the Padé integration scheme and numerical solution of the energy and species equations. Finally, results will be given which were used in the coupled shock layer-phenolic nylon ablator solutions of this study.

#### Development of the Char Zone Analysis

To predict the energy transferred in the char zone, the equations of change are written to apply to one-dimensional flow in the char zone. Using a quasi-steady state approximation, the point of view is taken of an observer moving with the negative of the surface recession velocity. This is illustrated in Figure A.1 which shows schematically the virgin plastic, pyrolysis gases, and char. The point at which thermal degradation of the plastic is initiated is taken as  $z = 0$ .

The material balance relating the virgin plastic flow with the flow of pyrolysis gases and degrading solid in the combined decomposition char zone was written by Stroud (Ref. A.2) as

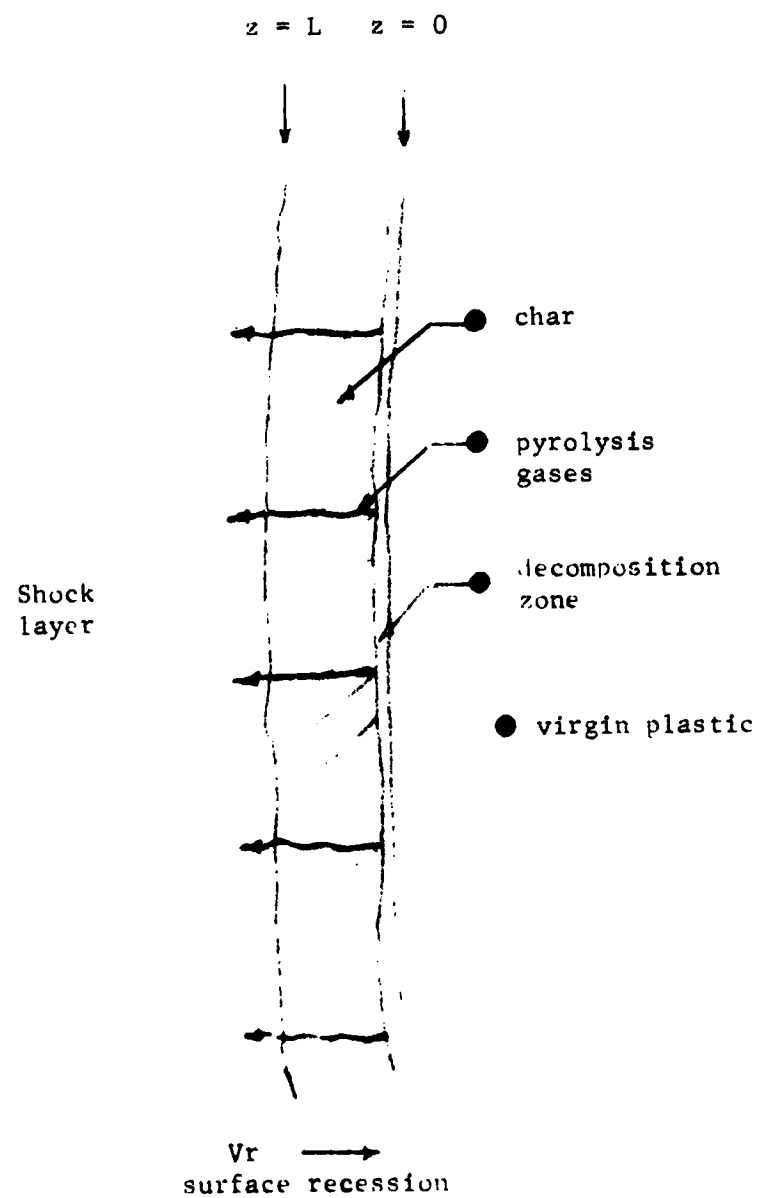


Figure A.1. Flow in the Char Zone



$$\rho_0 v = w + \rho v \quad [1]$$

where  $w$  is the total gas mass flux,  $v$  is the surface recession velocity, and  $\rho_0$  and  $\rho$  are the virgin and degrading solid densities respectively. This assumes one-dimensional flow, quasi-steady state, and that the bulk volume of the char is the same as that of the virgin plastic. All of these have experimental justification. The composition of the gases generated by the decomposition of the virgin plastic composites must be known to be able to accurately predict the energy absorbed due to chemical reactions in the char zone. As illustrated in Figure A.1, these pyrolysis products enter the char zone and exit through the front surface of the char, at  $Z = L$ . Changes in the mass flux of the various species within the char occur as a result of chemical reactions at finite chemical reaction rates.

The particular restrictions and assumptions made in the formation of the combined zone analysis are presented below and have been justified by del Valle in Ref. A.1.

- (1) Flow of pyrolysis gases is quasi-steady and one dimensional.
- (2) Pyrolysis products behave as a perfect gas mixture.
- (3) Thermal equilibrium between the pyrolysis gas products and the char is attained.
- (4) PV work and viscous dissipation are negligible.
- (5) Diffusional transport is negligible.
- (6) Virgin material, char and gas physical properties are variable.

- (7) For momentum transfer in the char zone the modified form of Darcy's Law was used.

The application of the above restrictions to the general equations for flow to the pyrolysis gases in the char zone is now discussed.

Species Continuity Equation: Referring to Figure A-1 the species continuity equation for the  $i$ th component of a gas mixture for flow through a porous medium is (Ref. A.3)

$$\frac{D\rho_i}{Dt} = -\rho_i (\nabla \cdot \mathbf{u}) - (\nabla \cdot \mathbf{J}_i) + R_i$$

where  $\rho_i$  is the concentration,  $\mathbf{J}_i$ , the mass flux by diffusion,  $R_i$ , the rate of generation of chemical species  $i$  and  $\mathbf{u}$  is the velocity of the pyrolysis products within the pores.

For steady, one dimensional flow of gases with the above restrictions Equation (2) reduces to

$$\frac{d}{dz} (\rho_i u) = R_i \quad [3]$$

If the mass flow rate of pyrolysis gases do not lose or gain by chemical reaction, Equation (3) becomes, summing over all the gas species:

$$\sum_{i=1}^n \frac{d}{dz} (\rho_i u) = 0 \quad [4]$$

However, if the mass flow rate of pyrolysis gases changes due to carbon deposition, Equation (4) becomes:

$$\sum_{i=1}^n \frac{d}{dz} (\rho_i u) = -R_c \quad [5]$$

where  $R_c$  is the amount of carbon deposition. If we define  $W_p$  as the mass flux of pyrolysis gases based on the cross-sectional area of voids in the char,  $\sum_{i=1}^n \rho_i u$ , (units of  $\text{lbs/ft}^2 \text{ voids} \text{ -sec}$ ) Equation (5) becomes:

$$\frac{dW_p}{dz} = \sum_{i=1}^n \frac{d}{dz} (\rho_i u) = -R_c \quad [6]$$

Equation (6) says that any change of the mass flux of pyrolysis gases is due to the loss or gain of carbon by chemical reaction of the pyrolysis gases with the char.

Momentum Equation: The momentum equation for flow through porous media was formulated by H. P. G. Darcy in 1856 (Ref. A.4). Darcy's Law is:

$$u = \frac{\alpha}{\epsilon \mu} (\nabla P - \rho g) \quad [7]$$

Applying this equation to a one-dimensional, horizontal flow through a porous char layer and solving for the pressure gradient gives:

$$-\frac{dP}{dz} = \frac{\mu}{\alpha} (u \epsilon) \quad [8]$$

This equation is valid at low gas flow velocities within the porous medium. However, at high gas velocities it is necessary to

to add a term to account for the inertial effects. This additional term leads to a modified form of Darcy's Law:

$$-\frac{dP}{dz} = \frac{\mu}{\alpha} (u\epsilon) + \beta \rho (u\epsilon)^2 \quad [9]$$

Multiplying both sides of Equation (9) by the gas density,  $\rho$ , followed by substitution of the ideal gas equation of state ( $\rho = PM_w/RT$ ) on the left hand side of the Equation (9), can be written in the following form

$$-\frac{M_w P}{RT} \frac{dP}{dz} = \rho(\mu/\alpha) (u\epsilon) + \beta \rho^2 (u\epsilon)^2 \quad [10]$$

If we define  $W$  as the total mass flux of pyrolysis gases based on the total area, we have that:

$$W = W_p = \epsilon \rho u \quad [11]$$

Therefore, substitution of Equation (11) into Equation (10) results after rearrangement:

$$-PdP = \frac{RT}{M_w} (\mu/\alpha) W + \beta (W)^2 \quad [12]$$

Integration of Equation (12) between the front surface pressure ( $P = P_L$  at  $Z = L$ ) and any point within the char layer, ( $P$  at  $z$ ), results in an integral equation for the pressure distribution in the char.

$$P = (-P_L^2 + 2R \int_L^z (\mu/\alpha) W (T/M_w) dz + \int_L^z \beta (T/M_w) (W)^2 dz)^{1/2} \quad [13]$$

In this equation all parameters that vary with temperature (hence, char distance) are left under the integral signs. These parameters are calculated by polynomials in temperature and from the solution of the energy equation.

Energy Equation. To formulate the equation that describes the energy transfer in the char zone, the energy equation for the gas and for the solid are written separately. They are later combined considering that the gas and solid are at the same temperature at any cross section in the flow; that is in thermal equilibrium.

The one dimensional, steady state energy equation for a gas mixture containing  $n$  species in reacting gas flow through porous media is

$$W \bar{C}_p \frac{dT}{dz} = \epsilon \frac{d}{dz} \left( k_g \frac{dT}{dz} \right) - \epsilon \sum_{i=1}^n H_i R_i \quad [14]$$

Similarly, the energy equation for the solid in the char zone simplified for the same restrictions to the following form:

$$(1-\epsilon) C_p \rho v \frac{dT}{dz} = (1-\epsilon) \frac{d}{dz} \left( k_c \frac{dT}{dz} \right) + (1-\epsilon) H_c R_c \quad [15]$$

The total energy transferred in the char zone is formed by adding Equations (14) and (15) and by using the definition of  $W$  of Equation (11):

$$[\epsilon \bar{C}_p W p + (1-\epsilon) C_p \rho v] \frac{dT}{dz} = \frac{d}{dz} \left( k_e \frac{dT}{dz} \right) - \sum_{i=1}^{n+1} H_i \bar{R}_i \quad [16]$$

where  $k_e$  represents an effective thermal conductivity defined as:

$$k_e = \epsilon \bar{k}_g + (1-\epsilon) k_c \quad [17]$$

where  $\bar{k}_g$  is the thermal conductivity of the flowing pyrolysis gases and  $k_c$  that of the solid matrix.

In addition the term  $\sum_{i=1}^{n+1} H_i \bar{R}_i$  is defined as:

$$\sum_{i=1}^{n+1} H_i \bar{R}_i = \epsilon \sum_{i=1}^n H_i R_i + (1-\epsilon) H_c R_c \quad [18]$$

The above represents the energy absorbed by the chemical reactions on a "total volume" basis. The solution of the energy equation, Equation (16), gives the temperature distribution in the char zone.

#### Padé Numerical Integration Technique

A system of  $n$  first order ordinary differential equations (ODE's) can be expressed in general form as:

$$\dot{y}(t) = f(t, y) ; \quad y(t_0) = y_0 \quad [19]$$

where  $y$  is a vector valued function of  $t$ , and  $f$  is a vector valued function of  $t$  and  $y$ . Given the initial vector  $y_0$ , the system of ODE's is to be numerically integrated over the time domain  $(t_0, t_f)$ .

A class of stable integration formulas can be derived by using known properties of linear ODE's. A numerical method of solution is termed stable if an error, once introduced, decreases from step to step. It can be shown that for a stable solution the eigenvalues of the matrix  $[\partial f_i / \partial y_j]$  have negative real parts (see Ref. A.5). These formulas, while directly applicable to a system of linear ODE's, may be applied to systems of non-linear differential

equations by linearizing at each step. Even considering the additional computer time necessary to linearize the equations, the total time needed to solve a complete problem is much less than required by any standard method (Ref. A.6).

Consider a system of  $n$  linear ODE's with constant coefficients:

$$\dot{y}(t) = A y(t) + b \quad [20]$$

the formal solution of this equation at  $t = h$  is

$$y(h) = \exp(hA) [y_0 + A^{-1} b] - A^{-1} b \quad [21]$$

where

$$\exp(hA) = I + hA + \frac{1}{2} (hA)^2 + \dots$$

In Equation 21, the exponential term can be approximated in numerous ways, and the form of the approximation will determine whether the procedure is stable. A convenient method of approximating any function is the rational approximation, which can be derived using the Padé transformation of the power series. If  $P$  and  $Q$  are polynomial of degree  $p$  and  $q$ , respectively, the exponential is represented as

$$e^{hA} = Q^{-1}P + E(hA)$$

The rational approximation agrees with the power series of  $\exp(hA)$  for at least  $p + q + 1$  terms and the residual error is denoted by  $E(hA)$ . The polynomials  $P$  and  $Q$  for the exponential function (Ref. A.7) are given by

$$P = \sum_{k=0}^p \frac{(p+q-k)! p!}{(p+q)! k! (p-k)!} (hA)^k$$

$$Q = \sum_{k=0}^q \frac{(p+q-k)! q!}{(p+q)! k! (q-k)!} (-hA)^k$$

Substituting the rational approximation into Equation (21) yields the desired integration formula:

$$y(h) = Q^{-1} P [y_0 + A^{-1} b] - A^{-1} b$$

Or rearranging the terms,

$$y(h) = Q^{-1} [Py_0 + (P - Q) A^{-1} b] \quad [22]$$

Equation (22) includes numerous types of single step integration methods and not all of these methods are stable. The Euler method, which has been shown to be partially stable is a special case of Equation (22) with  $Q = I$  and  $P = I + hA$  ( $q = 0$  and  $p = 1$ ). In this section only the stable procedures are of interest and for these methods  $q \geq p$ . Of particular interest are those methods where the diagonal ( $p = q$ ) of the Padé table is used. Such methods are called Padé integration procedures or rational approximation methods. For the cases  $p = q = 1$  and  $p = q = 2$ , the following expressions are obtained from Equation (22):

$$y(h) = (I - \frac{1}{2} hA)^{-1} [(I + \frac{1}{2} hA) y_0 + hb] + O(h^3) \quad [23a]$$



$$y(h) = [I - \frac{1}{2} hA + \frac{1}{12} (hA)^2]^{-1} [(I + \frac{1}{2} hA + \frac{1}{12} (hA)^2)y_0 + hb] + O(h^5) \quad [23b]$$

The truncation errors are  $O(h^3)$  and  $O(h^5)$ , respectively.

The Padé integration method is directly applicable to linear systems of ordinary differential equations. Some additional steps are required to apply the method to chemical kinetic equations, which are represented by Equation (19). The non-linear system of equations can be reduced to the form Equation (20) by expanding the function  $f$  in a Taylor's Series and neglecting terms which are second order and higher:

$$\dot{y} = f(t_0, y_0) + \frac{\partial f}{\partial y}(t_0, y_0)\Delta y + O(\Delta y^2) \quad [24]$$

where  $\Delta y = y - y_0$  and  $\frac{\partial f}{\partial y}(t_0, y_0)$  denotes the matrix  $A$  in Equation (20) with elements

$$a_{ij} = \frac{\partial f_i}{\partial y_j}$$

The indices  $i$  and  $j$  denote the  $i^{\text{th}}$  and  $j^{\text{th}}$  elements in the  $f$  and  $y$  vectors, respectively. Hence grouping terms as in Equation (20)

$$\dot{y} = \frac{\partial f}{\partial y}(t_0, y_0)y + (f(t_0, y_0) - \frac{\partial f}{\partial y}(t_0, y_0)y_0) + O(\Delta y^2) \quad [25]$$

The phenomenological kinetic rate expression represents the function  $f_i$  which must be expanded in a Taylor's Series as shown in Equation (24). In terms of the functions  $f_i$  the elements of the matrix,  $A$ , and vector,  $b$ , are

$$a_{ij} = \frac{\partial f_i}{\partial y_j} \quad [26]$$

$$b_i = f_i(t_0, y_0) - \sum_{k=1}^n a_{ik} (y_0)_k \quad [27]$$

where the index  $i$  denotes the species production equation, the index  $j$  denotes the species, and  $(y_0)_k$  is an element of the vector of species at time  $t_0$ .

The Padé integration routine used in this analysis assumes that  $y(0)=0$ . This assumption simplifies Eqs. (23) and the linearized system is then conveniently expressed in terms of  $\Delta y$  as

$$(I - \frac{1}{2} hA + KA^2) \Delta y = hb \quad [28]$$

where  $K = 0$  for a second order integration scheme and  $K=h^2/12$  for a fourth order integration scheme. By permitting  $K$  to be specified as an input, either the second order or fourth order methods can be conveniently selected. In cases where the higher order scheme offers no great advantages, the elimination of the calculation of the matrix  $A^2$  can save an appreciable amount of computer time.

The species equations are "stiff" and therefore need to be solved using the numerically stable Padé scheme instead of the more conventional Runge-Kutta method. The energy equation, however, is not "stiff" and its solution can be performed using Runge-Kutta as described below.

#### Energy Equation Solution

The solution of the energy equation was performed using a standard fourth order Runge-Kutta integration technique. Actually, the energy equation was solved as two simultaneous equations in

$\frac{dT}{dz}$  and  $T$  by taking Equation (16) and expanding as follows:

$$[\epsilon C_p W_p + (1-\epsilon) \bar{C}_p \rho v] \frac{dT}{dz} = k_e \frac{d^2 T}{dz^2} + \frac{dk_e}{dz} \frac{dT}{dz} - \sum_{i=1}^{\ell} H_i \bar{R}_i \quad [28]$$

where the term  $\frac{dk_e}{dz}$  can be further expanded as

$$\frac{dk_e}{dz} = \frac{dk_e}{dT} \frac{dT}{dz} \quad [29]$$

using the chain rule. This yields the following result

$$k_e \frac{d^2 T}{dz^2} = [\epsilon C_p W_p + (1-\epsilon) \bar{C}_p \rho v] \frac{dT}{dz} + \frac{dk_e}{dT} \frac{dT}{dz} + \sum_{i=1}^{\ell} H_i \bar{R}_i \frac{dT}{dz} \quad [30]$$

or when divided by  $k_e$

$$\frac{d^2 T}{dz^2} = \frac{1}{k_e} [\epsilon C_p W_p + (1-\epsilon) \bar{C}_p \rho v] \frac{dT}{dz} + \frac{dk_e}{dT} \frac{dT}{dz} + \sum_{i=1}^{\ell} H_i \bar{R}_i \frac{dT}{dz} \quad [31]$$

Equation (31) is a second order nonlinear differential equation in  $T$  which can be considered two first order in  $T$  and  $y = \frac{dT}{dz}$ . The two equations are

$$\frac{dy}{dz} = f(T, y) \quad [32]$$

and

$$\frac{dT}{dz} = y \quad [33]$$

The standard fourth order Runge-Kutta integration formula can then be applied to the simultaneous equations.

### Results of the Char Analysis

The results of the analysis in the char layer are presented in Fig. A.2. Heating rates are given as a function of ablator mass

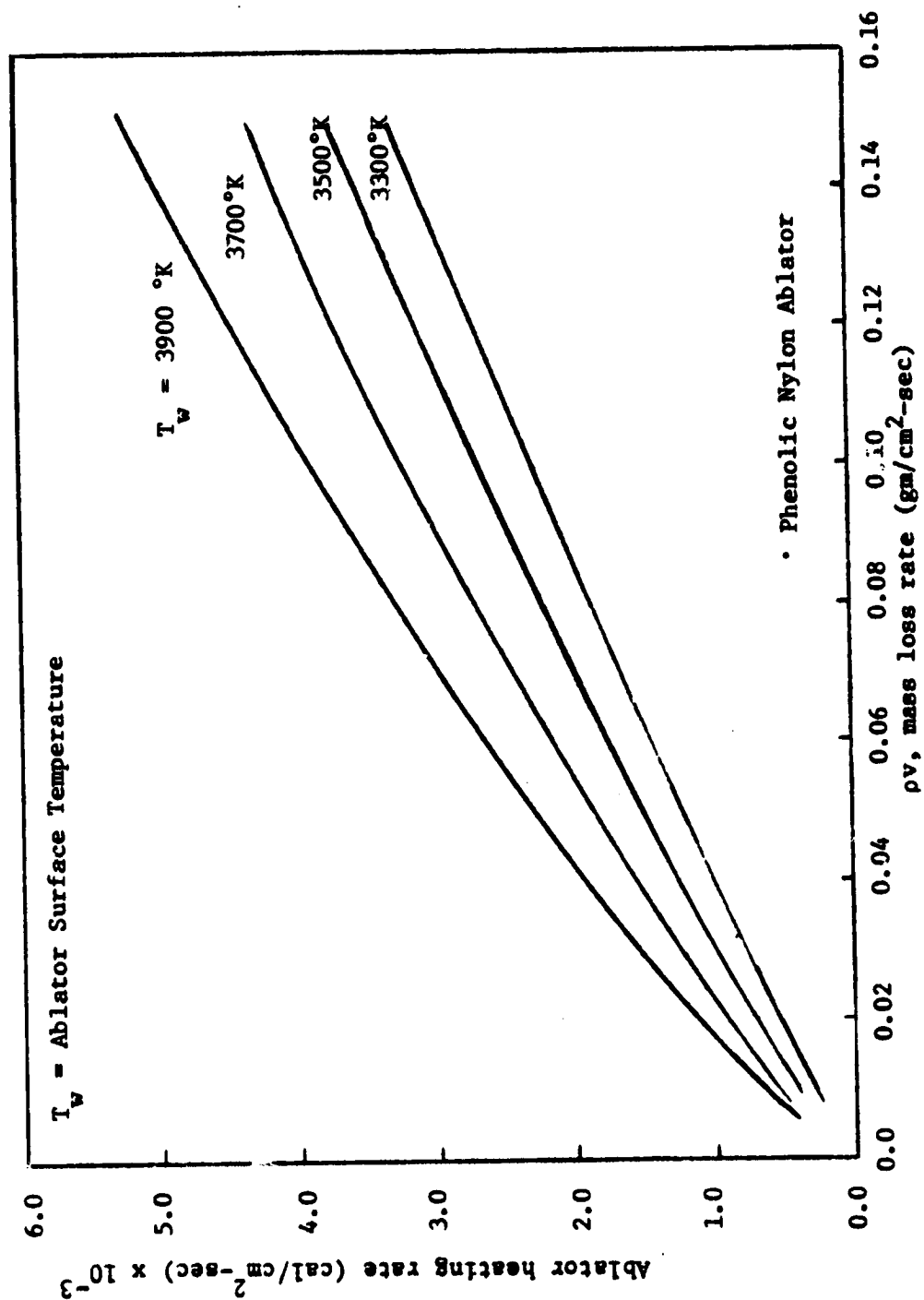


Figure A.2. Ablator Heating Response vs. Mass Loss Rate.

loss rates with ablator surface temperature as a parameter. The heating rates are typical of those encountered during reentry from planetary missions. These curves were curve fit using a linear regression program from SAS (Statistical Analysis System developed at North Carolina State University, Ref. A.8). The equation used to fit the curves was:

$$q = a_1 T^3 \rho v + a_2 T^2 \sqrt{\rho v} + a_3 T^3 \rho v + a_4 \rho v \quad [34]$$

where the coefficients  $a_1$  to  $a_4$  are given in Table A.1.

The energy and species equations needed boundary conditions for these results. For the energy equation, the heat flux from the decomposition zone was used (from an extension of Ref. A.1). This heat flux from the decomposition zone is a function of mass loss rate as depicted in Figure A.3. This heat flux was the value when the solution of the decomposition zone species equations showed the correct mass ratio of char to pyrolysis gases. The mass fractions of gases and char entering the char layer are given in Table A.2. By using this criteria for the specification of the interface, continuity was achieved between the decomposition zone and the char. The heat flux allows the initial specification of  $dT/dz$ . Also required in the solution of the energy equation is the temperature at decomposition zone-char zone interface (shown in Figure A.4).

Each species equation needs the specification of initial composition at the back surface. The values used in this analysis are the same as used in Ref. A.1. and are presented in Table A.2. The values of species compositions as they exit the char surface are shown in Table A.3 for conditions of mass loss rate of 0.04 lbm/sq

Table A.1

Curve Fit Coefficients for Heat Conducted  
through the Surface of a Phenolic Nylon Ablator

$$(q = a_1 T^3 \rho v + a_2 T^2 \sqrt{\rho v} + a_3 T^3 \sqrt{\rho v} + a_4 \rho v)$$

	$q$ in * <u>cal/sq. cm-sec</u>	$q$ in † <u>BTU/sq. ft-sec</u>
$a_1$	$0.9292 \times 10^{-8}$	$0.1674 \times 10^{-7}$
$a_2$	$-0.1292 \times 10^{-3}$	$-0.3330 \times 10^{-3}$
$a_3$	$0.4268 \times 10^{-7}$	$0.1099 \times 10^{-6}$
$a_4$	$0.1516 \times 10^4$	$0.2732 \times 10^4$

\* T is temperature in °K

$\rho v$  is mass loss rate in gm/sq cm-sec

† T is temperature in °K

$\rho v$  is mass loss rate in lbm/sq ft-sec

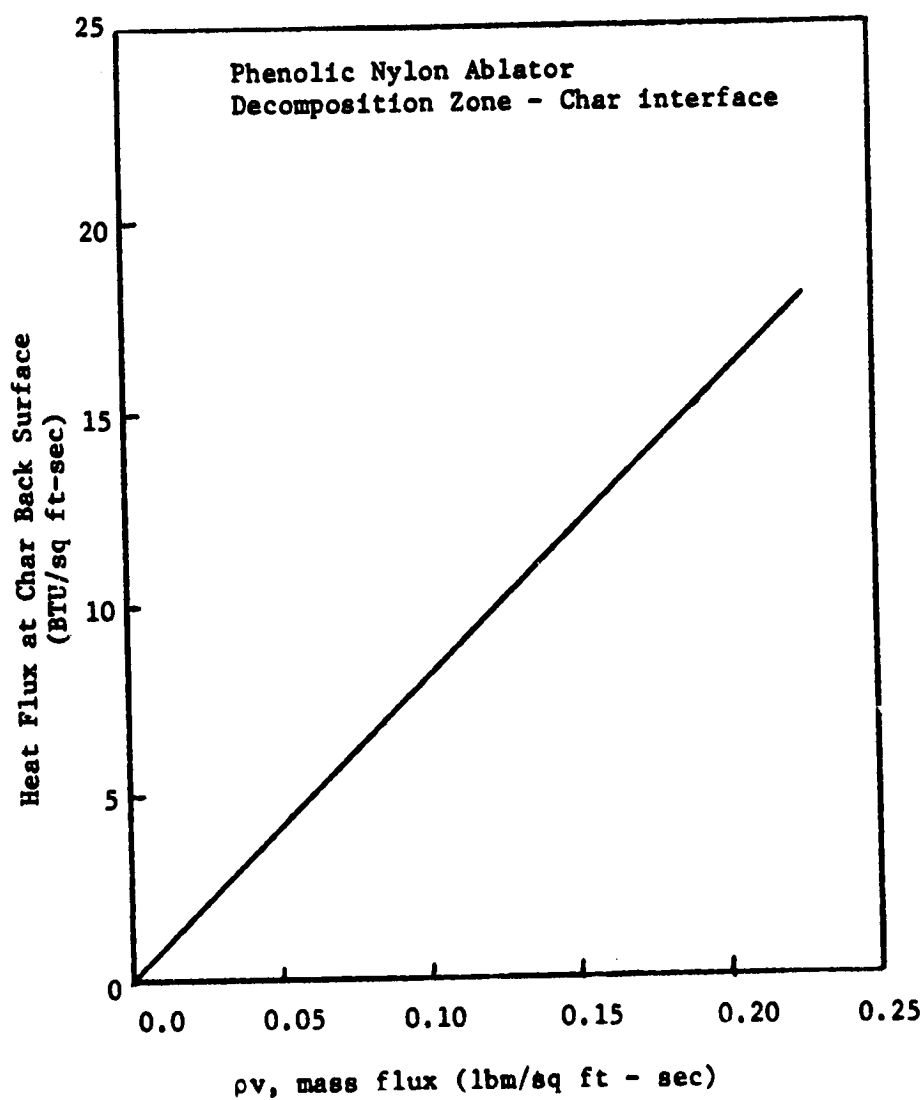


Figure A.3. Decomposition Zone Heat Absorption Rate.

Table A.2  
Species Compositions at the Char Zone -  
Decomposition Zone Interface

Component	Composition (Mass Percent)
H <sub>2</sub>	3.03
CH <sub>4</sub>	3.87
C <sub>2</sub> H <sub>2</sub>	3.89
C <sub>2</sub> H <sub>4</sub>	3.90
C <sub>2</sub> H <sub>6</sub>	0.65
C <sub>6</sub> H <sub>6</sub>	2.59
C <sub>6</sub> H <sub>5</sub> OH	23.18
CO	4.18
CO <sub>2</sub>	4.62
H <sub>2</sub> O	5.65
N <sub>2</sub>	4.96
Carbon (solid)	<u>39.48</u>
Total	100.00%



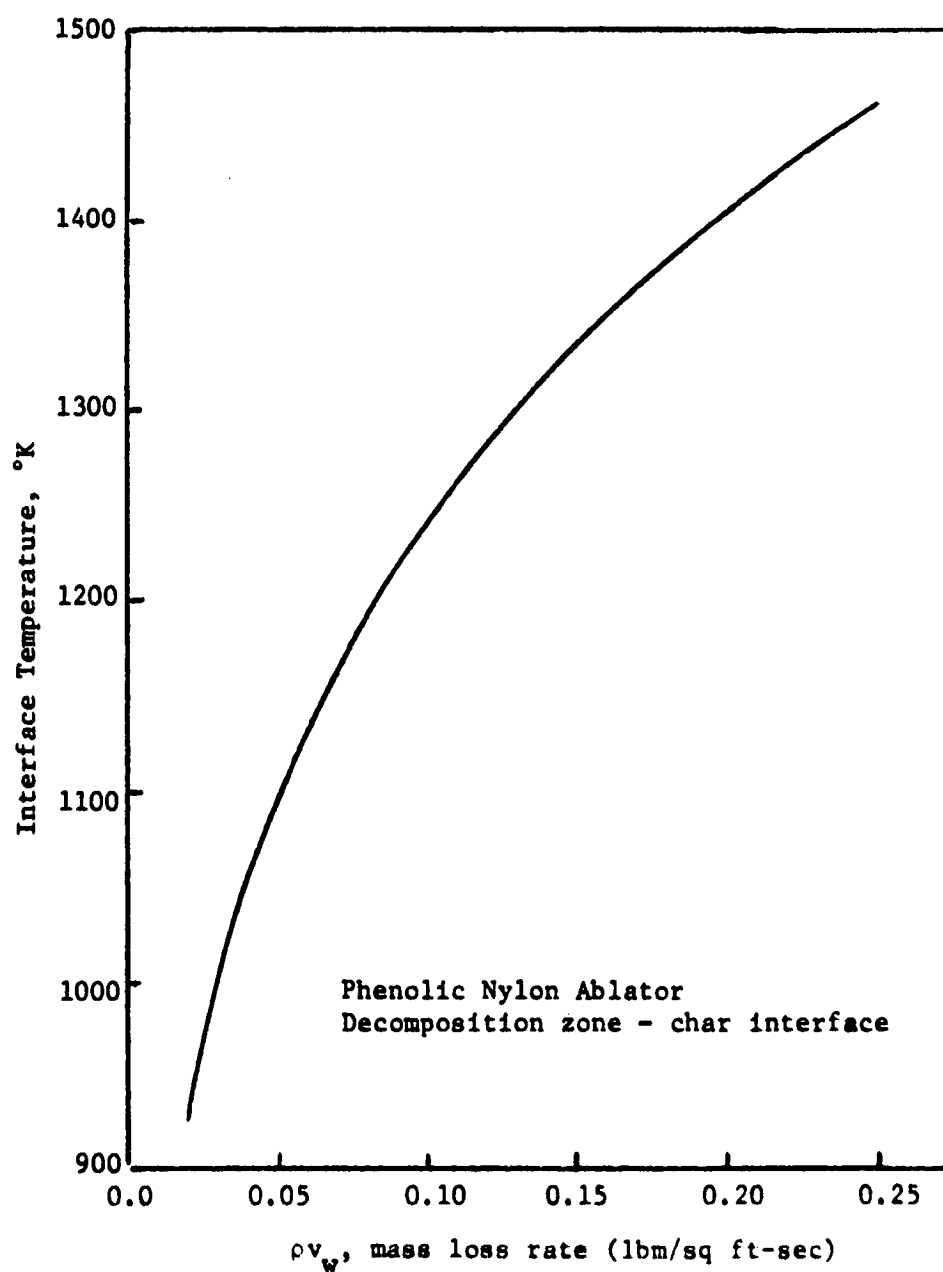


Figure A.4. Temperature at Char-Decomposition Zone Interface as a Function of Mass Loss Rate.

Table A.3  
Species Compositions at the Ablator Surface  
on the Char and Flowfield Sides

Component	Species Composition (Mass Percent)	
	Char Side	Flowfield Side
H <sub>2</sub>	3.81	4.14
C <sub>2</sub> H <sub>2</sub>	5.96	5.89
CO	25.70	25.70
N <sub>2</sub>	4.96	4.96
C <sub>2</sub> H	11.80	13.88
H	2.88	2.21
CH <sub>4</sub>	0.08	*
CH <sub>3</sub>	0.32	*
CH <sub>2</sub>	3.10	*
Carbon (solid)	42.11	*
C	*	36.95
C <sub>2</sub>	*	4.22
C <sub>3</sub>	*	2.11
C <sub>3</sub> H	*	1 x 10 <sup>-10</sup>
C <sub>4</sub> H	*	1 x 10 <sup>-10</sup>
O	*	1 x 10 <sup>-10</sup>
HCN	*	1 x 10 <sup>-10</sup>
CN	*	1 x 10 <sup>-10</sup>
N	*	1 x 10 <sup>-10</sup>
N <sup>+</sup>	*	1 x 10 <sup>-10</sup>
O <sup>+</sup>	*	1 x 10 <sup>-10</sup>
C <sup>+</sup>	*	1 x 10 <sup>-10</sup>
O <sub>2</sub>	*	1 x 10 <sup>-10</sup>
e <sup>-</sup>	*	1 x 10 <sup>-10</sup>

\* Species not considered present on this side.

ft - sec and surface temperature of 3600°K. These conditions are typical of those encountered by a re-entering vehicle from planetary missions. Also shown in Table A.3 are the species compositions modified to agree with the species considered in the flow field. The only species of any reasonable concentration that were omitted from ablator surface to flow field were  $\text{CH}_2$  and  $\text{CH}_3$ . The reason this was done was the equilibrium considerations show that these species do not exist in the shock layer. To reduce the number of species in the shock layer and thus reduce the computation time the species were assumed to instantaneously react to carbon solid,  $\text{C(s)}$ , and hydrogen,  $\text{H}_2$ . These values were used in the flow field analysis in order to eliminate the complexity of curve fitting all the compositions as functions of temperature and mass loss rate.

The reactions considered important in the char layer are given in Table A.4 along with kinetic data. A detailed discussion of the criteria for reaction selection and composition at the decomposition zone interface is given in Ref. A.1. The set of 15 reactions shown in Table A.4 are a result of applying this criteria for reaction selection. The procedure bases its selection on thermodynamic and kinetic consideration along with initial gas composition.

#### Summary

The conservation equations for the description of one dimensional char zone layer flow were presented in this section. The char was assumed to be degraded from a phenolic-nylon composition (40 wt. % nylon, 60 wt. % phenolic) at quasi-steady state. The pyrolysis gases were assumed to be in thermal equilibrium with the porous char

Table A.4  
Chemical Reactions Occurring  
in the Char Zone

---

$\text{CH}_4$	$\rightleftharpoons$	$\text{CH}_2 + \text{H}_2$
$\text{CH}_4$	$\rightleftharpoons$	$\text{CH}_3 + \text{H}$
$2\text{CH}_3$	$\rightleftharpoons$	$\text{C}_2\text{H}_6$
$\text{C}_2\text{H}_6$	$\rightleftharpoons$	$\text{C}_2\text{H}_4 + \text{H}_2$
$\text{C}_2\text{H}_4$	$\rightleftharpoons$	$\text{C}_2\text{H}_2 + \text{H}_2$
$\text{C}_2\text{H}_2$	$\rightleftharpoons$	$\text{C}_2\text{H} + \text{H}$
$\text{C(s)} + \text{H}_2\text{O}$	$\rightleftharpoons$	$\text{CO} + \text{H}_2$
$\text{C(s)} + \text{CO}_2$	$\rightleftharpoons$	$2\text{CO}$
$\text{H}_2 + \text{M}$	$\rightleftharpoons$	$2\text{H} + \text{M}$
$\text{H}_2\text{O} + \text{M}$	$\rightleftharpoons$	$\text{OH} + \text{H} + \text{M}$
$\text{CO}_2$	$\rightleftharpoons$	$\text{CO} + \text{O}$
$\text{H} + \text{CO}_2$	$\rightleftharpoons$	$\text{CO} + \text{OH}$
$\text{C}_6\text{H}_5\text{OH}$	$\rightleftharpoons$	$\text{H}_2\text{O} + \text{C}_6\text{H}_6$
$\text{C}_6\text{H}_6$	$\rightleftharpoons$	$3\text{C}_2\text{H}_2$

---

matrix and diffusional transport was considered negligible. The non-equilibrium species equations required numerical solution by an absolutely stable technique such as the Padé integration which was described in this section.

Heating rate results as a function of surface temperature and mass loss rate were curve fit using SAS (Statistical Analysis System). These heating rates can then be used as input to the flow field solution as a wall boundary condition. The other wall boundary conditions obtained from the char analysis are the species mass fractions. Due to the complexity of inputting compositions as a function of temperature and mass loss rate into the flow field, constant values of composition were used. These values were those obtained at a surface temperature of 3600°K and 0.04 lbm/sq ft sec -- conditions typical of planetary re-entry.

Fifteen chemical reactions and 19 species were included in the numerical solution. These reactions and species were identical to those used by del Valle (Ref. A.1). Selection of these was based on thermodynamic and kinetic considerations as discussed by del Valle.

Appendix B further specifies the wall boundary conditions by applying the Hertz-Knudsen equation to the char surface which relates surface temperature to wall pressure and mass loss rate.

## References

- A.1 del Valle, E. G., PhD dissertation, Louisiana State University (1975).
- A.2 Stroud, C. W., "A Study of the Chemical Reaction Zone in Charring Ablators During Thermal Degradation," Paper 33b, National Meeting of the AIChE, Dallas, Texas (February 6-9, 1966).
- A.3 Bird, R. B., W. E. Stewart, and E. N. Lightfoot, Transport Phenomena, John Wiley and Sons, Inc. New York, 828-30 (1960).
- A.4 Koh, J.C.Y., E. P. del Casal, R. W. Evans, and V. Deriugin, "Fluid Flow and Heat Transfer in High temperature Porous Matrices for Transpiration Cooling," Technical Report AFFDL-TR066-70, The Boeing Company, 207 pages (May 1966).
- A.5 Magnus, D. E. and H. S. Schechter, "Analysis of Error Growth and Stability for the Numerical Integration of the Equations of Chemical Kinetics." GASL TR-607, June 1966.
- A.6 Magnus, D., and H. Schechter, "Analysis and Application of the Padé Approximation for the Integration of Chemical Kinetic Equations," TR-642, GASL, (March 1967).
- A.7 Varga, R. S., "On Higher Order Stable Implicit Methods for Solving Parabolic Partial Differential Equations," J. Math. Physics, Vol. 40, pp. 220-331.
- A.8 Ban, A. J. and J. H. Goodnight, A Users Guide to the Statistical Analysis System, Department of Statistics, North Carolina State University, Raleigh, North Carolina (August, 1972).

APPENDIX B  
ON MODELING SUBLIMATION OF A CHARRING ABLATOR

Introduction

The determination of the sublimation rate of carbon for a char forming ablative heat shield is necessary to fully describe the material response to aerodynamic heating. This sublimation rate is required in the surface mass and energy balances which give the interaction of the flow field and the ablator.

The sublimation model that is used in this analysis and compared to other investigators is based on the Hertz-Knudsen equation. The basic equation originally given by Hertz in 1882 (Ref. B.1) relates in the mass of molecules striking a unit area per unit time to the pressure and temperature of the gas. The theoretical model is shown to compare quite favorably with arc-jet experimental data and demonstrates its usefulness by its accuracy and simplicity.

Model Description

The Hertz equation represents the condensation of a solid and is:

$$\dot{m}_c = \alpha \frac{A_r}{A} P \left( \frac{M}{2\pi RT} \right)^{\frac{1}{2}} \quad [1]$$

In the equation,  $\alpha$  is the fraction of molecules which strike the surface that condense and is referred to as an "accomodation coefficient." The ratio of available surface area,  $A_r$ , to the surface of total cross sectional area,  $A$ , is included in the equation, and for a perfect surface such as the face of a crystal,  $A_r$  equals  $A$ . In the case of rough or porous surfaces, the value of  $A_r$  will generally not be the same as  $A$ . Equation B.1 is a modification of the simple kinetic theory relation and gives the gross rate of condensation of the solid.

At equilibrium, the condensation rate is equal to the vaporization rate, and the pressure is equal to the equilibrium vapor pressure of the solid. This approach to solid vaporization was first suggested and experimentally verified by Knudsen (Ref. B.2) in 1915. The gross rate of evaporation is:

$$\dot{m}_v = \alpha \frac{A_r}{A} P_{\text{vap}} \left( \frac{M}{2\pi RT} \right)^{\frac{1}{2}} \quad [2]$$

For non-equilibrium such as the case where the subliming material is continually removed, the net rate of evaporation is given by the well known Hertz-Knudsen equation [Ref. B.3].

$$\dot{m} = \dot{m}_v - \dot{m}_c = \alpha \frac{A_r}{A} \left( \frac{M}{2\pi RT} \right)^{\frac{1}{2}} (P_{\text{vap}} - P) \quad [3]$$

The previous equation can be extended for vaporization of a solid into a mixture of gases by simply replacing the total pressure,  $P$ , with partial pressure of the vaporizing component,  $P_1$ .

$$\dot{m}_1 = \alpha_1 (1 - \epsilon) \left( \frac{M_1}{2\pi RT} \right)^{\frac{1}{2}} (P_{1, \text{vap}} - P_1) \quad [4]$$

Here the effective area ratio  $A_r/A$  is approximated by  $(1 - \epsilon)$  where  $\epsilon$  is the char porosity.

Carbon Species: Studies (Ref. B.4) have shown that the three primary constituents of vaporizing carbon are  $C_1$ ,  $C_2$ , and  $C_3$ . These species have Arrhenius dependence of vapor pressure on temperature of the form  $\log P_1 = a_1 + b_1/T$ . The accommodation coefficients,  $\alpha_1$ , along with the Arrhenius vapor pressure coefficients (Ref. B.5) are given in Table B.1.



Table B.1

Accommodation Coefficients and Arrhenius Coefficients  
for Vapor Pressure of  $C_1$ ,  $C_2$  and  $C_3$  (Ref. B.5)

Species	Accommodation Coefficient, $\alpha_1$	$\log P_{1,vap} = a_1 - b_1 \times 10^4/T$	
		$a_1$	$b_1$
$C_1$	0.37	8.14	3.72
$C_2$	0.34	9.69	4.23
$C_3$	0.08	9.81	4.03

Surface Mass Balance: The general species mass balance at the surface of the ablator includes convective flux to and away from the surface ( $\rho v C_i$ ), diffusive flux to and away from the surface ( $J_i$ ) and surface generation by chemical reaction ( $\omega_i$ ). The species mass balance is stated as follows:

$$\rho v C_i^- + \omega_i + J_i^- = \rho v C_i^+ + J_i^+ \quad [5]$$

In the region where the rate of sublimation is significant, reaction controlled or diffusion controlled surface chemical reactions will not occur (Ref. B.6). Also diffusion will not be important when the convective terms are large as is the case for the high heating rates encountered during manned return from planetary missions. The result is a porous char (solid carbon matrix) vaporizing into the gaseous species  $C_1$ ,  $C_2$ , and  $C_3$ . Also flowing through this porous matrix are the transpiring gases from the pyrolysis zone of the ablator. The composition of the pyrolysis gases was reported in error in Ref. B.6 for a phenolic-nylon ablator and corrected values are given in Table B.2. This table also includes a typical final composition of the pyrolysis gases at the surface after flowing through and reacting with the char as computed by a finite rate analysis, an extension of the work in Ref. B.7.

Combining the Hertz-Knudsen equation, Eq. B.4, with the surface mass balance for these previously described conditions for  $i$  being  $C_1$ ,  $C_2$ , or  $C_3$ , the mass flux of these gaseous carbon species injected into the shock layer is given by

$$\rho v C_i^+ = \alpha_i (1-\epsilon) \left( \frac{M_i}{2\pi RT} \right)^{\frac{1}{2}} (P_{i,vap} - P_i^+) \quad [6]$$

Table B.2  
Representative pyrolysis product composition of 40% (by weight)  
Nylon - 60% (by weight) phenolic resin ablative composite

Component	Composition of Pyrolysis Gases (Mass Percent)	
	Decomposition Zone	Ablator Surface
H <sub>2</sub>	3.03	3.81
CH <sub>4</sub>	3.87	0.08
C <sub>2</sub> H <sub>2</sub>	3.89	5.96
C <sub>2</sub> H <sub>4</sub>	3.90	*
C <sub>2</sub> H <sub>6</sub>	0.65	*
C <sub>6</sub> H <sub>6</sub>	2.59	*
C <sub>6</sub> H <sub>5</sub> OH	23.18	*
CO	4.18	25.70
CO <sub>2</sub>	4.62	*
H <sub>2</sub> O	5.65	*
N <sub>2</sub>	4.96	4.96
CH <sub>3</sub>	*	0.32
CH <sub>2</sub>	*	3.10
C <sub>2</sub> H	*	11.80
H	*	2.88
Carbon (solid)	<u>39.48</u>	<u>42.11</u>
Total	100.00%	100.00%

Element	Elemental Composition (mass percent)
C	73.03
H	7.29
N	4.96
O	<u>14.72</u>
Total	100.00%

\* Concentration is less than 0.01%.

or in terms of the total pressure at the surface

$$\rho v C_1^+ = \alpha_1 (1 - \epsilon) \left( \frac{M_1}{2\pi RT} \right)^{\frac{1}{2}} (P_{1,vap} - C_1^+ \bar{P} M_1) \quad [7]$$

Equation B.7 can be solved for the mass fraction of the gaseous carbon species  $C_1^+$  leaving the surface as:

$$C_1^+ = B P_{1,vap} / (\rho v + B \bar{P} M_1) \quad [8]$$

where

$$B = \alpha_1 (1 - \epsilon) \left( \frac{M_1}{2\pi RT} \right)^{\frac{1}{2}} \quad [9]$$

Finally a total mass balance on carbon says that all carbon coming to the surface as a solid leaves as a vapor.

$$C(s) = \sum_{i=1}^3 C_i^+ \quad [10]$$

This does not include the possibility of erosion as a surface removal mechanism, but erosion could be readily included if important.

Using Equations B.8, 9, and 10 and the data in Tables B.1 and B.2, the mass flux at the surface can be computed as a function of surface temperature. Also the composition of  $C_1$ ,  $C_2$  and  $C_3$  in the gases injected into the shock layer are determined.

### Results of the Analysis

In Figure B.1 a comparison is shown between our calculations using the Hertz-Knudsen equation and the results from Bishop and DiCristina (Ref. B.8) for a phenolic-carbon ablator. The porosity used in the Hertz-Knudsen equation was 0.2 which was the value given by Clayton et al. (Ref. B.9) for carbon-phenolic. The results

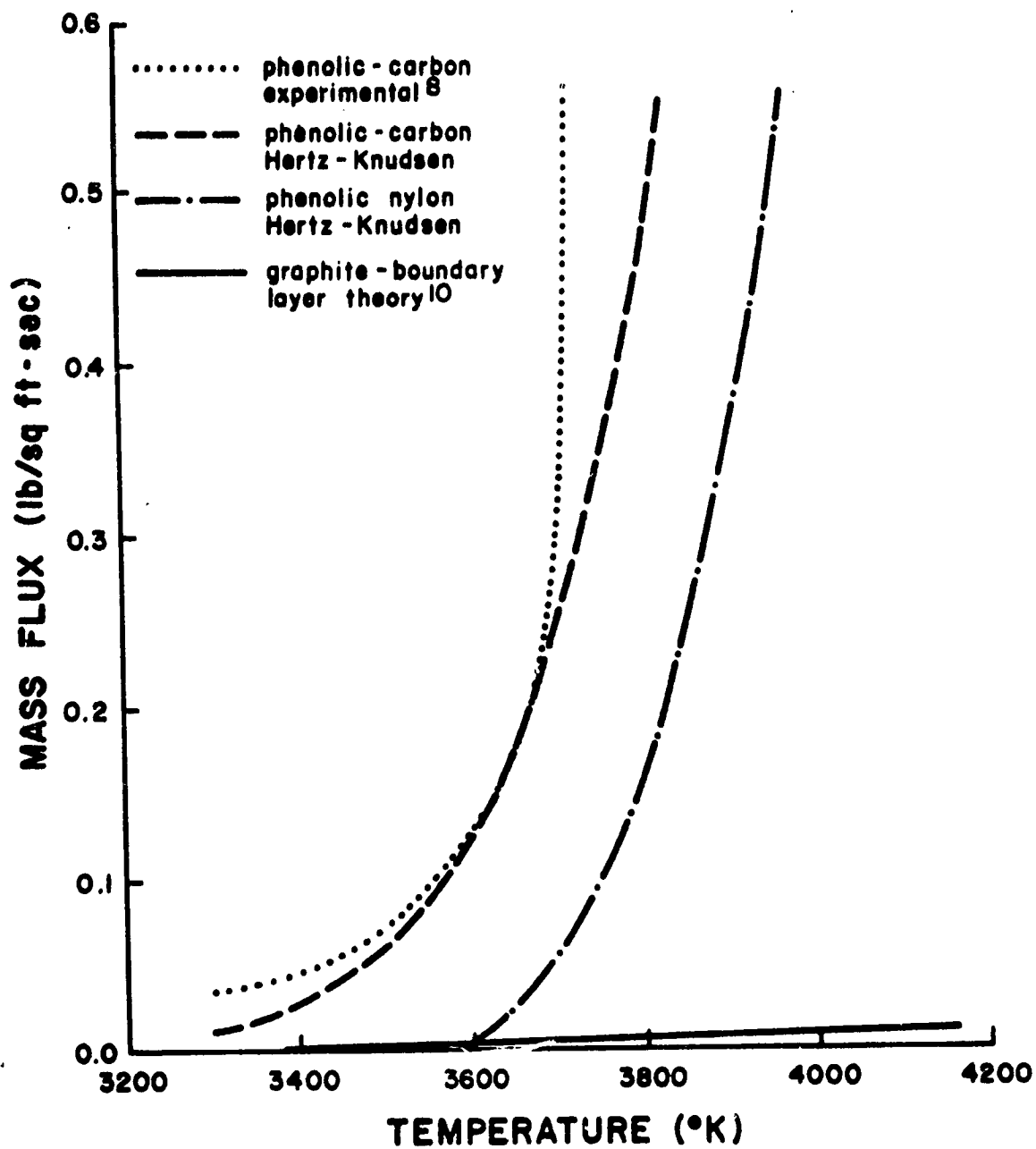


Figure B.1. Comparison of the Hertz-Knudsen Analysis with Other Investigators.

reported by Bishop and DiCristina (Ref. B.8) were obtained from arc jet experimental data for phenolic-carbon in the sublimation regime. The theoretical predictions of the Hertz-Knudsen sublimation model correspond quite well with the experimental curve except for some divergence at higher surface temperatures. Bishop and DiCristina stated in their paper that their curve would break less sharply if their surface temperature data had been corrected for the pyrometer error due to "significant nose curvature over the specimen area". Taking this fact into account, excellent agreement between the Hertz-Knudsen theoretical approach and the results from arc jet experimental work would be obtained.

Also shown in Figure B.1 are the results reported by Scala and Gilbert (Ref. B.10). These results were obtained using an equilibrium chemistry boundary layer analysis to predict the sublimation rate of graphite. These results differ greatly from both the Hertz-Knudsen predictions and the data of Bishop and DiCristina. The difference could be due to the fact that Scala and Gilbert's analysis assumed equilibrium between the graphite and the carbon vapor phase. Applying the Hertz-Knudsen analysis for graphite ablation gave essentially the same results as for the phenolic carbon. In the graphite analysis the porosity is zero, and there is no flow of pyrolysis products. These two effects shift the curve in opposite directions, and thus the graphite curve is essentially the same as that for carbon-phenolic which would correspond to Scala and Gilbert's results.

To demonstrate the difference that porosity can make on the mass flux, the Hertz-Knudsen analysis was applied to a phenolic-

nylon ablator with a porosity of 0.8. Referring to Figure B.1, increasing porosity has the effect of raising the surface temperature for a given mass flux.

In Figure B.2, the carbon species compositions for the phenolic-nylon ablator are given, and this shows that  $C_3$  is the predominant component that is obtained from the subliming char. This result agrees with Palmer and Shelef (Ref. B.4) who report that  $C_3$  becomes the primary constituent at these temperatures.

### Conclusions

The Hertz-Knudsen analysis has been shown to accurately predict the sublimation rate from a charring ablator. Porosity was shown to have a significant effect on the surface temperature - an increase in porosity results in an increase in surface temperature. The predominant carbon species in the vapor was  $C_3$  which agreed with previous investigations.

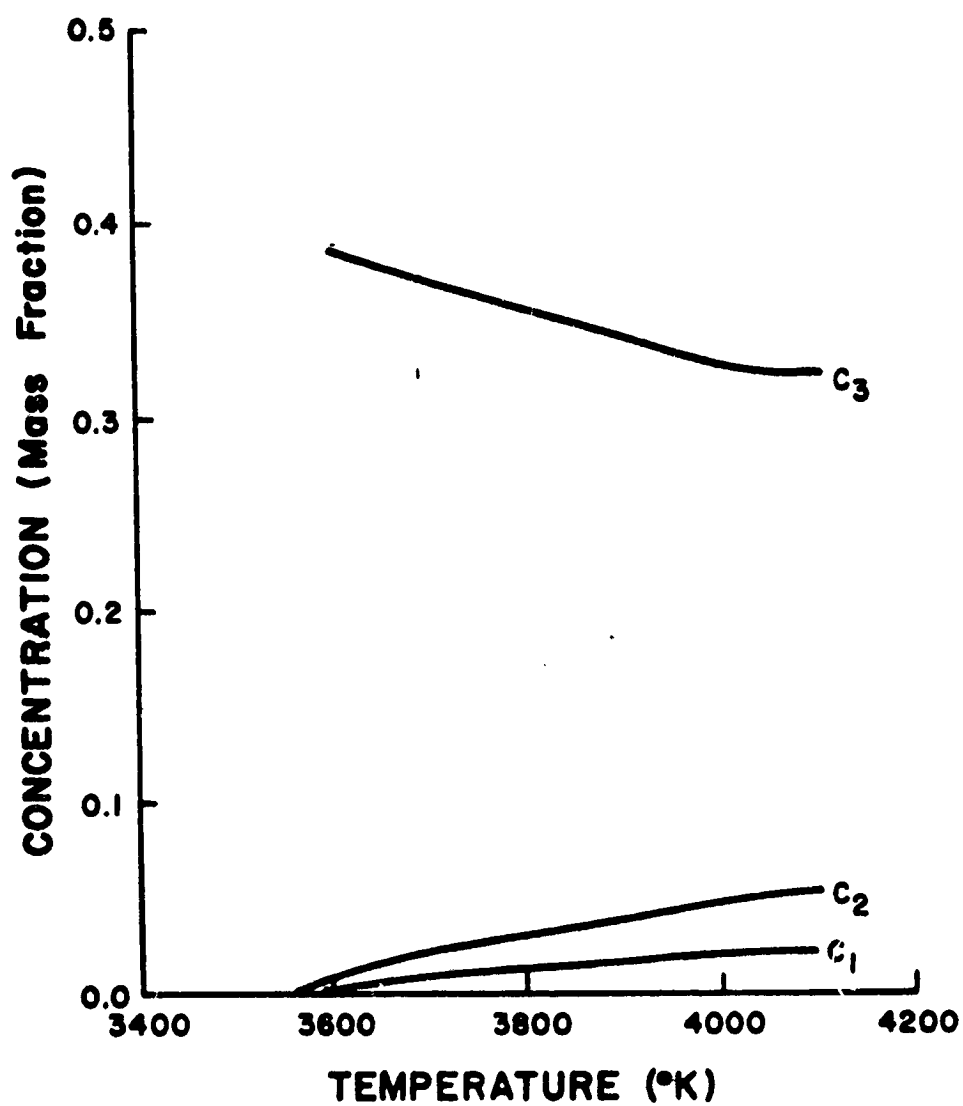


Figure B.2. Predominant Carbon Vapor Species from the Sublimation of a Phenolic Nylon Ablator.



## REFERENCES

- B.1. Somorhai, G.A., "Studies of the Sublimation Mechanism of Solids," in Advances in High Temperature Chemistry, ed. by Leroy Eyring. New York: Academic Press, 1969, Vol. 2, p. 205.
- B.2. Knudsen, Martin, "Die Maximale Verdampfungsgeschwindigkeit des Quicksilbers," Ann. Phy., 47, 697 (1915).
- B.3. Schrage, Robert W., A Theoretical Study of Interphase Mass Transfer, Columbia University Press, New York, p. 36, 1953.
- B.4. Palmer, H. B. and Mordecai Shelef, "Vaporization of Carbon," in Chemistry and Physics of Carbon, ed. by P. L. Walker, Jr., New York: Marcel Dekker, Inc., 1968, Vol. 4, p. 85 f.
- B.5. Thorn, R. J. and G. H. Winslow, "Vaporization Coefficient of Graphite and Composition of Equilibrium Vapor," J. Chem. Phys., 26, (1), 186 (1957).
- B.6. Pike, R. W., G. C. April, E. G. del Valle, and S. Hacker, "On Methods for Determining the Composition of Pyrolysis Products from Ablative Composites," JSR, 7 (10), pp. 1250-1253 (October 1970).
- B.7. April, G. C., R. W. Pike, E. G. del Valle, "Modeling Reacting Gas Flow in the Char Layer of an Ablator," AIAA Journal, 9 (6), pp. 1113-1119, (June 1971).
- B.8. Bishop, W. M. and V. DiCristina, "The Combustion and Sublimation of Carbon at Elevated Temperatures," AIAA Paper No. 68-759, Third Thermophysics Conf. (Los Angeles, California), June 24-26, 1968.
- B.9. Clayton, W. A., T. J. Fabish, J. F. Lagedrost, "Thermal Conductivity of Phenolic-Carbon Chars," AFML TR 69-313, (December 1969).
- B.10. Scala, S. M. and L. M. Gilbert, "Sublimation of Graphite at Hypersonic Speeds," AIAA Journal, 3 (9), 1635-1644, (1965).

## APPENDIX C

### POLYNOMIAL COEFFICIENTS FOR THERMODYNAMIC PROPERTIES

The thermodynamic properties of heat capacity,  $C_p^\circ$ , and enthalpy,  $H_T^\circ$ , are discussed in Chapter IV where it is stated that these variables are satisfactorily represented by a polynomial fit such as

Heat Capacity:

$$\frac{C_p^\circ}{R} = A_1 + A_2T + A_3T^2 + A_4T^3 + A_5T^4$$

Enthalpy:

$$\frac{H_T^\circ}{RT} = A_1 + \frac{A_2T}{2} + \frac{A_3T^2}{3} + \frac{A_4T^3}{4} + \frac{A_5T^4}{5} + \frac{A_6}{T}$$

The determination of these coefficients is discussed by Esch (Ref. C.1) and are presented in Table C.1. For each species two ranges of temperature were fit - 1000 - 7000°K and 5000°K - 18,000°K. The overlapping of the temperature ranges was necessary to overcome accuracy limitations at the extremes of the fit.

Table C.1  
POLYNOMIAL COEFFICIENTS FOR THERMODYNAMIC PROPERTY CORRELATIONS

Species	A1	A2	A3	A4	A5	A6	A7	T*
C*	0.2609E C1	-0.1393E-03	0.5959E-07	-0.1037E-10	0.6349E-15	0.2169E 06	0.3709E C1	L
	0.2528E C1	0.4469E-05	-0.7026E-09	0.1134E-11	-0.3476E-16	0.2169E 06	0.4139E C1	H
H	0.2500E C1	-0.8243E-06	0.4421E-09	-0.1720E-12	0.1457E-16	0.2547E 05	-0.4612E C1	L
	0.3334E C1	-0.1776E-02	0.6012E-06	-0.7919E-10	0.3482E-14	0.2547E 05	-0.4598E C1	F
N*	0.2727E C1	-0.2120E-03	0.1105E-06	-0.1591E-10	0.7447E-15	0.2254E 06	0.3645E C1	L
	0.2499E C1	-0.3725E-05	0.1147E-07	-0.1102E-11	0.3072E-16	0.2254E 06	0.4950E C1	H
O*	0.2491E C1	0.2762E-04	-0.1881E-07	0.3807E-11	-0.1028E-15	0.1879E 06	0.4424E C1	L
	0.2944E C1	-0.4108E-03	0.9156E-07	-0.5948E-11	0.1190E-15	0.1879E 06	0.1750E C1	F
E	0.2500E C1	0.3440E-06	-0.1454E-09	0.3937E-13	-0.2573E-17	-0.7450E 03	-0.1173E C2	L
	0.2508E C1	-0.6332E-05	0.1364E-08	-0.1094E-12	0.2934E-17	-0.7450E 03	-0.1203E C2	F
C	0.2612E C1	-0.2030E-03	0.1095E-06	-0.1039E-10	0.8550E-15	0.8542E 05	0.4144E C1	L
	0.2141E C1	0.3219E-03	-0.5498E-07	0.3604E-11	-0.5564E-16	0.8542E 05	0.6874E C1	F
CN	0.3411E C1	0.4097E-03	0.1005E-06	-0.3473E-10	0.2361E-14	0.4745E 05	0.4746E C1	L
	0.3473E C1	0.7337E-03	-0.9082E-07	0.4847E-11	-0.1018E-15	0.5820E 05	0.4152E C1	H
CO	0.3254E C1	0.9098E-03	-0.2647E-06	0.3037E-10	-0.1177E-14	-0.1434E 05	0.4875E C1	L
	0.3366E C1	0.8227E-03	-0.1768E-06	0.1940E-10	-0.5549E-15	-0.1434E 05	0.4203E C1	F
C2	0.4443E C1	-0.2485E-03	0.3036E-06	-0.6244E-10	0.3915E-14	0.9787E 05	-0.1030E C1	L
	0.4026E C1	0.4057E-03	-0.7026E-07	0.4666E-11	-0.1142E-15	0.9787E 05	0.1030E C1	F
C2H	0.3485E C1	0.3563E-02	-0.1237E-05	0.1866E-09	-0.1013E-13	0.5809E 05	0.4784E C1	L
	0.5307E C1	0.8966E-03	-0.1378E-06	0.3251E-11	-0.2278E-15	0.5809E 05	-0.5288E C1	F

\*Temperature Range L = 1000-6000°K, H = 6000-15,000°K

(continued)

Table C.1 (Cont'd)

Species	A1	A2	A3	A4	A5	A6	A7	T*
C <sub>2</sub> H <sub>2</sub>	0.3391E 01	0.5717E-02	-0.1937E-05	0.2931E-09	-0.1585E-13	0.2570E 05	0.6920E 00	L
	0.6789E 01	0.1503E-02	-0.2295E-06	0.1534E-10	-0.3763E-15	0.2590E 05	-0.1539E 02	H
C <sub>3</sub>	0.4002E 01	0.3541E-02	-0.1318E-05	0.2004E-09	-0.1144E-13	0.9423E 05	0.2020E 01	L
	0.2213E 02	-0.1759E-01	0.5565E-05	-0.6758E-09	0.2825E-13	0.9423E 05	-0.1021E 03	F
C <sub>3</sub> H	0.3365E 01	0.6200E-02	-0.2265E-05	0.3717E-09	-0.2262E-13	0.6283E 05	0.3467E 01	L
	0.3365E 01	0.6200E-02	-0.2265E-05	0.3717E-09	-0.2262E-13	0.6283E 05	0.3467E 01	F
C <sub>4</sub> H	0.5874E 01	0.7403E-02	-0.2729E-05	0.4437E-09	-0.2677E-13	0.7605E 05	-0.4010E 01	L
	0.5874E 01	0.7403E-02	-0.2729E-05	0.4437E-09	-0.2677E-13	0.7605E 05	-0.4010E 01	F
HCN	0.3654E 01	0.3444E-02	-0.1258E-05	0.2169E-09	-0.1430E-13	0.1442E 05	0.2373E 01	L
	0.3654E 01	0.3444E-02	-0.1258E-05	0.2169E-09	-0.1430E-13	0.1442E 05	0.2373E 01	F
H <sub>2</sub>	0.3358E 01	0.2794E-03	0.9372E-07	-0.2948E-10	0.2141E-14	-0.1018E 04	-0.3548E 01	L
	0.3358E 01	0.2794E-03	0.9372E-07	-0.2948E-10	0.2141E-14	-0.1018E 04	-0.3548E 01	F
N	0.3363E 01	0.4656E-03	-0.5127E-07	0.2802E-11	-0.4905E-16	0.5609E 05	0.4300E 01	L
	0.2474E 01	0.9097E-04	-0.7914E-07	0.2218E-10	-0.1485E-14	0.5609E 05	0.4300E 01	F
C	0.2746E 01	-0.3309E-03	0.1338E-06	-0.1191E-10	0.3369E-15	0.5609E 05	0.2872E 01	L
	0.2746E 01	-0.3309E-03	0.1338E-06	-0.1191E-10	0.3369E-15	0.5609E 05	0.2872E 01	F
O	0.2548E 01	-0.1970E-03	0.7153E-07	-0.8901E-11	0.4002E-15	0.2015E 05	0.4504E 01	L
	0.2548E 01	-0.1970E-03	0.7153E-07	-0.8901E-11	0.4002E-15	0.2015E 05	0.4504E 01	F
N <sub>2</sub>	0.3221E 01	0.5078E-03	-0.2907E-06	0.3932E-10	-0.2500E-14	0.2015E 05	0.5049E 01	L
	0.3221E 01	0.5078E-03	-0.2907E-06	0.3932E-10	-0.2500E-14	0.2015E 05	0.5049E 01	F
C <sub>2</sub>	0.3316E 01	0.1151E-02	-0.3726E-06	0.1154E-10	-0.3253E-15	-0.1043E 04	0.1254E 01	L
	0.3316E 01	0.1151E-02	-0.3726E-06	0.1154E-10	-0.3253E-15	-0.1043E 04	0.1254E 01	F
	0.3372E 01	0.4254E-03	-0.2835E-07	0.0000E-12	-0.5180E-17	-0.1044E 04	0.3254E 01	F

\*Temperature Range L = 1000-6000°K, H = 6000-15,000°K

## REFERENCES

- C.1 Esch, D. D., "Stagnation Region Heating of a Phenolic-Nylon Ablator During Return from Planetary Missions," Ph.D. Dissertation, Louisiana State University, Baton Rouge, Louisiana (1971).

## APPENDIX D

### USERS INSTRUCTIONS FOR CRAC COMPUTER PROGRAM

#### Introduction

The purpose of this appendix is to describe the procedure for using the Coupled, Radiating, and Chemically reacting shock layer computer program (CRAC). This program consists of a numerical solution of the viscous thin shock layer equations presented in Chapter III coupled to a phenolic nylon ablator by the equations presented in Appendix A. A general description of the overall logic and the details of the numerical solutions were given in Chapter V.

The program CRAC has been written in such a way that a minimum of input is required for the solution of any given problem. The required input is a specification of freestream velocity, freestream density, and body nose radius. From this input, the flow field profiles are calculated for a blowing rate of 5% of free-stream. A namelist option is provided in the input which allows the specification of a number of critical variables (such as blowing rate, surface temperature, etc.) and critical profiles if the user has a good idea of their values. If no value is input, the program internally estimates all variables and profiles.

The program outputs boundary condition information at wall and shock including ablator heating response for the assumed blowing rate. The shock layer solution provides temperature, pressure, density, velocity, species fractions, and radiative flux divergence

profiles along with flow field physical property profiles of thermal conductivity, viscosity and heat capacity. The heat flux from shock layer to ablator surface is also output for comparison with the ablator heating response. The ablator heating response is calculated by the program using the analysis and results presented in Appendices A and B. When these are equal, the solution has been coupled as discussed in Chapter V.

CRAC has been set up to be as versatile as possible with subroutines that can be altered or replaced to provide more sophistication, different ablators, different atmospheres, etc. For instance, if a different ablator is desired, subroutine BCWALL can be replaced to reflect the material response for the new ablator. If a different body shape is needed (the assumed shape is spherical), subroutine SHAPE can be replaced. The subroutines are listed below with their description which may require alteration for application to a different system.

BCWALL - Wall boundary conditions (a function of the ablator)

BLOCK DATA - Species, thermodynamic, and transport property fits, and initial guesses

FGH - Finite difference method

FG2 - Kinetic production rates of species (chemical reactions)

GAS - Freestream gas equilibrium model

LRAD - Radiation properties of flow field species

SHAPE - Body shape

The program solves the shock layer equations about a blunt body and always starts at  $\xi = 0$ , the stagnation line. The program will

handle solutions in an arbitrary freestream where the velocity, density, and species mass fraction define the environment.

In the iteration procedure employed to obtain the initial profiles, the stagnation line momentum is solved first, followed by an iteration on the species equations. The energy equation is not solved in this initialization step. This iteration is stopped when NOE iterations (internally set at NOE = 50 but can be changed by namelist option) have been performed. The solution then proceeds to the stagnation line iteration scheme. The radiative flux divergence is solved first using the initial profiles. Then MI iterations (internally set at MI = 5) are performed on the shock layer stagnation line equations using the initial profiles from above. An iteration of the species equations within this larger iteration is possible depending on the value of MIS (internally set at MIS = 5).

In the iteration procedure it is often necessary to weigh the calculated solution with the assumed solution as was discussed in Chapter V. The damping values used for this are HDAMP for enthalpy, TDAMP for temperature, CDAMP for mass fractions, FDAMP for the momentum solution, and DDAMP for the transformed shock stand off distance,  $\tilde{\delta}$ . The appropriate values to employ depend on the case being run and can be determined only from experience. Internally set values are given in Table D.1.

A description of input variables is given in Table D.1 and a brief description of each of the subprograms included in the computer code is given in Table D.2. Finally, a diagram of the program overlay structure used on the IBM system 360/65 is presented in Figure D.1



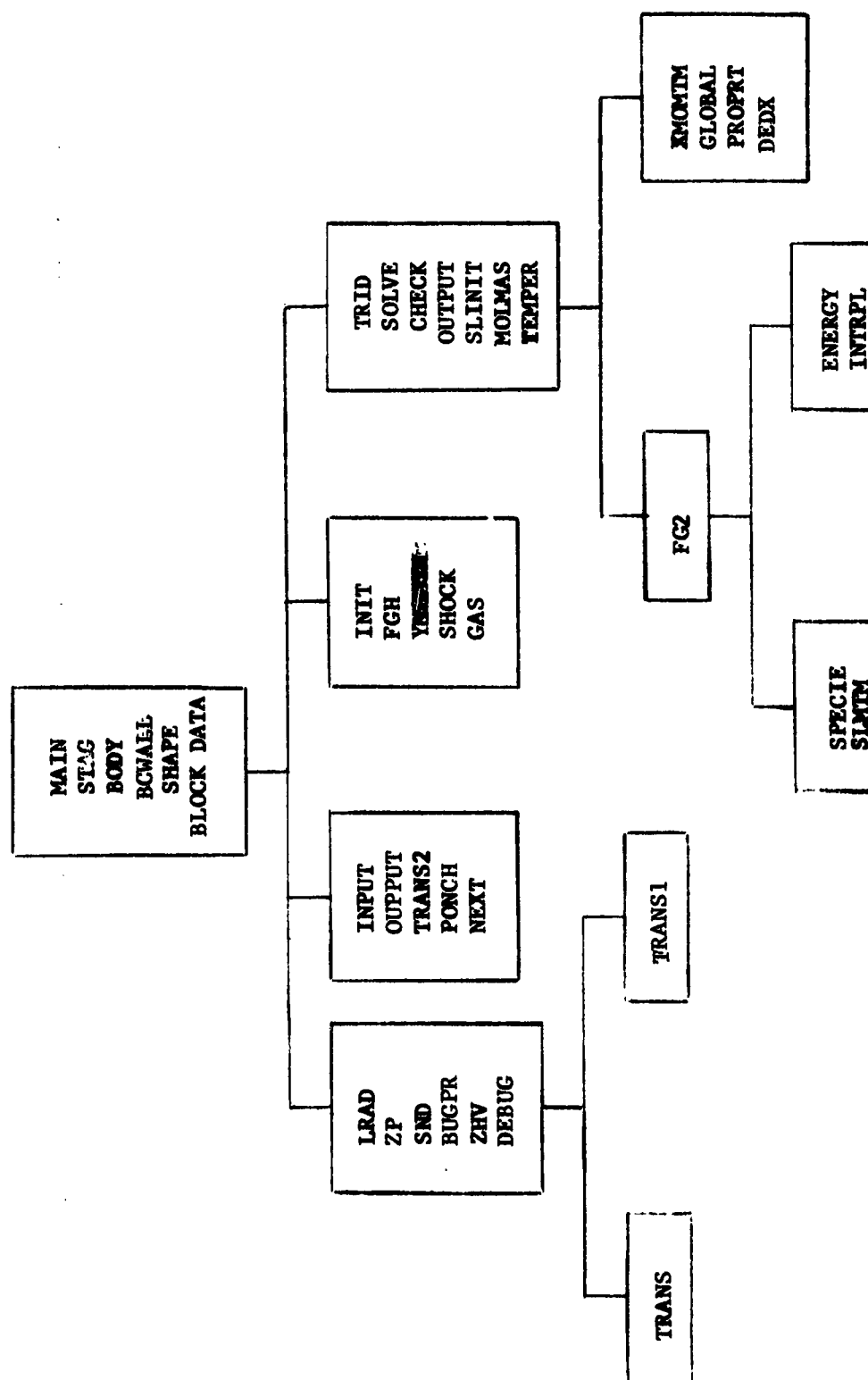


Figure D.1. CRAC overlay Structure

which reduces considerably the core requirements (from 250K bytes down to 186K bytes). This overlay allows the whole program to be run totally in high speed core.

Table D.1  
CRAC Input Variables

## CARD 1

## Title Card

<u>Fortran Variable</u>	<u>Format</u>	<u>Columns</u>	<u>Description</u>
TITLE	18A4	1-72	Any description information desired

## CARD 2

UINF	E10.1	1-10	Freestream velocity in ft/sec
RINF	E10.1	11-20	Freestream density in slugs/ft <sup>3</sup>
R	E10.1	21-30	Body radius, ft.
ISTART	I5	31-35	Program code 0 = initial start-up at stagnation line using only namelist input 1 = initial start-up at stagnation line using input deck from another run 2 = continuation of stagnation line solution using punched data deck 3 = around the body solution using punched, converged stagnation line solution data

Table D.1 (continued)

Namelist Option

<u>Variable</u>	<u>Variable Type</u>	
NETA	Integer	Number of points along the coordinate normal to the body [59]*
MI	Integer	Number of iterations made with conservation equations for each pass through LRAD radiation flux divergence calculation [5]
MIS	Integer	Number of internal iterations on the species equations [5]
MAXE	Integer	Maximum number of iterations for enthalpy-temperature calculation [10]
MAXM	Integer	Maximum number of iterations for the momentum equation [10]
NTIME	Integer	Minutes of CPU time remaining when user wishes to have program stop and punch output deck [0]
IDGD	Integer	Debug printout on $\delta$ [0] 0 = no output (default value) 2 = debug output
IDGF	Integer	Debug printout on x-momentum equation [0]
IDGH	Integer	Debug printout on enthalpy equation [0]
IDGT	Integer	Debug printout on temperature equation [0]
IDGS	Integer	Debug printout on species equation [0]
IDEBUG	Integer	Debug printout on stagnation line x-momentum equation [0]
CDAMP	Real	Damping coefficient for species [1.0]
DDAMP	Real	Damping coefficient for $\delta$ [0.5]
HDAMP	Real	Damping Coefficient for enthalpy [0.5]

\*[] Default value internally set

Table D.1 (continued)

<u>Variable</u>	<u>Variable Type</u>	
TDAMP	Real	Damping coefficient for temperature [0.25]
TPRCT	Real	Convergence tolerance on enthalpy- temperature calculation [0.01]
FPRCT	Real	Convergence tolerance on stagnation x-momentum equation [0.005]
PDTIL	Real	Convergence tolerance on $\tilde{\delta}$ [0.001]
TWOLD	Real	Estimate of surface temperature, °K [3600°K]
RVW	Real	Ablator blowing rate [0.05]
CWALL	Real Vector	Species mass fractions at ablator surface
RVN	Real Vector	Estimate of blowing rate around the body normalized to the stagnation value
DXI	Real	Tangential coordinate stepsize [0.05]
ETA	Real Vector	Normal coordinate nodal points
T	Real Vector	Estimate of temperature profile
IAB	Integer	Number of points in vectors RVN and EPSN [20]
EPSN	Real Vector	Estimate of $\epsilon$ profile around the body
XIN	Real Vector	Coordinate points for EPSN and RVN

Table D.2  
Description of CRAC Subroutines

<u>Subroutine</u>	<u>Description</u>
MAIN	Driver Program
INPUT	Reads in all necessary data
STAG	Driver program for stagnation line solution
SLINIT	Initialization routine for stagnation line
BODY	Driver program for around the body solution
SHAPE	Routine to calculate body and shock angles for a specified body shape (spherical)
BCWALL	Calculates ablator surface temperature, and ablator surface heating for a specified blowing rate and wall pressure
SHOCK	Calculates density ratio at the shock, shock temperature, pressure, and species mass fractions
GLOBAL	Routine to solve the global continuity equation around the body
SLMTM	Solution of the stagnation line x-momentum equation
XMOMTM	Solution of the around the body x-momentum equation
YMOMTM	Solution of the y-momentum equation
SPECIE	Solution of the species equations
ENERGY	Solution of the energy equation in enthalpy form
TEMPR	Temperature-enthalpy iteration to obtain temperature profile from the enthalpy profile
DEDX	Calculation of shock standoff distance, $\delta$ and $\partial\eta/\partial x$
MOLMAS	Changes species concentration from moles 1/mass total to mass fraction and vice versa
GAS	Calculates chemical equilibrium for air species along with thermodynamic and transport properties

Table D.2 (continued)

<u>Subroutine</u>	<u>Description</u>
FG2	Calculates chemical production terms and linearized chemical production terms for species conservation equation
LRAD, TRANS, SND, ZP, BUGPR, ZHV, & TRANS 1, TRANS 2	Calculates radiative flux divergence for specified temperature, pressure, and composition profiles
SOLVE	Takes $\alpha$ coefficients calculated in conservation subroutines XMOMTM, SPECIE, and ENERGY and sets up tridiagonal matrix for TRID
TRID	Solves the tridiagonal matrix
FGH	Finite difference coefficients
CHECK	Routine to check for converged solutions of profiles
INTRPL	Lagrange interpolation scheme
OUTPUT	Prints output for shock layer solution
OUTPUT	Debug output of intermediate profiles
PONCH	Punches solution on cards
BLOCK DATA	Contains thermodynamic and transport property data along with initialization data

**APPENDIX E**

**CRAC SOURCE DECK, SAMPLE INPUT, AND SAMPLE OUTPUT**



MAIN	1C
MAIN	2C
MAIN	3C
MAIN	4C
MAIN	5C
MAIN	6C
MAIN	7C
MAIN	8C
MAIN	9C
MAIN	10C
MAIN	11C
MAIN	12C
MAIN	13C
MAIN	14C

C R A C

COUPLED RADIATING AND CHEMICALLY REACTING SHOCK LAYER  
 BY JOHN FREDERICK BALHOFF FEB 1975

AROUND THE BODY SOLUTION

CALL INPUT  
 CALL INIT  
 CALL STAG  
 CALL BODY  
 CALL EXIT  
 END

C C C C C C C C

```

C
SUBROUTINE INIT
COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA
COMMON/PROPI/P1(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
COMMON/MIKE/ NETA1,NETA2,PNF,VNF,RHOD,DP,TS,UD,VD,RW,RWO,EPS,PHI
COMMON /BILL/ KODE(20),KOPI(10),CSHOCK(20)
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /DICK/ Q(60),OO(60)
COMMON/E02/AA(20,5),ICODE(20)
COMMON/E03/IA(20,5)

NETA1=NETA-1
NETA2=NETA-2
UINF2 = UINF+UINF
PNF=2116./RDZ/UINF2
VNF = R/UINF
RHOD=RDZ*32.174
DO 5 K=1,NSP
C(K,NETA) = CSHOCK(K)
5 CONTINUE
CALL RADIN
DO15 M=1,NE
DO15 I=1,NSP
AA(I,M) = IA(I,M)
15 CONTINUE
Q(1) = 0.0
Q(NETA) = 1.0
DO 35 J=2,NETA1
Q(J) = Z(J-1)*DTIL
CALL FGH (J,F1(J),G1(J),H1(J))
CALL FGH2(J,F2(J),G2(J),H2(J))
35 CONTINUE
INIT 10
INIT 20
INIT 30
INIT 40
INIT 50
INIT 60
INIT 70
INIT 80
INIT 90
INIT 100
INIT 110
INIT 120
INIT 130
INIT 140
INIT 150
INIT 160
INIT 170
INIT 180
INIT 190
INIT 200
INIT 210
INIT 220
INIT 230
INIT 240
INIT 250
INIT 260
INIT 270
INIT 280
INIT 290
INIT 300
INIT 310
INIT 320
INIT 330
INIT 340
INIT 350
INIT 360

```

C

INIT 370  
INIT 380

RETURN  
END

```

SUBROUTINE INPUT
** ROUTINE TO READ AND PRINT ALL INPUT DATA **

COMMON /CONV/ FPRCT,TPRCT,DDAMP,TDAMP,PDITL
COMMON/CONV1/HDAMP,FDAMP,CDAMP
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON /MAIN/KEEP,MAXE,MAXM,MAXO,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /NON/RDZ,MUDZ,RNDZ,AKNF,HNF,CPNF
COMMON/PROPI/PI(60),RHO(60), T(60),AKW(60),C (20,60),CC(5,60)
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/PROP3/CP5(20,60),H5(20,60),CP (60),HM(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON /RH/ DUD,DPH,TD,RZB,PD,HD,HTOTAL
COMMON/WALL/RVM,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /YL/ETA(60),YOND(60)
COMMON/TIT/TITLE(18)

COMMON/EQ1/AI(20), BI(20), CI(20), DI(20), EI(20), FI(20), GI(20),
X      AII(20),BII(20),CII(20),DII(20),EII(20),FII(20),GII(20)
COMMON/EQ2/AA(20,5), ICODE(20)
COMMON/EQ3/IA(20,5)
COMMON/ID/SP(20),EL(5)
COMMON/WT/SMW(20),AWT(5)
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/SP1/SS,TOL,NDEBUG
COMMON/JFB/ IPN
COMMON/MIKE/ META1,META2,PNF,WNF,RHOD,DP,TS,UD,VD,RW,RWD,EPS,PHI
COMMON/JOHN/ UO(60),U(60),DU(60),VO(60),RO(60),DMDX(60),DTLA,DTILOINRU
COMMON /BILL/ NODE(20),I START, IDGD, IDGF, IDGH, IDGT, IDGS.
X MI,NOE,NTIME,MIS

```

```

INPU 10
INPU 20
INPU 30
INPU 40
INPU 50
INPU 60
INPU 70
INPU 80
INPU 90
INPU 100
INPU 110
INPU 120
INPU 130
INPU 140
INPU 150
INPU 160
INPU 170
INPU 180
INPU 190
INPU 200
INPU 210
INPU 220
INPU 230
INPU 240
INPU 250
INPU 260
INPU 270
INPU 280
INPU 290
INPU 300
INPU 310
INPU 320
INPU 330
INPU 340
INPU 350
INPU 360

```

```

COMMON/GUESS/TG1(60),TG2(60)
COMMON/BOB/ XI,DXI,YL
COMMON /GEORGE/ EPSN(30),XIN(30),QWN(30)

REAL MU,MUDZ
LOGICAL MCONV,ECONV,DCONV
DATA END /'END '/
NAMELIST/ENTRY/ NETA,TS,RW,EPS,IRAD,IDEBUG,TG2,IDGD,IDSF,IDGH,
X IDST,IOGS,RVM,MAXE,TWOLD,CC,TOL,CWALL,ISTART,MI,NOE,NTIME,MIS
X ,HDAMP,TDAMP,CDAMP,XIF,XIN,EPSN,QWN,IAB,DXI,ETA

C ** INPUT FORMATS **
C
C 100 FORMAT (10A4,2I4)
C CARD 1 -----
READ (5,100) TITLE
READ(5,10) UINF,RINF,R,ISTART
10 FORMAT(3E10,3,I5)
IF (ISTART .EQ. 0) GO TO 1000
*** RE-START ***
C
C
READ(5,1001) FPRCT,TPRCT,DDAMP,TDAMP,PDITL,HDAMP
READ(5,1001) DF,TA,DTIL,DTILS
READ(5,1002) U INF, RINF, UINF2, R , RE, ITM, IEM, NETA
READ(5,1003) KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
READ(5,1001)RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
READ(5,1004) NSP,NNS,NE,NC
READ(5,1001) (PI(J),RHO(J),T(J),J=1,NETA)
READ (5,1001)(AMW(J),J=1,NETA)
READ(5,1001) ((C(I,J),J=1,NETA),I=1,NSP)
READ(5,1001)( HM(J),J=1,NETA)
READ(5,1001) (E(J),J=1,NETA)
READ(5,1004) IRAD,ITYPE
READ(5,1001)DUD,DPHI,TD,RZB,PD,HD,HTOTAL
READ (5,1001)(F(J),FC(J),Z(J),V(J),J=1,NETA)
READ(5,1001)RVM,PRN,TWOLD
INPU 370
INPU 380
INPU 390
INPU 400
INPU 410
INPU 420
INPU 430
INPU 440
INPU 450
INPU 460
INPU 470
INPU 480
INPU 490
INPU 500
INPU 510
INPU 520
INPU 530
INPU 540
INPU 550
INPU 560
INPU 570
INPU 580
INPU 590
INPU 600
INPU 610
INPU 620
INPU 630
INPU 640
INPU 650
INPU 660
INPU 670
INPU 680
INPU 690
INPU 700
INPU 710
INPU 720

```

```

READ(5,1801) (CWall(I),I=1,NSP)
READ(5,1801) (ECWall(K),K=1,NE)
READ(5,1801) (ETA(J),YOND(J),J=1,NETA)
  READ(5,715)NDBG, IAB, TOL
READ (5,1601)((CC(K,J),J=1,NETA),K=1,NE)
READ(5,1801) XI,RW,TS,U(NETA)

1001 FORMAT(6E13.5)
1002 FORMAT(5E12.5,3I3)
1003 FORMAT(5I5,3L3,2I5)
1004 FORMAT(4I4)
  715  FORMAT(2I5,E15.6)

C
1000 READ(5,ENTRY)
C
C
C
C
  *** PRINT DATA ***
  RETURN
END

```

STAG 10  
STAG 20  
STAG 30  
STAG 40  
STAG 50  
STAG 60  
STAG 70  
STAG 80  
STAG 90  
STAG 100  
STAG 110  
STAG 120  
STAG 130  
STAG 140  
STAG 150  
STAG 160  
STAG 170  
STAG 180  
STAG 190  
STAG 200  
STAG 210  
STAG 220  
STAG 230  
STAG 240  
STAG 250  
STAG 260  
STAG 270  
STAG 280  
STAG 290

# SUBROUTINE STAG

```

C
COMMON /FRSTRM/ U INF. RINF. UINF2. R . RE. LXI. IYM. IEM. NETA
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/BOB/ XI,DXI,YL
COMMON /BILL/ NNODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS.
X NI,NOE,NTIME,MIS

C
IF(ISTART .GT. 2) RETURN
CALL SHAPE
CALL SHOCK
CALL BCWALL
CALL SLINIT
CALL NEXT
1 CONTINUE
CALL PROPRT (NSP,1,NETA)
CALL LRAD
DO 100 IEM=1,M1
CALL SLWTM
DO 15 I=1,MIS
CALL SPECIE (KODS)
15 CONTINUE
CALL PROPRT (NSP,1,NETA)
CALL ENERGY (KODE)
IF(KODS.GT.3 .AND. KODE.GT.4) GO TO 1000
100 CONTINUE
1000 CALL GUPPUT (N)
GO TO 1
END

```

```

SUBROUTINE BODY
COMMON/BOB/ XI,DXI,YL
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON /FRSTRM/ U INF, RINF, UINF2, R , RE, LXI, ITM, IEM, META
COMMON /BILL/ NODE(20),ISTART,IDGD,IDCF,IDGH,IDGT,IDGS.
X MI,NOE,NTIME,MIS
1 CONTINUE
XI = XI + DXI
CALL SHAPE
CALL SHOCK
CALL BCWALL
CALL NEXT
CALL CEDX
CALL YMOMTM
2 CONTINUE
CALL PROPT (NSP,1,META)
CALL LRAD
DO 100 IEM=1,MI
6 CONTINUE
YL=0.5
CALL XMOMTM (KODE)
CALL YMOMTM
YL = 1.0
CALL GLOBAL (KOD )
IF(KOD .LT. 5 .OR. KODE .LT. 5) GO TO 6
YL=1.0
DO 25 I=1,MIS
CALL SPECIE (KODS)
25 CONTINUE
CALL PROPT (NSP,1,META)
CALL ENERGY (KODT)
IF(KODS.GT.3 .AND. KODT.GT.4) GO TO 1000
100 CONTINUE
C
C
1000 CALL OUPPUT (N)

```

```

BODY 10
BODY 20
BODY 30
BODY 40
BODY 50
BODY 60
BODY 70
BODY 80
BODY 90
BODY 100
BODY 110
BODY 120
BODY 130
BODY 140
BODY 150
BODY 160
BODY 170
BODY 180
BODY 190
BODY 200
BODY 210
BODY 220
BODY 230
BODY 240
BODY 250
BODY 260
BODY 270
BODY 280
BODY 290
BODY 300
BODY 310
BODY 320
BODY 330
BODY 340
BODY 350
BODY 360

```



BODY 370  
BODY 380

GO TO (1.2). N  
END

```

C
C
C
C
SUBROUTINE NEXT
ROUTINE TO DETERMINE GUESSES FOR PROFILES AT THE M + 1
BODY STATION USING THE PROFILES AT THE MTH STATION

COMMON/TOM/ SC(60),TO(60),PR(60),TE(60)
COMMON/JTB/ CO(60,20),RED,RED2
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/PROP1/PI(60),RHO(60), T(60),AM2(60),C (20,60),EC(5,60)
COMMON /RM/ DJD,DPHI,TD,RZB,PD,MD,KTOTAL
COMMON /FRSTRM/ U INF, RINF, UINF2, R , RE, LXI, ITM, IEM, META
COMMON/MIKE/ META1,META2,PNF,WNF,RHOD,DP,TS,UD,VD,RV,RWD,EPS,PHI
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON/JUM/ UO(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILO
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /YL/ETA(60),YOND(60),YO(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON/DON/ PO(60),EO(60),HMO(60)
COMMON/PROP3/CP(20,60),H(20,60),CPM(60),HM(60)
COMMON /DICK/ Q(60),QO(60)
COMMON/WALL/RVW,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)

DTILO = DTIL
DTLA=DTIL
T1 = TWOLD/TS
TDO = T(META)*TS
TW = T(1)
TWO = TW*TS
RHO1 = PI(1)*AMW(1)/TWOLD/1.3157/32.174
V1 = RVW*INF/RHO1
VDO = V(META)
VO1 = V(1)
FP = 0.0
DO 5 I=1,META
  TO(I)=T(1)
  OO(I) =O(I)

```

```

NEXT 10
NEXT 20
NEXT 30
NEXT 40
NEXT 50
NEXT 60
NEXT 70
NEXT 80
NEXT 90
NEXT 100
NEXT 110
NEXT 120
NEXT 130
NEXT 140
NEXT 150
NEXT 160
NEXT 170
NEXT 180
NEXT 190
NEXT 200
NEXT 210
NEXT 220
NEXT 230
NEXT 240
NEXT 250
NEXT 260
NEXT 270
NEXT 280
NEXT 290
NEXT 300
NEXT 310
NEXT 320
NEXT 330
NEXT 340
NEXT 350
NEXT 360

```

C

NEXT 370  
 NEXT 380  
 NEXT 390  
 NEXT 400  
 NEXT 410  
 NEXT 420  
 NEXT 430  
 NEXT 440  
 NEXT 450  
 NEXT 460  
 NEXT 470  
 NEXT 480  
 NEXT 490  
 NEXT 500  
 NEXT 510  
 NEXT 520  
 NEXT 530  
 NEXT 540

UD(I) = FP\*U(NETA)  
 U(I) = FP\*UD  
 FP = Z(I)\*DTIL  
 IF (1.EQ.NETA-1) FP = 1.0  
 T(I) = T1 + (T(I)-TW) \* (TD-TWOLD) / (TDQ-TWO)  
 TE(I)=T(I)  
 RO(I) = RMO(I)  
 VO(I) = V(I)  
 V(I) = V1 + (V(I)-V01)\*(VD-V1)/(V00-V01)  
 HMO(I) = HM(I)  
 PO(I) = PI(I)  
 EO(I) = E(I)  
 YO(I) = YOND(I)  
 DO 5 J=1,NSP  
 CO(I,J) = C(J,I)  
 5 CONTINUE  
 RETURN  
 END

```

SUBROUTINE SHAPE
COMMON/JOHN/ UD(60),U(60),DU(60),VD(60),RD(60),DNDX(60),DTLA,DTILO,SHAP
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /YL/ETA(60),YOND(60),YO(60)
COMMON/MIKE/ META1,META2,PNF,VNF,RHOD,DP,TS,UD,VD,RW,RWO,EPS,PHI
COMMON/BOB/ XI,DXI,YL
COMMON /MAIN/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /GEORGE/ EPSN(30),XIN(30),RVN(30)
DATA DELT /0.0/
EPS=0.0
IF(XI .EQ. 0.0) GO TO 20
DTHT = 0.025
DTHT = .01
XID = XI - DXI
IF(DELTA .EQ.0.0) DELT = DELTA
DELTO = DELT
1 XIDA = XID + DTHT/2.
CALL INTRPL (XIDA,XIN,EPSN,IAB,EPS)
DELT = DELT + DTHT*TAN(EPS)
XID = XID + DTHT
IF(XID .LE. XI) GO TO 1
DTHT = XI - XID
XIDA = XID + DTHT/2.
CALL INTRPL (XIDA,XIN,EPSN,IAB,EPS)
DELT = DELT + DTHT*TAN(EPS)
CALL INTRPL (XI ,XIN,EPSN,IAB,EPS)
WRITE (6,100) DELT,DELTA,EPS
100 FORMAT(5X,'DELTA (GEO) =',F10.7,5X,'DELTA (FF) =',F10.7,5X,'EPS (RS) =',F10.7)
XAD) =',F10.7)
DO 5 I=1,META
DNDX(I) = RHO(I)/DTIL*( YOND(I)*(DELT - DELTO)/DELTA/DXI)
5 CONTINUE
20 CONTINUE
PHI = XI - EPS
SHAP 10
SHAP 20
SHAP 30
SHAP 40
SHAP 50
SHAP 60
SHAP 70
SHAP 80
SHAP 90
SHAP 100
SHAP 110
SHAP 120
SHAP 130
SHAP 140
SHAP 150
SHAP 160
SHAP 170
SHAP 180
SHAP 190
SHAP 200
SHAP 210
SHAP 220
SHAP 230
SHAP 240
SHAP 250
SHAP 260
SHAP 270
SHAP 280
SHAP 290
SHAP 300
SHAP 310
SHAP 320
SHAP 330
SHAP 340
SHAP 350
SHAP 360

```

SHAP 370  
SHAP 380  
SHAP 390  
SHAP 400  
SHAP 410

THT = XI  
RWC = RW  
RW = SIN(THT)  
RETURN  
END

U U U U

**THIS SUBROUTINE DETERMINES THE BOUNDARY CONDITIONS AT THE SHOCK USING THE RANKINE HUGONIOY RELATIONS**

U U U

# CALCULATION OF PRESSURE AND ENTHALPY FROM R-H RELATIONS

U

1000 CONTINUE  
RZB = (RZB + RZB1)/2.  
CONVERGED ON DENSITY RATIO (RZB)

U U

SHOC 370  
 SHOC 380  
 SHOC 390  
 SHOC 400  
 SHOC 410  
 SHOC 420  
 SHOC 430  
 SHOC 440  
 SHOC 450  
 SHOC 460  
 SHOC 470  
 SHOC 480  
 SHOC 490  
 SHOC 500  
 SHOC 510  
 SHOC 520  
 SHOC 530  
 SHOC 540  
 SHOC 550  
 SHOC 560

```

C
C
C
C
C
C
      NOW CALCULATE VELOCITIES FROM R-H RELATIONS

      CALCULATE SPECIES COMPOSITION AT SHOCK THRU CHEMEO

      WRITE(6,130) TITLE
      WRITE(6,135) XI
      WRITE(6,140)
      WRITE(6,150) TD,VD,PD,UD,PHI,EPS
      130 FORMAT('1.////.18A4.6(//)')
      135 FORMAT(20X,'BOUNDARY CONDITIONS AT XI =',F6.2,////)
      140 FORMAT('0'.48X,'BOW SHOCK CONDITIONS',//)
      150 FORMAT(20X,'SHOCK TEMPERATURE (OK) =',F10.0,8X,
        X 'NORMAL VELOCITY =',F10.4,/,
        X '20X,'SHOCK PRESSURE (ATM) =',F10.4,8X,
        X 'TANGENTIAL VELOCITY =',F10.4,/,
        X '20X,'SHOCK ANGLE, PHI (RAD) =',F10.4,8X,
        X 'ANGLE EPS (RAD) =',F10.4,////////)
      RETURN
      END

```





CALCULATION OF SURFACE PRESSURE FROM YMCMTM EQ.

```

P = PI(NETA)
IF (ISTART .LT. 3) GO TO 1
FP = DTIL
DO 15 I=1,NETA1
  J = NETA - I
  DELTA = ETA(J+1) - ETA(J)
  P = P - DTIL * (FP*UD)**2*DELTA/PNF
  IF (J .GT. 1) FP = Z(J-1)*DTIL
15 CONTINUE

1 RVC = 0.0
PI(1) = P
RV = RVCC/.4328
DO 5 I=1,3
  PVP = 10.**(D(I) + E(I)/TW*1.E4)
  EVP = A(I)*(1. - FP)*SORT(XMW(I)/RR/TW)
  BT = AM*P/XMW(I)
  CG = FVP/(RV + BT*EVP)*PVP
  RVC = RVC + RV*CG
5 CONTINUE

```

CHECK TO SEE IF HERTZ-KNUDSEN EQ. IS CONVERGED

```

ABD = (RVCC + RVC)/RVCC
ABSD = ABS(ABD)
IF (ABSD .LT. 0.001) GO TO 20
IF (TL .GT. 0.0) GO TO 10
TL = TW
RVCL = RVC
TW = TW + 10.*ABD/ABSD
GO TO 1

10 TWOLD = TW
TW = TW + (TW-TL)/(RVC-RVCL)*(RVCC-RVC)
TL = TWOLD

```

BCWA 37C  
BCWA 380  
RCWA 39C  
BCWA 400  
BCWA 410  
RCWA 42C  
BCWA 43C  
BCWA 44C  
RCWA 450  
BCWA 460  
RCWA 47C  
RCWA 480  
BCWA 49C  
HCWA 500  
BCWA 51C  
BCWA 520  
RCWA 530  
BCWA 540  
BCWA 550  
RCWA 560  
BCWA 570  
RCWA 58C  
BCWA 590  
RCWA 60C  
BCWA 61C  
RCWA 620  
BCWA 630  
BCWA 64C  
BCWA 650  
RCWA 66C  
BCWA 67C  
BCWA 68C  
RCWA 69C  
BCWA 700  
BCWA 710  
BCWA 720

```

RVCL = RVC
GO TO 1
20 TWOLD = TW
C
C      CALCULATE HEAT ABSORBED & RERADIATED BY ABLATOR
C
C      T3 = TW*TW*TW
C      SRV = SORT(RV)
C      HEAT ABSORBED BY CARBON SUBLIMING (BTU/FT2-SEC)
C
C      QSUHL = 4440.*RV
C
C      HEAT CONDUCTED THRU CHAR (BTU/FT2-SEC)
C
C      QCOND = A1*RV*T3 + A2*SRV*TW*TW + A3*SRV*T3 + A4*RV
C
C      ENERGY RERADIATED BY CHAR SURFACE WITH EMISSIVITY OF 0.66
C
C      QRR = 3.3E-12*T3*TW
C
C      TOTAL HEAT FLUX ACCOMMODATED BY ABLATOR (BTU/FT2-SEC)
C
C      QT = QCOND + QRR + QSUHL
C
C      RVW = RV/RINF/UINF/32.174
C      WRITE (6,100)
C      WRITE(6,110) TW,QCOND,P,QSUHL,RVW,QRR,RV,QT
100 FORMAT(1H,50X,'ABLATOR SURFACE',/)
110 FORMAT(20X,'SURFACE TEMPERATURE (OK) =',F10.0,8X,
X      'QCOND (BTU/FT2-SFC) =',F10.0,/,
X      '20X,'SURFACE PRESSURE (ATM) =',F10.4,8X,
X      'QSUHL (BTU/FT2-SEC) =',F10.0,/,
X      '20X,'BLOWING RATE =',F10.4,8X,
X      'QRR (BTU/FT2-SEC) =',F10.0,/,
X      '20X,'MASS LOSS (LB/FT2-SEC) =',F10.4,8X,
X      'QTOTAL (BTU/FT2-SEC) =',F10.0)

```

BCWA 730  
 BCWA 740  
 BCWA 750  
 BCWA 760  
 BCWA 770  
 BCWA 780  
 BCWA 790  
 BCWA 800  
 BCWA 810  
 BCWA 820  
 BCWA 830  
 BCWA 840  
 BCWA 850  
 BCWA 860  
 BCWA 870  
 BCWA 880  
 BCWA 890  
 BCWA 900  
 BCWA 910  
 BCWA 920  
 BCWA 930  
 BCWA 940  
 BCWA 950  
 BCWA 960  
 BCWA 970  
 BCWA 980  
 BCWA 990  
 BCWA 1000  
 BCWA 1010  
 BCWA 1020  
 BCWA 1030  
 BCWA 1040  
 BCWA 1050  
 BCWA 1060  
 BCWA 1070  
 BCWA 1080

BCWAI090  
BCWAI100  
BCWAI11C

IF(PVC .EQ. C.O) CALL EXIT  
RETURN  
END

```

INCLUDE FOHLAB(GLOBAL,CHECK,INPUT2,XMOMTM,PONCH,MOLMAS,DEDX)
INCLUDE FOHLAB(GAS,FG2,TRID,DATA,SOLVE,SLMTM,OUPUT,YMOMTM)
INCLUDE FOHLAB(NEXT,SHOCK,SPECIE,INTRPL,TEMPR,PROPR,ENERGY)
INCLUDE FOHLAB(SLINIT,FGH,MAIN,SHAPE,CHEMEQ)
INCLUDE FOHLAB(WALL)
INCLUDE FOHLAB(LRAD)
ENTRY MAIN
OVERLAY ONE
INSERT TRID,SOLVE,CHECK,MOLMAS,SLINIT,OUTPUT,TEMPR
OVERLAY TWO
INSERT FG2
OVERLAY THREE
INSERT SPECIF,SLMTM
OVERLAY THREE
INSERT ENERGY,INTRPL
OVERLAY TWO
INSERT XMOMTM,GLOBAL,PROPR,DEDX
OVERLAY ONE
INSERT INPUT,OUPUT,TRANS2,PONCH,NEXT
OVERLAY ONE
INSERT SHOCK
OVERLAY FIVE
INSERT CHEMEQ,THERMC,MATINV
OVERLAY FIVE
INSERT GAS
OVERLAY ONE
INSERT LRAD,ZP,SND,BUGPR,ZHV
INSERT DBUG
OVERLAY FOUR
INSERT TRANS
OVERLAY FOUR
INSERT TRANS1

```

C A S F ? 53.00 FT/SEC 2.5E-7 SLUGS/CU FT  
2.5E-7 9. 0

57000.  
ENTRY  
F FNC

```

SUBROUTINE SLINIT
C
C ROUTINE TO GUESS INITIAL VALUES AT THE STAGNATION LINE FOR
C COLD START
C
COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
COMMON/JOHN/ U(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILOSLIN
COMMON/MIKE/ NETA1,NETA2,PNF,WNF,RHOD,DP,TS,UD,VD,RW,RWD,EPS,PHI
COMMON /YL/ETA(60),YOND(60)
COMMON /RH/ DJD,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON/WALL/RW,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/FROPI/PI(60),RHD(60), T(60),AMW(60),C (20,60),CC(5,60)
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
COMMON /NAIM/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /FRSTRM/ U INF, RINF, UINF2, R , RE, LXI, ITM, NETA
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON/BOB/ XI,DXI,YL
COMMON/GUESS/ TG1(60),TG2(60)
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON /BILL/ NODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS,
X MI,NOE,NTIME,KOPT
COMMON/WT/SMW(20),AWT(5)
REAL MU,MUDZ
LOGICAL MCONV,ECONV,DCONV
C
IF(ISTART.GT.1) RETURN
TS = TD
CALL PROPT (NSP,NETA,NETA)
RDZ = RHO(NETA)
AKNF = 1.0*778.28*TS/(RDZ*UINF*UINF2*R)
CPNF = 1.0*778.28*TS*32.174*2./UINF2
HNF = 2.0*778.28 *32.174 /UINF2
MUDZ = MU(NETA)

```

```

SLIN 10
SLIN 20
SLIN 30
SLIN 40
SLIN 50
SLIN 60
SLIN 70
SLIN 80
SLIN 90
SLIN 100
SLIN 110
SLIN 120
SLIN 130
SLIN 140
SLIN 150
SLIN 160
SLIN 170
SLIN 180
SLIN 190
SLIN 200
SLIN 210
SLIN 220
SLIN 230
SLIN 240
SLIN 250
SLIN 260
SLIN 270
SLIN 280
SLIN 290
SLIN 300
SLIN 310
SLIN 320
SLIN 330
SLIN 340
SLIN 350
SLIN 360

```

U U U                      U U U                      U

U

U U U U

U

C

```

15 CONTINUE
   AMW(I) = 1./AM
   5 CONTINUE
   DO 35 I=2,NETA1
     CALL FGH (I,F1(I),G1(I),H1(I))
   35 CONTINUE
     KODE=0
     CALL PROPRT (NSP,1,NETA)
     DO 25 I=1,NDE
       IF(KODE .LT. 2) GO TO 40
       CALL PROPRT (NSP,1,NETA)
       CALL SLMTH
     40 CALL SPECIE (KODE)
       IF(KODE .GT. 4) RETURN
     25 CONTINUE
     DO 45 I=2,NETA1
       CALL FGH (I,F1(I),G1(I),H1(I))
     45 CONTINUE
       RETURN
     END

```

```

SLIN 730
SLIN 740
SLIN 750
SLIN 760
SLIN 770
SLIN 780
SLIN 790
SLIN 800
SLIN 810
SLIN 820
SLIN 830
SLIN 840
SLIN 850
SLIN 860
SLIN 870
SLIN 880
SLIN 890
SLIN 900
SLIN 910
SLIN 920

```



SUBROUTINE GAS (KODE )  
 \*\* THERMODYNAMIC AND TRANSPORT PROPERTIES OF AIR \*\*  
 \*\* REFERENCE NASA TR R-50 \*\*  
 THE FOLLOWING PROPERTIES ARE CALCULATED  
 TEMPERATURE AT WHICH PROPERTIES ARE WANTED (T) IN DEG R  
 PRESSURE AT WHICH PROPERTIES ARE WANTED (P) IN LB/IN\*\*2  
 RATIO OF SPECIFIC HEATS (GAMMA) IN DIMENSIONLESS  
 SPECIFIC HEAT AT CONSTANT PRESSURE (CP) IN BTU/LB-DEG R  
 ABSOLUTE VISCOSITY (V) IN LB/FT-SEC  
 PRANDTL NUMBER (PR) IN DIMENSIONLESS  
 THERMAL CONDUCTIVITY (K) IN BTU/FT-SEC-DEG R  
 PRESSURE (P) IN ATMOSPHERES  
 DENSITY (DEN) IN LB/FT\*\*3  
 ENTHALPY (H) IN BTU/LB  
 ENTROPY (S) IN BTU/LB-DEG R  
 COMPRESSIBILITY (Z) IN DIMENSIONLESS  
 SPEED OF SOUND (SOS) IN FT/SEC  
 SPECIFIC HEAT AT CONSTANT VOLUME (CV) IN BTU/LB-DEG R  
 ENTROPY (S) IN FT\*\*2/SEC\*\*2  
 VELOCITY (VEL) IN FT/SEC  
 PRESSURE (P) IN LBS/FT\*\*2  
 MACH NUMBER (M) IN DIMENSIONLESS  
 NOMENCLATURE 1=OXYGEN MOLECULES, 2=NITROGEN MOLECULES, 3=OXYGEN ATMOSGAS  
 4=NITROGEN ATMOS, 5=OXYGEN IONS, 6=NITROGEN IONS  
 7=ELECTRONS  
 COMMON /FRSTRM/ U INF, R INF, U INF2, R , RE, LXI, ITM, IEM, META  
 COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF  
 COMMON/PROPI/PI(60),RHO(60),TI(60),AMW(60),C (20,60),CC(5,60)  
 COMMON/PROP2/ MU(60),RM(60), AK(60)  
 COMMON/PROP3/CPS(20,60),MS(20,60),CPT(60),HM(60)

GAS 10  
 GAS 20  
 GAS 30  
 GAS 40  
 GAS 50  
 GAS 60  
 GAS 70  
 GAS 80  
 GAS 90  
 GAS 100  
 GAS 110  
 GAS 120  
 GAS 130  
 GAS 140  
 GAS 150  
 GAS 160  
 GAS 170  
 GAS 180  
 GAS 190  
 GAS 200  
 GAS 210  
 GAS 220  
 GAS 230  
 GAS 240  
 GAS 250  
 GAS 260  
 GAS 270  
 GAS 280  
 GAS 290  
 GAS 300  
 GAS 310  
 GAS 320  
 GAS 330  
 GAS 340  
 GAS 350  
 GAS 360

```

COMMON /RH/ DUD,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON/WALL/RVW,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON/MIKE/ NETA1,NETA2,PNF,WNF,RHOD,DP,TZ,UD,VD,RW,RWD,EPS,PHI
REAL MU,MUDZ
LOGICAL MCONV,GCONV,SCONV
DATA GASC /49721.7/

C
C
DO 2000 I=KODE,NETA
  T = TI(I) * TZ
  P = PI(I)

C THE FOLLOWING PART OF PROGRAM USES PRESSURE IN ATMOSPHERES
C AND TEMPERATURE IN DEG K
C
  ITER=0
  ** TEMPERATURE - ENTHALPY ITERATION **

CONTINUE
ITER=ITER+1
IF(T.LT.100.) T=100.
A1=11390./T
A2=10990./T
A3=2270./T
A4=3390./T
A5=228./T
A6=326./T
A7=22800./T
A8=48600./T
A9=27700./T
A10=41500./T
A11=38600./T
A12=58200./T
A13=70.6/T
A14=188.9/T

```

GAS 370  
 GAS 380  
 GAS 390  
 GAS 400  
 GAS 410  
 GAS 420  
 GAS 430  
 GAS 440  
 GAS 450  
 GAS 460  
 GAS 470  
 GAS 480  
 GAS 490  
 GAS 500  
 GAS 510  
 GAS 520  
 GAS 530  
 GAS 540  
 GAS 550  
 GAS 560  
 GAS 570  
 GAS 580  
 GAS 590  
 GAS 600  
 GAS 610  
 GAS 620  
 GAS 630  
 GAS 640  
 GAS 650  
 GAS 660  
 GAS 670  
 GAS 680  
 GAS 690  
 GAS 700  
 GAS 710  
 GAS 720

```

A15=22000./T
A16=47000./T
A17=67900./T
A18=2270./(4.*T)
A19=TANH(A18)
A20=3390./(4.*T)
A21=TANH(A20)
TT=1./T
TSQ=T**2
TSQRT=T**.5
A22=112.222/T
A23=T/59000.
A24=T/113200.
A25=T/75400.
AA1=EXP(-A1)
AA2=EXP(-A2)
AA3=EXP(A3)
AA4=EXP(A4)
AA5=EXP(-A5)
AA6=EXP(-A6)
AA7=EXP(-A7)
AA8=EXP(-A8)
AA9=EXP(-A9)
AA10=EXP(-A10)
AA11=EXP(-A11)
AA12=EXP(-A12)
AA13=EXP(-A13)
AA14=EXP(-A14)
AA15=EXP(-A15)
AA16=EXP(-A16)
AA17=EXP(-A17)

C CALCULATING ENERGIES PER COMPONENT OF GAS MIXTURE ABOVE
C REFERENCE ENERGIES.
  E1=2.5+((2.*AA1*AA2*AA2)/(3.+2.*AA1*AA2))+((A3/(AA3-1.))
  ET=2.5+ (A4/(AA4-1.))
  E3=1.5+((3.*AA5*AA5+AA6*AA6+5.*AA7*AA7+AA8*AA8)/(5.+3.*AA5+AA6+5.*AA7+GAS
GAS 730
GAS 740
GAS 750
GAS 760
GAS 770
GAS 780
GAS 790
GAS 800
GAS 810
GAS 820
GAS 830
GAS 840
GAS 850
GAS 860
GAS 870
GAS 880
GAS 890
GAS 900
GAS 910
GAS 920
GAS 930
GAS 940
GAS 950
GAS 960
GAS 970
GAS 980
GAS 990
GAS 1000
GAS 1010
GAS 1020
GAS 1030
GAS 1040
GAS 1050
GAS 1060
GAS 1070
GAS 1080

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1AAB))
EA=1.5+((10.* AA9*AA9+6.*AA10*AA10)/(4.+10.*AA9+6.*AA10))
E5=1.5+((10.*AA11*AA11+6.*AA12*AA12)/(4.+10.*AA11+6.*AA12))
E6=1.5+((3.*AA13*AA13+5.*AA14*AA14+5.*AA15*AA15+AA16*AA16+5.*AA17*AA17))
1/(1.+3.*AA13+5.*AA14+5.*AA15+AA16+5.*AA17))
E7=1.5

C TOTAL ENERGY PER COMPONENT OF GAS MIXTURE
EN1=E1
EN2=E2
EN3=E3+29500./T
EN4=E4+56600./T
EN5=E5+187500./T
EN6=E6+225400./T
EN7=E7

C LOGS OF PARTITION FUNCTIONS
TL1=ALOG(T)*3.5
TL2=ALOG(T)*2.5
E01=TL1+.11+ALOG((3.+2.*AA1+AA2)/(1.--(1.0/AA3)))
E02=TL1-.42-ALOG((1.--(1.0/AA4)))
E03=TL2+.5+ALOG((5.+3.*AA5+AA6+5.*AA7+AA8))
E04=TL2+.3+ALOG((4.+10.*AA9+6.*AA10))
E05=TL2+.5+ALOG((4.+10.*AA11+6.*AA12))
E06=TL2+.3+ALOG((1.+3.*AA13+5.*AA14+5.*AA15+AA16+5.*AA17))
E07=TL2-14.24

C EQUILIBRIUM CONSTANTS FOR CHEMICAL REACTIONS
EK1=-59000./T+2.*E03-E01
EK2=-113200./T+2.*E04-E02
EK3=-158000./T+E05+E07-E03
EK4=-168800./T+E06+E07-E04
CCC=-79.9
IF(EK1.LE.CCC) EK1=-79.9
IF(EK2.LE.CCC) EK2=-79.9
IF(EK3.LE.CCC) EK3=-79.9
IF(EK4.LE.CCC) EK4=-79.9
XK1=EXP(EK1)
XK2=EXP(EK2)

GAS 1090
GAS 1100
GAS 1110
GAS 1120
GAS 1130
GAS 1140
GAS 1150
GAS 1160
GAS 1170
GAS 1180
GAS 1190
GAS 1200
GAS 1210
GAS 1220
GAS 1230
GAS 1240
GAS 1250
GAS 1260
GAS 1270
GAS 1280
GAS 1290
GAS 1300
GAS 1310
GAS 1320
GAS 1330
GAS 1340
GAS 1350
GAS 1360
GAS 1370
GAS 1380
GAS 1390
GAS 1400
GAS 1410
GAS 1420
GAS 1430
GAS 1440

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XK3=EXP(EK3)
XK4=EXP(EK4)
XK3A=-2*XK3+.8*XK4
EE1=(-0.8+(-.64+.8*(1.+(4.*P)/XK1)))*0.5)/(2.*(1.+4.*P/XK1))
EE2=(-0.4+(-.16+3.84*(1.+(4.*P)/(XK2)))*0.5)/(2.*(1.+4.*P/XK2))
EE3=1./((1.+P/XK3)*.5)
IF(EE1.GT..19999)EE1=.19999
IF(EE2.GT..79999)EE2=.79999
IF(EE3.GT..99999)EE3=.99999
C COMPRESSIBILITY (Z) DIMENSIONLESS
Z=1.+EE1+EE2+2.*EE3
C COMPONENT MOL FRACTIONS IN AIR
X1=(-.2-EE1)/Z
X2=(-.8-EE2)/Z
X3=(2.*EE1-.4*EE3)/Z
X4=(2.*EE2-1.6*EE3)/Z
X5=.4*EE3/Z
X6=1.6*EE3/Z
X7=2.*EE3/Z
IF(X1.LE.0.) X1=1.E-20
IF(X2.LE.0.) X2=1.E-20
IF(X3.LE.0.) X3=1.E-20
IF(X4.LE.0.) X4=1.E-20
IF(X5.LE.0.) X5=1.E-20
IF(X6.LE.0.) X6=1.E-20
IF(X7.LE.0.) X7=1.E-20
C ENERGY PER MOL OF INITIALLY UNDISSOCIATED AIR-DIMENSIONLESS
ER=2*(X1*EN1+X2*EN2+X3*EN3+X4*EN4+X5*EN5+X6*EN6+X7*EN7)
C ENTHALPY PER INITIAL MOL OF AIR-DIMENSIONLESS
HR=ER+Z
C ENTHALPY PER INITIAL MOL OF AIR (H) IN BTU/LB
H=HR*.12348
IF(KOCE.LT.NETA) GO TO 1000
HRATO=.5*(H-HD)/H
AMR=ABS(HRATO)
IF(AMR.LE.0.0010) GO TO 999
GAS 1450
GAS 1460
GAS 1470
GAS 1480
GAS 1490
GAS 1500
GAS 1510
GAS 1520
GAS 1530
GAS 1540
GAS 1550
GAS 1560
GAS 1570
GAS 1580
GAS 1590
GAS 1600
GAS 1610
GAS 1620
GAS 1630
GAS 1640
GAS 1650
GAS 1660
GAS 1670
GAS 1680
GAS 1690
GAS 1700
GAS 1710
GAS 1720
GAS 1730
GAS 1740
GAS 1750
GAS 1760
GAS 1770
GAS 1780
GAS 1790
GAS 1800

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IF(ITER .GT.1) GO TO 203
TP=T
HP=HRATO
T = T*(1. - HRATO)
IF(ITER .LT. 15) GO TO 900
203 CONTINUE
TS=T*(1.0-HRATO)
IF(HRATO*HP .LT.0.0) TS=.5*(T+TP)
TP=T
T=TS
HP=HRATO
IF(ITER .LT. 15) GO TO 900
WRITE(6,200) T,H,HT
200 FORMAT(39HTEMPERATURE-ENTHALPY DID NOT CONVERGE /3E15.6)
C CALL OUTPUT(4)
STOP
999 CONTINUE
TD = T
C
1000 CONTINUE
C ENTROPY PER INITIAL MOL OF AIR-DIMENSIONLESS
D1=EQ1+E1+1.
D2=EQ2+ET+1.
D3=EQ3+E3+1.
D4=EQ4+E4+1.
D5=EQ5+E5+1.
D6=EQ6+E6+1.
D7=EQ7+E7+1.
C TOTAL ENTROPY
SR=Z*(X1*D1+X2*D2+X3*D3+X4*D4+X5*D5+X6*D6+X7*D7)-Z*(X1*ALOG(X1) +
1X2*ALOG(X2)+X3*ALOG(X3)+X4*ALOG(X4)+X5*ALOG(X5)+X6*ALOG(X6)+X7*
2ALOG(X7))-Z*ALOG(P)
C ENTROPY PER INITIAL MOL OF AIR (S) IN BTU/LB-DEG R
S =SR*0.0686
C SPECIFIC HEAT AT CONSTANT VOLUME-CV
FF1=3.+2.*AA1+AA2

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GAS 1810  
 GAS 1820  
 GAS 1830  
 GAS 1840  
 GAS 1850  
 GAS 1860  
 GAS 1870  
 GAS 1880  
 GAS 1890  
 GAS 1900  
 GAS 1910  
 GAS 1920  
 GAS 1930  
 GAS 1940  
 GAS 1950  
 GAS 1960  
 GAS 1970  
 GAS 1980  
 GAS 1990  
 GAS 2000  
 GAS 2010  
 GAS 2020  
 GAS 2030  
 GAS 2040  
 GAS 2050  
 GAS 2060  
 GAS 2070  
 GAS 2080  
 GAS 2090  
 GAS 2100  
 GAS 2110  
 GAS 2120  
 GAS 2130  
 GAS 2140  
 GAS 2150  
 GAS 2160

CV1=2.5+((2.\*AA1\*AI+AA2\*A2+AA2)/FF1)-(((2.\*AA1\*AI+AA2\*A2)/FF1))\*GAS 2170  
 12.)+((.25\*A3\*A3)/(((2.\*A19)/(1.-A19\*A19))\*2)) GAS 2180  
 CV2=2.5+((.25\*AA4\*AA4)/(((2.\*A21)/(1.-A21\*A21))\*2)) GAS 2190  
 CV3=1.5+((3.\*AA5\*AA5+AA6\*AA6+5.\*AA7\*AA7+AA8\*AA8)/(5.+3.\*AA6AS 2200  
 15+AA6+5.\*AA7+AA8))-((E3-1.5)\*2.) GAS 2210  
 CV4=1.5+((10.\*AA9\*AA9+9\*AA10\*AA10)/(4.+10.\*AA9+6.\*AA10)) GAS 2220  
 1-((E4-1.5)\*2.) GAS 2230  
 CV5=1.5+((10.\*AA11\*AA11+6.\*AA12\*AA12)/(4.+10.\*AA11+6.\*AA12))GAS 2240  
 1-((E5-1.5)\*2.) GAS 2250  
 CV6=1.5+((3.\*AA13\*AA13+5.\*AA14\*AA14+5.\*AA15\*AA15+AA16\*AA16AS 2260  
 1\*AA16+5.\*AA17\*AA17\*2)/(1.+3.\*AA13+5.\*AA14+5.\*AA15+AA16+5.\*AA17))-((GAS 2270  
 2E6-1.5)\*2.) GAS 2280  
 CV7=1.5 GAS 2290  
 C LOGARITHMIC DERIVATIVES GAS 2300  
 CK1=TT\*(S9000./T+2.\*E3-E1) GAS 2310  
 CK2=TT\*(113200./T+2.\*E4-E1) GAS 2320  
 CK3=TT\*(158000./T+E5+E7-E3) GAS 2330  
 CK4=TT\*(168800./T+E6+E7-E4) GAS 2340  
 CK34=.2\*CK3+.8\*CK4 GAS 2350  
 PK1= CK1+TT GAS 2360  
 PK2= CK2+TT GAS 2370  
 PK3= CK3+TT GAS 2380  
 PK4= CK4+TT GAS 2390  
 PK34=0.2\*PK3+0.8\*PK4 GAS 2400  
 C PARTIAL DERIVATIVES REQUIRED FOR CP GAS 2410  
 DE1P=(PK1\*EE1\*(1.+EE1)\*(1.2-EE1))/(.8\*(1.5-EE1)) GAS 2420  
 DE2P=(PK2\*EE2\*(1.2+EE2)\*(1.8-EE2))/(.4\*(4.8-EF2)) GAS 2430  
 DE3P=.5\*PK34\*EE3\*(1.-EE3\*2) GAS 2440  
 DX1P=-DE1P GAS 2450  
 DX2P=-DE2P GAS 2460  
 DX3P=2.\*DE1P - .4\*DE3P GAS 2470  
 DX4P=2.\*DE2P-1.6\*DE3P GAS 2480  
 DX5P=.4\*DE3P GAS 2490  
 DX6P=1.6\*DE3P GAS 2500  
 DX7P=2.\* DE3P GAS 2510  
 C EQUATION FOR SPECIFIC HEAT AT CONSTANT PRESSURE GAS 2520

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GAS 2530
GAS 2540
GAS 2550
GAS 2560
GAS 2570
GAS 2580
GAS 2590
GAS 2600
GAS 2610
GAS 2620
GAS 2630
GAS 2640
GAS 2650
GAS 2660
GAS 2670
GAS 2680
GAS 2690
GAS 2700
GAS 2710
GAS 2720
GAS 2730
GAS 2740
GAS 2750
GAS 2760
GAS 2770
GAS 2780
GAS 2790
GAS 2800
GAS 2810
GAS 2820
GAS 2830
GAS 2840
GAS 2850
GAS 2860
GAS 2870
GAS 2880

CPF=Z*(X1*(CV1+1.)+X2*(CV2+1.)+X3*(CV3+1.)+X4*(CV4+1.)+X5*(CV5+1.
1)+X6*(CV6+1.)+X7*(CV7+1.))
CPF = CPF +
T*(DZX1P*(EN1+1.)+DZX2P*(EN2+1.)+DZX3P*(EN3+1.)+DZX4P*(EN4+1.)+DZX5P*(EN5+1.)+DZX6P*(EN6+1.)+DZX7P*(EN7+1.)+DZX8P*(EN8+1.)+DZX9P*(EN9+1.)+DZX10P*(EN10+1.)+DZX11P*(EN11+1.)+DZX12P*(EN12+1.)+DZX13P*(EN13+1.)+DZX14P*(EN14+1.)+DZX15P*(EN15+1.)+DZX16P*(EN16+1.)+DZX17P*(EN17+1.)+DZX18P*(EN18+1.)+DZX19P*(EN19+1.)+DZX20P*(EN20+1.)+DZX21P*(EN21+1.)+DZX22P*(EN22+1.)+DZX23P*(EN23+1.)+DZX24P*(EN24+1.)+DZX25P*(EN25+1.)+DZX26P*(EN26+1.)+DZX27P*(EN27+1.)+DZX28P*(EN28+1.)+DZX29P*(EN29+1.)+DZX30P*(EN30+1.)+DZX31P*(EN31+1.)+DZX32P*(EN32+1.)+DZX33P*(EN33+1.)+DZX34P*(EN34+1.)+DZX35P*(EN35+1.)+DZX36P*(EN36+1.)+DZX37P*(EN37+1.)+DZX38P*(EN38+1.)+DZX39P*(EN39+1.)+DZX40P*(EN40+1.)+DZX41P*(EN41+1.)+DZX42P*(EN42+1.)+DZX43P*(EN43+1.)+DZX44P*(EN44+1.)+DZX45P*(EN45+1.)+DZX46P*(EN46+1.)+DZX47P*(EN47+1.)+DZX48P*(EN48+1.)+DZX49P*(EN49+1.)+DZX50P*(EN50+1.)+DZX51P*(EN51+1.)+DZX52P*(EN52+1.)+DZX53P*(EN53+1.)+DZX54P*(EN54+1.)+DZX55P*(EN55+1.)+DZX56P*(EN56+1.)+DZX57P*(EN57+1.)+DZX58P*(EN58+1.)+DZX59P*(EN59+1.)+DZX60P*(EN60+1.)+DZX61P*(EN61+1.)+DZX62P*(EN62+1.)+DZX63P*(EN63+1.)+DZX64P*(EN64+1.)+DZX65P*(EN65+1.)+DZX66P*(EN66+1.)+DZX67P*(EN67+1.)+DZX68P*(EN68+1.)+DZX69P*(EN69+1.)+DZX70P*(EN70+1.)+DZX71P*(EN71+1.)+DZX72P*(EN72+1.)+DZX73P*(EN73+1.)+DZX74P*(EN74+1.)+DZX75P*(EN75+1.)+DZX76P*(EN76+1.)+DZX77P*(EN77+1.)+DZX78P*(EN78+1.)+DZX79P*(EN79+1.)+DZX80P*(EN80+1.)+DZX81P*(EN81+1.)+DZX82P*(EN82+1.)+DZX83P*(EN83+1.)+DZX84P*(EN84+1.)+DZX85P*(EN85+1.)+DZX86P*(EN86+1.)+DZX87P*(EN87+1.)+DZX88P*(EN88+1.)+DZX89P*(EN89+1.)+DZX90P*(EN90+1.)+DZX91P*(EN91+1.)+DZX92P*(EN92+1.)+DZX93P*(EN93+1.)+DZX94P*(EN94+1.)+DZX95P*(EN95+1.)+DZX96P*(EN96+1.)+DZX97P*(EN97+1.)+DZX98P*(EN98+1.)+DZX99P*(EN99+1.)+DZX100P*(EN100+1.)+DZX101P*(EN101+1.)+DZX102P*(EN102+1.)+DZX103P*(EN103+1.)+DZX104P*(EN104+1.)+DZX105P*(EN105+1.)+DZX106P*(EN106+1.)+DZX107P*(EN107+1.)+DZX108P*(EN108+1.)+DZX109P*(EN109+1.)+DZX110P*(EN110+1.)+DZX111P*(EN111+1.)+DZX112P*(EN112+1.)+DZX113P*(EN113+1.)+DZX114P*(EN114+1.)+DZX115P*(EN115+1.)+DZX116P*(EN116+1.)+DZX117P*(EN117+1.)+DZX118P*(EN118+1.)+DZX119P*(EN119+1.)+DZX120P*(EN120+1.)+DZX121P*(EN121+1.)+DZX122P*(EN122+1.)+DZX123P*(EN123+1.)+DZX124P*(EN124+1.)+DZX125P*(EN125+1.)+DZX126P*(EN126+1.)+DZX127P*(EN127+1.)+DZX128P*(EN128+1.)+DZX129P*(EN129+1.)+DZX130P*(EN130+1.)+DZX131P*(EN131+1.)+DZX132P*(EN132+1.)+DZX133P*(EN133+1.)+DZX134P*(EN134+1.)+DZX135P*(EN135+1.)+DZX136P*(EN136+1.)+DZX137P*(EN137+1.)+DZX138P*(EN138+1.)+DZX139P*(EN139+1.)+DZX140P*(EN140+1.)+DZX141P*(EN141+1.)+DZX142P*(EN142+1.)+DZX143P*(EN143+1.)+DZX144P*(EN144+1.)+DZX145P*(EN145+1.)+DZX146P*(EN146+1.)+DZX147P*(EN147+1.)+DZX148P*(EN148+1.)+DZX149P*(EN149+1.)+DZX150P*(EN150+1.)+DZX151P*(EN151+1.)+DZX152P*(EN152+1.)+DZX153P*(EN153+1.)+DZX154P*(EN154+1.)+DZX155P*(EN155+1.)+DZX156P*(EN156+1.)+DZX157P*(EN157+1.)+DZX158P*(EN158+1.)+DZX159P*(EN159+1.)+DZX160P*(EN160+1.)+DZX161P*(EN161+1.)+DZX162P*(EN162+1.)+DZX163P*(EN163+1.)+DZX164P*(EN164+1.)+DZX165P*(EN165+1.)+DZX166P*(EN166+1.)+DZX167P*(EN167+1.)+DZX168P*(EN168+1.)+DZX169P*(EN169+1.)+DZX170P*(EN170+1.)+DZX171P*(EN171+1.)+DZX172P*(EN172+1.)+DZX173P*(EN173+1.)+DZX174P*(EN174+1.)+DZX175P*(EN175+1.)+DZX176P*(EN176+1.)+DZX177P*(EN177+1.)+DZX178P*(EN178+1.)+DZX179P*(EN179+1.)+DZX180P*(EN180+1.)+DZX181P*(EN181+1.)+DZX182P*(EN182+1.)+DZX183P*(EN183+1.)+DZX184P*(EN184+1.)+DZX185P*(EN185+1.)+DZX186P*(EN186+1.)+DZX187P*(EN187+1.)+DZX188P*(EN188+1.)+DZX189P*(EN189+1.)+DZX190P*(EN190+1.)+DZX191P*(EN191+1.)+DZX192P*(EN192+1.)+DZX193P*(EN193+1.)+DZX194P*(EN194+1.)+DZX195P*(EN195+1.)+DZX196P*(EN196+1.)+DZX197P*(EN197+1.)+DZX198P*(EN198+1.)+DZX199P*(EN199+1.)+DZX200P*(EN200+1.)+DZX201P*(EN201+1.)+DZX202P*(EN202+1.)+DZX203P*(EN203+1.)+DZX204P*(EN204+1.)+DZX205P*(EN205+1.)+DZX206P*(EN206+1.)+DZX207P*(EN207+1.)+DZX208P*(EN208+1.)+DZX209P*(EN209+1.)+DZX210P*(EN210+1.)+DZX211P*(EN211+1.)+DZX212P*(EN212+1.)+DZX213P*(EN213+1.)+DZX214P*(EN214+1.)+DZX215P*(EN215+1.)+DZX216P*(EN216+1.)+DZX217P*(EN217+1.)+DZX218P*(EN218+1.)+DZX219P*(EN219+1.)+DZX220P*(EN220+1.)+DZX221P*(EN221+1.)+DZX222P*(EN222+1.)+DZX223P*(EN223+1.)+DZX224P*(EN224+1.)+DZX225P*(EN225+1.)+DZX226P*(EN226+1.)+DZX227P*(EN227+1.)+DZX228P*(EN228+1.)+DZX229P*(EN229+1.)+DZX230P*(EN230+1.)+DZX231P*(EN231+1.)+DZX232P*(EN232+1.)+DZX233P*(EN233+1.)+DZX234P*(EN234+1.)+DZX235P*(EN235+1.)+DZX236P*(EN236+1.)+DZX237P*(EN237+1.)+DZX238P*(EN238+1.)+DZX239P*(EN239+1.)+DZX240P*(EN240+1.)+DZX241P*(EN241+1.)+DZX242P*(EN242+1.)+DZX243P*(EN243+1.)+DZX244P*(EN244+1.)+DZX245P*(EN245+1.)+DZX246P*(EN246+1.)+DZX247P*(EN247+1.)+DZX248P*(EN248+1.)+DZX249P*(EN249+1.)+DZX250P*(EN250+1.)+DZX251P*(EN251+1.)+DZX252P*(EN252+1.)+DZX253P*(EN253+1.)+DZX254P*(EN254+1.)+DZX255P*(EN255+1.)+DZX256P*(EN256+1.)+DZX257P*(EN257+1.)+DZX258P*(EN258+1.)+DZX259P*(EN259+1.)+DZX260P*(EN260+1.)+DZX261P*(EN261+1.)+DZX262P*(EN262+1.)+DZX263P*(EN263+1.)
```



GAS 2890  
GAS 2900  
GAS 2910  
GAS 2920  
GAS 2930  
GAS 2940  
GAS 2950  
GAS 2960  
GAS 2970  
GAS 2980  
GAS 2990  
GAS 3000  
GAS 3010  
GAS 3020  
GAS 3030  
GAS 3040  
GAS 3050  
GAS 3060  
GAS 3070  
GAS 3080  
GAS 3090  
GAS 3100  
GAS 3110  
GAS 3120  
GAS 3130  
GAS 3140  
GAS 3150  
GAS 3160  
GAS 3170  
GAS 3180  
GAS 3190  
GAS 3200  
GAS 3210  
GAS 3220  
GAS 3230  
GAS 3240

X30D=2.\*EE1/F1  
X2ND=(.8-EE2)/F2  
X3ND=.4/F2  
X4ND=2.\*EE2/F2  
X4I=(1.-EE3)/F3  
X6I=EE3/F3  
  
C MEAN FREE PATH RATIOS  
SS1=S24/S2  
SS2=S4/S2  
SS3=S7/S2  
SS4=S47/S2  
FP10D=X10D+X20D+.9660918 +X30D\*SS1\*.8164966  
FP20D=X10D\*1.032796+X20D+X30D\*SS1\*.8528029  
FP30D=X10D\*1.154701\*SS1+X2ND\*SS1\*1.128152+X30D\*SS2  
FP2ND=X2ND+X4ND\*SS1\*.8164966+X3ND\*SS1\*.8528029  
FP3ND=X2ND\*SS1\*1.128152+X4ND\*SS2\*.9660918+X3ND\*SS2  
FP4ND=X2ND\*SS1\*1.154701+X4ND\*SS2+X3ND\*SS2\*1.032796  
FP4I=X4I\*SS2+X6I\*SS2  
FP6I=X4I\*SS2+X6I\*SS3  
FP7I=X4I\*SS4\*1.414186+X6I\*SS3\*1.414186+X6I\*SS3  
  
C VISCOSTITES OF THE COMPONENTS FOR THE DIFFERENT REACTIONS  
V10D=1.054093\*X10D\*1./FP10D  
V20D=.986013\*X20D\*1./FP20D  
V30D=.745356\*X30D\*1./FP30D  
V2ND=.986013\*X2ND\*1./FP2ND  
V3ND=.745356\*X3ND\*1./FP3ND  
V4ND=.6972167\*X4ND\*1./FP4ND  
V4I=.6972167\*X4I\*1./FP4I  
V6I=.6972167\*X6I\*1./FP6I  
V7I=.4367848\*1.0E-2\*X6I\*1./FP7I  
VR0D=V10D+V20D+V30D  
VRND=V2ND+V3ND+V4ND  
VRI=V4I+V6I+V7I  
F4=EE2/(.2-EE1+EE2)  
F5=2.\*EE3/(.8-EE2+2.\*EE3)  
VR=VR0D+(F4\*(VRND-VR0D))+(F5\*(VRI-VRND))

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C TOTAL VISCOSITY (V) IN LB/FT-SEC
  V=VR*.9841838*1.0E-6*TSORT/(1.+A22)
C CONDUCTIVITY DUE TO MOLECULAR COLLISIONS FOR DIFFERENT REACTIONS
  G1=.2105263*CV1+.4736842
  G2=.2105263*CV2+.4736842
  G3=.2105263*CV3+.4736842
  G4=.2105363*CV4+.4736842
  G5=.2105363*CV6+.4736842
  G6=.2105363*CV7+.4736842
  XKNOD=(V10D*.9*G1)+(V20D*1.028571*G2)+(V30D*1.8*G3)
  XKNND=(V2ND*1.028571*G2)+(V3ND*1.8*G3)+(V4ND*2.057143*G4)
  XKNI=(V4I*2.057143*G4)+(V6I*2.057143*G5)+(V7I*52416.0*G6)
  XKN=XKNOD+(F4*(XKNND-XKNOD))+(F5*(XKNI-XKNND))
C CONDUCTIVITY DUE TO CHEMICAL REACTIONS FOR THE DIFFERENT REACTIONS
  XKROD=(.178637*(T*PK1)**2)/((SP24/(1.732051*S2))*((X30D+2.*X10D)
1*2)/(X30D*X10D))+(4.*X20D/X30D))+(X20D/(1.414214*X10D)))
  XKRND=(.178637*(T*PK2)**2)/((SP24/(1.732051*S2))*((X4ND+2.*X2ND)GAS 3410
1*2)/(X4ND*X2ND)))+(X3ND/X2ND)))+(SP4*2.*X3ND/(S2*X4ND)))
  XKRI=(.178637*(T*PK34)**2)/((.5*SP4/S2)+(.4347826*1.0E-2*S47/S2))GAS 3430
  i*((X4I+X6I)**2)/(X4I*X6I)
  XKOD=XKNOD+XKROD
  XKNND=XKNND+XKROD
  XKI=XKNI+XKRI
  XKR=XKOD+(F4*(XKND-XKOD))+(F5*(XKI-XKND))
C TOTAL THERMAL CONDUCTIVITY (KK) IN BTU/FT-SEC-DEG R
  XK=XKR*(.3206522*1.0E-6*TSORT)/(1.+A22)
C PRANDTL NUMBER (PR) DIMENSIONLESS
  PRN = .2105263 * CPR * VR / XKR
  IF(1.EQ. 1) PRN = PRN
C FORM REQUIRED BY CALL STATEMENT
C
C ** RMO UNITS SLUGS/FT**3
C ** MU UNITS LBM/FT-SEC
C ** RM UNITS LBF**2 SEC**3/FT**6
C
  MU (1) = V

```

GAS 3250  
 GAS 3260  
 GAS 3270  
 GAS 3280  
 GAS 3290  
 GAS 3300  
 GAS 3310  
 GAS 3320  
 GAS 3330  
 GAS 3340  
 GAS 3350  
 GAS 3360  
 GAS 3370  
 GAS 3380  
 GAS 3390  
 GAS 3400  
 GAS 3410  
 GAS 3420  
 GAS 3430  
 GAS 3440  
 GAS 3450  
 GAS 3460  
 GAS 3470  
 GAS 3480  
 GAS 3490  
 GAS 3500  
 GAS 3510  
 GAS 3520  
 GAS 3530  
 GAS 3540  
 GAS 3550  
 GAS 3560  
 GAS 3570  
 GAS 3580  
 GAS 3590  
 GAS 3600

```

RHO(I)=DEN/32.174
RM(I)=RHO(I)*MU(I)/32.174
AK(I)=XK
CPT(I)=CPF
C *** CALCULATE THE MEAN MOLECULAR WT. ***
REAL = 25050.*S *Z / SR
AMW(I)= GASC / REAL
C MASS FRACTIONS
C(1,I) = X1 *32.00/AMW(I)
C(2,I) = X2 *28.00/AMW(I)
C(3,I) = X3 *16.00/AMW(I)
C(4,I) = X4 *14.00/AMW(I)
C(5,I) = X5 *16.00/AMW(I)
C(6,I) = X6 *14.00/AMW(I)
C(7,I) = X7 /(1820.*AMW(I))
C SPECIES ENTHALPY PER INITIAL MOLE OF AIR IN BTU/LB OF I
HS(1,I) = (Z*X1*EN1/C(1,I) +Z)*T*.12348
HS(2,I) = (Z*X2*EN2/C(2,I) +Z)*T*.12348
HS(3,I) = (Z*X3*EN3/C(3,I) +Z)*T*.12348
HS(4,I) = (Z*X4*EN4/C(4,I) +Z)*T*.12348
HS(5,I) = (Z*X5*EN5/C(5,I) +Z)*T*.12348
HS(6,I) = (Z*X6*EN6/C(6,I) +Z)*T*.12348
HS(7,I) = (Z*X7*EN7/C(7,I) +Z)*T*.12348
2000 CONTINUE
C
C
DO 40 I=CODE,NETA
C
C ** NONDIMENSIONALIZE RHO AND MU **
C
RHO(I) = RHO(I)/RDZ
MU(I) = MU(I)/MUOZ
RM(I) = RM(I)/RMOZ
AK(I) = AK(I)*AKNF
CPT(I) = CPT(I)*CPNF
C NONDIMENSIONAL SPECIES ENTHALPY

```

GAS 3610  
 GAS 3620  
 GAS 3630  
 GAS 3640  
 GAS 3650  
 GAS 3660  
 GAS 3670  
 GAS 3680  
 GAS 3690  
 GAS 3700  
 GAS 3710  
 GAS 3720  
 GAS 3730  
 GAS 3740  
 GAS 3750  
 GAS 3760  
 GAS 3770  
 GAS 3780  
 GAS 3790  
 GAS 3800  
 GAS 3810  
 GAS 3820  
 GAS 3830  
 GAS 3840  
 GAS 3850  
 GAS 3860  
 GAS 3870  
 GAS 3880  
 GAS 3890  
 GAS 3900  
 GAS 3910  
 GAS 3920  
 GAS 3930  
 GAS 3940  
 GAS 3950  
 GAS 3960

GAS 3970  
GAS 3980  
GAS 3990  
GAS 4000  
GAS 4010  
GAS 4020  
GAS 4030  
GAS 4040  
GAS 4050  
GAS 4060  
GAS 4070  
GAS 4080  
GAS 4090

HS(1,I) = HS(1,I)\*HNF  
HS(2,I) = HS(2,I)\*HNF  
HS(3,I) = HS(3,I)\*HNF  
HS(4,I) = HS(4,I)\*HNF  
HS(5,I) = HS(5,I)\*HNF  
HS(6,I) = HS(6,I)\*HNF  
HS(7,I) = HS(7,I)\*HNF

C  
C  
40 CONTINUE  
100 FORMAT(1X,9E14.6)  
RETURN  
END

## SUBROUTINE BCWALL

THIS SUBROUTINE CALCULATES BLOWING RATE AT THE ABLATOR SURFACE  
FOR A GIVEN SURFACE TEMPERATURE AND PRESSURE

COMMON /BILL/ NODE(20), ISTART, IDGO, IDGF, IDGH, IDGT, IDGS,

X MI, NOE, NTIME, KOPT

COMMON/MIKE/ NETA1, NETA2, PNF, WNF, RHOD, DP, TS, UD, VD, RW, RWO, EPS, PHI

COMMON /YL/ ETA(60), YOND(60)

COMMON/WALL/RVW, PRW, TWOLD, FLUX(20), CWALL(20), ECWALL(5)

COMMON /VEL/ F(60), FC(60), Z(60), V(60)

COMMON/PROPI/PI(60), RHD(60), T(60), AMW(60), C(20,60), CC(5,60)

COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA

COMMON /NON/RDZ, MUDZ, RMDZ, AKNF, HNF, CPNF

COMMON /DEL/ DELTA, DTIL, DTILS

COMMON /MAIM/KEEP, MAXE, MAXM, MAXD, IDEBUG, MCONV, ECONV, DCONV, LT, IAB

COMMON/BOB/ XI, DXI, YL

COMMON /GEORGE/ EPSN(30), XIN(30), RVN(30)

DIMENSION A(3), D(3), E(3), XMW(3)

DATA A /0.37, 0.34, 0.08/

DATA XMW/12., 24., 36./

DATA C /8.21, 9.64, 10.25/

DATA E /-3.75, -4.27, -4./

DATA EP, RR, AM/0.8, 1.215E-4, 13.3/

DATA A1, A2, A3, A4 /1.674065E-8, -.333005E-3, 10.99963E-8, 2731.659/

TL = 0.0

TW = TWOLD

NETA1 = NETA - 1

INITIAL GUESS FOR BLOWING RATE OF CARBON SOLID

IF (XI.EQ. 0.0) PRW = RVW

CALL INTRPL (XI, XIN, RVN, IAB, RVW)

RVW = RVW \* PRW

RVCC = RVW \* RINF \* UINF \* 32.174 \* 0.4328

BCWA 10  
BCWA 20  
BCWA 30  
BCWA 40  
BCWA 50  
BCWA 60  
BCWA 70  
BCWA 80  
BCWA 90  
BCWA  
BCW  
B/ 120  
JA 130  
BCWA 140  
BCWA 150  
BCWA 160  
BCWA 170  
BCWA 180  
BCWA 190  
BCWA 200  
BCWA 210  
BCWA 220  
BCWA 230  
BCWA 240  
BCWA 250  
BCWA 260  
BCWA 270  
BCWA 280  
BCWA 290  
BCWA 300  
BCWA 310  
BCWA 320  
BCWA 330  
BCWA 340  
BCWA 350  
BCWA 360

```

C
C
C      CALCULATION OF SURFACE PRESSURE FROM YHOMTM EQ.
C
      P = PI(NETA)
      IF (ISTART .LT. 3) GO TO 1
      FP = DTIL
      DO 15 I=1,NETA1
      J = NETA - I
      DETA = ETA(J+1) - ETA(J)
      P = P - DTIL * (FP*UD)**2*DETA/PNF
      IF (J .GT. 1) FP = Z(J-1)*DTIL
15 CONTINUE
C
      1 RVC = 0.0
      PI(1) = P
      RV = RVCC/.4328
      DO 5 I=1,3
      PVP = 10.**(D(I) + E(I)/TW*1.E4)
      EVP = A(I)*(1. - EP)*SQRT(XMW(I)/RR/TW)
      BT = AM*P/XMW(I)
      CG = EVP/(RV + BT*EVP)*PVP
      RVC = RVC + RV*CG
5 CONTINUE
C
C      CHECK TO SEE IF HERTZ-KNUDSEN EQ. IS CONVERGED
C
      ABD = (RVCC - RVC)/RVCC
      ABSD = ABS(ABD)
      IF(ABSD .LT. 0.001) GO TO 20
      IF ( TL .GT. 0.0 ) GO TO 10
      TL = TW
      RVCL = RVC
      TW = TW + 10.*ABD/ABSD
      GO TO 1
10 TWOLD = TW
      TW = TW + (TW-TL)/(RVC-RVCL)*(RVCC-RVC)
      TL = TWOLD

```

BCWA 370  
BCWA 380  
BCWA 390  
BCWA 400  
BCWA 410  
BCWA 420  
BCWA 430  
BCWA 440  
BCWA 450  
BCWA 460  
BCWA 470  
BCWA 480  
BCWA 490  
BCWA 500  
BCWA 510  
BCWA 520  
BCWA 530  
BCWA 540  
BCWA 550  
BCWA 560  
BCWA 570  
BCWA 580  
BCWA 590  
BCWA 600  
BCWA 610  
BCWA 620  
BCWA 630  
BCWA 640  
BCWA 650  
BCWA 660  
BCWA 670  
BCWA 680  
BCWA 690  
BCWA 700  
BCWA 710  
BCWA 720

```

RVCL = RVC
GO TO 1
20 TWOLD = TW
C
C      CALCULATE HEAT ABSORBED & RERADIATED BY ABLATOR
C
C      T3 = TW*TW*TW
C      SRV = SORT(RV)
C      HEAT ABSORBED BY CARBON SUBLIMING (BTU/FT2-SEC)
C
C      OSUBL = 4440.*RV
C
C      HEAT CONDUCTED THRU CHAR (BTU/FT2-SEC)
C
C      QCOND = A1*RV*T3 + A2*SRV*TW*TW + A3*SRV*T3 + A4*RV
C      ENERGY RERADIATED BY CHAR SURFACE WITH EMISSIVITY OF 0.66
C
C      ORR = 3.3E-12*T3*TW
C
C      TOTAL HEAT FLUX ACCOMMODATED BY ABLATOR (BTU/FT2-SEC)
C
C      QT = QCOND + ORR + OSUBL
C
C      RVW = RV/RINF/UINF/32.174
C      WRITE (6,100)
C      WRITE(6,110) TW,QCOND,P,OSUBL,RVW,ORR,RV,QT
100 FORMAT(1H,50X,'ABLATOR SURFACE',/)
110 FORMAT(20X,'SURFACE TEMPERATURE (OK) =',F10.0,8X,
X   'QCOND (BTU/FT2-SEC) =',F10.0,/,
X   '20X,'SURFACE PRESSURE (ATM) =',F10.4,8X,
X   'OSUBL (BTU/FT2-SEC) =',F10.0,/,
X   '20X,'BLOWING RATE =',F10.4,8X,
X   'ORR (BTU/FT2-SEC) =',F10.0,/,
X   '20X,'MASS LOSS (LB/FT2-SEC) =',F10.4,8X,
X   'QTOTAL (BTU/FT2-SEC) =',F10.0)
BCWA 730
BCWA 740
BCWA 750
BCWA 760
BCWA 770
BCWA 780
BCWA 790
BCWA 800
BCWA 810
BCWA 820
BCWA 830
BCWA 840
BCWA 850
BCWA 860
BCWA 870
BCWA 880
BCWA 890
BCWA 900
BCWA 910
BCWA 920
BCWA 930
BCWA 940
BCWA 950
BCWA 960
BCWA 970
BCWA 980
BCWA 990
BCWA 1000
BCWA 1010
BCWA 1020
BCWA 1030
BCWA 1040
BCWA 1050
BCWA 1060
BCWA 1070
BCWA 1080

```

BCVA1090  
BCVA1100  
BCVA1110

IF(RVC .EQ. 0.0) CALL EXIT  
RETURN  
END



```

SUBROUTINE SLMTM
C -----THIS SUBROUTINE SOLVES THE MOMENTUM EQUATION AS A
C SECOND ORDER EQUATION AND A FIRST ORDER EQUATION -----
COMMON /CONV/ FPRCT,TPRCT,DDAMP,TDAMP,PDIL
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA
COMMON /MAIN/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
COMMON /PROP1/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /PROP2/ MU(60),RM(60), AK(60)
COMMON /PROP3/CP5(20,60),HS(20,60),CP (60),HM(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /RH/ DUD,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON /VECTOR/ CA(60),CB(60),CC(60),B(60)
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON /WALL/RVW,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /YL/ETA(60),YOND(60)
LOGICAL MCONV,ECONV,DCONV
REAL*8 CA,CB,CC,B

C ----- INITIALIZED QUANTITIES -----
MCONV = .FALSE.
DTILS = DTIL
ITM = 1
N = NETA -2
L = NETA-1
AA3 = RZB*(1.-RZB)*DPHI**2/DUD
AAA3 = AA3
DTILS2 = DTILS
IF(IEM.GT.3) DTIL=.5*(DTIL+DTILS2)

C
C
C-----Z...+A1*Z...+A2*Z=A3
C COMPUTE A1,A2,A3

```

SLMT 10  
SLMT 20  
SLMT 30  
SLMT 40  
SLMT 50  
SLMT 60  
SLMT 70  
SLMT 80  
SLMT 90  
SLMT 100  
SLMT 110  
SLMT 120  
SLMT 130  
SLMT 140  
SLMT 150  
SLMT 160  
SLMT 170  
SLMT 180  
SLMT 190  
SLMT 200  
SLMT 210  
SLMT 220  
SLMT 230  
SLMT 240  
SLMT 250  
SLMT 260  
SLMT 270  
SLMT 280  
SLMT 290  
SLMT 300  
SLMT 310  
SLMT 320  
SLMT 330  
SLMT 340  
SLMT 350  
SLMT 360

```

14 CONTINUE
C----- BOUNDARY CONDITIONS -----
C
  RED = RE*DTIL
  RED2 = 2.*RED*DTIL*DUD
  DTIL2 = DTIL*DTIL
  FD = RZB/(2.*DUD*DTIL)
  FD = RZB/(2.*DUD*DTIL)*RHO(NETA)
  FW = -RVW*FD
  F(1) = FW
  B(L) = 1./DTIL
  ITER = 1
15 CONTINUE
  SUM = 0.0
  ZP = 1.0
  DO 35 I=1,N
    J = NETA-I
    ZA = (ZP + Z(J-1))/2.
    DETA = ETA(J+1) - ETA(J)
    SUM = SUM + DTIL2*Z(J-1)**2*DETA/      DUD*DTIL
    CC(J-1) = SUM
    ZP = Z(J-1)
35 CONTINUE
    II = 1
    DO 20 I=1,N
      AA3 = AA3 + CC(I)
      DET=ETA(I+1)-ETA(I)
      DETN=ETA(I+2)-ETA(I+1)
      D1 = DETN*(DETN+DET)
      D2 = DETN*DET
      D3 = DET*(DETN+DET)
      RMP = DET*RM(I+2)/D1 + (DETN-DET)*RM(I+1)/D2 -DETN*RM(I)/D3
      A1 = (RED2*F(I+1) +RMP)/RM(I+1)
      A2 = -RED2*DTIL*Z(I)/RM(I+1)
      A3 = -2.*RED*( AA3/(RHO(I+1)*RM(I+1))
        +DTIL2 *DUD*Z(I)**2/(2.*RM(I+1)) )
1
  
```

SLMT 370  
SLMT 380  
SLMT 390  
SLMT 400  
SLMT 410  
SLMT 420  
SLMT 430  
SLMT 440  
SLMT 450  
SLMT 460  
SLMT 470  
SLMT 480  
SLMT 490  
SLMT 500  
SLMT 510  
SLMT 520  
SLMT 530  
SLMT 540  
SLMT 550  
SLMT 560  
SLMT 570  
SLMT 580  
SLMT 590  
SLMT 600  
SLMT 610  
SLMT 620  
SLMT 630  
SLMT 640  
SLMT 650  
SLMT 660  
SLMT 670  
SLMT 680  
SLMT 690  
SLMT 700  
SLMT 710  
SLMT 720

```

C-----CA*Z(N-1)+CB*Z(N)+CC*Z(N+1)=B
C
C    COMPUTE CA,CB,CC
C    CA(11)=(2.-A1*DETN)/D3
C    CB(1)=A1*(DETN-DET)/D2-2./D2+A2
C    CC(1)=(2.+A1*DET)/D1
C    B(1)=A3
C    11 = 1
C    20 CONTINUE
C    B(N)=B(N)-CC(N)/DTIL
C
C    CALL TRID (N)
C
C-----INTEGRATE FIRST ORDER EQUATION-----
C    FC(1)=FW
C    SUM=FW+ (A(1))+FW)*((ETA(2)-ETA(1))*DTIL/2.
C    FC(2)=SUM
C    DO 30 K=3,NETA
C    SUM=SUM+DTIL*(B(K-1)+B(K-2))*((ETA(K) -ETA(K-1))/2.
C    30 FC(K) = SUM
C
C-----CHECK FOR CONVERGENCE
C
C    DO 40 K=2,NETA
C    PRCT=ABS((FC(K)-F(K))/F(K))
C    IF (PRCT.GT.FPRCT) GO TO 50
C    40 CONTINUE
C    GO TO 90
C    50 CONTINUE
C    ITER=ITER+1
C    DO 60 K=1,NETA
C    F(K)=FC(K)
C    DO 65 I=1,N
C    65 Z(I)=B(I)
C    IF(ITER.GE.MAXM ) GO TO 90
C    GO TO 15

```

SLMT 730  
 SLMT 740  
 SLMT 750  
 SLMT 760  
 SLMT 770  
 SLMT 780  
 SLMT 790  
 SLMT 800  
 SLMT 810  
 SLMT 820  
 SLMT 830  
 SLMT 840  
 SLMT 850  
 SLMT 860  
 SLMT 870  
 SLMT 880  
 SLMT 890  
 SLMT 900  
 SLMT 910  
 SLMT 920  
 SLMT 930  
 SLMT 940  
 SLMT 950  
 SLMT 960  
 SLMT 970  
 SLMT 980  
 SLMT 990  
 SLMT 1000  
 SLMT 1010  
 SLMT 1020  
 SLMT 1030  
 SLMT 1040  
 SLMT 1050  
 SLMT 1060  
 SLMT 1070  
 SLMT 1080

```

          90 CONTINUE
C
C----- COMPUTE NEW DTIL -----
C
      DTILC = (FD-FW)*DTIL/(F(NETA)-FW)
      PRCT = ABS((DTIL-DTILC)/DTIL)
      IF(ITM.GT.MAXM) GO TO 160
      ITM = ITM + 1
      IF(PRCT.LE.PDTIL) GO TO 150
      DTIL = DTIL + DDAMP*(DTILC-DTIL)
      GO TO 14
150 CONTINUE
      DTIL = DTIL + DDAMP*(DTILC -DTIL)
      MCONV = .TRUE.
C
C CHECK MOMENTUM-ENERGY CONVERGENCE
      PRCT = ABS((DTIL-DTILS)/DTILS)
      IF(PRCT.LE.PDTIL) DCONV = .TRUE.
160 CONTINUE
C
      DO 170 K=1,NETA
170 V(K) = -FC(K)*DTIL*2./RHD(K)
C DEBUG OUTPUT
      IF(IDEBUG.EQ. 0) RETURN
      WRITE(6,102) ITER,ITM
102 FORMAT(10X,2I3/)
      WRITE(6,100) DTIL,DTILC
      WRITE(6,101)
101 FORMAT(6X,'ETA',.12X,'F',.12X,'FC',.12X,'RHD',.12X,'RM',.12X,'VS',.12X,
1 'V')
      DO 120 K=1,NETA
120 K=1,NETA
      VS=-FC(K)*DTIL*UINF*2./RHD(K)
      WRITE(6,100) ETA(K),F(K),FC(K), RHD(K),RM(K),VS ,V(K)
100 FORMAT(1X,9E14.6)
120 CONTINUE
      WRITE(6,103)

```

SLMT1090  
 SLMT1100  
 SLMT1110  
 SLMT1120  
 SLMT1130  
 SLMT1140  
 SLMT1150  
 SLMT1160  
 SLMT1170  
 SLMT1180  
 SLMT1190  
 SLMT1200  
 SLMT1210  
 SLMT1220  
 SLMT1230  
 SLMT1240  
 SLMT1250  
 SLMT1260  
 SLMT1270  
 SLMT1280  
 SLMT1290  
 SLMT1300  
 SLMT1310  
 SLMT1320  
 SLMT1330  
 SLMT1340  
 SLMT1350  
 SLMT1360  
 SLMT1370  
 SLMT1380  
 SLMT1390  
 SLMT1400  
 SLMT1410  
 SLMT1420  
 SLMT1430  
 SLMT1440

SLMT1450  
SLMT1460  
SLMT1470  
SLMT1480  
SLMT1490  
SLMT1500  
SLMT1510

```
103 FORMAT(6X,'ETA',13X,'Z',13X,'B',12X,'2HF')
DO 121 I=1,N
  U=B(I)*DTIL
  WRITE(6,100) ETA(I+1),Z(I),B(I),U
121 CONTINUE
  RETURN
  END
```

```

SUBROUTINE DEDX
      THIS SUBROUTINE CALCULATES THE SHOCK SINDOFF DISTANCE, DELTA
      AND THE PARTIAL OF ETA WITH RESPECT TO X, DNDX

      COMMON/JOHNN/ UO(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILD
      COMMON/BOB/ XI,DXI,YL
      COMMON/FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA
      COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
      COMMON /DEL/ DELTA,DTIL,DTILS
      COMMON /YL/ETA(60),YOND(60),YO(60)

      YA = 0.0
      DO 15 J=2,NETA
        I=J-1
        DELTA = ETA(J) - ETA(J-1)
        YA = YA + DTIL*DELTA*(1./RHO(I) + 1./RHO(J))/2.
        YOND(J) = YA
      15 CONTINUE
      DELTA = YOND(NETA)
      RETURN
      END

```

```

SUBROUTINE SOLVE(BC1,SC2,W0)
C
C
C
C
C
      THIS SUBROUTINE SOLVES THE PARTIAL DIFFERENTIAL EQUATION
      
$$W'' + A1*W' + A2*W + A3 + A4*W*(X) = 0$$

      COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA
      COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
      COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
      DIMENSION WD(1)
      COMMON/BOB/ XI,DXI,YL
      REAL*8 SUB,DIAG,SUP,B
      NETA2=NETA-2
      DO5 I=1,NETA2
      J=I+1
      A1= SUB(I)
      A2= DIAG(I)
      A3= SUP(I)
      A4= B(I)
      CA = (H2(J)+A1*H1(J))*YL
      DIAG(I)= (G2(J)+A1*G1(J)+A2)*YL
      SUP(I) = (F2(J)+A1*F1(J))*YL
      B(I) = A4*W0(J)/DXI-(1-YL)/YL*(CA*W0(I)+ SUP(I)*W0(J+1)
      X
      DIAG(I)=DIAG(I)+A4/DXI
      IF(I.EQ.1) GO TO 4
      SUB(I-1)=CA
      GO TO 5
4 A = CA
5 CONTINUE
      B(1) = B(1) - A*BC1
      B(NETA2)=B(1,TA2) - SUP(NETA2)*BC2
      CALL TRIO (NETA2)
      RETURN
      END
SOLV 10
SOLV 20
SOLV 30
SOLV 40
SOLV 50
SOLV 60
SOLV 70
SOLV 80
SOLV 90
SOLV 100
SOLV 110
SOLV 120
SOLV 130
SOLV 140
SOLV 150
SOLV 160
SOLV 170
SOLV 180
SOLV 190
SOLV 200
SOLV 210
SOLV 220
SOLV 230
SOLV 240
SOLV 250
SOLV 260
SOLV 270
SOLV 280
SOLV 290
SOLV 300
SOLV 310
SOLV 320
SOLV 330
SOLV 340
SOLV 350

```

```

C
C
C
C
SUBROUTINE YNOMTM
      THIS ROUTINE SOLVES THE THIN SHOCK LAYER Y-MOMENTUM EQUATION
      FOR THE PRESSURE

      COMMON/MIKE/ NETA1,NETA2,PNF,VNF,RHOD,DP,TS,UD,VD,RV,RWO,EPS,PHI
      COMMON /YL/ETA(60),YOND(60),YO(60)
      COMMON/JOHN/ UD(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILOYNOM
      COMMON /FRSTRM/ U INF, R INF, UINF2, R , RE, LXI, ITM, IEM, NETA
      COMMON/PROPI/PI(60),RHQ(60), T(60),AMW(60),C (20,60),EC(5,60)
      COMMON /DEL/ DELTA,DTIL,DTILS
      COMMON /VEL/ F(60),FC(60),Z(60),V(60)
      COMMON/BOB/ XI,DXI,YL

      DO 15 I=1,NETA1
      J = NETA - I
      DETA = ETA(J+1) - ETA(J)
      UA = (U(J)+U(J+1))/2.
      PI(J) = PI(J+1) - DTIL*UA+UA*DETA/PNF
15 CONTINUE
      RETURN
      END
C

```

```

YNOM 10
YNOM 20
YNOM 30
YNOM 40
YNOM 50
YNOM 60
YNOM 70
YNOM 80
YNOM 90
YNOM 100
YNOM 110
YNOM 120
YNOM 130
YNOM 140
YNOM 150
YNOM 160
YNOM 170
YNOM 180
YNOM 190
YNOM 200
YNOM 210
YNOM 220

```



```

SUBROUTINE XMOMTM(KODE)
C
C      THIS ROUTINE SOLVES THE X-MOMENTUM EQUATION FOR THE TANGENTIAL
C      VELOCITY (U) FOR THE RADIATING, CHEMICALLY REACTING, VISCOUS
C      FLOW IN THE SHOCK LAYER ABOUT A BLUNT BODY
C
COMMON /FRSTRM/ U INF, R INF, UINF2, R, RE, LXI, ITM, IEM, NETA
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/JOHN/ UO(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILO
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
COMMON/JTB/ CO(60,20),RED,RED2
COMMON/BOB/ XI,DXI,YL
COMMON/MIKE/ NETA1,NETA2,PNF,VNF,RHOD,DP,TS,UD,VD,RW,RWD,EPS,PHI
COMMON/CON/ PD(60)
COMMON/CONV1/HDAMP,FDAMP,CDAMP
COMMON /BILL/ MODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS
REAL*8 SUB,DIAG,SUP,B

DTLA = YL*DTIL + (1-YL)*DTILO
RED = RE*DTLA
RED2 = RED*DTLA
DO 5 I=1,NETA2
  J=I+1
  RHOA= YL*RHO(J) + (1-YL)*RO(J)
  VA = YL*V(J) + (1-YL)*VO(J)
  DP = (PI(J)-PO(J))*PNF/DXI
  DRH=H1(J)*RM(J-1)+G1(J)*RM(J)+F1(J)*RM(J+1)
  UA = YL*U(J) + (1-YL)*UO(J)
  DU(J)=(U(J)-UO(J))/DXI
  DUN = H1(J)*U(I)+G1(J)*U(J)+F1(J)*U(J+1)
  DUN= YL*DUN + (1-YL)*(H1(J)*UO(I)+G1(J)*UO(J)+F1(J)*UO(J+1))

```

```

XMOM 10
XMOM 20
XMOM 30
XMOM 40
XMOM 50
XMOM 60
XMOM 70
XMOM 80
XMOM 90
XMOM 100
XMOM 110
XMOM 120
XMOM 130
XMOM 140
XMOM 150
XMOM 160
XMOM 170
XMOM 180
XMOM 190
XMOM 200
XMOM 210
XMOM 220
XMOM 230
XMOM 240
XMOM 250
XMOM 260
XMOM 270
XMOM 280
XMOM 290
XMOM 300
XMOM 310
XMOM 320
XMOM 330
XMOM 340
XMOM 350
XMOM 360

```

```

C          CALCULATION OF A1
SUB(I) =DRM/RM(J)-RED*RHOA*VA/RM(J)
X          -RED2*UA*DNDX(J)/RM(J)
C
C          CALCULATION OF A2
DIAG(I)=-RED2*DU(J)/RM(J)
X          -RED2*DNDX(J)*DUN/RM(J)
X          -2.*RED2*DTLA*UA*DNDX(J)/RHOA/RM(J)
C
C          CALCULATION OF A3
SUP(I) =RED2/RM(J)*(UA *DU(J)-DP/RHOA)
X          +RED2*DNDX(J)*UA*(DUN + UA*DTLA/RHOA)/RM(J)
C
C          CALCULATION OF A4
B(I) =-RED2*UA /RM(J)
5 CONTINUE
CALL SOLVE(0.0,U(NETA),UO)
CALL CHECK(U,B,KODE,IDGF,'U' ,.FDAMP)
RETURN
END

```

```

XMOM 370
XMOM 380
XMOM 390
XMOM 400
XMOM 410
XMOM 420
XMOM 430
XMOM 440
XMOM 450
XMOM 460
XMOM 470
XMOM 480
XMOM 490
XMOM 500
XMOM 510
XMOM 520
XMOM 530
XMOM 540
XMOM 550
XMOM 560

```

```

SUBROUTINE GLOBAL (KODE)
C
C      SOLUTION TO THE GLOBAL CONTINUITY EQUATION FOR 'V' PROFILE
C      USING THE 'U' FORM OF THE GLOBAL CONT. EQUATION
C      THE TRANSFORMED STANDOFF DISTANCE. DTIL. IS ALSO CALCULATED
C
COMMON/WALL/RVW,PRW,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,MNF,CPNF
COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /FRSTRM/ U INF, RINF, UINF2, R , RE, LXI, ITM, IEM, META
COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
COMMON/JOHN/ UO(60),U(60),DU(60),VD(60),RO(60),DNDX(60),DTLA,DTILO
COMMON/JTB/ CO(60,20),RED,RED2
COMMON/BOB/ XI,DXI,YL
COMMON /YL/ETA(60),YOND(60),YO(60)
COMMON/MIKE/ META1,META2,PNF,MNF,RHOD,DP,TS,UD,VD,RV,RWO,EPS,PHI
COMMON /CONV/ FPRCT,TFRCT,DDAMP,TDAMP,POTIL
COMMON /BILL/ NODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS
REAL *8 SUB,SUP,DIAG,B

DTILC = DTIL
RWA = YL*RVW      + (1-YL)*RWO
IDN = 0
1 SUM=0.0
TERM1 = -RHO(1)*V(1)/DTIL
TERM1 = 0.0
TERM2 = 0.0
TERM3 = 0.0
DO 15 J=1,META1
I = J+1
DETA= ETA(1) - ETA(J)
UA = YL*UO(J)      + (1-YL)*UO(J)

```

GLOB 10  
 GLOB 20  
 GLOB 30  
 GLOB 40  
 GLOB 50  
 GLOB 60  
 GLOB 70  
 GLOB 80  
 GLOB 90  
 GLOB 100  
 GLOB 110  
 GLOB 120  
 GLOB 130  
 GLOB 140  
 GLOB 150  
 GLOB 160  
 GLOB 170  
 GLOB 180  
 GLOB 190  
 GLOB 200  
 GLOB 210  
 GLOB 220  
 GLOB 230  
 GLOB 240  
 GLOB 250  
 GLOB 260  
 GLOB 270  
 GLOB 280  
 GLOB 290  
 GLOB 300  
 GLOB 310  
 GLOB 320  
 GLOB 330  
 GLOB 340  
 GLOB 350  
 GLOB 360

```

      VA = YL*V(J) + (1-YL)*V(J)
      RHOA = YL*RHO(J) + (1-YL)*RHO(J)
      ROA = YL*RHO(I) + (1-YL)*RHO(I)
      ROAA = (RHOA + ROA)/2.
      DRU = (ROA*(YL*U(I) + (1-YL)*U(I)) - RHOA*UA)/DETA
      UAP = YL*U(I) + (1-YL)*U(I)
      UAA = (UA + UAP)/2.
      DNX = (DNX(I) + DNX(J))/2.
      TERM1 = TERM1 + DETA/ROAA +
1      (RHO(J)*U(J)-RO(J)*U(I) + RHO(I)*U(I)-RO(I)*UO(I))/DXI/2.
      TERM2 = TERM2 + DETA*UAA*(RW-RWO)/RWA/DXI
      TERM3 = TERM3 + DNX*DRU*DETA/ROAA
      SUM1 = SUM
      SUM = (TERM1 + TERM2 + TERM3)
      IF(ION.EQ.0) GO TO 15
      TERM = SUM - SUM1
      V(I) = (-(1-YL)*V(J) + (1-YL)*V(I) + (RHOA*VA - DTIL*TERM)/ROA)/YL
15 CONTINUE
      IF(ION.EQ.1) GO TO 25
      DTIL = -(RHO(NETA)*VD - RHO(I)*V(I))/SUM
      TEST = (DTIL - DTILC)/DTIL
      KODE = 5
      IF(ABS(TEST).GT. PDIL) KODE = 0
      DTIL = DDAMP*DTIL + (1.-DDAMP)*DTILC
      IDN = 1
      GO TO 1
25 CONTINUE
      IF(IDGD.GT.0)
        XWRITE(6,100) DTIL,DTILO,DELTA
      IF(IDGD.GT.0)
        XWRITE(6,100) TERM1,TERM2,TERM3
100 FORMAT(1H0,15X,'DTIL =',3E16.6)
      RETURN
      END

```

GLOB 370  
 GLOB 380  
 GLOB 390  
 GLOB 400  
 GLOB 410  
 GLOB 420  
 GLOB 430  
 GLOB 440  
 GLOB 450  
 GLOB 460  
 GLOB 470  
 GLOB 480  
 GLOB 490  
 GLOB 500  
 GLOB 510  
 GLOB 520  
 GLOB 530  
 GLOB 540  
 GLOB 550  
 GLOB 560  
 GLOB 570  
 GLOB 580  
 GLOB 590  
 GLOB 600  
 GLOB 610  
 GLOB 620  
 GLOB 630  
 GLOB 640  
 GLOB 650  
 GLOB 660  
 GLOB 670  
 GLOB 680  
 GLOB 690  
 GLOB 700

## SUBROUTINE SPECIE (KOD)

THIS ROUTINE SOLVES THE SPECIES EQUATIONS FOR RADIATING.  
 CHEMICAL NONEQUILIBRIUM, VISCOUS FLOW IN A SHOCK LAYER ABOUT  
 A BLUNT BODY

COMMON/JOHN/ UD(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DT ILO  
 COMMON/HOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)  
 COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NETA  
 COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)  
 COMMON /VEL/ F(60),FC(60),Z(60),V(60)  
 COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)  
 COMMON/NUMBER/NSP,MNS,NE,NC  
 COMMON/CK/ISN(20),XWT(19)  
 COMMON/FOG/ AII,BB,CC  
 COMMON/JTB/ CO(60,20),RED,RED2  
 COMMON /RM/ DUP,DPHI,TD,RZB,PD,HD,HTOTAL  
 COMMON/ID/SP(20),EL(5)  
 COMMON /NON/RDZ,NUDZ,RMDZ,AKNF,HNF,CPNF  
 COMMON/TOM/ SC(60),TO(60)  
 COMMON/NIKE/ NETA1,NETA2,PNF,WNF,RHOD,DP,TS,UD,VD,RW,RWO,EPS,PHI  
 COMMON/BOB/ XI,DXI,YL  
 COMMON /DEL/ DELTA,DTIL,DTILS  
 DIMENSION DRO(60),Y(20),FT(60)  
 COMMON /BILL/ KODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS  
 COMMON/CONVI/HDAMP,FDAMP,CDAMP  
 REAL\*8 SUB,DIAG,SUP,B  
 REAL NUDZ

KOD = 5  
 DTLA = YL\*DTIL + (1-YL)\*DTILO  
 RED = RE\*DTLA  
 RED2 = RED\*DTLA  
 CALL NASMDL (NSP,NETA)  
 DO 15 K=1,NSP  
 KK=ISN(K)

SPEC 10  
 SPEC 20  
 SPEC 30  
 SPEC 40  
 SPEC 50  
 SPEC 60  
 SPEC 70  
 SPEC 80  
 SPEC 90  
 SPEC 100  
 SPEC 110  
 SPEC 120  
 SPEC 130  
 SPEC 140  
 SPEC 150  
 SPEC 160  
 SPEC 170  
 SPEC 180  
 SPEC 190  
 SPEC 200  
 SPEC 210  
 SPEC 220  
 SPEC 230  
 SPEC 240  
 SPEC 250  
 SPEC 260  
 SPEC 270  
 SPEC 280  
 SPEC 290  
 SPEC 300  
 SPEC 310  
 SPEC 320  
 SPEC 330  
 SPEC 340  
 SPEC 350  
 SPEC 360

```

IF(IEM.GT.1 .AND. KODE(KK).GT.4) GO TO 15
DO 5 I=1,META2
  J=I+1
  DO 25 L=1,NSP
    C(ISN(L),J)= ABS(C(ISN(L),J))
    Y(L)= YL+C(ISN(L),J) + (1-YL)*CO(J,ISN(L))
  25 CONTINUE
  TA = YL*Y(J) + (1-YL)*TO(J)
  UA = YL*U(J) + (1-YL)*UD(J)
  VA = YL*V(J) + (1-YL)*VD(J)
  RMOA= YL*RHO(J) + (1-YL)*RO(J)
  CALL FG2(Y,RMOA,RMOA,TS+TA,K)
  DSC = F1(J)*SC(J+1) + G1(J)*SC(J) + H1(J)*SC(1)
C
C      CALCULATION OF A1
SUB(I) =( DSC-RED*RMOA+VA - RED2*UA+DNDX(J))/SC(J)
C
C      CALCULATION OF A2
DIAG(I)=RED2*A11*WNF/SC(J)
C
C      CALCULATION OF A3
SUP(I) =RED2*(CC-A11*Y(K))*WNF/SC(J)
C
C      CALCULATION OF A4
B(I) =-RED2*UA/SC(J)
5 CONTINUE
C
CALL SOLVE (C(KK,1),C(KK,META),CO(1,KK))
DO 45 I=1,META
  FT (I)=C(KK,I)
45 CONTINUE
CALL CHECK (FT,B,KODE(KK),IDGS,SP(KK),CDAMP)
DO 55 I=1,META
  C(KK,I)= FT(I)
55 CONTINUE
KOD = AMINO (KOD,KODE(KK))
SPEC 370
SPEC 380
SPEC 390
SPEC 400
SPEC 410
SPEC 420
SPEC 430
SPEC 440
SPEC 450
SPEC 460
SPEC 470
SPEC 480
SPEC 490
SPEC 500
SPEC 510
SPEC 520
SPEC 530
SPEC 540
SPEC 550
SPEC 560
SPEC 570
SPEC 580
SPEC 590
SPEC 600
SPEC 610
SPEC 620
SPEC 630
SPEC 640
SPEC 650
SPEC 660
SPEC 670
SPEC 680
SPEC 690
SPEC 700
SPEC 710
SPEC 720

```

SPEC 730  
SPEC 740  
SPEC 750  
SPEC 760  
SPEC 770  
SPEC 780  
SPEC 790

15 CONTINUE  
IF(IDGS.GT.0)  
XWRITE (6,100) (SP(1),KODE(1),I=1,NSP)  
100 FORMAT((2X,10(A4,I2.4X)))  
CALL MOLMAS (NSP,NETA)  
RETURN  
END

MOLM 10  
 MOLM 20  
 MOLM 30  
 MOLM 40  
 MOLM 50  
 MOLM 60  
 MOLM 70  
 MOLM 80  
 MOLM 90  
 MOLM 100  
 MOLM 110  
 MOLM 120  
 MOLM 130  
 MOLM 140  
 MOLM 150  
 MOLM 160  
 MOLM 170  
 MOLM 180  
 MOLM 190  
 MOLM 200  
 MOLM 210  
 MOLM 220  
 MOLM 230  
 MOLM 240  
 MOLM 250  
 MOLM 260  
 MOLM 270

```

SUBROUTINE MOLMAS (K,N)
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/WT/SW(20),AWT(5)
COMMON/JTB/ CO(60,20),RED,RED2
DO 5 J=1,N
AM=0.0
SUM=0.0
DO 15 I=1,K
AM = AM + C(I,J)
C(I,J)=C(I,J)*SW(I)
SUM=SUM + C(I,J)
15 CONTINUE
DO 35 I=1,K
C(I,J)=C(I,J)/SUM
CO(J,I)=CO(J,I)*SW(I)
35 CONTINUE
AMW(J)=1./AM
5 CONTINUE
RETURN
ENTRY MASMOL (K,N)
DO 25 J=1,N
DO 25 I=1,K
C(I,J)=C(I,J)/SW(I)
CO(J,I)=CO(J,I)/SW(I)
25 CONTINUE
RETURN
END
  
```



```

C
C
C
C
C
C
SUBROUTINE ENERGY (KODE)
    THIS ROUTINE SOLVES THE ENTHALPY FORM OF THE ENERGY EQUATION
    FOR THE RADIATING, CHEMICALLY REACTING, VISCOUS FLOW IN A
    SHOCK LAYER ABOUT A BLUNT BODY

    COMMON/DON/ PO(60),EO(60),HMO(60)
    COMMON/BOB/ XI,DXI,YL
    COMMON/MIKE/ NETA1,NETA2,PNF,VNF,RHOD,DP,TS,UD,VD,RV,RWD,EPS,PHI
    COMMON/PROP2/VISM(60),RM(60),TCM(60)
    COMMON/JTB/ CO(60,20),RED,RED2
    COMMON/VEL/ F(60),FC(60),Z(60),V(60)
    COMMON/MOFF1/F1(60),F2(60),G1(60),G2(60),H1(60),H2(60)
    COMMON/TOM/ SC(60),TO(60),PR(60),TE(60)
    COMMON/JOHN/ UO(60),U(60),DU(60),VD(60),RO(60),DNDX(60),DTLA,DTILO
    COMMON/NUMBER/NSP,NNS,NE,NC
    COMMON/NOB/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
    COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
    COMMON/PROPI/PI(60),RHO(60),T(60),AMW(60),C(20,60),EC(5,60)
    COMMON/FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
    COMMON/DEL/ DELTA,DTIL,DTILS
    COMMON/PROP3/CP(20,60),H(20,60),CPM(60),HM(60)
    COMMON/RFLUX/ E(60),IRAD,ITYPE
    COMMON/CONV1/HDAMP,FDAMP,CDAMP
    COMMON/BILL/ NODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS
    REAL *8 SUB,SUP,DIAG,B

    IF ( KODE .GT. 4 .AND. IEM .GT. 1 ) RETURN
    DTLA = YL*DTIL + (1-YL)*DTILO
    DTIL2 = DTLA*DTLA

    EMA = YL*E(1) + (1-YL)*EO(1)
    EP = E(2)
    EA = YL*EP + (1-YL)*EO(2)
    DO 15 I=1,NETA2
        J = I+1

```

ENER 10  
ENER 20  
ENER 30  
ENER 40  
ENER 50  
ENER 60  
ENER 70  
ENER 80  
ENER 90  
ENER 100  
ENER 110  
ENER 120  
ENER 130  
ENER 140  
ENER 150  
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ENER 350  
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ENER 370  
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ENER 390  
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ENER 410  
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ENER 570  
ENER 580  
ENER 590  
ENER 600  
ENER 610  
ENER 620  
ENER 630  
ENER 640  
ENER 650  
ENER 660  
ENER 670  
ENER 680  
ENER 690  
ENER 700  
ENER 710  
ENER 720

```

EP = E(J+1)
EPA = YL*EP + (1-YL)*EO(J+1)
HMM = YL*MM(I) + (1-YL)*MMO(I)
HMP = YL*MM(J+1) + (1-YL)*HMO(J+1)
VA = YL*V(J) + (1-YL)*VO(J)
UA = YL*U(J) + (1-YL)*UO(J)
RHOA = YL*RHO(J) + (1-YL)*RO(J)
S1 = 0.0
S2 = 0.0
S3 = 0.0
DO 25 K=1,NSP
DH = F1(J)*H(K,J+1) + G1(J)*H(K,J) + H1(J)*H(K,I)
DC1 = F1(J)*C(K,J+1) + G1(J)*C(K,J) + H1(J)*C(K,I)
DC2 = F2(J)*C(K,J+1) + G2(J)*C(K,J) + H2(J)*C(K,I)
S1 = S1 + DC1*DH
S2 = S2 + H(K,J)*DC1
S3 = S3 + H(K,J)*DC2
25 CONTINUE
DPR = F1(J)*PR(J+1) + G1(J)*PR(J) + H1(J)*PR(I)
DSC = F1(J)*SC(J+1) + G1(J)*SC(J) + H1(J)*SC(I)
DRM = F1(J)*RM(J+1) + G1(J)*RM(J) + H1(J)*RM(I)
DVN = F1(J)*V(J+1) + G1(J)*V(J) + H1(J)*V(I)
DUN = F1(J)*U(J+1) + G1(J)*U(J) + H1(J)*U(I)
DUN2 = F2(J)*U(J+1) + G2(J)*U(J) + H2(J)*U(I)
DVN = YL*DVN + (1-YL)*(F1(J)*VO(J+1)+G1(J)*VO(J)+H1(J)*VO(I))
DUN = YL*DUN + (1-YL)*(F1(J)*UO(J+1)+G1(J)*UO(J)+H1(J)*UO(I))
DUN2 = YL*DUN2 + (1-YL)*(F2(J)*UO(J+1)+G2(J)*UO(J)+H2(J)*UO(I))
UP = (U(J) - UO(J))/DX1
DV = (V(J) - VO(J))/DX1
SC(J) = RM/SC
S1 = S1*(SC(J)/RE/2. - PR(J))
S3 = S3*(SC(J)/RE/2. - PR(J))
S2 = S2*(DSC/RE/2. - DPR)
S4 = ((RM(J)*DUN + UA*DRM)*DUN + RM(J)*UA*DUN2)/RE
S5 = DTIL2*(UA*UA*UP + (UA*UA*DNDX(J)+RHOA*VA*UA/DTLA)*DUN
      + VA*UA*DV + (VA*UA*DNDX(J) + RHOA*VA*VA/DTLA)*DVN)
X

```

C

ENER 730  
ENER 740  
ENER 750  
ENER 760  
ENER 770  
ENER 780  
ENER 790  
ENER 800  
ENER 810  
ENER 820  
ENER 830  
ENER 840  
ENER 850  
ENER 860  
ENER 870  
ENER 880  
ENER 890  
ENER 900  
ENER 910  
ENER 920  
ENER 930  
ENER 940  
ENER 950  
ENER 960

```

C
C
C      CALCULATION OF A1
      SUB(I) = (DPR - UA*DTIL2*DNDX(J))/2. - RHOA*VA*DTLA/2.)/PR(J)
C
C
C      CALCULATION OF A2
      DIAG(I) = (EPA - EMA)/(HMP - HMM)/PR(J)*(- DTIL2/RHOA)
      DIAG(I) = 0.0
C
C
C      CALCULATION OF A3
      SUP(I) = ( S1 + S2 + S3 - DTIL2*EA/RHOA + S4 - S5 )/PR(J)
      X = DIAG(I)*HMM(J)
C
C
C      CALCULATION OF A4
      B(I) = -DTIL2*UA/PR(J)/2.
      EMA = EA
      EA = EPA
      IS CONTINUE
C
C
C      CALL SOLVE (MM(1),MM(NETA),HNO)
      CALL CHECK (MM,B,KODE,IDGH,'ENTH',MDAMP)
      CALL TEMPR
      RETURN
      END

```

```

C
C
C
C
SUBROUTINE TEMPR
C
C      THIS ROUTINE DETERMINES THE TEMPERATURE PROFILE GIVEN
C      A STYTIC ENTHALPY PROFILE AND SPECIES PROFILES
C
COMMON/PROP3/CP(20,60),H(20,60),CPM(60),HM(60)
COMMON /FRSTRM/ U INF, RINF, UINF2, R , RE, LXI, ITM, IEM, META
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
COMMON/MIKE/ META1,META2,PNF,MNF,RHOD,DP,TS,UD,VD,RV,RWD,EPS,PHI
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON /MAIM/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /CONV/ FPRCT,TPRCT,DDAMP,TDAMP,PD TIL
COMMON /BILL/ NODE(20),ISTART,IDGD,IDGF,IDGH,IDGT,IDGS
COMMON /YL/ETA(60),YOND(60)
REAL *8 SUB,SUP,DIAG,B

DO 15 J=2,META1
TC = T(J)
B(J-1) = HM(J)

K = 0
10 CALL PROPT (NSP,J,J)
K = K + 1
DTT = (B(J-1) - HM(J))/CPM(J)
T(J) = T(J) + DTT
DTT = DTT/T(J)

IF(K .GT. MAXE) GO TO 30
IF(ABS(DTT).GT. TPRCT) GO TO 10
GO TO 20

30 WRITE(6,200) MAXE,J,TC,T(J)
200 FORMAT(2X,*** MAXIMUM NO. OF ITERATIONS('13.') EXCEEDED FOR TEMPR
XERATURE - ENTHALPY AT ETA POINT,'13./,5X,'TOLD='F10.4,3X,'TNEW='F10.4)
X F10.4)
20 B(J-1) = T(J)
TEMP 10
TEMP 20
TEMP 30
TEMP 40
TEMP 50
TEMP 60
TEMP 70
TEMP 80
TEMP 90
TEMP 100
TEMP 110
TEMP 120
TEMP 130
TEMP 140
TEMP 150
TEMP 160
TEMP 170
TEMP 180
TEMP 190
TEMP 200
TEMP 210
TEMP 220
TEMP 230
TEMP 240
TEMP 250
TEMP 260
TEMP 270
TEMP 280
TEMP 290
TEMP 300
TEMP 310
TEMP 320
TEMP 330
TEMP 340
TEMP 350
TEMP 360

```

TEMP 370  
TEMP 380  
TEMP 390  
TEMP 400  
TEMP 410  
TEMP 420  
TEMP 430

DIFF = (T(J) - TC)/TC  
IF(ABS(DIFF) .GT. 0.03) B(J-1) = TC + .03\*TC\*ABS(DIFF)/DIFF  
T(J) = TC  
15 CONTINUE  
CALL CHECK (T,B,KODE,IDGT,TEMP,TDAMP)  
RETURN  
END

```

10 PROP
20 PROP
30 PROP
40 PROP
50 PROP
60 PROP
70 PROP
80 PROP
90 PROP
100 PROP
110 PROP
120 PROP
130 PROP
140 PROP
150 PROP
160 PROP
170 PROP
180 PROP
190 PROP
200 PROP
210 PROP
220 PROP
230 PROP
240 PROP
250 PROP
260 PROP
270 PROP
280 PROP
290 PROP
300 PROP
310 PROP
320 PROP
330 PROP
340 PROP
350 PROP
360 PROP

SUBROUTINE PROPRT(NSP,NI,NF)
* * * * *
SUBROUTINE FOR THE CALCULATION OF THERMODYNAMIC AND TRANSPORT
PROPERTIES

NOMENCLATURE....
Y(I)...MOLE FRACTION OF COMPONENT I
C(I,N)...MASS FRACTION OF COMPONENT I AT POINT N
T1.....TEMPERATURE, DEG. K
CP(I)....SPECIES SPECIFIC HEAT, CAL/GMOLE OF I-K
CPM(N)....MIXTURE SPECIFIC HEAT, CAL/GMOLE OF M-K
H(I)....SPECIES ENTHALPY, CAL/GMOLE OF I
HM(N)....MIXTURE ENTHALPY, CAL/GMOLE OF M
VIS(I)....SPECIES VISCOSITY, LBM/FT-SEC
VISM(N)....MIXTURE VISCOSITY, LBM/FT-SEC
TC(I)....SPECIES THERMAL CONDUCTIVITY, BTU/FT-SEC-DEG. R
TCM(N)....MIXTURE THERMAL CONDUCTIVITY, BTU/FT-SEC-DEG. R

COMMON /BLOCK1/V1(20),V2(20),V3(20)
COMMON /BLOCK3/K1(20),K2(20)
COMMON /EQ1/A1(20), B1(20), C1(20), D1(20), E1(20), F1(20), G1(20),
X A11(20),B11(20),C11(20),D11(20),E11(20),F11(20),G11(20)
COMMON /FRSTRM/ U INF, RINF, UINF2,RAD, RE, LXI, ITM, IEM,META
COMMON /PROPI/PI(60),RHO(60),Y(60),AMW(60),C (20,60),CC(5,60)
COMMON /PROP2/VISM(60),RM(60),TCM(60)
COMMON /PROP3/CP(20,60),H(20,60),CPM(60),HM(60)
COMMON /WT/SMW(20),AWT(5)
COMMON /RH/ DUD,DPHI,TD,RZB,PD,MD,HTOTAL
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,MNF,CPNF
COMMON /TOM/ SC(60),TO(60),PR(60)
COMMON /NIKE/ META1,META2,PNF,MNF,RHOD,DP,TS=UD,VD,RV,RMO,EPS
COMMON /JOHN/ UD(60),U(60),DU(60),VO(60),FO(60),DNDX(60),DTILU,
COMMON /BOB/ XI,DXI,YL
DIMENSION PHI(20,20),YPHI(20),Y(20),TC(20),VIS(20)
DIMENSION SAMW(2)
REAL MU,MUDZ,K1,K2

```

```

C
DATA R /1.98716/
DO500N=NI,NF
TA = YL*Y(N) + (1-YL)*Y(N)
T1=TA*TS
T2=T1*T1
T3=T2*T1
T4=T3*T1
T5=T4*T1
DO150I=1,NSP
IF(T1.GT.6000.)GOTO 50
CP(I,N)=( AI(I)+ BI(I)*T1+ CI(I)*T2+ DI(I)*T3+ EI(I)*T4)*R
H(I,N)=( AI(I)*T1+ BI(I)*T2/2.+ CI(I)*T3/3.+ DI(I)*T4/4.
      + EI(I)*T5/5.+ FI(I))*R
      X
      GOTO 100
50 CP(I,N)=(AI(I)+BI(I)*T1+CI(I)*T2+DI(I)*T3+EI(I)*T4)*R
H(I,N)=(AI(I)*T1+BI(I)*T2/2.+CI(I)*T3/3.+DI(I)*T4/4.
      +EI(I)*T5/5.+FI(I))*R
      X
100 IF(NF.NE.NETA) GO TO 150
Y(I)=C(I,N)*AMW(N)/SMW(I)
VIS(I)=(V1(I) + V2(I)*T1 + V3(I)*T2)*1.0E-05
TC(I)=(K1(I)+K2(I)*T1)*1.0E-5
150 CONTINUE
IF(NF.NE.NETA) GO TO 500
C-----CALCULATE PHI(I,J) PARAMETERS FOR MIXTURE PROPERTIES....
C
DO200 I=1,NSP
DO200 J=1,NSP
VIS12=SQRT(VIS(I)/VIS(J))
SMW14=((SMW(J)/SMW(I))**.25
PHI(I,J)=.35*((1.+VIS12*SMW14)**2)/SORT(1.+SMW(I)/SMW(J))
200 CONTINUE
C
C-----CALCULATION OF MIXTURE PROPERTIES....
C
DO250 I=1,NSP
PROP 370
PROP 380
PROP 390
PROP 400
PROP 410
PROP 420
PROP 430
PROP 440
PROP 450
PROP 460
PROP 470
PROP 480
PROP 490
PROP 500
PROP 510
PROP 520
PROP 530
PROP 540
PROP 550
PROP 560
PROP 570
PROP 580
PROP 590
PROP 600
PROP 610
PROP 620
PROP 630
PROP 640
PROP 650
PROP 660
PROP 670
PROP 680
PROP 690
PROP 700
PROP 710
PROP 720

```

```

YPHI(I)=0.0
DO250 J=1,NSP
  250 YPHI(I)=YPHI(I) + Y(J)*PHI(I,J)
  VISM(N)=0.0
  TCM(N) = 0.
  DO300 I=1,NSP
    IF(Y(I).LT.1.E-5) GO TO 300
    VISM(N)=VISM(N)+Y(I)*VIS(I)/YPHI(I)
    300 TCM(N) = TCM(N) + Y(I)*TC(I)/YPHI(I)
    RHO(N) = PI(N)*AMW(N)/T(N)/TS/RDZ/1.3157/32.174
    RHOA= YL*RHO(N) + (1-YL)*RO(N)
    D=8.128E-8*T1*1.659/PD
    SC(N) = VISM(N)/RHOA/RHOD/O
  500 CONTINUE
C
C----- NONDIMENSIONALIZE QUANTITIES -----
C
DO 550 N=N1,NF
  CPM(N) = 0.
  HM(N) = 0.
  DO 551 I=1,NSP
    CP(I,N) = CP(I,N)*CPNF/SW(I)
    H(I,N) = H(I,N)*HNF*1.80/SW(I)
    CPM(N) = CPM(N) + CP(I,N)*C(I,N)
    HM(N) = HM(N) + H(I,N)*C(I,N)
  555 CONTINUE
  IF(NF.NE.NETA) GO TO 550
  VISM(N) = VISM(N)/MUDZ
  TCM(N) = TCM(N)*AKNF
  RHOA= YL*RHO(N) + (1-YL)*RO(N)
  RM(N)=RHOA*VISM(N)
  SC(N) = RM(N)/SC(N)
  PR(N) = RHOA*TCM(N)/CPM(N)
  550 CONTINUE
  RETURN
  END

```

PROP 730  
 PROP 740  
 PROP 750  
 PROP 760  
 PROP 770  
 PROP 780  
 PROP 790  
 PROP 800  
 PROP 810  
 PROP 820  
 PROP 830  
 PROP 840  
 PROP 850  
 PROP 860  
 PROP 870  
 PROP 880  
 PROP 890  
 PROP 900  
 PROP 910  
 PROP 920  
 PROP 930  
 PROP 940  
 PROP 950  
 PROP 960  
 PROP 970  
 PROP 980  
 PROP 990  
 PROP1000  
 PROP1010  
 PROP1020  
 PROP1030  
 PROP1040  
 PROP1050  
 PROP1060  
 PROP1070  
 PROP1080



```

SUBROUTINE TRID (M)
C*** TRID --TRIDIAGONAL EQUATION SOLVER OBTAINED FROM CONTE P-184 ***
C SUBROUTINE SOLVES AX = B FOR THE VECTOR X (WHERE A IS TRIDIAGONAL)
C M = ORDER OF SYSTEM
C SUP = SUPER DIAGONAL OF A
C SUB = SUB DIAGONAL OF A
C DIAG = MAIN DIAGONAL OF A
C B = CONSTANT VECTOR
C SUP AND DIAG ARE DESTROYED
C SOLUTION VECTOR IS RETURNED IN B
C
COMMON/VECTOR/ SUB(60),DIAG(60),SUP(60),B(60)
REAL*8 SUB,DIAG,SUP,B
C
N = M
NN = N - 1
SUP(1) = SUP(1)/DIAG(1)
B(1) = B(1)/DIAG(1)
DO 10 I=2,N
  II = I - 1
  C-----DECOMPOSE A TO FORM A = LU WHERE L IS LOWER TRIANGULAR,
  C AND U IS UPPER TRIANGULAR -----
  DIAG(I) = DIAG(I) - SUP(II)*SUB(II)
  IF(I.EQ. N) GO TO 10
  SUP(I) = SUP(I) / DIAG(II)
  C-----COMPUTE Z WHERE LZ = B
  10 B(I) = (B(I) - SUB(II) *B(II))/ DIAG(I)
  C-----COMPUTE X BY BACK SUBSTITUTION WHERE UX = Z
  DO 20 K =1,NN
    I = N - K
    20 B(I) = B(I) -SUP(I) *B(I+1)
  RETURN
END

```

TRID 10  
 TRID 20  
 TRID 30  
 TRID 40  
 TRID 50  
 TRID 60  
 TRID 70  
 TRID 80  
 TRID 90  
 TRID 100  
 TRID 110  
 TRID 120  
 TRID 130  
 TRID 140  
 TRID 150  
 TRID 160  
 TRID 170  
 TRID 180  
 TRID 190  
 TRID 200  
 TRID 210  
 TRID 220  
 TRID 230  
 TRID 240  
 TRID 250  
 TRID 260  
 TRID 270  
 TRID 280  
 TRID 290  
 TRID 300  
 TRID 310  
 TRID 320  
 TRID 330

U U U U U U U U U U U U U

OUTP 10  
 OUTP 20  
 OUTP 30  
 OUTP 40  
 OUTP 50  
 OUTP 60  
 OUTP 70  
 OUTP 80  
 OUTP 90  
 OUTP 100  
 OUTP 110  
 OUTP 120  
 OUTP 130  
 OUTP 140  
 OUTP 150  
 OUTP 160  
 OUTP 170

```

SUBROUTINE OUTPUT(W,B,SP,N)
C
C   DEBUG OUTPUT FOR GUESSED(W) AND CALCULATED(B) VALUES OF *SP*
C
  DIMENSION W(1),B(1)
  REAL*8 B
C
  WRITE(6,110) SP
  110 FORMAT(1H0,5X,A4,/)
  I=1
  WRITE(6,100)I,W(I)
  M=N+1
  L=M+1
  WRITE(6,100) (J,W(J),B(J-1),J=2,M),L,W(L)
  100 FORMAT(10X,3(15,2E16.7))
  RETURN
END

```

```

SUBROUTINE FG2 (Y,RHOA,TN,K)
C-----ORDER OF SPECIES ISO
C      C      C      C      C      C      C      C      C      C      C      C      C      C      C      C      C      C      C      C
C      1      2      3      4      5      6      7      8      9      10     11     12     13     14     15     16     17     18     19     20
C      CN      C2H      C+      C3      C4      C5      C6      C7      C8      C9      C10     C11     C12     C13     C14     C15     C16     C17     C18     C19
C      7      6      5      4      3      2      1      0      14     13     12     11     10     9      8      7      6      5      4      3
C      HCN      H2      H      N2      N      O+      E-      C2H2      C2H      C3H      C4H      C5H      C6H      C7H      C8H      C9H      C10H     C11H     C12H
C      15      16      17      18      19      20      21      22      23      24      25      26      27      28      29      30      31      32      33      34
C      COMMON/FOG/ALL .B .C .
C      DIMENSION Y(19),T1(5)
C      DIMENSION RHO(2),OK(16,2),PL(16,10),RATES(3,16),EQK(16)
C      DATA R/82.05/
C      DATA RATES/
C      KINETIC COEFFICIENTS FOR CASE 1
C      1 .4352597E 02. .5000000E 00. .3500000E 05. .2683251E 02.
C      2 .5000000E 00. .1761263E 05. .4835429E 02. -.1500000E 01.
C      3 .1131738E 05. .4272747E 02. -.8200000E 00. .5193211E 05.
C      4 .2866064E 02. .5000000E 00. .1575073E 06. .2870602E 02.
C      5 .5000000E 00. .1675716E 06. .4684016E 02. -.1000000E 01.
C      6 .6632414E 05. .4357476E 02. .5000000E 00. .6038617E 05.
C      7 .4369782E 02. .5000000E 00. .7045053E 05. .4369782E 02.
C      8 .5000000E 00. .5887652E 05. .4588918E 02. -.1000000E 01.
C      9 .1297799E 06. .4374912E 02. .5000000E 00. .9561144E 05.
C      A .4367655E 02. .5000000E 00. .7052980E 05. .4368724E 02.
C      1 .5000000E 00. .1307138E 06. .4369782E 02. .5000000E 00.
C      2 .8303098E 05. .4369782E 02. .5000000E 00. .7296662E 05/
C      DIMENSION E1(95),E2(95),E3(34),E(16,7,2)
C      EQUIVALENCE (E(1,1,1),E1(1))
C      DATA E1/
C      1 0.3530E 00. 0.1816E 01. 0.1727E 01. 0.1642E 01. 0.2321E 01.
C      2 0.2753E 01. 0.1675E 01. 0.2257E 01. 0.3458E 01. 0.2094E 01.
C      3 0.2028E 01. 0.3053E 01. 0.7810E 00. 0.2497E 01. 0.2132E 01.
C      4 0.7030E 00. -0.7681E -03. -0.3571E -02. -0.8059E -03. -0.2810E -03.
C      5 0.2250E -03. -0.3726E -03. -0.6017E -03. -0.2955E -02. -0.3852E -02.
C      6 -0.2155E -02. -0.1370E -02. -0.4032E -02. -0.1175E -03. 0.6404E -04.
C      7 -0.2840E -02. -0.1406E -02. 0.5153E -06. 0.1634E -05. 0.1344E -06.

```

FG2 10  
 FG2 20  
 FG2 30  
 FG2 40  
 FG2 50  
 FG2 60  
 FG2 70  
 FG2 80  
 FG2 90  
 FG2 100  
 FG2 110  
 FG2 120  
 FG2 130  
 FG2 140  
 FG2 150  
 FG2 160  
 FG2 170  
 FG2 180  
 FG2 190  
 FG2 200  
 FG2 210  
 FG2 220  
 FG2 230  
 FG2 240  
 FG2 250  
 FG2 260  
 FG2 270  
 FG2 280  
 FG2 290  
 FG2 300  
 FG2 310  
 FG2 320  
 FG2 330  
 FG2 340  
 FG2 350  
 FG2 360

8-0.9244E-07.-0.9094E-07. 0.1884E-06.-0.6914E-07. 0.1359E-05.	FG2	370
9 0.1541E-05. 0.7206E-06. 0.4461E-06. 0.1731E-05.-0.8460E-07.	FG2	380
A-0.5011E-07. 0.1138E-05. 0.5735E-06.-0.9618E-10.-0.2783E-09.	FG2	390
1 0.4980E-11. 0.2914E-10. 0.1275E-10.-0.3765E-10. 0.3996E-10.	FG2	400
2-0.2518E-09.-0.2492E-09.-0.1067E-09.-0.5622E-10.-0.2058E-09.	FG2	410
3 0.2854E-10. 0.6619E-11.-0.2020E-09.-0.8895E-10. 0.5427E-14.	FG2	420
4 0.1617E-13.-0.9780E-15.-0.2112E-14.-0.5056E-15. 0.2271E-14.	FG2	430
5-0.2991E-14. 0.1668E-13. 0.1406E-13. 0.5735E-14. 0.2436E-14.	FG2	440
6 0.1621E-13.-0.2197E-14.-0.2271E-15. 0.1335E-13. 0.4609E-14.	FG2	450
7 0.3485E 05. 0.1329E 05. 0.1132E 06. 0.5196E 05. 0.1580E 06.	FG2	460
8 0.1686E 06. 0.9406E 05. 0.5850E 05. 0.6525E 05. 0.5766E 05.	FG2	470
9 0.1289E 06. 0.8906E 05. 0.7297E 05. 0.1306E 06. 0.8068E 05/	FG2	480
EQUIVALENCE (E(16.6.1).E2(1))		
DATA E2/		
1 0.7220E 05. 0.7500E-01.-0.8961E 01. 0.4274E 01. 0.2626E 01.	FG2	500
2-0.1181E 02.-0.1238E 02. 0.3698E 01. 0.1912E 01.-0.6335E 01.	FG2	510
3 0.3671E 01. 0.3773E 01. 0.1034E 01. 0.9378E 01.-0.1216E 02.	FG2	520
4 0.5461E 01. 0.1162E 02.-0.9100E-01.-0.1852E 01. 0.1765E 01.	FG2	530
5 0.4505E 01. 0.2904E 01. 0.2261E 01. 0.1414E 01. 0.3753E 01.	FG2	540
6 0.2653E 01. 0.2452E 01. 0.1323E 01.-0.1596E 02. 0.2560E 00.	FG2	550
7 0.2895E 01. 0.3483E 01. 0.2320E 00. 0.2624E-03. 0.1831E-02.	FG2	560
8-0.1250E-02.-0.4018E-02.-0.3576E-03. 0.3808E-03.-0.8027E-03.	FG2	570
9-0.4486E-02.-0.2187E-02.-0.2382E-02.-0.5403E-03. 0.1840E-01.	FG2	580
A 0.1581E-03.-0.3234E-03.-0.4981E-02.-0.6811E-03.-0.8699E-09.	FG2	590
1-0.5850E-06. 0.3816E-06. 0.1254E-05. 0.6591E-07.-0.1210E-06.	FG2	600
2 0.1697E-06. 0.1768E-05. 0.6688E-06. 0.6930E-06. 0.1688E-06.	FG2	610
3-0.5690E-05.-0.3970E-07. 0.4932E-07. 0.2072E-05. 0.4090E-06.	FG2	620
4-0.5441E-11. 0.7641E-10.-0.3536E-10.-0.1592E-09.-0.3159E-11.	FG2	630
5 0.1070E-10.-0.1315E-10.-0.2902E-09.-0.8277E-10.-0.8428E-10.	FG2	640
6-0.1859E-10. 0.6841E-09. 0.2542E-11.-0.2579E-11.-0.3588E-09.	FG2	650
7-0.6840E-10. 0.2100E-15.-0.3417E-14. 0.1003E-14. 0.7013E-14.	FG2	660
8 0.2813E-16.-0.3032E-15. 0.3831E-15. 0.1768E-13. 0.3596E-14.	FG2	670
9 0.3630E-14. 0.5931E-15.-0.2842E-13. 0.2920E-17. 0.2381E-16/	FG2	680
EQUIVALENCE (E(15.5.2).E3(1))		
DATA E3/		
1 0.2234E-13. 0.3694E-14. 0.4160E 05. 0.1329E 05. 0.1132E 06.	FG2	690
	FG2	700
	FG2	710
	FG2	720

```

2 0.5196E 05. 0.1580E 06. 0.1686E 06. 0.8731E 05. 0.6525E 05.
3 0.6525E 05. 0.5766E 05. 0.1289E 06. 0.8906E 05. 0.7297E 05.
4 0.1306E 06. 0.8068E 05. 0.7220E 05. 0.2066E 01. 0.1126E 02.
5 0.4450E 01.-0.1348E 02.-0.1538E 02.-0.1000E 02. 0.5594E 01.
6-0.6819E 01.-0.2220E 01. 0.1504E 01. 0.7660E 01. 0.1101E 03.
7 0.1266E 02.-0.1481E 02.-0.1881E 01. 0.1435E 02/
  IF (K.LT. 20) GO TO 5
  A11=0.0
  C =0.0
  RETURN
5 CONTINUE
  C      RHOA= DENSITY (LBM/FT3)
  RHOB= .031085*RHOA
  RHO(1)=.515362*RHOB
  RHO(2)=RHO(1)*RHO(1)
  X=0.
  DO 6 I=1. 19
    6 X=X+Y(I)
  IF(ABS((TN-T)/TN)--.0001)20.20.10
10 T=TN
  T1(1) = T
  DO12M=2.4
12 T1(M) = T1(1)*T1(M-1)
  T1(5) = ALOG(T1(1))
  RT = R*T
  L=1
  IF(T1(1).GT.6000.)L=2
  DO 18 I=1. 16
    OK(1,1)=T*RATES(2,1)*EXP(RATES(1,1)-RATES(3,1)/T)
    OK(1,2) = E(1,1,L)*(1.-T1(5))-E(1,2,L)*T1(1)/2.-E(1,3,L)*T1(2)/6.
    X-E(1,4,L)*T1(3)/12.-E(1,5,L)*T1(4)/20.+E(1,6,L)/T1(1)-E(1,7,L)
    EOK(1) = EXP(-OK(1,2))
    OK(1,2) = OK(1,1)/ EOK(1)
    IF(1.GT.2)OK(1,2) = RT*OK(1,2)
18 CONTINUE
20 PL( 1. 1) = OK( 1. 1)*RHO(1)*Y( 12)

```

FG2 730  
 FG2 740  
 FG2 750  
 FG2 760  
 FG2 770  
 FG2 780  
 FG2 790  
 FG2 800  
 FG2 810  
 FG2 820  
 FG2 830  
 FG2 840  
 FG2 850  
 FG2 860  
 FG2 870  
 FG2 880  
 FG2 890  
 FG2 900  
 FG2 910  
 FG2 920  
 FG2 930  
 FG2 940  
 FG2 950  
 FG2 960  
 FG2 970  
 FG2 980  
 FG2 990  
 FG2 1000  
 FG2 1010  
 FG2 1020  
 FG2 1030  
 FG2 1040  
 FG2 1050  
 FG2 1060  
 FG2 1070  
 FG2 1080

```

IND = 0
PL( 1. = OK( 1. 1)*RHO(1)*Y( 1)
PL( 1. = OK( 1. 2)*RHO(1)*Y( 14)
PL( 1. = OK( 1. 2)*RHO(1)*Y( 7)
PL( 2. = OK( 2. 1)*RHO(1)*Y( 10)
PL( 2. = OK( 2. 1)*RHO(1)*Y( 6)
PL( 2. = OK( 2. 2)*RHO(1)*Y( 9)
PL( 2. = OK( 2. 2)*RHO(1)*Y( 3)
PL( 3. = OK( 3. 1)*RHO(1)*X
PL( 3. = 2.*OK( 3. 2)*RHO(2)*Y( 12)*X
PL( 3. = OK( 3. 1)*RHO(1)*Y( 11)
PL( 3. = OK( 3. 2)*RHO(2)*Y( 12)*Y( 12)
PL( 4. = OK( 4. 1)*RHO(1)*X
PL( 4. = 2.*OK( 4. 2)*RHO(2)*Y( 10)*X
PL( 4. = OK( 4. 1)*RHO(1)*Y( 9)
PL( 4. = OK( 4. 2)*RHO(2)*Y( 10)*Y( 10)
PL( 5. = OK( 5. 1)*RHO(1)*X
PL( 5. = OK( 5. 2)*RHO(2)*Y( 16)*X
PL( 5. = OK( 5. 2)*RHO(2)*Y( 15)*X
PL( 5. = OK( 5. 1)*RHO(1)*Y( 14)
PL( 5. = OK( 5. 2)*RHO(2)*Y( 15)*Y( 16)
PL( 6. = OK( 6. 1)*RHO(1)*X
PL( 6. = OK( 6. 2)*RHO(2)*Y( 16)*X
PL( 6. = OK( 6. 2)*RHO(2)*Y( 13)*X
PL( 6. = OK( 6. 1)*RHO(1)*Y( 12)
PL( 6. = OK( 6. 2)*RHO(2)*Y( 13)*Y( 16)
PL( 7. = OK( 7. 1)*RHO(1)*X
PL( 7. = OK( 7. 2)*RHO(2)*Y( 12)*X
PL( 7. = OK( 7. 2)*RHO(2)*Y( 2)*X
PL( 7. = OK( 7. 1)*RHO(1)*Y( 7)
PL( 7. = OK( 7. 2)*RHO(2)*Y( 2)*Y( 12)
PL( 8. = OK( 8. 1)*RHO(1)*X
PL( 8. = OK( 8. 2)*RHO(2)*Y( 10)*X
PL( 8. = OK( 8. 2)*RHO(2)*Y( 7)*X
PL( 8. = OK( 8. 1)*RHO(1)*Y( 8)
PL( 8. = OK( 8. 2)*RHO(2)*Y( 7)*Y( 10)

```

```

FG2 1090
FG2 1100
FG2 1110
FG2 1120
FG2 1130
FG2 1140
FG2 1150
FG2 1160
FG2 1170
FG2 1180
FG2 1190
FG2 1200
FG2 1210
FG2 1220
FG2 1230
FG2 1240
FG2 1250
FG2 1260
FG2 1270
FG2 1280
FG2 1290
FG2 1300
FG2 1310
FG2 1320
FG2 1330
FG2 1340
FG2 1350
FG2 1360
FG2 1370
FG2 1380
FG2 1390
FG2 1400
FG2 1410
FG2 1420
FG2 1430
FG2 1440

```

PL( 9. 1) = OK( 9. 1)*RHO(1)*X	FG2 1450
PL( 9. 6) = OK( 9. 2)*RHO(2)*Y( 10)*X	FG2 1460
PL( 9. 7) = OK( 9. 2)*RHO(2)*Y( 3)*X	FG2 1470
PL( 9. 5) = OK( 9. 1)*RHO(1)*Y( 6)	FG2 1480
PL( 9. 10) = OK( 9. 2)*RHO(2)*Y( 3)*Y( 10)	FG2 1490
PL( 10. 1) = OK( 10. 1)*RHO(1)*X	FG2 1500
PL( 10. 6) = OK( 10. 2)*RHO(2)*Y( 10)*X	FG2 1510
PL( 10. 7) = OK( 10. 2)*RHO(2)*Y( 6)*X	FG2 1520
PL( 10. 5) = OK( 10. 1)*RHO(1)*Y( 17)	FG2 1530
PL( 10. 10) = OK( 10. 2)*RHO(2)*Y( 6)*Y( 10)	FG2 1540
PL( 11. 1) = OK( 11. 1)*RHO(1)*X	FG2 1550
PL( 11. 6) = OK( 11. 2)*RHO(2)*Y( 14)*X	FG2 1560
PL( 11. 7) = OK( 11. 2)*RHO(2)*Y( 2)*X	FG2 1570
PL( 11. 5) = OK( 11. 1)*RHO(1)*Y( 1)	FG2 1580
PL( 11. 10) = OK( 11. 2)*RHO(2)*Y( 2)*Y( 14)	FG2 1590
PL( 12. 1) = OK( 12. 1)*RHO(1)*X	FG2 1600
PL( 12. 6) = OK( 12. 2)*RHO(2)*Y( 2)*X	FG2 1610
PL( 12. 7) = OK( 12. 2)*RHO(2)*Y( 3)*X	FG2 1620
PL( 12. 5) = OK( 12. 1)*RHO(1)*Y( 4)	FG2 1630
PL( 12. 10) = OK( 12. 2)*RHO(2)*Y( 3)*Y( 2)	FG2 1640
PL( 13. 1) = OK( 13. 1)*RHO(1)*X	FG2 1650
PL( 13. 6) = 2.*OK( 13. 2)*RHO(2)*Y( 2)*X	FG2 1660
PL( 13. 5) = OK( 13. 1)*RHO(1)*Y( 3)	FG2 1670
PL( 13. 10) = OK( 13. 2)*RHO(2)*Y( 2)*Y( 2)	FG2 1680
PL( 14. 1) = OK( 14. 1)*RHO(1)*X	FG2 1690
PL( 14. 6) = OK( 14. 2)*RHO(2)*Y( 16)*X	FG2 1700
PL( 14. 7) = OK( 14. 2)*RHO(2)*Y( 5)*X	FG2 1710
PL( 14. 5) = OK( 14. 1)*RHO(1)*Y( 2)	FG2 1720
PL( 14. 10) = OK( 14. 2)*RHO(2)*Y( 5)*Y( 16)	FG2 1730
PL( 15. 1) = OK( 15. 1)*RHO(1)*X	FG2 1740
PL( 15. 6) = OK( 15. 2)*RHO(2)*Y( 2)*X	FG2 1750
PL( 15. 7) = OK( 15. 2)*RHO(2)*Y( 6)*X	FG2 1760
PL( 15. 5) = OK( 15. 1)*RHO(1)*Y(18)	FG2 1770
PL( 15. 10) = OK( 15. 2)*RHO(2)*Y( 6)*Y( 2)	FG2 1780
PL( 16. 1) = OK( 16. 1)*RHO(1)*X	FG2 1790
PL( 16. 6) = OK( 16. 2)*RHO(2)*Y( 2)*X	FG2 1800



```

PL( 16. 7) = OK( 16. 2)*RHO(2)*Y(18)*X
PL( 16. 5) = OK( 16. 1)*RHO(1)*Y(19)
PL( 16. 10) = OK( 16. 2)*RHO(2)*Y(18)*Y( 2)
GO TO (31,32,33,34,35,36,37,38,39,40,41,42,43,44,45,46,47,48,49).KF62 1840
31 AA = PL( 11. 10) - PL( 11. 5)
   AII = AA
   GO TO 90
32 BB= PL( 7. 6)+ PL( 7. 10)+ PL( 11. 6)+ PL( 11. 10)
   BB=BB+ PL( 12. 7)+ PL( 12. 10)+2*PL( 13. 6)+2*PL( 13. 10)
   AA= PL( 7. 5)+ PL( 11. 5)+ PL( 12. 5)+2*PL( 13. 5)
   AA=AA+ PL( 14. 10)+ PL(15. 5)+ PL(16. 5)
   BB=BB+ PL( 14. 1)+ PL( 14. 5)+ PL(15.10)+ PL(15. 7)
   BB=BB+ PL( 16.10)+ PL(16. 7)
   AII = AA - BB
   GO TO 90
33 AA= PL( 9. 5)+ PL( 12. 5)+ PL( 13. 10)
   BB= PL( 9. 10)+ PL( 12. 10)+ PL( 13. 5)
   AA = AA - BB
   AII = AA
   1 - PL( 13. 1)
   GO TO 90
34 AA= PL( 12. 10)
   BB= PL( 12. 5)
   AA = AA - BB
   AII = AA
   - PL( 12. 1)
   GO TO 90
35 BB= PL( 14. 10)
   AA= PL( 14. 5)
   AA = AA - BB
   AII = AA
   - PL( 14. 6)
   GO TO 90
36 AA= PL( 9. 10)+ PL( 10. 5)+ PL( 15. 5)
   BB= PL( 9. 5)+ PL( 10. 10)+ PL( 15. 10)+ PL( 15. 6)
   AA = AA - BB
   AII = AA
   - PL( 2. 1)- PL( 9. 1)- PL( 10. 6)
   GO TO 90

```

```

FG2 1810
FG2 1820
FG2 1830
KF62 1840
FG2 1850
FG2 1860
FG2 1870
FG2 1880
FG2 1890
FG2 1900
FG2 1910
FG2 1920
FG2 1930
FG2 1940
FG2 1950
FG2 1960
FG2 1970
FG2 1980
FG2 1990
FG2 2000
FG2 2010
FG2 2020
FG2 2030
FG2 2040
FG2 2050
FG2 2060
FG2 2070
FG2 2080
FG2 2090
FG2 2100
FG2 2110
FG2 2120
FG2 2130
FG2 2140
FG2 2150
FG2 2160

```



```

44 AA= PL( 5. 10)+ PL( 11. 5) FG2 2530
BB= PL( 5. 5)+ PL( 11. 10) FG2 2540
AA = AA - BB FG2 2550
AII = AA - PL( 1. 7) - PL( 5. 1) - PL( 11. 7) FG2 2560
GO TO 90 FG2 2570
45 AA= PL( 5. 5) FG2 2580
BB= PL( 5. 10) FG2 2590
AA = AA - BB FG2 2600
AII = AA - PL( 5. 6) FG2 2610
GO TO 90 FG2 2620
46 AA= PL( 5. 5)+ PL( 6. 5)+ PL( 14. 5) FG2 2630
BB= PL( 5. 10)+ PL( 6. 10)+ PL( 14. 10) FG2 2640
AA = AA - BB FG2 2650
AII = AA - PL( 5. 7) - PL( 6. 7) - PL( 14. 7) FG2 2660
GO TO 90 FG2 2670
47 AA = PL( 10. 10) - PL( 10. 5) FG2 2680
AII = AA - PL( 10. 1) FG2 2690
GO TO 90 FG2 2700
48 AA = PL(15.10) - PL(15. 5) + PL(16. 5) - PL(16.10) FG2 2710
AII = AA - PL(15. 1) - PL(16. 6) FG2 2720
GO TO 90 FG2 2730
49 AA = PL(16.10) - PL(16. 5) FG2 2740
AII = AA - PL(16. 1) FG2 2750
90 PL( 1. 1)=Y( 1)*PL( 1. 1) FG2 2760
PL( 1. 6)=Y( 7)*PL( 1. 6) FG2 2770
PL( 2. 1)=Y( 6)*PL( 2. 1) FG2 2780
PL( 2. 6)=Y( 3)*PL( 2. 6) FG2 2790
PL( 3. 1)=Y( 11)*PL( 3. 1) FG2 2800
PL( 3. 6)=Y( 12)*PL( 3. 6)/2. FG2 2810
PL( 4. 1)=Y( 9)*PL( 4. 1) FG2 2820
PL( 4. 6)=Y( 10)*PL( 4. 6)/2. FG2 2830
PL( 5. 1)=Y( 14)*PL( 5. 1) FG2 2840
PL( 5. 6)=Y( 15)*PL( 5. 6) FG2 2850
PL( 6. 1)=Y( 12)*PL( 6. 1) FG2 2860
PL( 6. 6)=Y( 13)*PL( 6. 6) FG2 2870
PL( 7. 1)=Y( 7)*PL( 7. 1) FG2 2880

```

PL( 7. 6)=Y( 2)\*PL( 7. 6)  
 PL( 8. 1)=Y( 8)\*PL( 8. 1)  
 PL( 8. 6)=Y( 7)\*PL( 8. 6)  
 PL( 9. 1)=Y( 6)\*PL( 9. 1)  
 PL( 9. 6)=Y( 3)\*PL( 9. 6)  
 PL( 10. 1)=Y( 17)\*PL( 10. 1)  
 PL( 10. 6)=Y( 6)\*PL( 10. 6)  
 PL( 11. 1)=Y( 1)\*PL( 11. 1)  
 PL( 11. 6)=Y( 2)\*PL( 11. 6)  
 PL( 12. 1)=Y( 4)\*PL( 12. 1)  
 PL( 12. 6)=Y( 3)\*PL( 12. 6)  
 PL( 13. 1)=Y( 3)\*PL( 13. 1)  
 PL( 13. 6)=Y( 2)\*PL( 13. 6)  
 PL( 14. 1)=Y( 2)\*PL( 14. 1)  
 PL( 14. 6)=Y( 5)\*PL( 14. 6)  
 PL( 15. 1)=Y( 18)\*PL( 15. 1)  
 PL( 15. 6)=Y( 6)\*PL( 15. 6)  
 PL( 16. 1)=Y( 19)\*PL( 16. 1)  
 PL( 16. 6)=Y( 18)\*PL( 16. 6)  
 191 GO TO (51,52,53,54,55,56,57,58,59,60,61,62,63,64,65,66,67,68,69),KF62 3080  
 51 AA= PL( 1. 6)+ PL( 11. 6)  
 BB= PL( 1. 1)+ PL( 11. 1)  
 B = AA - BB  
 GO TO 95  
 52 BB= PL( 7. 6)+ PL( 11. 6)+ PL( 12. 6)+2.\*PL( 13. 6)  
 AA= PL( 7. 1)+ PL( 11. 1)+ PL( 12. 1)+2.\*PL( 13. 1)  
 AA=AA+ PL( 14. 6)+ PL( 15. 1)+ PL( 16. 1)  
 BB=BB+ PL( 14. 1)+ PL( 15. 6)+ PL( 16. 6)  
 B = AA - BB  
 GO TO 95  
 53 AA= PL( 2. 1)+ PL( 9. 1)+ PL( 12. 1)+ PL( 13. 6)  
 BB= PL( 2. 6)+ PL( 9. 6)+ PL( 12. 6)+ PL( 13. 1)  
 B = AA - BB  
 GO TO 95  
 54 AA= PL( 12. 6)  
 BB= PL( 12. 1)

FG2 2890  
 FG2 2900  
 FG2 2910  
 FG2 2920  
 FG2 2930  
 FG2 2940  
 FG2 2950  
 FG2 2960  
 FG2 2970  
 FG2 2980  
 FG2 2990  
 FG2 3000  
 FG2 3010  
 FG2 3020  
 FG2 3030  
 FG2 3040  
 FG2 3050  
 FG2 3060  
 FG2 3070  
 FG2 3080  
 FG2 3090  
 FG2 3100  
 FG2 3110  
 FG2 3120  
 FG2 3130  
 FG2 3140  
 FG2 3150  
 FG2 3160  
 FG2 3170  
 FG2 3180  
 FG2 3190  
 FG2 3200  
 FG2 3210  
 FG2 3220  
 FG2 3230  
 FG2 3240



```

63 AA= PL( 6, 1)
BB= PL( 6, 6)
B = AA - BB
GO TO 95
64 AA= PL( 1, 1)+ PL( 5, 6)+ PL( 11, 1)
BB= PL( 1, 6)+ PL( 5, 1)+ PL( 11, 6)
B = AA - BB
GO TO 95
65 AA= PL( 5, 1)
BB= PL( 5, 6)
B = AA - BB
GO TO 95
66 AA= PL( 5, 1)+ PL( 6, 1)+ PL( 14, 1)
BB= PL( 5, 6)+ PL( 6, 6)+ PL( 14, 6)
B = AA - BB
GO TO 95
67 AA=PL(10,6)
BB=PL(10,1)
B = AA - BB
GO TO 95
68 AA= PL(15, 6) + PL(16,1)
BB= PL(15, 1) + PL(16, 6)
B = AA - BB
GO TO 95
69 AA= PL(16, 6)
BB= PL(16, 1)
B = AA - BB
95 IF(IND)110,101,110
101 IND = 1
GAMMA = 0.
DO102 I=1,16
IF(1.67.2)GAMMA = 1.
Q = (RATES(2,1) + RATES(3,1))/T)/T
PL(1,1) = Q*PL(1,1)
Q = GAMMA/RT + Q - E(1,1.L)/T1(1) - E(1,2.L)/2. - E(1,3.L)*T1(1)/3F62 3950
X. - E(1,4.L)*T1(2)/4. - E(1,5.L)*T1(3)/5. - E(1,6.L)/T1(2) F62 3960

```

```

F62 3610
F62 3620
F62 3630
F62 3640
F62 3650
F62 3660
F62 3670
F62 3680
F62 3690
F62 3700
F62 3710
F62 3720
F62 3730
F62 3740
F62 3750
F62 3760
F62 3770
F62 3780
F62 3790
F62 3800
F62 3810
F62 3820
F62 3830
F62 3840
F62 3850
F62 3860
F62 3870
F62 3880
F62 3890
F62 3900
F62 3910
F62 3920
F62 3930
F62 3940
F62 3950
F62 3960

```

FG2 3970  
FG2 3980  
FG2 3990  
FG2 4000  
FG2 4010

102 PL(I,6) = Q\*PL(I,6)  
C = B  
GO TO 191  
110 RETURN  
END

```

SUBROUTINE LRAD
9999
** THIS IS A DRIVER PROGRAM FOR SUBROUTINE TRANS WHICH CALCULATES **
THE RADIATIVE FLUX DIVERGENCE THROUGH A ONE-DIMENSIONAL SLAB
FOR A GIVEN TEMPERATURE AND SPECIES DISTRIBUTION
COMMON /SFLUX/ ORI(3)
COMMON /TRN/ NUT(60), FMC(12,60), FPC(12,60),
FM(9,60), FP(9,60), LINES
COMMON /TEST/ETZ(60), IEZ
COMMON /VL/ETA(60), YOND(60)
COMMON /PROPI/PI(60), RHO(60), T(60), AMW(60), C(20,60), EC(5,60)
COMMON /SIS/SY(20,60)
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON /DEL/ DELTA, DTIL, DTILS
COMMON /NON/RDZ, MUDZ, RMDZ, AKNF, HNF, CPNF
COMMON /MAIN/KEEP, MAXE, MAXM, MAXD, IDG, MCONV, ECONV, DCONV, LT, IAB
COMMON /RFLUX/ E(60), IRAD, ITYPE
COMMON /NUMDEN/ SNDO2(60), SNDN2(60), SNDO(60), SNDN(60),
SNDH(60), SNDC2(60), SNDE(60), SNDC(60),
SNDC3(60), SNDC2H(60)
COMMON /SPEC/ MF, XMOL
COMMON /TOM/ SC(60), TD(60), PR(60), TE(60)
DO100 I=1, 20
DO100 J=1, NETA
100 SY(I,J) = C(I,J)
DO200 I=1, 20
DO200 J=1, NETA
200 IF(C(I,J).LT.0.)C(I,J) = 1.E-20
** NETA = NUMBER OF ETA POINTS
MF = 1 IF SPECIE MOLE FRACTIONS ARE INPUT AND NUMBER DENSITY
TO BE COMPUTED
0 IF SPECIE NUMBER DENSITIES ARE INPUT

```

LRAD 10  
LRAD 20  
LRAD 30  
LRAD 40  
LRAD 50  
LRAD 60  
LRAD 70  
LRAD 80  
LRAD 90  
LRAD 100  
LRAD 110  
LRAD 120  
LRAD 130  
LRAD 140  
LRAD 150  
LRAD 160  
LRAD 170  
LRAD 180  
LRAD 190  
LRAD 200  
LRAD 210  
LRAD 220  
LRAD 230  
LRAD 240  
LRAD 250  
LRAD 260  
LRAD 270  
LRAD 280  
LRAD 290  
LRAD 300  
LRAD 310  
LRAD 320  
LRAD 330  
LRAD 340  
LRAD 350  
LRAD 360



```

C      NS = NUMBER OF SPECIES TO BE INPUT
C      LINES= 1 IF LINE CALCULATION IS TO BE DONE
C      0 IF ONLY CONTINUUM CALCULATION IS TO BE DONE
C      IDG = 0 ONLY FINAL PRINT IS GIVEN
C      1 PRINT IS GIVEN FOR EACH ETA
C      99 COMPLETE PRINT
C
C      ** R = BODY RADIUS (FT)
C      DELTA = NONDIMENSIONAL STAND-OFF DISTANCE
C      DTIL = TRANSFORMED STAND-OFF DISTANCE
C      XMOL = 1.0 FOR RUN WITH MOLECULES
C      0.0 FOR RUN WITHOUT MOLECULES
C
C      XMOL = 0.0
C      XMOL= 1.0
C
C      ** DETERMINE ETZ POINTS
C      N2 = META-2
C      K = 0
C      IEZ = 0
C      DO 20 I=1,N2.2
C      K=K+1
C      ETZ(K) = ETA(I)
C      ETZ(K+1) = ETA(META-1)
C      ETZ(K+2) = ETA(META)
C      IEZ = K +1
C
C      ** COMPUTE THE Y COORDINATE **
C      YOND(1) = 0.0
C      SUM = 0.0
C      DO 30 K=2,META
C      DELTA = ETA(K) -ETA(K-1)
C      SUM = SUM +DELTA*(1./RHO(K) +1./RHO(K-1) )/2.0
C      YOND(K) = DTIL +SUM
C      CONTINUE
C
C      20
C      30

```

LRAD 730  
 LRAD 740  
 LRAD 750  
 LRAD 760  
 LRAD 770  
 LRAD 780  
 LRAD 790  
 LRAD 800  
 LRAD 810  
 LRAD 820  
 LRAD 830  
 LRAD 840  
 LRAD 850  
 LRAD 860  
 LRAD 870  
 LRAD 880  
 LRAD 890  
 LRAD 900  
 LRAD 910  
 LRAD 920  
 LRAD 930  
 LRAD 940  
 LRAD 950  
 LRAD 960  
 LRAD 970  
 LRAD 980  
 LRAD 990  
 LRAD1000  
 LRAD1010  
 LRAD1020  
 LRAD1030  
 LRAD1040  
 LRAD1050

DELTA = YOND(NETA)  
 DO 40 K=1,NETA  
 YOND(K) = YOND(K)/YOND(NETA)  
 T(K) = 0.65 \* T(K) + 0.35 \* TE(K)  
 40 CONTINUE

LINES= 1  
 IDGS = IDG  
 IDG = 0  
 CALL TRANS(1)  
 CALL TRANS(1)  
 IDG= IDGS

\*\* INDEX IS NUMBER GIVEN SPECIE FOR USE IN STORING ARRAYS \*\*

1	=	O2
2	=	N2
3	=	O
4	=	N
5	=	E-
6	=	C
7	=	H
8	=	C2
9	=	H2
10	=	CO
11	=	C3
12	=	C2H

DO300I=1,20  
 DO300J=1,NETA  
 300 C(I,J) = SY(I,J)  
 RETURN  
 END

10  
 20  
 30  
 40  
 50  
 60  
 70  
 80  
 90  
 100  
 110  
 120  
 130  
 140  
 150  
 160  
 170  
 180  
 190  
 200  
 210  
 220  
 230  
 240  
 250  
 260  
 270  
 280  
 290  
 300  
 310  
 320  
 330  
 340  
 350  
 360

```

SUBROUTINE RADIN
COMMON /DBG/ QLC(60), QCL(60), QLL(60), DON(60), OCC(60),
BEEC(12,60), FMUC(12,60), EM(12,60),
EP(12,60), TAUC(12,60), BEEL(9,60),
OCCP(12), WMM(9,60), GNM(9,60),
EEM(9,60), XLMM(9,60), QLCP(9),
OCLP(9), OLLP(9), DELTA, IY, IYY,
WPP(9,60), GPP(9,60), EEP(9,60),
XLPP(9,60), FG(9,4), GP(9,4),
WN(9,4), FMUL(9,60), SSM(9,4,60),
GGM(9,4,60), ETAM(9,4,60), SBM(9,4,60),
TAUL(9,60)

C ** GROUP 1 **
WN(1,1)=0.
FG(1,1)=0.
GP(1,1)=0.
WN(1,2)=18.
WN(1,3)=15.
WN(1,4)=5.

C ** GROUP 2 **
WN(2,1)=3.0
WN(2,2)=5.0
WN(2,3)=11.0
WN(2,4)=10.

C ** GROUP 3 **
WN(3,1)=0.
FG(3,1)=0.
GP(3,1)=0.
WN(3,2)=2.0
WN(3,3)=0.
FG(3,3)=0.
GP(3,3)=0.
WN(3,4)=0.
FG(3,4)=0.
GP(3,4)=0.

C ** GROUP 4 **

```

RADI 370  
 RADI 380  
 RADI 390  
 RADI 400  
 RADI 410  
 RADI 420  
 RADI 430  
 RADI 440  
 RADI 450  
 RADI 460  
 RADI 470  
 RADI 480  
 RADI 490  
 RADI 500  
 RADI 510  
 RADI 520  
 RADI 530  
 RADI 540  
 RADI 550  
 RADI 560  
 RADI 570  
 RADI 580  
 RADI 590  
 RADI 600  
 RADI 610  
 RADI 620  
 RADI 630  
 RADI 640  
 RADI 650  
 RADI 660  
 RADI 670  
 RADI 680  
 RADI 690  
 RADI 700  
 RADI 710  
 RADI 720

WN(4.1)=0.  
 FG(4.1)=0.  
 GP(4.1)=0.  
 WN(4.2)=8.0  
 WN(4.3)=2.0  
 WN(4.4)=0.  
 FG(4.4)=0.  
 GP(4.4)=0.  
 C \*\* GROUP 5 \*\*  
 WN(5.1)=0.  
 FG(5.1)=0.  
 GP(5.1)=0.  
 WN(5.2)=14.  
 WN(5.3)=4.0  
 WN(5.4)=1.0  
 C \*\* GROUP 6 \*\*  
 WN(6.1)=1.0  
 WN(6.2)=4.0  
 WN(6.3)=13.0  
 WN(6.4)=2.0  
 C \*\* GROUP 7 \*\*  
 WN(7.1)=0.  
 FG(7.1)=0.  
 GP(7.1)=0.  
 WN(7.2)=6.0  
 WN(7.3)=14.0  
 WN(7.4)=3.0  
 C \*\* GROUP 8 \*\*  
 WN(8.1)=2.0  
 WN(8.2)=2.0  
 WN(8.3)=11.  
 WN(8.4)=15.  
 C \*\* GROUP 9 \*\*  
 WN(9.1)=0.  
 FG(9.1)=0.  
 GP(9.1)=0.

RADI 730  
RADI 740  
RADI 750  
RADI 760  
RADI 770

WN(9.2)=1.0  
WN(9.3)=11.  
WN(9.4)=10.  
RETURN  
END

## SUBROUTINE TRANS (ISW)

```

C-----THIS IS A MODIFIED VERSION OF SUBROUTINE TRANS FROM K WILSON
C TRANS IS DOCUMENTED IN LMSC-687209 APRIL 69 -----
C
COMMON /ZPI/ ZPO(6), ZPN(6),ZPH(2), ZPC(7)
COMMON /FINV/ NHVL,NHVC,FHVC(12),DJ(9),HVJ(9),ZKZ
COMMON /SFLUX/ ORI(3)
COMMON /TRN/ NUT(60), FMC(12,60), FPC(12,60),
1 FM(9,60), FP(9,60), LINES
COMMON /YL/ETA(60), YD(60)
COMMON/PROPI/P (60), R(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /FRSTRM/ U INF, RINF, UINF2,XL , RE, LXI, ITM, IEM, NES
COMMON /DEL/ W(1),DTIL,DTILS
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,MNF,CPNF
COMMON /NAIN/KEEP,MAXE,MAXM,MAXD, IDG,MCONV,ECONV,DCONV,LT,IAB
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /RM/ DUJ,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON/WT/SMW(20),AWT(5)
COMMON /TEST/ETZ(60),IEZ
COMMON /NUMDEN/ SNO2(60), SNO2H(60), SNO2(60), SNO2(60), SNO2(60),
1 SNO2H(60), SNO2C(20), SNO2H(60)
2 SNO2C(60),SNO2C2H(60)
3 COMMON /DEBUG/ QLC(60), QCL(60), QLL(60), DON(60), OCC(60),
1 BEEC(12,60), FMUC(12,60), EM(12,60),
2 EP(12,60), TAU(12,60), BEEL(9,60),
3 OCCP(12), WMW(9,60), GMW(9,60),
4 EEM(9,60), XLWM(9,60), QLCP(9),
5 QCLP(9), DELTA, IY, IYY,
6 WPP(9,60), GPP(9,60), EEP(9,60),
7 XLPP(9,60), FG(9,4), GP(9,4),
8 WN(9,4), FMUL(9,60), SSM(9,4,60),
9 GGM(9,4,60), ETAM(9,4,60), SBM(9,4,60),
A TAU(9,60)
COMMON /SPEC/ MF, XMDL

```

TRAN 10  
 TRAN 20  
 TRAN 30  
 TRAN 40  
 TRAN 50  
 TRAN 60  
 TRAN 70  
 TRAN 80  
 TRAN 90  
 TRAN 100  
 TRAN 110  
 TRAN 120  
 TRAN 130  
 TRAN 140  
 TRAN 150  
 TRAN 160  
 TRAN 170  
 TRAN 180  
 TRAN 190  
 TRAN 200  
 TRAN 210  
 TRAN 220  
 TRAN 230  
 TRAN 240  
 TRAN 250  
 TRAN 260  
 TRAN 270  
 TRAN 280  
 TRAN 290  
 TRAN 300  
 TRAN 310  
 TRAN 320  
 TRAN 330  
 TRAN 340  
 TRAN 350  
 TRAN 360

```

COMMON/MIKE/ NETA1,NETA2,PNF,WNF,RHOD,DP,TS,UD,VD,RW,RWJ,EPS,PHI
DIMENSION XKT(60), DQ(60)

C ** BAND AVERAGE ABSORPTION CROSS SECTION (EQ.A2) **
C
C
      SIGMA(ZH,ZA,ZB,ZG)= ((5.0E+03*TI*ZG*ZKZ)/BE) * (EXP(ZDL/TI)
1      *ZH*(ZA+ZB*(ZH**2)/3.0) +
2      *TI * (ZA+2.0*ZB*TI2) -TI*EXP((ZH-ZHVP)/TI))
3      *(ZA+ZB*(ZHVP-ZH)**2) -TI*EXP((ZH-ZHVP)/TI))
4      *2.0*ZB*TI*(ZHVP-ZH+TI))
      SIGMA2(ZH,ZG,ZE,ZY)=7.26E-16*TI*ZG*EXP((-ZE+ZY+ZDL)/TI)/ZH**3
      GAMMA(ZX)=(1.0+(1.5707963*ZX)**1.25)**(-0.4)
      XLAMB(ZX)=(1.0+ZX*EXP(-ZX))/SORT(1.0+6.283185 *ZX)

C ** W(GROUP)/D CORRELATION (EQ.88) **
C
C
      PHI1(ZX)=(ATAN(1.570796 *ZX)/1.570796 )

C ** FLUX DIVERGENCE OVERLAPPING FUNCTION (EQ.92) **
C
C
      PHI2(ZX)=EXP(-ZX)

C
      DO 400 I=1,NES
      T(I)=T(I)*TS
      ZHVP=5.0
      YI=0.0
      CONVER = 3.10375E+23 *R (I) *RDZ
      SNDE(NES) = CONVER * C( 7,NES)/SMW(7)
      XNE=SNDE(NES)
      FNE=(4.71E-6 * XNE**((2.0/7.0))/((T(NES)/11606.))**((1.0/7.0)))
      ZDL=AMIN1(0.20,FNE)

C
C ** DEBUG PRINT **
C
      IF (IDG.NE.0) CALL BUGPR (I)
      DELTA=W(I) * XL * 30.48006

```

```

TRAN 370
TRAN 380
TRAN 390
TRAN 400
TRAN 410
TRAN 420
TRAN 430
TRAN 440
TRAN 450
TRAN 460
TRAN 470
TRAN 480
TRAN 490
TRAN 500
TRAN 510
TRAN 520
TRAN 530
TRAN 540
TRAN 550
TRAN 560
TRAN 570
TRAN 580
TRAN 590
TRAN 600
TRAN 610
TRAN 620
TRAN 630
TRAN 640
TRAN 650
TRAN 660
TRAN 670
TRAN 680
TRAN 690
TRAN 700
TRAN 710
TRAN 720

```

```

        CALL BUGPR (2)
6001 CONTINUE
      DO 91 L=1,NES
        XKT(L)=Y(L)/11606.
        T1=XKT(L)
        CALL SND(L)

C ** PARTITION FUNCTIONS FOR H. C. N. O **
C
C 94 IF(T(L).GT.15000.) GO TO 6
C
C ** LOW TEMPERATURE **
C
      SUMH=2.0
      SUMC=9.0 + 5.0 * EXP(-1.264/T1) + EXP(-2.684/T1) +
        5.0 * EXP(-4.183/T1)
1      SUMN=4.0 + 10.0 * EXP(-2.384/T1) + 6.0 * EXP(-3.576/T1)
      SUMO= 9.0 + 5.0 * EXP(-1.975/T1)
      GO TO 7

C ** HIGH TEMPERATURE **
C
C 6 SUMH=2.0
      SUMC=2.71818 + 6.40677 * Y(L)/1.0E4 - 0.45466 * (T(L)/1.0E4)**2
      SUMN=5.938216 - 0.225593 * Y(L)/1.0E3 + 0.015408 * (T(L)/1.0E3)**2
      SUMO=11.79563 - 0.317964 * Y(L)/1.0E3 + 0.013765 * (T(L)/1.0E3)**2
7 CONTINUE
      T12=T1**2
      GH = 6.4994
      DO 5 K=1.12
        GF=FHVC(K)/T1
        GH=GH
        GH=EXP(-GF) * GF * (GF**2 + 3.0 * GF + 6.0 + 6.0/GF)

C ** PLANK MEAN ABSORPTION COEFFICIENT FOR BAND INTERVALS (EQ.A3) **
C
C

```

```

TRAN 730
TRAN 740
TRAN 750
TRAN 760
TRAN 770
TRAN 780
TRAN 790
TRAN 800
TRAN 810
TRAN 820
TRAN 830
TRAN 840
TRAN 850
TRAN 860
TRAN 870
TRAN 880
TRAN 890
TRAN 900
TRAN 910
TRAN 920
TRAN 930
TRAN 940
TRAN 950
TRAN 960
TRAN 970
TRAN 980
TRAN 990
TRAN1000
TRAN1010
TRAN1020
TRAN1030
TRAN1040
TRAN1050
TRAN1060
TRAN1070
TRAN1080

```



```

BEEC(K,L)=5.04E3 * (T12**2) * (GHM-GH)
BE=BEEC(K,L)

C ** ABSORPTION CROSS SECTIONS **
C SOECIES -- N N2 O2 CO
C C C C C2H C3
C
C SGM=0.
C SGN=0.
C SGC=0.
C SGO=0.
C SGC0=0.
C SGC2=0.
C SGO2=0.
C SGN2=0.
C SGM2=0.
C SGC3=0.
C SGC2H = 0.0
GO TO (581,582,583,584,585,586,587,588,589,590,591,592) * K
581 SGM=SIGMA(2.4,1.0,0.0,1.0) * EXP(-13.56/T1)
SGC=SIGMA(3.78, 0.3, 0.0488, 1.33) * EXP(-11.26/T1)
SGN=SIGMA(4.22, 0.24, 0.0426, 4.5) * EXP(-14.54/T1)
SGO=SIGMA(4.22, 0.24, 0.0426, 2888889) * EXP(-13.61/T1)
GO TO 38
582 ZZHV=5.5
SGC2=8.0E-18 * EXP(-0.5/T1) + 3.0E-18
SGC3=4.0E-18
583 CALL ZHV(ZZHV,ZZO,ZZN,ZZI,ZZC)
SGC=SIGMA2(ZZHV, 1.33, 11.26, 3.78) * ZZC + SGC
SGN=SIGMA2(ZZHV, 4.50, 14.54, 4.22) * ZZN
SGO=SIGMA2(ZZHV, .889, 13.61, 4.22) * ZZO
SGM=SIGMA2(ZZHV, 1.00, 13.56, 2.40)
GO TO 38
TRAN1090
TRAN1100
TRAN1110
TRAN1120
TRAN1130
TRAN1140
TRAN1150
TRAN1160
TRAN1170
TRAN1180
TRAN1190
TRAN1200
TRAN1210
TRAN1220
TRAN1230
TRAN1240
TRAN1250
TRAN1260
TRAN1270
TRAN1280
TRAN1290
TRAN1300
TRAN1310
TRAN1320
TRAN1330
TRAN1340
TRAN1350
TRAN1360
TRAN1370
TRAN1380
TRAN1390
TRAN1400
TRAN1410
TRAN1420
TRAN1430
TRAN1440

```

```

583 ZZHV=6.5
   SGC2=1.0E-18
   SGC0=3.0E-18 * EXP(-0.7/T1)
   GO TO 593

584 ZZHV=7.5
   SGC=5.0E-17 * EXP(-4.18/T1)/SUMC
   SGC0=1.9E-17 * EXP(-0.5/T1)
   SGC2=6.0E-19
   SGC2H = 1.3E-18
   GO TO 593

585 ZZHV=8.5
   SGC=5.0E-17 *EXP(-4.18/T1)/SUMC +
   1 2.2E-17* EXP(-2.68/T1)/SUMC
   SGC0=2.5E-17
   SGC2=2.0E-19
   SGC2H = 8.5E-19
   GO TO 593

586 ZZHV=9.5
   SGC=5.0E-17 * EXP(-4.18/T1)/SUMC +
   1 2.2E-17 * EXP(-2.68/T1)/SUMC
   SGC0=5.0E-18
   SGC2=1.0E-18
   GO TO 593

587 SGN=3.2E-18 *T1 *EXP(-10.2/T1)/SUMN
   SGC2=6.0E-19
   ZZHV=10.4
   CALL ZHV(ZZHV,ZZ0,ZZN,ZZ1,ZZC)
596 SGC=(8.5E-17 *EXP(-1.26/T1) + 2.2E-17 * EXP(-2.75/T1)
   1 + 5.0E-17 * EXP(-4.18/T1))/SUMC
   GO TO 594

588 ZZHV=10.9
   CALL ZHV(ZZHV,ZZ0,ZZN,ZZ1,ZZC)
   SGN=(5.16E-17 *EXP(-3.50/T1))/SUMN
   GO TO 596

589 ZZHV=11.6
   CALL ZHV(ZZHV,ZZ0,ZZN,ZZ1,ZZC)

```

```

TRAN1450
TRAN1460
TRAN1470
TRAN1480
TRAN1490
TRAN1500
TRAN1510
TRAN1520
TRAN1530
TRAN1540
TRAN1550
TRAN1560
TRAN1570
TRAN1580
TRAN1590
TRAN1600
TRAN1610
TRAN1620
TRAN1630
TRAN1640
TRAN1650
TRAN1660
TRAN1670
TRAN1680
TRAN1690
TRAN1700
TRAN1710
TRAN1720
TRAN1730
TRAN1740
TRAN1750
TRAN1760
TRAN1770
TRAN1780
TRAN1790
TRAN1800

```

```

TRAN1810
TRAN1820
TRAN1830
TRAN1840
TRAN1850
TRAN1860
TRAN1870
TRAN1880
TRAN1890
TRAN1900
TRAN1910
TRAN1920
TRAN1930
TRAN1940
TRAN1950
TRAN1960
TRAN1970
TRAN1980
TRAN1990
TRAN2000
TRAN2010
TRAN2020
TRAN2030
TRAN2040
TRAN2050
TRAN2060
TRAN2070
TRAN2080
TRAN2090
TRAN2100
TRAN2110
TRAN2120
TRAN2130
TRAN2140
TRAN2150
TRAN2160

SGN2=1.0E-18
SGN=(5.16E-17 * EXP(-3.50))/SUMN
598 SGC=(9.9E-17 * 8.5E-17 * EXP(-1.26/T1) +2.2E-17 * EXP(-2.75/T1))
1 * 5.0E-17 * EXP(-4.18/T1))/SUMC
IF (K.LT.11) GO TO 594
GO TO 38
590 ZZW=12.7
CALL ZHV (ZZHV.ZZO.ZZN.ZZ1.ZZC)
SGN2=2.0E-18
SGH2 = 2.7E-17
599 SGN=(6.4E-17 * EXP(-2.30/T1) + 5.16E-17 * EXP(-3.50/T1))/SUMN
1 * SGN
GO TO 598
591 SGH=1.18E-17/SUMH
SGO=3.6E-17/SUMO
SGN2=1.0E-17
SGH2 = 2.7E-17
GO TO 599
592 SGN=3.6E-17/SUMN
SGN2=1.0E-18
GO TO 599
38 CONTINUE
FMUC(K,L)= SMDH(L)*SGH + SDC(L)*SGC + SMDN(L)*SGN + SDD(L)*SGO
+ XMDL * (SMDN2(L)*SGN2 + SDD2(L)*SGO2 +
1 SDC2(L)*SGC2 + SDD2(L)*SGH2 + SDDC(L)*SGCO +
2 SDDC3(L)*SGC3 +SDDC2H(L)*SGC2H )
3 IF (L.GT.1) GO TO 8
TAUC(K,L)=0.
GO TO 5
8 TAUC(K,L)=TAUC(K,L-1)+(YD(L)-YD(L-1))*
1 (FMUC(K,L-1)+FMUC(K,L)) * DELTA
5 CONTINUE
IF (LINES.EQ.0) GO TO 91
C ** FRACTIONAL POPULATION STATES FOR H. N. O. C **
C
C

```

```

C ** CALL ZP (T1,SUMN,SUMO,SUMH,SUMC)
C ** CALCULATION OF PARAMETERS FOR 9 LINE GROUPES **
C   WN -- NUMBER OF LINES
C   FG -- EFFECTIVE F-NUMBER
C   GP -- EFFECTIVE HALF-WIDTH
C
C GROUP 1
  FG(1,2)=(1.02 * ZPC(5) + .795 * ZPC(6) + 0.114 * ZPC(7))
  1 /WN(1,2)
  GP(1,2)=(8.16E-11 * SORT(ZPC(5)) + 1.25E-10 * SORT(ZPC(6))
  1 +2.55E-10 * SORT(ZPC(7)))**2 / (FG(1,2) * WN(1,2)**2)
  FG(1,3)=(1.040 * ZPN(4) + 1.29 * ZPN(5) + 0.00 * ZPN(6))
  1 /WN(1,3)
  GP(1,3)=(6.65E-11 * SORT(ZPN(4)) + 1.71E-10 * SORT(ZPN(5))
  1 + 0.20E-10 * SORT(ZPN(6)))**2 / (FG(1,3) * WN(1,3)**2)
  FG(1,4)=(1.00 * ZPO(5) + .978 * ZPO(6)) / WN(1,4)
  GP(1,4)=(3.90E-11 * SORT(ZPO(5)) + 9.68E-11 * SORT(ZPO(6)))**2
  1 / (FG(1,4) * WN(1,4)**2)
  FMUL(1,L)=FMUC(1,L)
C GROUP 2
  FG(2,1)=0.805 * ZPH(2) / WN(2,1)
  GP(2,1)=2.37E-10 * 2.37E-10 * ZPH(2) / (FG(2,1) * WN(2,1)**2)
  FG(2,2)=(0.00E-2 * ZPC(5) + 6.71E-2 * ZPC(6)) / WN(2,2)
  GP(2,2)=(0.00E-12 * SORT(ZPC(5)) + 7.15E-11 * SORT(ZPC(6)))**2
  1 / (FG(2,2) * WN(2,2)**2)
  FG(2,3)=(0.047 * ZPN(4) + 2.85E-2 * ZPN(5)) / WN(2,3)
  GP(2,3)=(1.11E-10 * SORT(ZPN(4)) + 6.07E-11 * SORT(ZPN(5)))**2
  1 / (FG(2,3) * WN(2,3)**2)
  FG(2,4)=(.6217 * ZPO(4) + 8.25E-2 * ZPO(5)) / WN(2,4)
  GP(2,4)=(2.61E-11 * SORT(ZPO(4)) + 7.19E-11 * SORT(ZPO(5)))**2
  1 / (FG(2,4) * WN(2,4)**2)
  FMUL(2,L)=FMUC(1,L)
C GROUP 3
  FG(3,2)=(7.29E-2 * ZPC(2) + 6.76E-2 * ZPC(3)) / WN(3,2)
  GP(3,2)=(9.08E-12 * SORT(ZPC(2)) + 8.75E-12 * SORT(ZPC(3)))**2
  1 / (FG(3,2) * WN(3,2)**2)

```

TRAN2170  
 TRAN2180  
 TRAN2190  
 TRAN2200  
 TRAN2210  
 TRAN2220  
 TRAN2230  
 TRAN2240  
 TRAN2250  
 TRAN2260  
 TRAN2270  
 TRAN2280  
 TRAN2290  
 TRAN2300  
 TRAN2310  
 TRAN2320  
 TRAN2330  
 TRAN2340  
 TRAN2350  
 TRAN2360  
 TRAN2370  
 TRAN2380  
 TRAN2390  
 TRAN2400  
 TRAN2410  
 TRAN2420  
 TRAN2430  
 TRAN2440  
 TRAN2450  
 TRAN2460  
 TRAN2470  
 TRAN2480  
 TRAN2490  
 TRAN2500  
 TRAN2510  
 TRAN2520

```

FMUL(3,L)=FMUC(2,L)
C GROUP 4
  FG(4,2)=(1.05 * ZPC(1) + 1.10E-2 * ZPC(2) + 0.150 * ZPC(3))
  1 /WN(4,2)
  GP(4,2)=(9.57E-12 * SORT(ZPC(1)) + 4.86E-12 * SORT(ZPC(2))
  1 + 5.93E-10 * SORT(ZPC(3)))**2/(FG(4,2) * WN(4,2))**2
  FG(4,3)=( 7.40E-2 * ZPN(2) + 6.34E-2 * ZPN(3))/WN(4,3)
  GP(4,3)=(8.22E-12 * SORT(ZPN(2)) + 7.60E-12 * SORT(ZPN(3)))**2
  1 /(FG(4,3) * WN(4,3))**2
  FMUL(4,L)=FMUC(4,L)
C GROUP 5
  FG(5,2)=(0.329 * ZPC(1) + 0.118 * ZPC(2) + 0.226 * ZPC(4))
  1 /WN(5,2)
  GP(5,2)=(3.65E-11 * SORT(ZPC(1)) + 5.77E-10 * SORT(ZPC(2))
  1 + 6.56E-11 * SORT(ZPC(4)))**2/(FG(5,2) * WN(5,2))**2
  FG(5,3)=0.108 * ZPN(3)/WN(5,3)
  GP(5,3)=3.09E-11 * 3.09E-11 * ZPN(3)/(FG(5,3) * WN(5,3))**2
  FG(5,4)=4.71E-2 * ZPO(1)/WN(5,4)
  GP(5,4)=5.08E-12 * 5.08E-12 * ZPO(1)/(FG(5,4) * WN(5,4))**2
  FMUL(5,L)=FMUC(6,L)
C GROUP 6
  FG(6,1)=0.416 * ZPH(1)/WN(6,1)
  GP(6,1)=3.02E-11 * 3.02E-11 * ZPH(1)/(FG(6,1) * WN(6,1))**2
  FG(6,2)=8.65E-2 * ZPC(1)/WN(6,2)
  GP(6,2)=2.35E-10 * 2.35E-10 * ZPC(1)/(FG(6,2) * WN(6,2))**2
  FG(6,3)=(0.184 * ZPN(1) + 0.290 * ZPN(2) + 8.52E-2 * ZPN(3))
  1 /WN(6,3)
  GP(6,3)=(1.07E-11 * SORT(ZPN(1)) + 4.28E-11 * SORT(ZPN(2))
  1 + 2.09E-10 * SORT(ZPN(3)))**2/(FG(6,3) * WN(6,3))**2
  FG(6,4)=(.120 * ZPO(2) + 0.151 * ZPO(3))/WN(6,4)
  GP(6,4)=(8.85E-12 * SORT(ZPO(2)) + 9.93E-12 * SORT(ZPO(3)))**2
  1 /(FG(6,4) * WN(6,4))**2
  FMUL(6,L)=FMUC(7,L)
C GROUP 7
  FG(7,2)=(4.51E-2 * ZPC(1) + 0.705 * ZPC(2))/WN(7,2)
  GP(7,2)=(6.07E-10 * SORT(ZPC(1)) + 2.10E-10 * SORT(ZPC(2)))**2

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TRAN2530  
 TRAN2540  
 TRAN2550  
 TRAN2560  
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 TRAN2580  
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 TRAN2790  
 TRAN2800  
 TRAN2810  
 TRAN2820  
 TRAN2830  
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 TRAN2860  
 TRAN2870  
 TRAN2880

```

1      /((FG(7.2) * WN(7.2)**2)
FG(7.3)=(0.454 * ZPN(1) + 9.66E-2 * ZPN(2)
+ 0.178 * ZPN(3))/WN(7.3)
1      GP(7.3)=(2.71E-12 * SORT(ZPN(1)) + 2.34E-10 * SORT(ZPN(2))
+ 2.46E-11 * SORT(ZPN(3))**2/(FG(7.3) * WN(7.3)**2)
1      FG(7.4)=4.23E-2 * ZPO(3)/WN(7.4)
GP(7.4)=2.52E-11 * 2.52E-11 * ZPO(3)/(FG(7.4) * WN(7.4)**2)
FMUL(7.L)=FMUC(9.L)

C GROUP 8
FG(8.1)=0.108 * ZPH(1)/WN(8.1)
GP(8.1)=1.32E-10 + 1.32E-10 * ZPH(1)
/((FG(8.1) * WN(8.1)**2)
1      FG(8.2)=(0.379 * ZPC(1) + 1.05 * ZPC(3))/WN(8.2)
GP(8.2)=(1.95E-11 * SORT(ZPC(1)) + 1.27E-10 * SORT(ZPC(3))**2
/((FG(8.2) * WN(8.2)**2)
1      FG(8.3)=(0.155 * ZPN(1) + 0.142*ZPN(2) + 3.75E-2 * ZPN(3))
/WN(8.3)
1      GP(8.3)=(2.98E-11 * SORT(ZPN(1)) + 7.08E-11 * SORT(ZPN(2))
+ 1.33E-10 * SORT(ZPN(3))**2/(FG(8.3) * WN(8.3)**2)
1      FG(8.4)=(0.146 * ZPO(1) + 8.61E-2*ZPO(2)
+ 9.33E-2 * ZPO(3))/WN(8.4)
1      GP(8.4)=(1.97E-10 * SORT(ZPO(1)) + 1.80E-11 * SORT(ZPO(2))
+ 8.13E-11 * SORT(ZPO(3))**2/(FG(8.4) * WN(8.4)**2)
1      FMUL(8.L)=FMUC(10.L)

C GROUP 9
FG(9.2)=2.95 * ZPC(2)/WN(9.2)
GP(9.2)=5.85E-12 * 5.85E-12 * ZPC(2)/(FG(9.2) * WN(9.2)**2)
FG(9.3)=(0.224 * ZPN(1) + 2.92E-2 * ZPN(2))/WN(9.3)
GP(9.3)=(3.41E-10 * SORT(ZPN(1)) + 1.48E-10 * SORT(ZPN(2))**2
/((FG(9.3) * WN(9.3)**2)
1      FG(9.4)=(5.24E-2 * ZPO(1) + 7.22E-2 * ZPO(2)
+ 6.04E-2 * ZPO(3))/WN(9.4)
1      GP(9.4)=(5.76E-12 * SORT(ZPO(1)) + 7.20E-11 * SORT(ZPO(2))
+ 8.05E-11 * SORT(ZPO(3))**2/(FG(9.4) * WN(9.4)**2)
1      FMUL(9.L)=FMUC(11.L)

C

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TRAN2890  
 TRAN2900  
 TRAN2910  
 TRAN2920  
 TRAN2930  
 TRAN2940  
 TRAN2950  
 TRAN2960  
 TRAN2970  
 TRAN2980  
 TRAN2990  
 TRAN3000  
 TRAN3010  
 TRAN3020  
 TRAN3030  
 TRAN3040  
 TRAN3050  
 TRAN3060  
 TRAN3070  
 TRAN3080  
 TRAN3090  
 TRAN3100  
 TRAN3110  
 TRAN3120  
 TRAN3130  
 TRAN3140  
 TRAN3150  
 TRAN3160  
 TRAN3170  
 TRAN3180  
 TRAN3190  
 TRAN3200  
 TRAN3210  
 TRAN3220  
 TRAN3230  
 TRAN3240

```

C ** PLANK FUNCTION **
C
DO 9 J=1,NHVL
  BEEL(J,L)=5.04E3 * HVJ(J)**3 / (EXP(HVJ(J)/T1) - 1.0)
C
C ** INDUCED EMISSION FACTOR (EQ 81) **
C
SSM(J,1,L)=1.10E-16*SNDH (L)*(1.0-EXP(-HVJ(J)/T1)) * FG(J,1)
SSM(J,2,L)=1.10E-16*SNDH (L)*(1.0-EXP(-HVJ(J)/T1)) * FG(J,2)
SSM(J,3,L)=1.10E-16*SNDH (L)*(1.0-EXP(-HVJ(J)/T1)) * FG(J,3)
SSM(J,4,L)=1.10E-16*SNDH (L)*(1.0-EXP(-HVJ(J)/T1)) * FG(J,4)
DO 10 M=1,4
  GGM(J,M,L)=GP(J,M) * SNDE(L) * (T(L)/1.0E4)**0.25
  1 * 1.0E-6
  IF(L.GT.1) GO TO 11
  ETAM(J,M,1)=0.
  SBM (J,M,1)=0.
  GO TO 10
11 ETAM(J,M,L)=ETAM(J,M,L-1)+ (YD(L)-YD(L-1))
  1 * (SSM(J,M,L-1) * GGM(J,M,L-1) + SSM(J,M,L) * GGM(J,M,L))
  2 * DELTA/3.14159265
  SBM(J,M,L)=SBM(J,M,L-1) + (YD(L)-YD(L-1))
  1 * (SSM(J,M,L-1)+SSM(J,M,L)) * DELTA
10 CONTINUE
  IF (L.GT.1) GO TO 12
  TAU1(J,1)=0.
  GO TO 9
12 TAU1(J,L)=TAU1(J,L-1) + (YD(L)-YD(L-1))
  1 * (FMUL(J,L-1)+FMUL(J,L)) * DELTA
  9 CONTINUE
  IF (ICG.NE.99) GO TO 91
  CALL BUGPR (?)
C
91 CONTINUE
  IEZ=IEZ+1
  ETZ(IEZ)=1.0

```

TRAN3250  
 TRAN3260  
 TRAN3270  
 TRAN3280  
 TRAN3290  
 TRAN3300  
 TRAN3310  
 TRAN3320  
 TRAN3330  
 TRAN3340  
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 TRAN3360  
 TRAN3370  
 TRAN3380  
 TRAN3390  
 TRAN3400  
 TRAN3410  
 TRAN3420  
 TRAN3430  
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 TRAN3470  
 TRAN3480  
 TRAN3490  
 TRAN3500  
 TRAN3510  
 TRAN3520  
 TRAN3530  
 TRAN3540  
 TRAN3550  
 TRAN3560  
 TRAN3570  
 TRAN3580  
 TRAN3590  
 TRAN3600

TRAN3610  
TRAN3620  
TRAN3630

RETURN  
END

C



```

C
C
C
SUBROUTINE TRANS1 ( ISW)
C
C **      CONTINUUM - CONTINUUM FLUX DIVERGENCE CALCULATION **
C
COMMON /ZPI/ ZPD(6), ZPN(6),ZPH(2), ZPC(7)
COMMON /FINV/ NHVL,NHVC,FHVC(12),DJ(9),HVJ(9),ZKZ
COMMON /SFLUX/ ORI(3)
COMMON /TRN/ NUT(60), FMC(12,60), FPC(12,60),
1 FM(9,60), FP(9,60), LINES
COMMON /YL/ETA(60), YD(60)
COMMON /PROP1/P (60), R(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /FRSTRM/ U INF, RINF, UINF2,XL , RE, LXI, ITM, IEM, NES
COMMON /DEL/ W(11),DTIL,D7ILS
COMMON /NON/RDZ,MUDZ,RNDZ,AKNF,HNF,CPNF
COMMON /MAIN/KEEP,MAXE,MAXM,MAXD, IDG,MCONV,ECONV,DCONV,LT,IAB
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /RH/ DUD,DPH1,TD,RZB,PD,HD,HTOTAL
COMMON /WT/SMW(20),AWT(5)
COMMON /TEST/ETZ(60),IEZ
COMMON /NUNDEN/ SND02(60), SNDN2(60), SNDG(60), SNDN(60),
2 SNDE(60), SNDCC(60),
3 SNDC(60), SNDC2(60), SNDC2H(60)
COMMON /DEBUG/ QLC(60), QCL(60), QLL(60), DON(60), QCC(60),
1 BEEC(12,60), FMUC(12,60), EM(12,60),
2 EP(12,60), TAUC(12,60), BEEL(9,60),
3 OCCP(12), WMW(9,60), GMM(9,60),
4 EEW(9,60), XLMM(9,60), QLCP(9),
5 QCLP(9), QLLP(9), DELTA, IY, IYY,
6 WPP(9,60), GPP(9,60), EEP(9,60),
7 XLPP(9,60), FG(9,4), GP(9,4),
8 WN(9,4), FMUL(9,60), SSM(9,4,60),
9 GGM(9,4,60), ETAM(9,4,60), SBM(9,4,60),
A TAU(9,60)
COMMON /SPEC/ MF, XMOL
COMMON/MIKE/ META1,META2,PNF,MNF,RHOD,DP,TS,UD,VD,RV,RWQ,EPS,PHI
C
C
C

```



```

1612 FP(J,L)=0.
DO 1613 L=1,IEZ
  QCL(L)=0.
  QLC(L)=0.
  QLL(L)=0.
1613 DO 49 IY=1,IEZ
  IY=NUT(IY)
  DO 20 K=1,12
    FMC(K,IY)=0.
    FPC(K,IY)=0.
    IF (IY.EQ.1) GO TO 44
    DO 40 L=1,IY
      C ** MINUS EMISSIVITY FUNCTION (EQ 47) *
      C
      EM(K,L)=1.0 - EXP(TAUC(K,L)-TAUC(K,IY))
      IF (L.EQ.1) GO TO 40
      C ** MINUS CONTINUUM FLUX (EQ 46) **
      C
      FMC(K,IY)=FMC(K,IY) - (EM(K,L)-EM(K,L-1))
      I * (BEEC(K,L-1)+BEEC(K,L))/2.
      40 CONTINUE
      44 IF (IY.EQ.NES) GO TO 41
      DO 42 L=IY,NES
      C ** POSITIVE EMISSIVITY FUNCTION (EQ 47) **
      C
      EP(K,L)=1.0 - EXP(TAUC(K,IY)-TAUC(K,L))
      IF (L.EQ.IY) GO TO 42
      C ** POSITIVE EMISSIVITY CONTINUUM FLUX (EQ 46) **
      C
      FPC(K,IY)=FPC(K,IY) + (EP(K,L)-EP(K,L-1))
      I * (BEEC(K,L-1)+BEEC(K,L))/2.
      42 CONTINUE

```

TRAN 730  
 TRAN 740  
 TRAN 750  
 TRAN 760  
 TRAN 770  
 TRAN 780  
 TRAN 790  
 TRAN 800  
 TRAN 810  
 TRAN 820  
 TRAN 830  
 TRAN 840  
 TRAN 850  
 TRAN 860  
 TRAN 870  
 TRAN 880  
 TRAN 890  
 TRAN 900  
 TRAN 910  
 TRAN 920  
 TRAN 930  
 TRAN 940  
 TRAN 950  
 TRAN 960  
 TRAN 970  
 TRAN 980  
 TRAN 990  
 TRAN1000  
 TRAN1010  
 TRAN1020  
 TRAN1030  
 TRAN1040  
 TRAN1050  
 TRAN1060  
 TRAN1070  
 TRAN1080

```

C ** POSITIVE EMISSIVITY CONTINUUM FLUX DIVERGENCE (EQ 51) **
C
C
C 41 OCCP(K)=6.2831853 * FMC(K,IY) *
      1 (FMC(K,IY) + FPC(K,IY) - 2.0* BEEC(K,IY))
      FMC(K,IY)=FMC(K,IY) * 3.14159265
      FPC(K,IY)=FPC(K,IY) * 3.14159265
C 20 CONTINUE
C
C ** DEBUG PRINT **
C
C IF (IDG.NE.99) GO TO 21
CALL BUGPR (3)
C 21 OCC(IY)=0.
DO 24 K=1.12
C
C ** LINE AND CROSS TERM FLUX DIVERGENCE CALCULATION **
C
C 24 OCC(IY)=OCC(IY) + OCCP(K)
IF (LINES.EQ.0) GO TO 1614
C
C ** INTEGRATION FROM 1 TO IY **
C
C IF (IY.EQ.1) GO TO 68
DO 65 J=1.9
DO 66 L=1.IY
WIM=0.
SUM1=0.
SUM2=0.
DO 67 M=1.4
DIF=ETAM(J,M,IY) - ETAM(J,M,L)
DIFSBN = SBN(J,M,IY)-SBN(J,M,L)
IF (ABS(DIFSBN).LT.1.E-10) DIFSBN = 1.E-10
BETAN=DIF / ( DIFSBN
IF (L.EQ.IY) BETAN=GM(J,M,L)
IF (ABS(DIF).GT.1.E-10) GO TO 9001

```

```

TRAN1090
TRAN1100
TRAN1110
TRAN1120
TRAN1130
TRAN1140
TRAN1150
TRAN1160
TRAN1170
TRAN1180
TRAN1190
TRAN1200
TRAN1210
TRAN1220
TRAN1230
TRAN1240
TRAN1250
TRAN1260
TRAN1270
TRAN1280
TRAN1290
TRAN1300
TRAN1310
TRAN1320
TRAN1330
TRAN1340
TRAN1350
TRAN1360
TRAN1370
TRAN1380
TRAN1390
TRAN1400
TRAN1410
TRAN1420
TRAN1430
TRAN1440

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```

          TM = 1.E-10
          GO TO 9002
        9001 CONTINUE
          TM=DI/2.0/BETAM**2
          9002 RRM=DI/2.0/GGM(J.M.IY)**2
              WWM=6.2831853 * WN(J.M) * BETAM * GAMMA(TM) * TM
              SUM1=SUM1 + GAMMA(TM) * WN(J.M) * SSM(J.M.IY)
              SUM2=SUM2 + XLAMB(RRM) * WN(J.M) * SSM(J.M.IY)
        67 WIM=WIM + WWM
          ALPHAM=WIM/DJ(J)

          C ** OVERLAPPING LINE CALCULATIONS **
          C
          C ** GROUP EQUIVALENT WIDTHS (EQ.88) **
          C
          WWM(J.L)=DJ(J) * PHI1(ALPHAM) * EXP(TAUL(J.L)-TAUL(J.IY))
          C
          C ** GROUP GAMMA -- LINE TRANSPORT FUNCTION (EQ.92) **
          C
          GMM(J.L)=PHI2(ALPHAM) * SUM1
          C
          C ** MINUS EMISSIVITY FUNCTION FOR LINES (EQ.47) **
          C
          EEM(J.L)=1.0 - EXP(TAUL(J.L)-TAUL(J.IY))
          66 XLMM(J.L)=PHI2(ALPHAM) * SUM2
          65 CONTINUE
              IF (IDG.EC.99) CALL BUGPR(1)
              IF (IDG.EQ.99) CALL BUGPR(4)
          68 IF (IY.EQ.NES) GO TO 70
          C
          C ** INTEGRATION FROM IY TO NES **
          C
          DO 69 J=1.9
          DO 70 L=IY.NES
          VIP=0.

```

```

TRAN1450
TRAN1460
TRAN1470
TRAN1480
TRAN1490
TRAN1500
TRAN1510
TRAN1520
TRAN1530
TRAN1540
TRAN1550
TRAN1560
TRAN1570
TRAN1580
TRAN1590
TRAN1600
TRAN1610
TRAN1620
TRAN1630
TRAN1640
TRAN1650
TRAN1660
TRAN1670
TRAN1680
TRAN1690
TRAN1700
TRAN1710
TRAN1720
TRAN1730
TRAN1740
TRAN1750
TRAN1760
TRAN1770
TRAN1780
TRAN1790
TRAN1800

```

```

SUM1=0.
SUM2=0.
DO 71 N=1,4
DIF=ETAM(J,M,L) - ETAM(J,M,IY)
DIFSBM = SEM(J,M,L)-SEM(J,M,IY)
IF(ABS(DIFSBM).LT.1.E-10) DIFSBM = 1.E-10
BETAP=DIF / ( DIFSBM ) * 3.14159265
IF (L.EQ.IY) BETAP=GGM(J,M,L)
IF(ABS(DIF).GT.1.E-10) GO TO 9003
TP = 1.E-10
GO TO 9004

9003 CONTINUE
TP=DIF/2.0/BETAP**2
9004 RRP=DIF/2.0/GGM(J,M,IY)**2
WIP=6.2831853 * WN(J,M) * BETAP * GAMMA(TP) * TP
SUM1=SUM1 + GAMMA(TP) * WN(J,M) * SSM(J,M,IY)
SUM2=SUM2 + XLAMB(RRP) * WN(J,M) * SSM(J,M,IY)
71 WIP=WIP+WIP
ALPHAP=WIP/DJ(J)
WPP(J,L)=DJ(J) * PHI1(ALPHAP) * EXP(TAUL(J,IY)-TAUL(J,L))
GPP(J,L)=PHI2(ALPHAP) * SUM1

C ** POSITIVE EMISSIVITY FUNCTION FOR LINES (EQ.47) **
C
C
EEP(J,L)=1.0 - EXP(TAUL(J,IY)-TAUL(J,L))
70 XLPP(J,L)=PHI2(ALPHAP) * SUM2
69 CONTINUE
C
C ** DEBUG PRINT **
C IF (IDG.EQ.99) CALL BUGPR (5)
C
72 DO 80 J=1,9
ASM1=0.
ASM2=0.
FM(J,IY)=0.
IF (IY.EQ.1) GO TO 81

```

TRAN1810  
 TRAN1820  
 TRAN1830  
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 TRAN1870  
 TRAN1880  
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 TRAN1920  
 TRAN1930  
 TRAN1940  
 TRAN1950  
 TRAN1960  
 TRAN1970  
 TRAN1980  
 TRAN1990  
 TRAN2000  
 TRAN2010  
 TRAN2020  
 TRAN2030  
 TRAN2040  
 TRAN2050  
 TRAN2060  
 TRAN2070  
 TRAN2080  
 TRAN2090  
 TRAN2100  
 TRAN2110  
 TRAN2120  
 TRAN2130  
 TRAN2140  
 TRAN2150  
 TRAN2160

```

DO 82 L=2,IY
FM(J,IY)=FM(J,IY) - (WM(J,L)-WM(J,L-1))
1  * (BEEL(J,L-1)+BEEL(J,L)) * 1.5707963
IF (L.EQ.IY) GO TO 82
ASM1=ASM1 - (EEM(J,L)-EEM(J,L-1))
1  * (BEEL(J,L-1) * XLMM(J,L-1) + BEEL(J,L) * XLMM(J,L))/2.
ASM2=ASM2 - (XLMM(J,L)-XLMM(J,L-1))
1  * (BEEL(J,L-1) * EXP(TAUL(J,L-1)-TAUL(J,IY)) + BEEL(J,L)
2  * EXP(TAUL(J,L)-TAUL(J,IY)))/2.0
82 CONTINUE
81 ASP1=0.
ASP2=0.
IYP=IY+1
IF (IY.EQ.NES) GO TO 83
DO 84 L=IYP,NES
FP(J,IY)=FP(J,IY) + (WPP(J,L)-WPP(J,L-1))
1  * (BEEL(J,L-1)+BEEL(J,L)) * 1.5707963
IF (L.EQ.IYP) GO TO 84
ASP1=ASP1 + (EEP(J,L)-EEP(J,L-1))
1  * (BEEL(J,L-1) * XLPP(J,L-1) + BEEL(J,L) * XLPP(J,L))/2.0
ASP2=ASP2 + (XLPP(J,L)-XLPP(J,L-1)) *
1  (BEEL(J,L-1) * EXP(TAUL(J,IY)-TAUL(J,L-1)) + BEEL(J,L)
2  * EXP(TAUL(J,IY)-TAUL(J,L)))/2.0
84 CONTINUE
83 QLCP(J)=2.0 * FMUL(J,IY) * (FM(J,IY)+FP(J,IY))
SUMS=1.0
SUMT=.
DO 86 M=1.4
SUMT=SUMT + SSM(J,M,IY) * WN(J,M)
ATM1=0.
IF (IY.NE.1) ATM1=(BEEL(J,IY-1)+BEEL(J,IY)) /2.0 * EEM(J,IY-1)
1  * XLMM(J,IY-1)
ATP1=0.
IF (IY.NE.NES) ATP1=(BEEL(J,IY+1)+BEEL(J,IY))/2.0 * EEP(J,IY+1)
1  * XLPP(J,IY+1)
1  QCLP(J)=6.2831853 * SUMS * (ASM1+ASP1+ATM1+ATP1)

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TRAN2170  
 TRAN2180  
 TRAN2190  
 TRAN2200  
 TRAN2210  
 TRAN2220  
 TRAN2230  
 TRAN2240  
 TRAN2250  
 TRAN2260  
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 TRAN2280  
 TRAN2290  
 TRAN2300  
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 TRAN2500  
 TRAN2510  
 TRAN2520

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TRAN2530
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TRAN2770
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TRAN2790
TRAN2800
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TRAN2860
TRAN2870
TRAN2880

      IF (IY.EQ.1) ATM2=-BEEL(J,IY) * SUMT
      IF (IY.NE.1) ATM2=(BEEL(J,IY-1)-BEEL(J,IY)) * GMM(J,IY-1)
      1 - BEEL(J,IY-1) * XLMM(J,IY-1)
      IF (IY.EQ.NES) ATP2=-BEEL(J,IY) * SUMT
      IF (IY.NE.NES) ATP2=(BEEL(J,IY+1)-BEEL(J,IY)) * GPP(J,IY+1)
      1 - BEEL(J,IY+1) * XLPP(J,IY+1)
      QLLP(J)=6.2831853 * SUMS*(-ASM2-ASP2+ATM2+ATP2)
      80 CONTINUE
      QCL(IY)=0.
      QLC(IY)=0.
      QLL(IY)=0.
      DO 85 J=1,9
      QCL(IY)=QCL(IY) + QCLP(J)
      QLC(IY)=QLC(IY) + QLCP(J)
      85 QLL(IY)=QLL(IY) + QLLP(J)
      1614 CONTINUE
      DON(IY)=-((QCL(IY)+QCL(IY)+QLC(IY)+QLL(IY))
C
C **      DEBUG PRINT **
C
      IF (IDG.EQ.0) GO TO 49
      CALL BUGPR(6)
      49 CONTINUE
      IEZ=IEZ-1
      DO(1)=DON(1)
      L=2
      DO 1 N=2,NES
      DO 2 I=2,IEZ
      NP=1
      IF (ETZ(1).GT.ETA(N)) GO TO 3
      2 CONTINUE
      3 NN=NP-1
      AA=0.0
      ZB=(DON(NN)-DON(NP)) / (ETZ(NN)-ETZ(NP))
      CC=DON(NN) - ZB * ETZ(NN)
      DO(N)=AA * ETA(N)**2 + ZB * ETA(N) + CC

```



TRAN2890  
 TRAN2900  
 TRAN2910  
 TRAN2920  
 TRAN2930  
 TRAN2940  
 TRAN2950  
 TRAN2960  
 TRAN2970  
 TRAN2980

```

      GO TO 1
      4 DO(N)=DON(NN)
      1 CONTINUE
C
C ** NON-DIMENSIONALIZE E(I) **
DO 250 I=1,NES
  T(I) = T(I)/TS
  250 E(I) = ((DO(I)*XL)/(RDZ *UINF**3))**20866.0
      RETURN
      END

```

```

SUBROUTINE TRANS2
COMMON /SFLUX/ ORI(3)
COMMON /YL/ETA(60),YOND(60)
COMMON /FRSTRM/ U INF, R INF, UINF2, R, RE, LXI, ITM, IEM, NES
COMMON /TRN/
1      FM(9,60), FP(9,60), LINES
COMMON /FINV/ NHVL,NHVC,FHVC(12),DJ(9),HVJ(9),ZKZ
COMMON /TEST/ETZ(60),IEZ
COMMON /NUMDEN/ SNDO2(60), SNDN2(60), SNDO(60), SNDN(60),
2      SNDE(60), SNDC(60),
3      SNDH(60), SNDC2(60), SNDH2(60), SNDC2H(60)
COMMON /MAIN/KEEP,MAXE,MAXM,MAXD, IDG,MCONV,ECONV,DCONV,LT,IAB
COMMON /SPEC/ M7, XMDL
DIMENSION ETOUT(3)
META=NES
ETOUT(1)=0.0
ETOUT(2)=0.5
ETOUT(3)=1.0
NOUT=3

C  OUTPUT FLUX
C
C      IF (IDG .GT. 0)
C      XWRITE (6,600)
C      IF (IDG .GT. 0)
C      XWRITE (6,603) (ETA(1),SNDN2(1),SNDO2(1),SNDN(1),SNDO(1),
1      SNDE(1), SNDH(1),
2      SNDC(1),SNDC2(1),SNDH2(1),SNDC2H(1),SNDN2(1),SNDN(1),SNDC(1),
3      SNDC2H(1),
C **      CONTINUUM CONTRIBUTION TO THE SPECTRAL FLUX **
C
C      IF (IDG .GT. 0)
C      XWRITE (6,4103)
C      DO 8040 K=1,NOUT
C      DO 8041 LK=1,NES

```

```

TRAN 10
TRAN 20
TRAN 30
TRAN 40
TRAN 50
TRAN 60
TRAN 70
TRAN 80
TRAN 90
TRAN 100
TRAN 110
TRAN 120
TRAN 130
TRAN 140
TRAN 150
TRAN 160
TRAN 170
TRAN 180
TRAN 190
TRAN 200
TRAN 210
TRAN 220
TRAN 230
TRAN 240
TRAN 250
TRAN 260
TRAN 270
TRAN 280
TRAN 290
TRAN 300
TRAN 310
TRAN 320
TRAN 330
TRAN 340
TRAN 350
TRAN 360

```

```

      NUT(K)=LK
      IF (ABS(ETOUT(K)-ETA(LK)) - 1.0E-05) 8040,8040,8041
8041 CONTINUE
8040 CONTINUE
      L1=NUT(1)
      L2=NUT(2)
      L3=NUT(3)
      IF (IDG .GT. 0)
        XWRITE (6,8037)(ETOUT(IL),IL=1,3)
      FM1=0.0
      FP1=0.0
      FM2=0.0
      FP2=0.0
      FM3=0.0
      FP3=0.0
      DO 4104 KL=1,NHVC
        IF (IDG .GT. 0)
          XWRITE (6,8042) KL, FHV(KL), FMC(KL,L1), FPC(KL,L1),
1          FMC(KL,L2), FPC(KL,L2), FMC(KL,L3), FPC(KL,L3)
          FM1=FM1 + FMC(KL,L1)
          FP1=FP1 + FPC(KL,L1)
          FM2=FM2 + FMC(KL,L2)
          FP2=FP2 + FPC(KL,L2)
          FM3=FM3 + FMC(KL,L3)
          FP3=FP3 + FPC(KL,L3)
4104 CONTINUE
      IF (IDG .GT. 0)
        XWRITE (6,8045) FM1, FP1, FM2, FP2, FM3, FP3
        ORI(1)=FM1+FP1
        ORI(2)=FM2+FP2
        ORI(3)=FM3+FP3
C
C ** LINE CONTRIBUTION TO THE SPECTRAL FLUX **
C
      IF (LINES.EQ.0) RETURN
      IF (IDG .GT. 0)

```

TRAN 370  
 TRAN 380  
 TRAN 390  
 TRAN 400  
 TRAN 410  
 TRAN 420  
 TRAN 430  
 TRAN 440  
 TRAN 450  
 TRAN 460  
 TRAN 470  
 TRAN 480  
 TRAN 490  
 TRAN 500  
 TRAN 510  
 TRAN 520  
 TRAN 530  
 TRAN 540  
 TRAN 550  
 TRAN 560  
 TRAN 570  
 TRAN 580  
 TRAN 590  
 TRAN 600  
 TRAN 610  
 TRAN 620  
 TRAN 630  
 TRAN 640  
 TRAN 650  
 TRAN 660  
 TRAN 670  
 TRAN 680  
 TRAN 690  
 TRAN 700  
 TRAN 710  
 TRAN 720

```

C
C ** TOTAL FLUX CALCULATION **
C
      DO 8043 KL=1,NHVL
      IF (IDG.GT. 0)
      XWRITE (6,8037) (ETOUT(IL),IL=1,3)
      FM1=0.0
      FP1=0.0
      FM2=0.0
      FP2=0.0
      FM3=0.0
      FP3=0.0

      C
      C ** TOTAL FLUX CALCULATION **
      C
      DO 8043 KL=1,NHVL
      IF (IDG.GT. 0)
      XWRITE (6,8042) KL, HVJ(KL), FM(KL,L1), FP(KL,L1),
      1 FM(KL,L2), FP(KL,L2), FM(KL,L3), FP(KL,L3)
      FM1=FM1 + FM(KL,L1)
      FP1=FP1 + FP(KL,L1)
      FM2=FM2 + FM(KL,L2)
      FP2=FP2 + FP(KL,L2)
      FM3=FM3 + FM(KL,L3)
      FP3=FP3 + FP(KL,L3)
      8043 CONTINUE
      IF (IDG.GT. 0)
      XWRITE (6,8045) FM1, FP1, FM2, FP2, FM3, FP3
      ORI(1)=ORI(1) + FM1 + FP1
      ORI(2)=ORI(2) + FM2 + FP2
      ORI(3)=ORI(3) + FM3 + FP3

      C
      600 FORMAT (1H1,33NUMBER DENSITIES (PARTICLES/CM3) ///5X,3META, 8X,
      1 2MN2, 8X,2MO2, 8X,1MN, 8X,1MO, 8X, 2ME-.8X,
      2 1H1,8X,1HC, 8X,2MC2, 8X,2MN2, 8X,2MO2, 8X,2MC3,8X,3HC2H///)
      603 FORMAT (1P13E10.2)
      4103 FORMAT (44H1CONTINUUM CONTRIBUTION TO THE SPECTRAL FLUX)
      8035 FORMAT (39H1LINE CONTRIBUTION TO THE SPECTRAL FLUX)
      8037 FORMAT (/22X,5META =F7.3,13X,5META =F7.3,13X,5META =F7.3//3X,1H1,

```

```

TRAN 730
TRAN 740
TRAN 750
TRAN 760
TRAN 770
TRAN 780
TRAN 790
TRAN 800
TRAN 810
TRAN 820
TRAN 830
TRAN 840
TRAN 850
TRAN 860
TRAN 870
TRAN 880
TRAN 890
TRAN 900
TRAN 910
TRAN 920
TRAN 930
TRAN 940
TRAN 950
TRAN 960
TRAN 970
TRAN 980
TRAN 990
TRAN1000
TRAN1010
TRAN1020
TRAN1030
TRAN1040
TRAN1050
TRAN1060
TRAN1070
TRAN1080

```

TRAN1090  
TRAN1100  
TRAN1110  
TRAN1120  
TRAN1130  
TRAN1140

1 3X.3HNU.. 6HMINUS.7X.5HPLUS.8X.6HMINUS.7X.5HPLUS.8X.

2 6HMINUS.7X.5HPLUS/)

8042 FORMAT (14.F8.3.1P8E13.3)

8045 FORMAT (12H0TOTAL FLUX .1P8E13.3)

RETURN

END

```

SUBROUTINE SND(I)
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,MNF,CPNF
COMMON/WT/SNW(20),AWT(5)
COMMON /NUMDEN/ SNDO2(60), SOND2(60), SOND(60), SOND(60), SOND(60),
1
2      SNDH(60), SNDC2(60), SOND2H(60)
3      SNDC3(60),SNDC2H(60)
** CALCULATE SPECIE NUMBER DENSITIES BASED ON MOLE FRACTIONS **
CONVER = 3.10375E+23 *RHO(I) *RDZ
C
C
C
SNDO2(I) = CONVER * C( 1,I)/SNW( 1)
SOND2(I) = CONVER * C( 2,I)/SNW( 2)
SNDO(I) = CONVER * C( 3,I)/SNW( 3)
SOND(I) = CONVER * C( 4,I)/SNW( 4)
SNDE(I) = CONVER * C( 7,I)/SNW( 7)
SNDC(I) = CONVER * C( 8,I)/SNW( 8)
SONDH(I) = CONVER * C( 9,I)/SNW( 9)
SONDH2(I) = CONVER * C(10,I)/SNW(10)
SNDCO(I) = CONVER * C(11,I)/SNW(11)
SNDC3(I) = CONVER * C(12,I)/SNW(12)
SNDC2(I) = CONVER * C(19,I)/SNW(19)
SNCC2H(I) = CONVER * C(14,I)/SNW(14)
RETURN
END
SND( 10
SND( 20
SND( 30
SND( 40
SND( 50
SND( 60
SND( 70
SND( 80
SND( 90
SND( 100
SND( 110
SND( 120
SND( 130
SND( 140
SND( 150
SND( 160
SND( 170
SND( 180
SND( 190
SND( 200
SND( 210
SND( 220
SND( 230
SND( 240
SND( 250
SND( 260
SND( 270

```

```

SUBROUTINE ZP(T1,SUMN,SUMD,SUMH,SUMC)
      ** FRACTIONAL POPULATION STATES FOR N. O. H. C **
      COMMON /ZP1/ ZPD(6), ZPN(6),ZPH(2), ZPC(7)
      ZPH(1)=2.0/SUMH
      ZPH(2)=8.0 * EXP(-10.20/T1)/SUMH
      ZPC(1)=9.0/SUMC
      ZPC(2)=5.0 * EXP(-1.264/T1)/SUMC
      ZPC(3)=EXP(-2.684/T1)/SUMC
      ZPC(4)=5.0 * EXP(-4.183/T1)/SUMC
      ZPC(5)=12.0 * EXP(-7.532/T1)/SUMC
      ZPC(6)=36.0*EXP(-8.722/T1)/SUMC
      ZPC(7)=60.0 * EXP(-9.724/T1)/SUMC
      ZPN(1)=4.0/SUMN
      ZPN(2)=10.0* EXP(-2.384/T1)/SUMN
      ZPN(3)=6.0 * EXP(-3.576/T1)/SUMN
      ZPN(4)=18.0 * EXP(-10.452/T1)/SUMN
      ZPN(5)=54.0 * EXP(-11.877/T1)/SUMN
      ZPN(6)=90.0 * EXP(-13.002/T1)/SUMN
      ZPD(1)=9.0/SUMD
      ZPD(2)=5.0 * EXP(-1.967/T1)/SUMD
      ZPD(3)=EXP(-4.188/T1)/SUMD
      ZPD(4)=8.0 * EXP(-9.283/T1)/SUMD
      ZPD(5)=24.0 * EXP(-10.830/T1)/SUMD
      ZPD(6)=40.0 * EXP(-12.077/T1)/SUMD

      RETURN
      END

```

```

ZP(T  10
ZP(T  20
ZP(T  30
ZP(T  40
ZP(T  50
ZP(T  60
ZP(T  70
ZP(T  80
ZP(T  90
ZP(T 100
ZP(T 110
ZP(T 120
ZP(T 130
ZP(T 140
ZP(T 150
ZP(T 160
ZP(T 170
ZP(T 180
ZP(T 190
ZP(T 200
ZP(T 210
ZP(T 220
ZP(T 230
ZP(T 240
ZP(T 250
ZP(T 260
ZP(T 270
ZP(T 280
ZP(T 290

```

```

SUBROUTINE BUGPR (IDGSW)
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, NES
COMMON /YL/ETA(60), YD(60)
COMMON /TRN/ NUT(60), FMC(12,60), FPC(12,60),
1 FM(9,60), FP(9,60), LINES
COMMON /DEBUG/ OLC(60), QCL(60), QLL(60), DON(60), OCC(60),
1 BEEC(12,60), FMUC(12,60), EM(12,60),
2 EP(12,60), TAUC(12,60), BEEL(9,60),
3 OCCP(12), WM(9,60), GMM(9,60),
4 EEM(9,60), XLMM(9,60), OLC(9),
5 OCLP(9), DELTA, IY, IYY,
6 WPP(9,60), GPP(9,60), EEP(9,60),
7 XLPP(9,60), FG(9,4), GP(9,4),
8 WN(9,4), FMUL(9,60), SSM(9,4,60),
9 GGM(9,4,60), ETAM(9,4,60), SBM(9,4,60),
A TAU(9,60)
GO TO (10,20,30,40,50,60,70), IDGSW
10 WRITE (6,194)
194 FORMAT (1H1)
RETURN
20 WRITE (6,7182) DELTA
7182 FORMAT (7HDELTA=1PE14.7,3H CM)
RETURN
30 WRITE (6,190) IY, YD(IY)
190 FORMAT (4H1Y=13.2X,3HYD=1PE12.5//2X,1HK,2X,1HL,7X,3META,13X,2HYD,
1 13X,2MMU,11X,3HTAU,14X,1HE,11X,3MBEE//)
DO 22 K=1,12
IF (IY.EQ.1) GO TO 23
WRITE(6,191) (K, L, ETA(L), YD(L), FMUC(K,L), TAUC(K,L),
1 EM(K,L), BEEC(K,L), L=1,IY)
191 FORMAT (213,1P6E15.5)
WRITE (6,192)
192 FORMAT (//)
23 IF (IY.EQ.NES) GO TO 22
WRITE (6,191) (K, L, ETA(L), YD(L), FMUC(K,L),
1 TAUC(K,L), EP(K,L), BEEC(K,L), L=IY,NES)

```

BUGP 10  
 BUGP 20  
 BUGP 30  
 BUGP 40  
 BUGP 50  
 BUGP 60  
 BUGP 70  
 BUGP 80  
 BUGP 90  
 BUGP 100  
 BUGP 110  
 BUGP 120  
 BUGP 130  
 BUGP 140  
 BUGP 150  
 BUGP 160  
 BUGP 170  
 BUGP 180  
 BUGP 190  
 BUGP 200  
 BUGP 210  
 BUGP 220  
 BUGP 230  
 BUGP 240  
 BUGP 250  
 BUGP 260  
 BUGP 270  
 BUGP 280  
 BUGP 290  
 BUGP 300  
 BUGP 310  
 BUGP 320  
 BUGP 330  
 BUGP 340  
 BUGP 350  
 BUGP 360



```

22 WRITE (6,193) FMC(K,IY), FPC(K,IY), OCCP(K)
193 FORMAT (5HOFIM=1PE12.5,2X,4HFIP=E12.5,2X,5HOCPP=E12.5)
RETURN
40 WRITE (6,195) IY, YD(IY), ((J, L, YD(L),
1 WMH(J,L), GMM(J,L), XLMM(J,L), EEM(J,L),
2 BEEL(J,L), L=1,IY), J=1,9)
195 FORMAT (4H0IY=13,2X,3HYI=1PE12.5//2X,1MJ,2X,1ML,7X,2HYD,12X,3HWMH,BUGP 370
1 12X,3HGMH,11X,4HXLMM,13X,3HEEM,13X,3HBEEL//2I3,6E16.5) BUGP 380
RETURN BUGP 390
50 WRITE (6,196) IY, YD(IY), ((J, L, YD(L),
1 WPP(J,L), GPP(J,L), XLPP(J,L), EEP(J,L),
2 BEEL(J,L), L=1,IY), J=1,9) BUGP 400
196 FORMAT (4H0IY=13,2X,3HYI=1PE12.5//2X,1MJ,2X,1ML,7X,2HYD,12X,3HWMH,BUGP 410
1 2X,3HGMH,11X,4HXLPP,13X,3HEEP,13X,3HBEEL//2I3,6E16.5) BUGP 420
RETURN BUGP 430
60 WRITE (6,198) IY, ETA(IY), YD(IY)
198 FORMAT (4H0IY=13,2X,4META=1PE12.5,2X,3HYI=E12.5//2X,1MJ,5X,3HOCCL,BUGP 440
1 11X,3HFMH,11X,3HFP,11X,3HOCCL,11X,3HOLC,11X,3HOLL,12X,2HFM,12X, BUGP 450
2 2HFP,11X,3HOCCL//)
198 FORMAT (4H0IY=13,2X,3HYI=1PE12.5//2X,1MJ,2X,1ML,7X,2HYD,12X,3HWMH,BUGP 460
1 2X,3HGMH,11X,4HXLPP,13X,3HEEP,13X,3HBEEL//2I3,6E16.5) BUGP 470
RETURN BUGP 480
60 WRITE (6,199) (J, OCCP(J), FMC(J,IY), FPC(J,IY),
1 QCLP(J), OLCP(J), OLLP(J), FM(J,IY), FP(J,IY),
2 J=1,9) BUGP 490
199 FORMAT (13,1P8E14.5) BUGP 500
WRITE (6,8069) (J, OCCP(J), FMC(J,IY), FPC(J,IY), J=10,12) BUGP 510
8069 FORMAT (13,1P3E14.5) BUGP 520
WRITE (6,200) OCC(IY), OCL(IY), OLC(IY), OLL(IY),
1 DON(IY)
200 FORMAT (1H0,2X,1PE14.5,20X,3E14.5,20X,E14.5) BUGP 530
RETURN BUGP 540
70 WRITE (6,197) L, ETA(L), YD(L), ((J, M, WN(J,M),
1 FG(J,M), GP(J,M), FMUL(J,L), TAU(L,J,L),
2 SSM(J,M,L), GGM(J,M,L), ETAM(J,M,L), SBN(J,M,L), BUGP 550
3 M=1,4),J=1,9) BUGP 560
197 FORMAT (3HOL=13,2X,4META=1PE12.5,2X,3HYI=E12.5//2X,1MJ,2X,1ML,7X, BUGP 570
1 1MN,13X,1HF,13X,1HG,11X,3HFMU,11X,3HTAU,11X,3HSSM,11X,3HGMH,10X, BUGP 580
2 4META,11X,3HSSM//2I3,9E14.5) BUGP 590

```

BUGP 730  
BUGP 740

398

RETURN  
END

## SUBROUTINE ZHV(HV,ZO,ZN,ZI,ZC)

\*\* THIS SUBROUTINE CALCULATES THE QUANTUM MECHANICAL CORRECTION  
FACTORS GIVEN A FREQUENCY (HV) \*\*

```

X= HV
X2 =X*X
X3 =X2*X
X4 =X3*X
X5 =X4*X
X6 =X5*X
X7 =X6*X
IF (Y -9.82) 1,1.2
1  Z0 = .9999795
2  +6.677328 E-03*X3
   -7.708637 E-05*X6
   +2.668133 E-06*X7
   -0.3155480*X
   +2.824548 E-02*X2
   +8.058070 E-04*X5
   GO TO 3
2  Z0 = (X/9.82)**3
3  IF (X -8.35) 4,4.5
4  ZN = 1.000148
   -9.779458 E-02*X3
   +4.515535E-04*X6
   -0.4183535 *X
   +3.354635 E-02*X4
   -1.403585 E-05*X7
   +.1680359 *X2
   -5.609353 E-03*X5
   GO TO 6
5  ZN = (X/8.35)**3
6  Y = X/4.0
   IF (Y-6.6) 9,9.10
9  Y2 =Y*Y
   Y3 =Y2*Y
   Y4 =Y3*Y
   Y5 =Y4*Y
   Y6 =Y5*Y
   Y7 =Y6*Y
   ZI = 1.000379
   -0.2964767 *Y
   +7.505242 E-02*Y2
   +3.279554 E-03*Y4
   -2.128469 E-04*Y5
   GO TO 11

```

ZHV( 10  
ZHV( 20  
ZHV( 30  
ZHV( 40  
ZHV( 50  
ZHV( 60  
ZHV( 70  
ZHV( 80  
ZHV( 90  
ZHV( 100  
ZHV( 110  
ZHV( 120  
ZHV( 130  
ZHV( 140  
ZHV( 150  
ZHV( 160  
ZHV( 170  
ZHV( 180  
ZHV( 190  
ZHV( 200  
ZHV( 210  
ZHV( 220  
ZHV( 230  
ZHV( 240  
ZHV( 250  
ZHV( 260  
ZHV( 270  
ZHV( 280  
ZHV( 290  
ZHV( 300  
ZHV( 310  
ZHV( 320  
ZHV( 330  
ZHV( 340  
ZHV( 350  
ZHV( 360

```

10  ZI = (Y/6.6)**3
11  IF (X-7.37) 12,12,13
12  ZC = .9974367
    1  -1.393917 E-02**X3
    2  2.812126 E-05**X6
      GO TO 14
13  ZC = (X/7.37)**3
14  RETURN
    END

      - .43418*2 *X
      +4.038545 E-03**X4
      -3.883530 E-07**X7

      +8.531314 E-02**X2
      -5.426425 E-04**X5

ZHV( 370
ZHV( 380
ZHV( 390
ZHV( 400
ZHV( 410
ZHV( 420
ZHV( 430
ZHV( 440
ZHV( 450

```

```

SUBROUTINE FGH(N,F,G,H)
COMMON /VEL/ FZ(60),FC(60),Z(60),V(60)
C ** ** FIRST DERIVATIVE ** **
C-----EVALUATES COEFFICIENTS FOR THREE POINT DIFFERENCE APPROXIMATION
C OF DAVIS (AIAA JR. VOL. 8, NO. 5, MAY 1970).
COMMON /YL/ETA(60),YOND(60)
DN=ETA(N+1)-ETA(N)
DNM1=ETA(N)-ETA(N-1)
DEL=DN+DNM1
D1=DN*DEL
D2=DN*DNM1
D3=DNM1*DEL
F=DNM1/D1
G=((DN-DNM1)/D2
H=-DN/D3
RETURN
ENTRY FGH2(N,F,G,H)
C ** ** SECOND DERIVATIVE ** **
C-----EVALUATES COEFFICIENTS FOR THREE POINT DIFFERENCE APPROXIMATION
C OF DAVIS (AIAA JR. VOL. 8, NO. 5, MAY 1970).
DN=ETA(N+1)-ETA(N)
DNM1=ETA(N)-ETA(N-1)
F=2./((DN*(DN+DNM1))
G=-2./((DN*DNM1)
H=2./((DNM1*(DN+DNM1))
RETURN
END

```

```

FGH( 10
FGH( 20
FGH( 30
FGH( 40
FGH( 50
FGH( 60
FGH( 70
FGH( 80
FGH( 90
FGH( 100
FGH( 110
FGH( 120
FGH( 130
FGH( 140
FGH( 150
FGH( 160
FGH( 170
FGH( 180
FGH( 190
FGH( 200
FGH( 210
FGH( 220
FGH( 230
FGH( 240
FGH( 250
FGH( 260
FGH( 270

```

```

SUBROUTINE INTRPL(VAR,X,F,IMAX,SOM)
C
C-----THIS PROGRAM PERFORMS LAGRANGIAN INTERPOLATION
C WITH NON-EQUAL STEP SIZE BETWEEN POINTS
C F=DEPENDANT VARIABLE
C X=INDEPENDENT VARIABLE
C VAR=VALUE OF X FOR WHICH CORRESPONDING VALUE OF
C F IS DESIRED BY INTERPOLATION
C IMAX=NUMBER OF POINTS IN ARRAY X OR F
C SOM=VALUE OF INTERPOLATED DEPENDENT VARIABLE
C NPTS=NUMBER OF POINTS USED FOR INTERPOLATION
C
C DIMENSION X(100),F(100),XN(300),FN(300)
C NPTS=3
C XUP=1.E30
C DO611 I=1,IMAX
C T=VAR-X(I)
C IF(T.GE.0.)GOTO609
C T=-T
C IF(T.GE.XUP)GOTO611
C IP=1
C YP=T
C CONTINUE
C IN=1
C NPP=NPTS+1
C DO618 I=1,NPP
C FN(I)=F(IP)
C XN(I)=X(IP)
C IF(IN.GT.0)GOTO613
C IO=IP-1
C GOTO615
C IO=IP+1
C IF(IMAX.GE.IO)GOTO615
C IP=IP-1
C GOTO618
607 XUP=1.E30
608 T=-T
609 IF(T.GE.XUP)GOTO611
610 IP=1
611 CONTINUE
612 IO=IP-1
613 IO=IP+1
614 IP=IP-1
GOTO618
INTR 10
INTR 20
INTR 30
INTR 40
INTR 50
INTR 60
INTR 70
INTR 80
INTR 90
INTR 100
INTR 110
INTR 120
INTR 130
INTR 140
INTR 150
INTR 160
INTR 170
INTR 180
INTR 190
INTR 200
INTR 210
INTR 220
INTR 230
INTR 240
INTR 250
INTR 260
INTR 270
INTR 280
INTR 290
INTR 300
INTR 310
INTR 320
INTR 330
INTR 340
INTR 350
INTR 360

```

```

615 IF(IQ.GT.0)GOTO617
616 IP=IP+1
    GOTO618
617 IP=IQ
    IN=-IN
618 CONTINUE
    SOM=0.
    FACT=1.
    DO620J=1,NPTS
    SOM=SOM+FACT*FN(I)
    DO619I=J,NPTS
    IQ=I-J+1
619 FN(IQ)=(FN(IQ+1)-FN(IQ))/(XN(I+1)-XN(I))
620 FACT=FACT*(VAR-XN(J))
    RETURN
    END
INTR 370
INTR 380
INTR 390
INTR 400
INTR 410
INTR 420
INTR 430
INTR 440
INTR 450
INTR 460
INTR 470
INTR 480
INTR 490
INTR 500
INTR 510
INTR 520

```

QU	10
QU	20
QU	30
QU	40
QU	50
QU	60
QU	70
QU	80
QU	90

```

FUNCTION QUAD (X,FX,I)
** TRAPEZOIDAL QUADRATURE FUNCTION **
    DIMENSION X(60),FX(60)
    DX=X(I)-X(I-1)
    QUAD = (DX/2.0) * (FX(I) + FX(I-1))
    RETURN
END

```

C  
C  
C



```

SUBROUTINE OUPPUT (N)
** ROUTINE TO PRINT SHOCK LAYER SOLUTION **

COMMON/ID/SP(20),EL(5)
COMMON /CONV/ FPRCT,TPRCT,DDAMP,TDAMP,PDIL
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON /MAIM/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,LT,IAB
LOGICAL MCONV,ECONV,DCONV
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,C.PNF
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/PROP1/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/PROP3/CPS(20,60),MS(20,60),CP (60),HM(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /RH/ DUD,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON /SFLUX/QR1(3)
COMMON/VECTOR/ CA(60),CB(60),CC(60),B(60)
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON/WALL/RVW,PRW,TEOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /YL/ETA(60),YOND(60)
COMMON /DD/ D(60)
COMMON/SP1/SS,TOL,NDEBUG
COMMON/MIKE/ META1,META2,PNF,WNF,RHOD,DP,TS,UD,VD,RV,RVO,EPS,PHI
COMMON/JOHN/ UO(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILOU
COMMON/BOB/ XI,DXI,YL,QLAST
COMMON/TIT/TITLE(18)
COMMON /BILL/ MODE(20),ISTART,IOGD,IDGF,IDGH,IDGT,IDGS,
X MI,NOE,NTIME,MIS
DIMENSION BOUT(60),DQR(30)
REAL MU,MUDZ
DATA HEAD1,'WALL',HEAD2,' ',HEAD3,'SHOC' /
D(1) = 8.128E-8*(T(1)*TS)**1.659/PD/R/UINF

```

OUPP 10  
 OUPP 20  
 OUPP 30  
 OUPP 40  
 OUPP 50  
 OUPP 60  
 OUPP 70  
 OUPP 80  
 OUPP 90  
 OUPP 100  
 OUPP 110  
 OUPP 120  
 OUPP 130  
 OUPP 140  
 OUPP 150  
 OUPP 160  
 OUPP 170  
 OUPP 180  
 OUPP 190  
 OUPP 200  
 OUPP 210  
 OUPP 220  
 OUPP 230  
 OUPP 240  
 OUPP 250  
 OUPP 260  
 OUPP 270  
 OUPP 280  
 OUPP 290  
 OUPP 300  
 OUPP 310  
 OUPP 320  
 OUPP 330  
 OUPP 340  
 OUPP 350  
 OUPP 360



```

445 CONTINUE
C BTU/FT**2-SEC
QRP=QR*0.88
C ** COMPUTE DIFFUSIVE FLOW TO SURFACE **
C WATTS/CM**2
OD = 0.
DO 1768 I=1,NSP
1768 QD = QD + MS(I,1)*(C(I,2) - C(I,1))
QD = - RHO(1)*D(1)*QD/(2. *YOND(2))
CD = RDZ *UINF*UINF2*QD/(.88*778.28)
C BTU/FT**2-SEC
QDP = .88*QD
QTOTAL=QC+QR+QD
QTOTP=QTOTAL*.88
C
C *** CHECK FOR CONVERGENCE ***
N = 2
ABD = ABS((QTOTAL-QLAST)/QTOTAL)
QLAST = QTOTAL
IF(XI .EQ. 0.0) GO TO 300
IF(ABD .GT. 0.005) RETURN
N = 1
300 CONTINUE
WRITE(6,203)
203 FORMAT(1H1)
DO 50 K=1,NETA
YOND(K)= YOND(K)/DELTA
50 CONTINUE
C
C
C ** DIMENSIONALIZE RHO,MU,P,AND E **
DO 450 I = 1 , NETA
RHO(I)=RHO(I)*RDZ
MU (I)=MU(I)*MUZ

```

OUPP 730  
 OUPP 740  
 OUPP 750  
 OUPP 760  
 OUPP 770  
 OUPP 780  
 OUPP 790  
 OUPP 800  
 OUPP 810  
 OUPP 820  
 OUPP 830  
 OUPP 840  
 OUPP 850  
 OUPP 860  
 OUPP 870  
 OUPP 880  
 OUPP 890  
 OUPP 900  
 OUPP 910  
 OUPP 920  
 OUPP 930  
 OUPP 940  
 OUPP 950  
 OUPP 960  
 OUPP 970  
 OUPP 980  
 OUPP 990  
 OUPP1000  
 OUPP1010  
 OUPP1020  
 OUPP1030  
 OUPP1040  
 OUPP1050  
 OUPP1060  
 OUPP1070  
 OUPP1080

```

RM (I)=RM(I)*RMDZ
AK(I) = AK(I)/AKNF
E(I) = E(I) * RDZ
CP(I) = CP(I)/CPNF
450 CONTINUE
C
WRITE(6,200) TITLE
200 FORMAT(10X,18A4,/)
C
** PRINT SHOCK QUANTITIES AND HEATING RATE **
C
WRITE(6,204) X1, DELTA,DTIL
204 FORMAT(1H0,' X1 =',F9.4,10X,9H DELTA = ,1PE14.6,10X,7HDTIL = ,
1 E15.6)
C
WRITE (6,210) EPS, QC,QCP
210 FORMAT ( 1H0,' EPS =',F9.4,10X,5H QC = E15.6,2X,13H(WATTS/CM**2),
1 2X,1H=E15.6,2X,17H(BTU/FT**2 - SEC) ,
WRITE(6,212) RZB,QR,QRP
212 FORMAT ( 1H0,6H RB =,F9.4,10X,5H QR = E15.6,2X,13H(WATTS/CM**2),
1 2X,1H=E15.6,2X,17H(BTU/FT**2 - SEC) ,
WRITE(6,213) RVW,QD,ODP
213 FORMAT ( 1H0,6H RVW =,F9.4,10X,5H OD = E15.6,2X,13H(WATTS/CM**2),
1 2X,1H=E15.6,2X,17H(BTU/FT**2 - SEC) ,
C
WRITE(6,215) QTOTAL,Q1,TP
215 FORMAT(1H0,16HTOTAL MEA,1NG = E15.6,2X,13H(WATTS/CM**2),
1 2X,1H=E15.6,2X,17H(BTU/FT**2 - SEC) ,
C
** PRINT Y/D , F AND T PROFILES **
C
C
WRITE(6,205)
205 FORMAT(1H0,7X, 4H ETA, 5X, 4HY/DZ, 8X, 2HF', 8X, 3H RV, 8X ,
1 4HT/TO, 4X ,13H E(WATTS/CM3),4X,2H V, 7X,12H V (FT/SEC) ,
2 5X,2H G,6X,12H H (STATIC) , /)
FP = 0.0
OUPP1090
OUPP1100
OUPP1110
OUPP1120
OUPP1130
OUPP1140
OUPP1150
OUPP1160
OUPP1170
OUPP1180
OUPP1190
OUPP1200
OUPP1210
OUPP1220
OUPP1230
OUPP1240
OUPP1250
OUPP1260
OUPP1270
OUPP1280
OUPP1290
OUPP1300
OUPP1310
OUPP1320
OUPP1330
OUPP1340
OUPP1350
OUPP1360
OUPP1370
OUPP1380
OUPP1390
OUPP1400
OUPP1410
OUPP1420
OUPP1430
OUPP1440

```

```

C
NS=NSP
DO 100 I=1,NETA
C COMPUTE ENTHALPIES
HSTAT = 0.0
DO 99 J=1,NS
HSTAT = HSTAT + HS(J,I)*C(J,I)
G = HSTAT + V(I)**2
C
HEAD=HEAD2
IF(I.EQ.1) HEAD=HEAD1
IF(I.EQ.NETA) HEAD=HEAD3
YDZ = YOND(I)
IF(I.EQ.NETA) FP = 1.0
RV = -FC(I)*DTIL*2.
RV = RMD(I)*V(I)/RDZ
VS = V(I)*VINF
TP = T(I)/T(NETA)
WRITE(6,208) HEAD,ETA(I),YDZ,FP,RV,TP,E(I),V(I),VS,G,HSTAT
208 FORMAT(1H,A4,F6.3,I10E12.3)
IF(I.LT.NETA-1) FP = Z(I)*DTIL
100 CONTINUE
C
** WRITE OUT SHOCK LAYER GAS PROPERTIES **
C
WRITE(6,44)
44 FORMAT(1H1.48X,28H-SHOCK LAYER GAS PROPERTIES-
)
C
WRITE(6,206)
206 FORMAT(1H0.3X,3HETA,8X,4H Y/D,12X,2HP,12X,2H T,11X,3HRMD,11X,2HMMU,
1,12X,3HRMU,11X,2H K)
C
WRITE(6,207)
207 FORMAT(1H,27X,6H(ATM.),6X,13H (DEG.KEL.),12H(SLUGS/FT3),2X,
1,28H(LBM/FT-SEC) (LBM2-SEC3/FT6),16H (BTU/FT-SEC-R),//)
C

```

OUPP1450  
 OUPP1460  
 OUPP1470  
 OUPP1480  
 OUPP1490  
 OUPP1500  
 OUPP1510  
 OUPP1520  
 OUPP1530  
 OUPP1540  
 OUPP1550  
 OUPP1560  
 OUPP1570  
 OUPP1580  
 OUPP1590  
 OUPP1600  
 OUPP1610  
 OUPP1620  
 OUPP1630  
 OUPP1640  
 OUPP1650  
 OUPP1660  
 OUPP1670  
 OUPP1680  
 OUPP1690  
 OUPP1700  
 OUPP1710  
 OUPP1720  
 OUPP1730  
 OUPP1740  
 OUPP1750  
 OUPP1760  
 OUPP1770  
 OUPP1780  
 OUPP1790  
 OUPP1800

```

C
DO 101 I=1,NETA
TP = T(I)*TS
WRITE(6,8)ETA(I),YOND(I),PI(I),TP ,RHO(I),MU(I),RM (I),AK(I)
C
8 FORMAT(1H OPF7.4,1P8E14.4)
9 FORMAT(1H OPF7.4,1P7E14.4)
C
101 CONTINUE
C
** WRITE SPECIES MASS FRACTIONS **
C
WRITE(6,230)
230 FORMAT (1H1.48X.26H-SPECIES MASS FRACTIONS- )
WRITE(6,231)
231 FORMAT (3X,ETA,8X.3H 02.11X.2HN2.11X.3H 0 .11X.3H N .11X.3H 0+.
1 11X.3H N+.11X.3H E-./)
DO 102 I=1,NETA
WRITE(6,8) ETA(I),C(1,I),C(2,I),C(3,I),C(4,I),C(5,I).
1 C(6,I),C(7,I)
102 CONTINUE
WRITE(6,230)
WRITE(6,233) (SP(I),I=8,15)
233 FORMAT(2X.4H ETA.1X.8(10X.A4)/)
WRITE(6,8) (ETA(I), (C(J,I),J= 8,15),I=1,NETA)
WRITE(6,230)
WRITE(6,234) (SP(I),I=16,20)
234 FORMAT(2X.4H ETA.7X.3H CP.5(11X.A4).8X.4H AMW./)
WRITE(6,9) (ETA(I),CP(I),C(J,I),J=16,20),AMW(I),I=1,NETA)
C NONDIMENSIONALIZE
DO100 I=1,NETA
YOND(I) = YOND(I)*DELTA
RHO(I) = RHO(I)/RDZ
MU(I) = MU(I)/MUDZ
RM(I) = RM(I)/RMDZ
E(I) = ((E(I)*R)/(RDZ *UINF**3))*20866.0

```

OUPP1810  
 OUPP1820  
 OUPP1830  
 OUPP1840  
 OUPP1850  
 OUPP1860  
 OUPP1870  
 OUPP1880  
 OUPP1890  
 OUPP1900  
 OUPP1910  
 OUPP1920  
 OUPP1930  
 OUPP1940  
 OUPP1950  
 OUPP1960  
 OUPP1970  
 OUPP1980  
 OUPP1990  
 OUPP2000  
 OUPP2010  
 OUPP2020  
 OUPP2030  
 OUPP2040  
 OUPP2050  
 OUPP2060  
 OUPP2070  
 OUPP2080  
 OUPP2090  
 OUPP2100  
 OUPP2110  
 OUPP2120  
 OUPP2130  
 OUPP2140  
 OUPP2150  
 OUPP2160

OUPP2170  
 OUPP2180  
 OUPP2190  
 OUPP2200  
 OUPP2210  
 OUPP2220  
 OUPP2230  
 OUPP2240  
 OUPP2250  
 OUPP2260  
 OUPP2270  
 OUPP2280  
 OUPP2290  
 OUPP2300  
 OUPP2310  
 OUPP2320

```

CP(I)=CP(I)*CPNF
1001 AK(I)=AK(I)*AKNF
217  FORMAT(6E12.5)
1000 CONTINUE
      CALL CPUTIME(ITIME)
      WRITE(6,110) ITIME
110  FORMAT('0.40X,TIME LEFT =.112. MSEC'////)
      IF(ITIME/60000 .LT. NTIME) GO TO 301
      IF(X1 .GT. 0.0) RETURN
      IF(ABD .LT. 0.005) GO TO 301
      RETURN
301  CALL PUNCH
      CALL EXIT
C
      RETURN
      END
  
```

```

SUBROUTINE PUNCH
COMMON /CONV/ FPRCT,TPRCT,DDAMP,TDAMP,PD,TIL
COMMON/CONV1/HDAMP
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON /MAIM/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/PROP1/PI(60),RMD(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/PROP3/CPS(20,60),HS(20,60),CP (60),HM(60)
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /RH/ DUD,DPHI,TD,RZB,PD,HD,HTOTAL
COMMON /VEL/ F(60),FC(60),Z(60),V(60)
COMMON/WALL/RW,RWB,PRB,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /YL/ETA(60),YOND(60)
COMMON/TIT/TITLE(16)
COMMON/SP1/SS,TOL,NDBUG
COMMON/BOB/ XI,DXI,YL
COMMON/MIKE/ META1,META2,PNF,MNF,RHOD,DP,TS,UD,VD,RW,RWD,EPS,PHI
COMMON/JFB/ IPN
REAL MU,MUDZ

*** PUNCH-OUT CARDS FOR RESTART ***

WRITE(7,566)TITLE,IEM
566 FORMAT(18A4,I8)
WRITE(7,1801) FPRCT,TPRCT,DDAMP,TDAMP,PD,TIL,HDAMP
WRITE(7,1801)DELTA,DTIL,DTILS
WRITE(7,1802) U INF, RINF, UINF2, R, RE, ITM, IEM, META
WRITE(7,1803)KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,IAB
PUNCH 1801, RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
WRITE(7,1804) NSP,NNS,NE,NC
WRITE(7,1801)(PI(J),RMD(J),T(J),J=1,META)
WRITE(7,1801)(AMW(J),J=1,META)

```

PONC 10  
 PONC 20  
 PONC 30  
 PONC 40  
 PONC 50  
 PONC 60  
 PONC 70  
 PONC 80  
 PONC 90  
 PONC 100  
 PONC 110  
 PONC 120  
 PONC 130  
 PONC 140  
 PONC 150  
 PONC 160  
 PONC 170  
 PONC 180  
 PONC 190  
 PONC 200  
 PONC 210  
 PONC 220  
 PONC 230  
 PONC 240  
 PONC 250  
 PONC 260  
 PONC 270  
 PONC 280  
 PONC 290  
 PONC 300  
 PONC 310  
 PONC 320  
 PONC 330  
 PONC 340  
 PONC 350  
 PONC 360



```

WRITE(7,1801)((C(I,J),J=1,NETA),I=1,NSP)
WRITE(7,1801)((HM(J),J=1,NETA)
WRITE(7,1801)((E(J),J=1,NETA)
WRITE(7,1804)IRAD,ITYPE
PUNCH 1801. DUD,DPHI,TD,RZB,PD,MD,HTOTAL
WRITE(7,1801)((F(J),FC(J),Z(J),V(J),J=1,NETA)
PUNCH 1801. RVW,PRV,TWOLD
WRITE(7,1801)((CWALL(I),I=1,NSP)
WRITE(7,1801)((ECWALL(K),K=1,NE)
WRITE(7,1801)((ETA(J),YOND(J),J=1,NETA)
WRITE(7,715)NDEBUG,IAB,TOL
WRITE(7,1801)((EC(K,J),J=1,NETA),K=1,NE)
WRITE(7,1801)XI,RV,TS,UD
1801 FORMAT(6E13.5)
1802 FORMAT(5E12.5,3I3)
1803 FORMAT(5I5,3L3,2I5)
1804 FORMAT(4I4)
715 FORMAT(2I5,E15.6)
RETURN
END

```

```

PUNC 370
PUNC 380
PUNC 390
PUNC 400
PUNC 410
PUNC 420
PUNC 430
PUNC 440
PUNC 450
PUNC 460
PUNC 470
PUNC 480
PUNC 490
PUNC 500
PUNC 510
PUNC 520
PUNC 530
PUNC 540
PUNC 550
PUNC 560

```

```

BLOCK DATA
DATA 10
DATA 20
DATA 30
DATA 40
DATA 50
DATA 60
DATA 70
DATA 80
DATA 90
DATA 100
DATA 110
DATA 120
DATA 130
DATA 140
DATA 150
DATA 160
DATA 170
DATA 180
DATA 190
DATA 200
DATA 210
DATA 220
DATA 230
DATA 240
DATA 250
DATA 260
DATA 270
DATA 280
DATA 290
DATA 300
DATA 310
DATA 320
DATA 330
DATA 340
DATA 350
DATA 360

DATA INITIALIZATION ROUTINE

COMMON/FINV/ NHVL,NIHVC,FHVC(12),DJ(9),HVJ(9),ZKZ
COMMON /FRSTRM/ U INF, RINF, UINF2, R, RE, LXI, ITM, IEM, META
COMMON/GUESS/TG1(60),TG2(60)
COMMON/PROPI/PI(60),RHO(60), T(60),AMW(60),C (20,60),EC(5,60)
COMMON/PROP2/ MU(60),RM(60), AK(60)
COMMON/PROP3/CPS(20,60),NS(20,60),CP (60),HM(60)
COMMON/NUMBER/NSP,NNS,NE,NC
COMMON/ELSP/LSP(5)
COMMON/ID/SP(20),EL(5)
COMMON/WT/SMW(20),AWT(5)
COMMON /BLOCK1/V1(20),V2(20),V3(20)
COMMON/BLOCK3/K1(20),K2(20)
COMMON/EQ1/A1(20), B1(20), C1(20), DI(20), EI(20), FI(20), GI(20), H1(20), I1(20), J1(20), K1(20), L1(20), M1(20), N1(20), O1(20), P1(20), Q1(20), R1(20), S1(20), T1(20), U1(20), V1(20), W1(20), X1(20), Y1(20), Z1(20)
COMMON/EQ2/AA(20,5), ICODE(20)
COMMON/EG3/IA(20,5)
REAL K1,K2
COMMON/CK/ISN(20),MWT(19)
REAL MWT

COMMON/JTB/ CO(60,20),RED,RED2
COMMON /YL/EIA(60),YOND(60),YO(60)
COMMON /RH/ DUD,DPHI,TD,RZB,PD,MO,HTOTAL
COMMON/WALL/RWV,PRV,TWOLD,FLUX(20),CWALL(20),ECWALL(5)
COMMON /NAIM/KEEP,MAXE,MAXM,MAXD,IDEBUG,MCONV,ECONV,DCONV,LT,1AB
COMMON/CONVI/HDAMP,FDAMP,CDAMP
COMMON /CONV/ FPRCT,TPRCT,DDAMP,YDAMP,POTIL
COMMON /RFLUX/ E(60),IRAD,ITYPE
COMMON /DEL/ DELTA,DTIL,DTILS
COMMON /NON/RDZ,MUDZ,RMDZ,AKNF,HNF,CPNF
COMMON/JOHN/ UD(60),U(60),DU(60),VO(60),RO(60),DNDX(60),DTLA,DTILODATA
COMMON/BOB/ XI,DXI,YL,QLAST

```

C  
C  
C

C

```

COMMON/MIKE/ NETA1,NETA2,PNF,WNF,RHOD,DP,TS,UD,VD,RW,RWD,EPS,PHI DATA 370
COMMON /BILL/ KODE(20),ISTART,IDGO,IDGF,IDGH,IDGT,IDGS, DATA 380
X MI,NOE,NTIME,MIS,CSMDCK(20) DATA 390
COMMON /GEORGE/ EPSN(30),XIN(30),OWN(30) DATA 400
LOGICAL MCONV,ECONV,DCONV DATA 410
REAL MU,MUDZ DATA 420
DATA 430
DATA 440
DATA 450
DATA 460
DATA 470
DATA 480
DATA 490
DATA 500
DATA 510
DATA 520
DATA 530
DATA 540
DATA 550
DATA 560
DATA 570
DATA 580
DATA 590
DATA 600
DATA 610
DATA 620
DATA 630
DATA 640
DATA 650
DATA 660
DATA 670
DATA 680
DATA 690
DATA 700
DATA 710
DATA 720

ORDER OF SPECIES USED IN CHEMICAL PRODUCTION ROUTINE (FG2)

DATA ISN/
1 11. 8. 19. 12. 20. 14. 13. 18. 10. 9. 2. 4. 6.
2 3. 5. 7. 15. 16. 17. 1/

SPECIE MOLECULAR WEIGHTS (FG2 ORDER)

DATA MWT/
1 28.011. 12.011. 24.022. 36.033. 12.011. 25.030. 26.019. 27.027.
2 02.016. 01.008. 28.016. 14.008. 14.008. 16.000. 16.000. .0005486.
3 26.038. 37.041. 49.052/
DATA NETA/59/

INITIAL GUESS FOR DENSITY PROFILE

DATA RHO /25.1,14.3,8.85,6.50,4.37,3.01,2.49,2.17,1.90,1.67,1.46.
1 1.29,1.16,1.08,1.03,1.00,4*1.0/

INITIAL GUESS FOR RM PROFILE

DATA RM /10.0,7.71,5.89,5.10,4.18,3.54,3.31,3.10,2.83,2.48,2.09.
1 1.72,1.42,1.22,1.09,1.02,4*1.0/

INITIAL GUESS FOR NO BLOWING TEMPERATURE PROFILE

DATA TGI / .1033,.2294,.3531,.4719,.5777,.6531,.6867,.7034,.7145.
1 .7236,.7321,.7401,.7479,.7554,.7628,.7699,.7769,.7836,.7902.
2 .7967,.8030,.8092,.8153,.8213,.8272,.8331,.8389,.8447,.8504.

```

3 .8562..8619..8676..8734..8791..8850..8908..8968..9028..9089.  
 4 .9151..9215..9280..9347..9417..9488..9563..9641..9723..9809.  
 5 .9901.10\*1.0/

# INITIAL GUESS FOR BLOWING TEMPERATURE PROFILE

DATA TG2 / .3325..3325..3325..3325..3325..3325..3325..3325..3326..3328.  
 1 .3331..3336..3344..3357..3378..3408..3452..3515..3601..3718.  
 2 .3873..4076..4335..4665..5075..5560..6054..6487..6857..7161.  
 3 .7404..7595..7749..7878..7993..8100..8203..8302..8399..8496.  
 4 .8594..8693..8797..8904..9019..9142..9278..9476..9609..9757.  
 5 .9877.10\*1.0/

DATA NSP.NNS.NE.NC/20.0.5.20/

## SPECIES PRESENT

DATA SP/	O2	N2	H2	CO	HCN	CH4	OH	H	CH2	CH2H2	C+
1	.	.	.	.	.	.	.	.	.	.	.
2	.	.	.	.	.	.	.	.	.	.	.
3	.	.	.	.	.	.	.	.	.	.	.

## SPECIES MOLECULAR WEIGHTS (REGULAR ORDER)

DATA SMW/	32.000	28.016	16.000	14.008	16.000
1	14.038	5.486E-4	12.011	1.008	2.016
2	28.011	36.033	26.019	25.030	26.038
3	37.041	49.052	27.027	24.022	12.011/

## CONSTANTS FOR POLYNOMIAL FITS OF PURE COMPONENT VISCOSITY

DATA VI/	0.1693E	01.0.9704E	00.0.1519E	01.0.2534E	00.0.0
1	0.0	0.0	0.1997E <td>01.0.2941E <td>00.-.7944E-01.</td> </td>	01.0.2941E <td>00.-.7944E-01.</td>	00.-.7944E-01.
2	0.2404E <td>01.0.2019E <td>01.0.2404E <td>01.0.2404E <td>01.0.1396E 01.</td> </td></td></td>	01.0.2019E <td>01.0.2404E <td>01.0.2404E <td>01.0.1396E 01.</td> </td></td>	01.0.2404E <td>01.0.2404E <td>01.0.1396E 01.</td> </td>	01.0.2404E <td>01.0.1396E 01.</td>	01.0.1396E 01.
3	0.2019E <td>01.0.2019E <td>01.0.1378E <td>01.0.1931E <td>01.0.0 /</td> </td></td></td>	01.0.2019E <td>01.0.1378E <td>01.0.1931E <td>01.0.0 /</td> </td></td>	01.0.1378E <td>01.0.1931E <td>01.0.0 /</td> </td>	01.0.1931E <td>01.0.0 /</td>	01.0.0 /

DATA 730  
 DATA 740  
 DATA 750  
 DATA 760  
 DATA 770  
 DATA 780  
 DATA 790  
 DATA 800  
 DATA 810  
 DATA 820  
 DATA 830  
 DATA 840  
 DATA 850  
 DATA 860  
 DATA 870  
 DATA 880  
 DATA 890  
 DATA 900  
 DATA 910  
 DATA 920  
 DATA 930  
 DATA 940  
 DATA 950  
 DATA 960  
 DATA 970  
 DATA 980  
 DATA 990  
 DATA1000  
 DATA1010  
 DATA1020  
 DATA1030  
 DATA1040  
 DATA1050  
 DATA1060  
 DATA1070  
 DATA1080

DATA V2/0.1496E-02,0.1613E-02,0.1875E-02,0.2206E-02,0.5000E-03,  
 1 0.5000E-03,0.5000E-03,0.1772E-02,0.8893E-03,0.7907E-03,  
 2 0.1363E-02,0.1179E-02,0.1363E-02,0.1363E-02,0.8423E-03,  
 3 0.1179E-02,0.1179E-02,0.9651E-03,0.1393E-02,0.5000E-03/  
 C

DATA V3/-0.2276E-07,-0.1916E-07,-0.2228E-07,-0.3737E-07,-0.1000E-07,  
 1 -0.1000E-07,-0.1000E-07,-0.3378E-07,-0.8111E-08,-0.8864E-08,  
 2 -0.2184E-07,-0.1655E-07,-0.2184E-07,-0.2184E-07,-0.6939E-08,  
 3 -0.1655E-07,-0.1655E-07,-0.9481E-08,-0.2575E-07,-0.1000E-07/  
 C

CONSTANTS FOR POLYNOMIAL FITS OF THERMODYNAMIC DATA

DATA A1/0.3316E 01,0.3221E 01,0.2670E 01,0.2474E 01,0.2491E 01,  
 1 0.2727E 01,0.2500E 01,0.2612E 01,0.2500E 01,0.3358E 01,  
 2 0.3254E 01,0.4002E 01,0.3411E 01,0.3485E 01,0.3891E 01,  
 3 0.3965E 01,0.5874E 01,0.3654E 01,0.4443E 01,0.2609E 01/  
 C

DATA B1/0.1151E-02,0.9878E-03,-0.1970E-03,0.9097E-04,0.2762E-04,  
 1 -0.2820E-03,0.3440E-06,-0.2030E-03,-0.8243E-06,0.2794E-03,  
 2 0.9698E-03,0.3541E-02,0.4897E-03,0.3563E-02,0.5717E-02,  
 3 0.6200E-02,0.7403E-02,0.3444E-02,-0.2885E-03,-0.1393E-03/  
 C

DATA C1/-0.3726E-06,-0.2907E-06,0.7193E-07,-0.7814E-07,-0.1881E-07,  
 1 0.1105E-06,-0.1954E-09,0.1095E-06,0.6421E-09,0.9372E-07,  
 2 -0.2647E-06,-0.1318E-05,0.1005E-06,-0.1237E-05,-0.1957E-05,  
 3 -0.2265E-05,-0.2729E-05,-0.1258E-05,0.3036E-06,0.5959E-07/  
 C

DATA D1/0.6186E-10,0.3938E-10,-0.8901E-11,0.2218E-10,0.3807E-11,  
 1 -0.1551E-10,0.3937E-13,-0.1695E-10,-0.1720E-12,-0.2948E-10,  
 2 0.3037E-10,0.2064E-09,-0.3473E-10,0.1866E-09,0.2931E-09,  
 3 0.3717E-09,0.4437E-09,0.2169E-09,-0.6244E-10,-0.1037E-10/  
 C

DATA E1/-0.3666E-14,-0.2000E-14,0.4002E-15,-0.1489E-14,-0.1028E-15,  
 1 0.7847E-15,-0.2573E-17,0.8590E-15,0.1457E-16,0.2141E-14,  
 2 -0.1177E-14,-0.1144E-13,0.2361E-14,-0.1013E-13,-0.1585E-13,  
 3 -0.2262E-13,-0.2637E-13,-0.1430E-13,0.3915E-14,0.6345E-15/  
 C

DATA1090  
 DATA1100  
 DATA1110  
 DATA1120  
 DATA1130  
 DATA1140  
 DATA1150  
 DATA1160  
 DATA1170  
 DATA1180  
 DATA1190  
 DATA1200  
 DATA1210  
 DATA1220  
 DATA1230  
 DATA1240  
 DATA1250  
 DATA1260  
 DATA1270  
 DATA1280  
 DATA1290  
 DATA1300  
 DATA1310  
 DATA1320  
 DATA1330  
 DATA1340  
 DATA1350  
 DATA1360  
 DATA1370  
 DATA1380  
 DATA1390  
 DATA1400  
 DATA1410  
 DATA1420  
 DATA1430  
 DATA1440

DATA FI/-0.1044E 04.-.1043E 04.0.2915E 05.0.5609E 05.0.1879E 06.  
 1 0.2254E 06.-.7450E 03.0.8542E 05.0.2547E 05.-.1018E 04.  
 2 -.1434E 05.0.9423E 05.0.4745E 05.0.5809E 05.0.2590E 05.  
 3 0.6283E 05.0.7605E 05.0.1442E 05.0.9787E 05.0.2168E 06/  
 DATA GI/0.5393E 01.0.4326E 01.0.4504E 01.0.4300E 01.0.4424E 01.  
 1 0.3645E 01.-.1173E 02.0.4144E 01.-.4612E 00.-.3548E 01.  
 2 0.4875E 01.0.2020E 01.0.4746E 01.0.4784E 01.0.6520E 00.  
 3 0.3467E 01.-.4010E 01.0.2373E 01.-.1090E 01.0.3709E 01/  
 DATA AI1/0.3721E 01.0.3727E 01.0.2540E 01.0.2746E 01.0.2944E 01.  
 1 0.2499E 01.0.2508E 01.0.2141E 01.0.3934E 01.0.3363E 01.  
 2 0.3366E 01.0.2213E 02.0.3473E 01.0.5307E 01.0.6789E 01.  
 3 0.3965E 01.0.5874E 01.0.3654E 01.0.4026E 01.0.2528E 01/  
 DATA BI1/0.4254E-03.0.4684E-03.-.5952E-04.-.3909E-03.-.4108E-03.  
 1 -.3725E-05.-.6332E-05.0.3219E-03.-.1776E-02.0.4656E-03.  
 2 0.8027E-03.-.1759E-01.0.7337E-03.0.8966E-03.0.1503E-02.  
 3 0.6200E-02.0.7403E-02.0.3444E-02.0.4857E-03.0.4869E-05/  
 DATA CI1/-0.2835E-07.-.1140E-06.0.2701E-07.0.1338E-06.0.9156E-07.  
 1 0.1147E-07.0.1364E-08.-.5498E-07.0.6013E-06.-.5127E-07.  
 2 -.1968E-06.0.5565E-05.-.9088E-07.-.1378E-06.-.2295E-06.  
 3 -.2265E-05.-.2729E-05.-.1258E-05.-.7026E-07.-.7026E-08/  
 DATA DI1/0.6050E-12.0.1154E-10.-.2798E-11.-.1191E-10.-.5848E-11.  
 1 -.1102E-11.-.1094E-12.0.3604E-11.-.7819E-10.0.2802E-11.  
 2 0.1940E-10.-.6758E-09.0.4847E-11.0.9251E-11.0.1534E-10.  
 3 0.3717E-09.0.4437E-09.0.2169E-09.0.4666E-11.0.1134E-11/  
 DATA EI1/-0.5186E-17.-.3293E-15.0.9380E-16.0.3369E-15.0.1190E-15.  
 1 0.3078E-16.0.2934E-17.-.5564E-16.0.3482E-14.-.4905E-16.  
 2 -.5549E-15.0.2625E-13.-.1018E-15.-.2278E-15.-.3763E-15.  
 3 -.2262E-13.-.2637E-13.-.1430E-13.-.1142E-15.-.3476E-16/  
 DATA I450  
 DATA I460  
 DATA I470  
 DATA I480  
 DATA I490  
 DATA I500  
 DATA I510  
 DATA I520  
 DATA I530  
 DATA I540  
 DATA I550  
 DATA I560  
 DATA I570  
 DATA I580  
 DATA I590  
 DATA I600  
 DATA I610  
 DATA I620  
 DATA I630  
 DATA I640  
 DATA I650  
 DATA I660  
 DATA I670  
 DATA I680  
 DATA I690  
 DATA I700  
 DATA I710  
 DATA I720  
 DATA I730  
 DATA I740  
 DATA I750  
 DATA I760  
 DATA I770  
 DATA I780  
 DATA I790  
 DATA I800

DATA F11/-1044E 04.-1043E 04.0.2915E 05.0.5609E 05.0.1879E 06.  
 1 0.2254E 06.-7450E 03.0.8542E 05.0.2547E 05.-1018E 04.  
 2 -1434E 05.0.9423E 05.0.5420E 05.0.5809E 05.0.2590E 05.  
 3 0.6283E 05.0.7605E 05.0.1442E 05.0.9787E 05.0.2168E 06/  
 DATA G11/0.3254E 01.0.1294E 01.0.5049E 01.0.2872E 01.0.1750E 01.  
 1 0.4950E 01.-1208E 02.0.6874E 01.-8598E 01.-3716E 01.  
 2 0.4263E 01.-1021E 03.0.4152E 01.-5288E 01.-1539E 02.  
 3 0.3467E 01.-4010E 01.0.2373E 01.0.1090E 01.0.4139E 01/  
 CONSTANTS FOR LINEAR FITS OF PURE COMPONENT THERMOCONDUCTIVITY  
 DATA K1/0.1019E 01.0.6541E 00.0.1250E 01.0.1281E 01.2.6000E 01.  
 1 2.6000E 01.2.6000E 01.0.2506E 01.0.2496E 01.0.3211E 01.  
 2 0.8585E 00.0.6304E 00.0.8589E 00.0.1126E 01.0.1126E 01.  
 3 0.6304E 00.0.6304E 00.0.4855E 00.0.8589E 00.0.1000E-04/  
 DATA K2/0.4901E-03.0.5457E-03.0.7092E-03.0.8593E-03.0.0000E-03.  
 1 0.0000E-03.0.0000E-03.0.7479E-03.0.5129E-02.0.5344E-02.  
 2 0.6233E-03.0.5804E-03.0.6233E-03.0.7439E-03.0.7439E-03.  
 3 0.5804E-03.0.5804E-03.0.8714E-03.0.6233E-03.0.7350E-03/  
 DATA ICODE/20\*0/  
 ELEMENTS  
 DATA EL/ 'C'. 'H'. 'N'. 'O'. 'E' /  
 ATOMIC WEIGHTS  
 DATA AW1/ 12.011, 1.008, 14.008, 16.000, 5.486E-4 /  
 DATA IA/ 0.0.0.0.0.0.1.0.0.1.3.1.2.2.3.1.1.2.1.  
 H 0.0.0.0.0.0.0.1.2.0.0.0.1.2.1.1.1.0.0.  
 N 0.2.0.1.0.1.0.0.0.0.0.1.0.0.0.0.1.0.0.  
 O 2.0.1.0.1.0.0.0.0.0.1.0.0.0.0.0.0.0.0.0.

```

E      2.2,1.1,0.0,1.1,0.0,2.3,2.2,2.3,4.2,2.0/

DATA LSP/20.9,6.5,7/
DATA NWVL /9/. NIHVC /12/
DATA FHVC /5.0, 6.0, 7.0, 8.0, 9.0, 10.0, 10.8, 11.1,
1 12.0, 13.4, 14.3, 20.0/
DATA CJ /0.6, 2.2, 1.5, 1.65, 1.4, 1.0, 1.2, 1.4, 1.0/
DATA HWJ /1.3, 2.7, 5.75, 7.57, 9.1, 10.4, 11.4, 12.7, 13.9/
DATA ZKZ /7.26E-16/
C * * * * *
VALUES FOR COLD START OF STANDARD CASE
C
DATA TWLD,TS,RVM,RZB /3450.,14000.-,0.05,0.06/
DATA CELTA,DTIL,DPHI,XI,RV,PWF,PO,EPS,DXI /8*0.0,0.05/
DATA FDAMP,DDAMP,HDAMP,TDAMP,CDAMP/3*0.5,0.25,1.0/
DATA TPRCT,FPRCT,PDIL/2*0.005,0.001/
DATA RDZ,MUDZ,RWDZ,AKNF,CPNF,HMF,YL /7*1.0/
DATA IRAD,I TYPE /3.0/
DATA KEEP,MAXE,MAXM,MAXD,IDEBUG,LTY/0.10,15.30,0.2/
DATA MCONVE,ECONV,DCONV /F.F.F/
DATA ITM,LTY,KODE,ISTART,IDGD,IDGF,IDGH,IDGT,IDGS,NTIME /29*0/
DATA MI,MIS,NDE /2*5.50/
DATA U,CNDX /120*0.0/
DATA ETA/
X   0.000,.0050,..0111,.0150..0200..0231..0263..0294..0325..0356.
X   .0368,.0450..0538..0625..0713..0800..0850..0925..1000..1100.
X   .1200..1300..1400..1500..1600..1700..1800..2000..2100..2200.
X   .2300..2400..2500..2600..2800..3100..3400..3700..4000..4300.
X   .4600..4900..5200..5500..5800..6100..6400..6700..7000..7300.
X   .7600..7900..8200..8500..8800..9100..9400..9700..1.0,1.0/
DATA CSMOCK /
X   1.110E-05,2.735E-05,9.583E-02,3.355E-01,1.259E-01,4.407E-01.
X   2.162E-05,1.295E-04,1.010E-03,3.957E-11,1.188E-09,1.329E-38.
X   1.265E-09,8.679E-22,4.401E-43,2.569E-43,1.717E-50,3.672E-18.
X   1.768E-13,8.811E-04/

```



```

DATA CVAL/
X .1000E-09..2389E-01..1000E-09..2000E-04..1000E-09..1000E-09.
X .1000E-09..1570E-02..1708E-01..3277E-01..2578E+00..1767E-01.
X .3100E-02..1500E-00..8216E-01..2037E+00..1589E+00..4710E-01.
X .4130E-02..1000E-09/
DATA RMOD.T /6181.0/
DATA EC.CO /1500*0.0/
DATA C /1200*1.E-20/
DATA EPSN.XIN /31*0.0..05..1..15..2..25..3..35..4..45..5..55.
X .6..65..7..75..8..85..9..95..1..1.05..1.1.15..1.2.1.25..1.3.
X 1.35..1.4.1.45/
DATA QWN /1...98..95..91..85..76..68..57..5..42..35..28..22.
X .17..18..12.14*.1/
DATA QLST/0.0/
DATA IAB /30/
END
DATA2530
DATA2540
DATA2550
DATA2560
DATA2570
DATA2580
DATA2590
DATA2600
DATA2610
DATA2620
DATA2630
DATA2640
DATA2650
DATA2660
DATA2670
DATA2680

```

ENTRY MAIN  
OVERLAY ONE  
INSERT FRID, SOLVE, CHECK, MOLMAS, SLINI, I, OUTPUT, TEMPR  
OVERLAY TWO  
INSERT FG2  
OVERLAY THREE  
INSERT SPECIE, SLM TM  
OVERLAY THREE  
INSERT ENERGY, INTRPL  
OVERLAY TWO  
INSERT XDM TM, GLOBAL, PROPRT, DEDX  
OVERLAY ONE  
INSERT INPUT, OUPPUT, TRANS2, PUNCH, NEXT  
OVERLAY ONE  
INSERT SHOCK  
OVERLAY FIVE  
INSERT GAS  
OVERLAY ONE  
INSERT LRAD, ZP, SND, BUGPR, ZHV  
INSERT DEBUG  
OVERLAY FOUR  
INSERT TRANS  
OVERLAY FOUR  
INSERT TRANS1

C A S E 2 53.000 FT/SEC 2.5E-7 SLUGS/CU FT  
2.5E-7 9. 0

53000.  
CENTRY  
RVW = 0.0605  
END

C A S E 2      53,000 FT/SEC      2.5E-7 SLUGS/CU FT

BOUNDARY CONDITIONS AT       $XI = 0.0$

BOW SHOCK CONDITIONS

SHOCK TEMPERATURE (OK)      =      14511.  
SHOCK PRESSURE (ATM)      =      0.3132  
SHOCK ANGLE, PHI (RAD)      =      0.0  
NORMAL VELOCITY      =      -0.0563  
TANGENTIAL VELOCITY      =      0.0  
ANGLE EPS (RAD)      =      0.0

ABLATOR SURFACE

SURFACE TEMPERATURE (OK)      =      3577.  
SURFACE PRESSURE (ATM)      =      0.3132  
FLOWING RATE      =      0.0612  
MASS LOSS (LB/FT2-SEC)      =      0.0261  
QCOND (BTU/FT2-SEC)      =      216.  
QSUPL (BTU/FT2-SEC)      =      116.  
QRR (BTU/FT2-SEC)      =      541.  
QTOTAL (BTU/FT2-SEC)      =      873.

ORIGINAL PAGE IS  
OF POOR QUALITY

NFLTA	=	3.96750E-02		DTIL	=	5.820692E-02
QC	=	-0.17450ME+03	(WATTS/CM#02)		=	-0.153547E+03 (BTU/FY#02 - SEC)
QR	=	-0.76111E+03	(WATTS/CM#02)		=	-0.671803E+03 (BTU/FY#02 - SEC)
QD	=	-0.544773E+02	(WATTS/CM#02)		=	-0.479400E+02 (BTU/FY#02 - SEC)
TOTAL WATTING	=	-0.402397E+03	(WATTS/CM#02)		=	-0.873310E+03 (BTU/FY#02 - SEC)

PTA	W <sup>1/2</sup>	P <sup>0</sup>	DV	T/TD	E (WATTS/CM3)	V	V (FT/SEC)	G	M (STATIC)
0.0	0.0	0.0	3.402E-03	2.460E-01	-1.631E+02	5.448E-04	2.808E+01	2.048E-01	2.048E-01
0.05	2.127E-03	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	5.581E-04	2.808E+01	2.076E-01	2.076E-01
0.1	4.274E-03	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	5.719E-04	2.808E+01	2.110E-01	2.110E-01
0.15	6.501E-03	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	5.861E-04	2.808E+01	2.148E-01	2.148E-01
0.2	8.712E-03	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	5.914E-04	2.808E+01	2.188E-01	2.188E-01
0.25	1.091E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	5.965E-04	2.808E+01	2.233E-01	2.233E-01
0.3	1.313E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.016E-04	2.808E+01	2.283E-01	2.283E-01
0.35	1.534E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.067E-04	2.808E+01	2.338E-01	2.338E-01
0.4	1.755E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.118E-04	2.808E+01	2.398E-01	2.398E-01
0.45	1.976E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.169E-04	2.808E+01	2.463E-01	2.463E-01
0.5	2.197E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.220E-04	2.808E+01	2.533E-01	2.533E-01
0.55	2.418E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.271E-04	2.808E+01	2.608E-01	2.608E-01
0.6	2.639E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.322E-04	2.808E+01	2.688E-01	2.688E-01
0.65	2.860E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.373E-04	2.808E+01	2.773E-01	2.773E-01
0.7	3.081E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.424E-04	2.808E+01	2.863E-01	2.863E-01
0.75	3.302E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.475E-04	2.808E+01	2.958E-01	2.958E-01
0.8	3.523E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.526E-04	2.808E+01	3.058E-01	3.058E-01
0.85	3.744E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.577E-04	2.808E+01	3.163E-01	3.163E-01
0.9	3.965E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.628E-04	2.808E+01	3.273E-01	3.273E-01
0.95	4.186E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.679E-04	2.808E+01	3.388E-01	3.388E-01
1.0	4.407E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.730E-04	2.808E+01	3.508E-01	3.508E-01
1.05	4.628E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.781E-04	2.808E+01	3.633E-01	3.633E-01
1.1	4.849E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.832E-04	2.808E+01	3.763E-01	3.763E-01
1.15	5.070E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.883E-04	2.808E+01	3.898E-01	3.898E-01
1.2	5.291E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.934E-04	2.808E+01	4.038E-01	4.038E-01
1.25	5.512E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	6.985E-04	2.808E+01	4.183E-01	4.183E-01
1.3	5.733E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.036E-04	2.808E+01	4.333E-01	4.333E-01
1.35	5.954E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.087E-04	2.808E+01	4.488E-01	4.488E-01
1.4	6.175E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.138E-04	2.808E+01	4.643E-01	4.643E-01
1.45	6.396E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.189E-04	2.808E+01	4.798E-01	4.798E-01
1.5	6.617E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.240E-04	2.808E+01	4.953E-01	4.953E-01
1.55	6.838E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.291E-04	2.808E+01	5.108E-01	5.108E-01
1.6	7.059E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.342E-04	2.808E+01	5.263E-01	5.263E-01
1.65	7.280E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.393E-04	2.808E+01	5.418E-01	5.418E-01
1.7	7.501E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.444E-04	2.808E+01	5.573E-01	5.573E-01
1.75	7.722E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.495E-04	2.808E+01	5.728E-01	5.728E-01
1.8	7.943E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.546E-04	2.808E+01	5.883E-01	5.883E-01
1.85	8.164E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.597E-04	2.808E+01	6.038E-01	6.038E-01
1.9	8.385E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.648E-04	2.808E+01	6.193E-01	6.193E-01
1.95	8.606E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.699E-04	2.808E+01	6.348E-01	6.348E-01
2.0	8.827E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.750E-04	2.808E+01	6.503E-01	6.503E-01
2.05	9.048E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.801E-04	2.808E+01	6.658E-01	6.658E-01
2.1	9.269E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.852E-04	2.808E+01	6.813E-01	6.813E-01
2.15	9.490E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.903E-04	2.808E+01	6.968E-01	6.968E-01
2.2	9.711E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	7.954E-04	2.808E+01	7.123E-01	7.123E-01
2.25	9.932E-02	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.005E-04	2.808E+01	7.278E-01	7.278E-01
2.3	1.0153E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.056E-04	2.808E+01	7.433E-01	7.433E-01
2.35	1.0374E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.107E-04	2.808E+01	7.588E-01	7.588E-01
2.4	1.0595E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.158E-04	2.808E+01	7.743E-01	7.743E-01
2.45	1.0816E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.209E-04	2.808E+01	7.898E-01	7.898E-01
2.5	1.1037E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.260E-04	2.808E+01	8.053E-01	8.053E-01
2.55	1.1258E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.311E-04	2.808E+01	8.208E-01	8.208E-01
2.6	1.1479E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.362E-04	2.808E+01	8.363E-01	8.363E-01
2.65	1.1699E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.413E-04	2.808E+01	8.518E-01	8.518E-01
2.7	1.1920E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.464E-04	2.808E+01	8.673E-01	8.673E-01
2.75	1.2141E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.515E-04	2.808E+01	8.828E-01	8.828E-01
2.8	1.2362E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.566E-04	2.808E+01	8.983E-01	8.983E-01
2.85	1.2583E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.617E-04	2.808E+01	9.138E-01	9.138E-01
2.9	1.2804E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.668E-04	2.808E+01	9.293E-01	9.293E-01
2.95	1.3025E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.719E-04	2.808E+01	9.448E-01	9.448E-01
3.0	1.3246E-01	3.402E-03	3.402E-03	2.460E-01	-1.631E+02	8.770E-04	2.808E+01	9.603E-01	9.603E-01

0.550	4.124E-01	5.740E-01	-1.427E-02	8.318E-01	2.178E+02	-1.120E-02	-5.935E+02	8.598E-01
0.550	4.677E-01	6.571E-01	-1.620E-02	8.162E-01	2.245E+02	-1.302E-02	-6.892E+02	8.718E-01
0.610	5.018E-01	6.344E-01	-1.846E-02	8.403E-01	2.311E+02	-1.491E-02	-7.995E+02	8.828E-01
0.650	5.144E-01	6.014E-01	-2.073E-02	8.441E-01	2.377E+02	-1.693E-02	-8.974E+02	8.927E-01
0.670	5.264E-01	6.014E-01	-2.310E-02	8.477E-01	2.454E+02	-1.906E-02	-1.010E+03	9.010E-01
0.700	5.314E-01	7.837E-01	-2.556E-02	8.512E-01	2.531E+02	-2.129E-02	-1.128E+03	9.104E-01
0.730	5.427E-01	7.871E-01	-2.812E-02	8.546E-01	2.622E+02	-2.363E-02	-1.253E+03	9.189E-01
0.760	5.540E-01	7.740E-01	-3.078E-02	8.581E-01	2.712E+02	-2.609E-02	-1.383E+03	9.258E-01
0.790	5.654E-01	8.010E-01	-3.354E-02	8.616E-01	2.808E+02	-2.867E-02	-1.520E+03	9.328E-01
0.820	5.769E-01	8.210E-01	-3.639E-02	8.661E-01	3.004E+02	-3.139E-02	-1.664E+03	9.406E-01
0.850	5.884E-01	8.501E-01	-3.934E-02	8.711E-01	3.361E+02	-3.425E-02	-1.815E+03	9.475E-01
0.880	6.014E-01	8.877E-01	-4.239E-02	8.780E-01	3.718E+02	-3.732E-02	-1.978E+03	9.532E-01
0.910	6.144E-01	9.144E-01	-4.554E-02	8.870E-01	5.417E+02	-4.063E-02	-2.153E+03	9.603E-01
0.940	6.274E-01	9.434E-01	-4.879E-02	9.054E-01	7.117E+02	-4.455E-02	-2.361E+03	9.721E-01
0.970	6.404E-01	9.718E-01	-5.213E-02	9.342E-01	1.357E+03	-4.923E-02	-2.609E+03	9.850E-01
1.000	1.072E+00	1.070E+00	-5.557E-02	1.000E+00	2.003E+03	-5.627E-02	-2.982E+03	1.010E+00

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ORIGINAL PAGE IS  
OF POOR QUALITY

## -SHOCK LAYER GAS PROPERTIES-

ETA	V/D	P (ATM.)	T (DEG.K/°C)	RHO (SLUGS/FT <sup>3</sup> )	MU (LBM/FT-SEC)	RMU (LBM <sup>2</sup> -SEC/FT <sup>6</sup> )	K (BTU/FT-SEC-R)
0.0	0.0	3.1320E-01	3.5775E+03	2.7404E-05	6.0764E-05	1.6773E-09	1.0847E-04
0.0050	1.111E-01	3.1314E-01	3.6144E+03	2.7044E-05	6.1562E-05	1.6649E-09	1.0969E-04
0.0111	2.222E-01	3.1311E-01	3.6562E+03	2.6394E-05	6.2278E-05	1.6437E-09	1.1094E-04
0.0150	3.333E-01	3.1310E-01	3.6981E+03	2.5981E-05	6.2706E-05	1.6291E-09	1.1180E-04
0.0200	4.444E-01	3.1310E-01	3.7180E+03	2.5551E-05	6.3249E-05	1.6098E-09	1.1294E-04
0.0231	5.555E-01	3.1310E-01	3.7401E+03	2.5122E-05	6.3590E-05	1.5976E-09	1.1371E-04
0.0263	6.666E-01	3.1310E-01	3.7613E+03	2.4743E-05	6.3948E-05	1.5848E-09	1.1449E-04
0.0294	7.777E-01	3.1310E-01	3.7826E+03	2.4353E-05	6.4300E-05	1.5723E-09	1.1527E-04
0.0325	8.888E-01	3.1310E-01	3.8037E+03	2.3962E-05	6.4660E-05	1.5598E-09	1.1605E-04
0.0356	9.999E-01	3.1310E-01	3.8248E+03	2.3570E-05	6.5027E-05	1.5471E-09	1.1683E-04
0.0387	1.111E-01	3.1310E-01	3.8459E+03	2.3178E-05	6.5391E-05	1.5339E-09	1.1777E-04
0.0418	1.222E-01	3.1310E-01	3.8670E+03	2.2786E-05	6.5751E-05	1.5201E-09	1.1872E-04
0.0449	1.333E-01	3.1310E-01	3.8881E+03	2.2394E-05	6.6111E-05	1.5063E-09	1.1952E-04
0.0480	1.444E-01	3.1310E-01	3.9092E+03	2.1999E-05	6.6471E-05	1.4925E-09	1.2032E-04
0.0511	1.555E-01	3.1310E-01	3.9303E+03	2.1604E-05	6.6831E-05	1.4787E-09	1.2112E-04
0.0542	1.666E-01	3.1310E-01	3.9514E+03	2.1209E-05	6.7191E-05	1.4649E-09	1.2192E-04
0.0573	1.777E-01	3.1310E-01	3.9725E+03	2.0814E-05	6.7551E-05	1.4511E-09	1.2272E-04
0.0604	1.888E-01	3.1310E-01	3.9936E+03	2.0419E-05	6.7911E-05	1.4373E-09	1.2352E-04
0.0635	1.999E-01	3.1310E-01	4.0147E+03	2.0024E-05	6.8271E-05	1.4235E-09	1.2432E-04
0.0666	2.111E-01	3.1310E-01	4.0358E+03	1.9629E-05	6.8631E-05	1.4097E-09	1.2512E-04
0.0697	2.222E-01	3.1310E-01	4.0569E+03	1.9234E-05	6.8991E-05	1.3959E-09	1.2592E-04
0.0728	2.333E-01	3.1310E-01	4.0780E+03	1.8839E-05	6.9351E-05	1.3821E-09	1.2672E-04
0.0759	2.444E-01	3.1310E-01	4.0991E+03	1.8444E-05	6.9711E-05	1.3683E-09	1.2752E-04
0.0790	2.555E-01	3.1310E-01	4.1202E+03	1.8049E-05	7.0071E-05	1.3545E-09	1.2832E-04
0.0821	2.666E-01	3.1310E-01	4.1413E+03	1.7654E-05	7.0431E-05	1.3407E-09	1.2912E-04
0.0852	2.777E-01	3.1310E-01	4.1624E+03	1.7259E-05	7.0791E-05	1.3269E-09	1.2992E-04
0.0883	2.888E-01	3.1310E-01	4.1835E+03	1.6864E-05	7.1151E-05	1.3131E-09	1.3072E-04
0.0914	2.999E-01	3.1310E-01	4.2046E+03	1.6469E-05	7.1511E-05	1.2993E-09	1.3152E-04
0.0945	3.111E-01	3.1310E-01	4.2257E+03	1.6074E-05	7.1871E-05	1.2855E-09	1.3232E-04
0.0976	3.222E-01	3.1310E-01	4.2468E+03	1.5679E-05	7.2231E-05	1.2717E-09	1.3312E-04
0.1007	3.333E-01	3.1310E-01	4.2679E+03	1.5284E-05	7.2591E-05	1.2579E-09	1.3392E-04
0.1038	3.444E-01	3.1310E-01	4.2890E+03	1.4889E-05	7.2951E-05	1.2441E-09	1.3472E-04
0.1069	3.555E-01	3.1310E-01	4.3101E+03	1.4494E-05	7.3311E-05	1.2303E-09	1.3552E-04
0.1100	3.666E-01	3.1310E-01	4.3312E+03	1.4099E-05	7.3671E-05	1.2165E-09	1.3632E-04
0.1131	3.777E-01	3.1310E-01	4.3523E+03	1.3704E-05	7.4031E-05	1.2027E-09	1.3712E-04
0.1162	3.888E-01	3.1310E-01	4.3734E+03	1.3309E-05	7.4391E-05	1.1889E-09	1.3792E-04
0.1193	3.999E-01	3.1310E-01	4.3945E+03	1.2914E-05	7.4751E-05	1.1751E-09	1.3872E-04
0.1224	4.111E-01	3.1310E-01	4.4156E+03	1.2519E-05	7.5111E-05	1.1613E-09	1.3952E-04
0.1255	4.222E-01	3.1310E-01	4.4367E+03	1.2124E-05	7.5471E-05	1.1475E-09	1.4032E-04
0.1286	4.333E-01	3.1310E-01	4.4578E+03	1.1729E-05	7.5831E-05	1.1337E-09	1.4112E-04
0.1317	4.444E-01	3.1310E-01	4.4789E+03	1.1334E-05	7.6191E-05	1.1199E-09	1.4192E-04
0.1348	4.555E-01	3.1310E-01	4.4999E+03	1.0939E-05	7.6551E-05	1.1061E-09	1.4272E-04
0.1379	4.666E-01	3.1310E-01	4.5210E+03	1.0544E-05	7.6911E-05	1.0923E-09	1.4352E-04
0.1410	4.777E-01	3.1310E-01	4.5421E+03	1.0149E-05	7.7271E-05	1.0785E-09	1.4432E-04
0.1441	4.888E-01	3.1310E-01	4.5632E+03	9.7544E-06	7.7631E-05	1.0647E-09	1.4512E-04
0.1472	4.999E-01	3.1310E-01	4.5843E+03	9.3649E-06	7.7991E-05	1.0509E-09	1.4592E-04
0.1503	5.111E-01	3.1310E-01	4.6054E+03	8.9754E-06	7.8351E-05	1.0371E-09	1.4672E-04
0.1534	5.222E-01	3.1310E-01	4.6265E+03	8.5859E-06	7.8711E-05	1.0233E-09	1.4752E-04
0.1565	5.333E-01	3.1310E-01	4.6476E+03	8.1964E-06	7.9071E-05	1.0095E-09	1.4832E-04
0.1596	5.444E-01	3.1310E-01	4.6687E+03	7.8069E-06	7.9431E-05	9.957E-10	1.4912E-04
0.1627	5.555E-01	3.1310E-01	4.6898E+03	7.4174E-06	7.9791E-05	9.819E-10	1.4992E-04
0.1658	5.666E-01	3.1310E-01	4.7109E+03	7.0279E-06	8.0151E-05	9.681E-10	1.5072E-04
0.1689	5.777E-01	3.1310E-01	4.7320E+03	6.6384E-06	8.0511E-05	9.543E-10	1.5152E-04
0.1720	5.888E-01	3.1310E-01	4.7531E+03	6.2489E-06	8.0871E-05	9.405E-10	1.5232E-04
0.1751	5.999E-01	3.1310E-01	4.7742E+03	5.8594E-06	8.1231E-05	9.267E-10	1.5312E-04
0.1782	6.111E-01	3.1310E-01	4.7953E+03	5.4699E-06	8.1591E-05	9.129E-10	1.5392E-04
0.1813	6.222E-01	3.1310E-01	4.8164E+03	5.0804E-06	8.1951E-05	8.991E-10	1.5472E-04
0.1844	6.333E-01	3.1310E-01	4.8375E+03	4.6909E-06	8.2311E-05	8.853E-10	1.5552E-04
0.1875	6.444E-01	3.1310E-01	4.8586E+03	4.3014E-06	8.2671E-05	8.715E-10	1.5632E-04
0.1906	6.555E-01	3.1310E-01	4.8797E+03	3.9119E-06	8.3031E-05	8.577E-10	1.5712E-04
0.1937	6.666E-01	3.1310E-01	4.9008E+03	3.5224E-06	8.3391E-05	8.439E-10	1.5792E-04
0.1968	6.777E-01	3.1310E-01	4.9219E+03	3.1329E-06	8.3751E-05	8.301E-10	1.5872E-04
0.1999	6.888E-01	3.1310E-01	4.9430E+03	2.7434E-06	8.4111E-05	8.163E-10	1.5952E-04
0.2030	6.999E-01	3.1310E-01	4.9641E+03	2.3539E-06	8.4471E-05	8.025E-10	1.6032E-04
0.2061	7.111E-01	3.1310E-01	4.9852E+03	1.9644E-06	8.4831E-05	7.887E-10	1.6112E-04
0.2092	7.222E-01	3.1310E-01	5.0063E+03	1.5749E-06	8.5191E-05	7.749E-10	1.6192E-04
0.2123	7.333E-01	3.1310E-01	5.0274E+03	1.1854E-06	8.5551E-05	7.611E-10	1.6272E-04
0.2154	7.444E-01	3.1310E-01	5.0485E+03	1.0000E-06	8.5911E-05	7.473E-10	1.6352E-04
0.2185	7.555E-01	3.1310E-01	5.0696E+03	8.1455E-07	8.6271E-05	7.335E-10	1.6432E-04
0.2216	7.666E-01	3.1310E-01	5.0907E+03	6.2900E-07	8.6631E-05	7.197E-10	1.6512E-04
0.2247	7.777E-01	3.1310E-01	5.1118E+03	4.4349E-07	8.6991E-05	7.059E-10	1.6592E-04
0.2278	7.888E-01	3.1310E-01	5.1329E+03	2.5798E-07	8.7351E-05	6.921E-10	1.6672E-04
0.2309	7.999E-01	3.1310E-01	5.1540E+03	7.1119E-08	8.7711E-05	6.783E-10	1.6752E-04
0.2340	8.111E-01	3.1310E-01	5.1751E+03	5.2568E-08	8.8071E-05	6.645E-10	1.6832E-04
0.2371	8.222E-01	3.1310E-01	5.1962E+03	3.4017E-08	8.8431E-05	6.507E-10	1.6912E-04
0.2402	8.333E-01	3.1310E-01	5.2173E+03	1.5466E-08	8.8791E-05	6.369E-10	1.6992E-04
0.2433	8.444E-01	3.1310E-01	5.2384E+03	9.6119E-09	8.9151E-05	6.231E-10	1.7072E-04
0.2464	8.555E-01	3.1310E-01	5.2595E+03	7.7568E-09	8.9511E-05	6.093E-10	1.7152E-04
0.2495	8.666E-01	3.1310E-01	5.2806E+03	5.9017E-09	8.9871E-05	5.955E-10	1.7232E-04
0.2526	8.777E-01	3.1310E-01	5.3017E+03	4.0466E-09	9.0231E-05	5.817E-10	1.7312E-04
0.2557	8.888E-01	3.1310E-01	5.3228E+03	2.1915E-09	9.0591E-05	5.679E-10	1.7392E-04
0.2588	8.999E-01	3.1310E-01	5.3439E+03	3.3364E-10	9.0951E-05	5.541E-10	1.7472E-04
0.2619	9.111E-01	3.1310E-01	5.3650E+03	1.4813E-10	9.1311E-05	5.403E-10	1.7552E-04
0.2650	9.222E-01	3.1310E-01	5.3861E+03	9.9589E-11	9.1671E-05	5.265E-10	1.7632E-04
0.2681	9.333E-01	3.1310E-01	5.4072E+03	8.1038E-11	9.2031E-05	5.127E-10	1.7712E-04
0.2712	9.444E-01	3.1310E-01	5.4283E+03	6.2487E-11	9.2391E-05	4.989E-10	1.7792E-04
0.2743	9.555E-01	3.1310E-01	5.4494E+03	4.3936E-11	9.2751E-05	4.851E-10	1.7872E-04
0.2774	9.666E-01	3.1310E-01	5.4705E+03	2.5385E-11	9.3111E-05	4.713E-10	1.7952E-04
0.2805	9.777E-01	3.1310E-01	5.4916E+03	6.8119E-12	9.3471E-05	4.575E-10	1.8032E-04
0.2836	9.888E-01	3.1310E-01	5.5127E+03	4.9568E-12	9.3831E-05	4.437E-10	1.8112E-04
0.2867	9.999E-01	3.1310E-01	5.5338E+03	3.1017E-12	9.4191E-05	4.299E-10	1.8192E-04
0.2898	1.011E-01	3.1310E-01	5.5549E+03	1.2466E-12	9.4551E-05	4.161E-10	1.8272E-04
0.2929	1.022E-01	3.1310E-01	5.5760E+03	9.6119E-13	9.4911E-05	4.023E-10	1.8352E-04
0.2960	1.033E-01	3.1310E-01	5.5971E+03	7.7568E-13	9.5271E-05	3.885E-10	1.8432E-04
0.2991	1.044E-01	3.1310E-01	5.6182E+03	5.9017E-13	9.5631E-05	3.747E-10	1.8512E-04
0.3022	1.055E-01	3.1310E-01	5.6393E+03	4.0466E-13	9.5991E-05	3.609E-10	1.8592E-04
0.3053	1.066E-01	3.1310E-01	5.6604E+03	2.1915E-13	9.6351E-05	3.471E-10	1.8672E-04
0.3084	1.077E-01	3.1310E-01	5.6815E+03	3.3364E-14	9.6711E-05	3.333E-10	1.8752E-04
0.3115	1.088E-01	3.1310E-01	5.7026E+03	1.4813E-14	9.7071E-05	3.195E-10	1.8832E-04
0.3146	1.099E-01	3.1310E-01	5.7237E+03	9.9589E-15	9.7431E-05	3.057E-10	1.8912E-04
0.3177	1.110E-01	3.1310E-01	5.7448E+03	8.1038E-15	9.7791E-05	2.919E-10	1.8992E-04
0.3208	1.121E-01	3.1310E-01	5.7659E+03	6.2487E-15	9.8151E-05	2.781E-1	

0.9170	4.7671E-01	3.1111E-01	1.2808E+04	4.9727E-04	7.7631E-05	3.6404E-10	2.2727E-04
0.9400	4.1607E-01	3.1111E-01	1.3166E+04	4.8533E-04	7.8201E-05	3.7992E-10	2.2826E-04
0.9700	4.5644E-01	3.1111E-01	1.3585E+04	4.8081E-04	7.9405E-05	3.7385E-10	2.2931E-04
1.0000	1.5060E+00	3.1120E-01	1.4941E+04	4.3820E-04	8.2562E-05	3.6179E-10	2.3121E-04

ORIGINAL PAGE IS  
OF POOR QUALITY





2.4996E-05  
2.5102E-05

5.1058E-01  
5.1168E-01

1.4448E-01  
1.4620E-01

2.6558E-01  
2.6447E-01

7.7273E-02  
7.5560E-02

2.7186E-05  
2.7353E-05

1.1104E-04  
1.1101E-04

0.9720  
1.0000

ORIGINAL PAGE IS  
OF POOR QUALITY

PTA	C	H	H2	-SPECIES MASS FRACTIONS- CO	C3-G	CN	C2H	C2H2
0.0	2.1021-02	2.2092-02	4.1334-02	2.5699-01	3.6913-01	9.9911-11	1.3871-01	5.8879-02
0.005	2.1710-02	2.4521-02	3.7937-02	2.5690-01	3.6916-01	7.3831-05	1.6368-01	5.8051-02
0.011	2.2631-02	2.5624-02	3.5624-02	2.5692-01	3.6915-01	1.7276-05	1.7111-01	5.6961-02
0.015	2.3149-02	2.7047-02	3.4362-02	2.5694-01	3.6877-01	2.4176-04	1.7226-01	5.6212-02
0.020	2.3627-02	2.8662-02	3.2870-02	2.5696-01	3.6855-01	3.3736-04	1.7202-01	5.5190-02
0.025	2.4114-02	3.0451-02	3.1522-02	2.5697-01	3.6840-01	4.0111-04	1.7149-01	5.4519-02
0.031	2.4514-02	3.1500-02	3.0500-02	2.5699-01	3.6824-01	4.7075-04	1.7074-01	5.3794-02
0.035	2.4934-02	3.2737-02	2.9050-02	2.5700-01	3.6807-01	5.4202-04	1.6930-01	5.3061-02
0.040	2.5364-02	3.4127-02	2.8055-02	2.5702-01	3.6790-01	6.1733-04	1.6799-01	5.2294-02
0.045	2.5804-02	3.5727-02	2.7129-02	2.5704-01	3.6771-01	6.9722-04	1.6651-01	5.1493-02
0.050	2.6244-02	3.7497-02	2.6196-02	2.5706-01	3.6751-01	7.8559-04	1.6511-01	5.0629-02
0.055	2.6684-02	3.9447-02	2.5197-02	2.5710-01	3.6731-01	8.7977-04	1.6368-01	4.9844-02
0.060	2.7124-02	4.1597-02	2.4197-02	2.5711-01	3.6711-01	9.7977-04	1.6211-01	4.9044-02
0.065	2.7564-02	4.3947-02	2.3197-02	2.5712-01	3.6691-01	1.0731-03	1.6044-01	4.8244-02
0.070	2.8004-02	4.6497-02	2.2197-02	2.5713-01	3.6671-01	1.1622-03	1.5871-01	4.7444-02
0.075	2.8444-02	4.9147-02	2.1197-02	2.5714-01	3.6651-01	1.2522-03	1.5694-01	4.6644-02
0.080	2.8884-02	5.1897-02	2.0197-02	2.5715-01	3.6631-01	1.3422-03	1.5517-01	4.5844-02
0.085	2.9324-02	5.4747-02	1.9197-02	2.5716-01	3.6611-01	1.4322-03	1.5340-01	4.5044-02
0.090	2.9764-02	5.7697-02	1.8197-02	2.5717-01	3.6591-01	1.5222-03	1.5163-01	4.4244-02
0.095	3.0204-02	6.0747-02	1.7197-02	2.5718-01	3.6571-01	1.6122-03	1.4986-01	4.3444-02
0.100	3.0644-02	6.3897-02	1.6197-02	2.5719-01	3.6551-01	1.7022-03	1.4809-01	4.2644-02
0.105	3.1084-02	6.7147-02	1.5197-02	2.5720-01	3.6531-01	1.7922-03	1.4632-01	4.1844-02
0.110	3.1524-02	7.0497-02	1.4197-02	2.5721-01	3.6511-01	1.8822-03	1.4455-01	4.1044-02
0.115	3.1964-02	7.3947-02	1.3197-02	2.5722-01	3.6491-01	1.9722-03	1.4278-01	4.0244-02
0.120	3.2404-02	7.7497-02	1.2197-02	2.5723-01	3.6471-01	2.0622-03	1.4101-01	3.9444-02
0.125	3.2844-02	8.1147-02	1.1197-02	2.5724-01	3.6451-01	2.1522-03	1.3924-01	3.8644-02
0.130	3.3284-02	8.4897-02	1.0197-02	2.5725-01	3.6431-01	2.2422-03	1.3747-01	3.7844-02
0.135	3.3724-02	8.8747-02	0.9197-02	2.5726-01	3.6411-01	2.3322-03	1.3570-01	3.7044-02
0.140	3.4164-02	9.2697-02	0.8197-02	2.5727-01	3.6391-01	2.4222-03	1.3393-01	3.6244-02
0.145	3.4604-02	9.6747-02	0.7197-02	2.5728-01	3.6371-01	2.5122-03	1.3216-01	3.5444-02
0.150	3.5044-02	10.0897-02	0.6197-02	2.5729-01	3.6351-01	2.6022-03	1.3039-01	3.4644-02
0.155	3.5484-02	10.5147-02	0.5197-02	2.5730-01	3.6331-01	2.6922-03	1.2862-01	3.3844-02
0.160	3.5924-02	10.9497-02	0.4197-02	2.5731-01	3.6311-01	2.7822-03	1.2685-01	3.3044-02
0.165	3.6364-02	11.3947-02	0.3197-02	2.5732-01	3.6291-01	2.8722-03	1.2508-01	3.2244-02
0.170	3.6804-02	11.8497-02	0.2197-02	2.5733-01	3.6271-01	2.9622-03	1.2331-01	3.1444-02
0.175	3.7244-02	12.3147-02	0.1197-02	2.5734-01	3.6251-01	3.0522-03	1.2154-01	3.0644-02
0.180	3.7684-02	12.7897-02	0.0197-02	2.5735-01	3.6231-01	3.1422-03	1.1977-01	2.9844-02
0.185	3.8124-02	13.2747-02		2.5736-01	3.6211-01	3.2322-03	1.1800-01	2.9044-02
0.190	3.8564-02	13.7697-02		2.5737-01	3.6191-01	3.3222-03	1.1623-01	2.8244-02
0.195	3.9004-02	14.2747-02		2.5738-01	3.6171-01	3.4122-03	1.1446-01	2.7444-02
0.200	3.9444-02	14.7897-02		2.5739-01	3.6151-01	3.5022-03	1.1269-01	2.6644-02
0.205	3.9884-02	15.3147-02		2.5740-01	3.6131-01	3.5922-03	1.1092-01	2.5844-02
0.210	4.0324-02	15.8497-02		2.5741-01	3.6111-01	3.6822-03	1.0915-01	2.5044-02
0.215	4.0764-02	16.3947-02		2.5742-01	3.6091-01	3.7722-03	1.0738-01	2.4244-02
0.220	4.1204-02	16.9497-02		2.5743-01	3.6071-01	3.8622-03	1.0561-01	2.3444-02
0.225	4.1644-02	17.5147-02		2.5744-01	3.6051-01	3.9522-03	1.0384-01	2.2644-02
0.230	4.2084-02	18.0897-02		2.5745-01	3.6031-01	4.0422-03	1.0207-01	2.1844-02
0.235	4.2524-02	18.6747-02		2.5746-01	3.6011-01	4.1322-03	1.0030-01	2.1044-02
0.240	4.2964-02	19.2697-02		2.5747-01	3.5991-01	4.2222-03	0.9853-01	2.0244-02
0.245	4.3404-02	19.8747-02		2.5748-01	3.5971-01	4.3122-03	0.9676-01	1.9444-02
0.250	4.3844-02	20.4897-02		2.5749-01	3.5951-01	4.4022-03	0.9499-01	1.8644-02
0.255	4.4284-02	21.1147-02		2.5750-01	3.5931-01	4.4922-03	0.9322-01	1.7844-02
0.260	4.4724-02	21.7497-02		2.5751-01	3.5911-01	4.5822-03	0.9145-01	1.7044-02
0.265	4.5164-02	22.3947-02		2.5752-01	3.5891-01	4.6722-03	0.8968-01	1.6244-02
0.270	4.5604-02	23.0497-02		2.5753-01	3.5871-01	4.7622-03	0.8791-01	1.5444-02
0.275	4.6044-02	23.7147-02		2.5754-01	3.5851-01	4.8522-03	0.8614-01	1.4644-02
0.280	4.6484-02	24.3897-02		2.5755-01	3.5831-01	4.9422-03	0.8437-01	1.3844-02
0.285	4.6924-02	25.0747-02		2.5756-01	3.5811-01	5.0322-03	0.8260-01	1.3044-02
0.290	4.7364-02	25.7697-02		2.5757-01	3.5791-01	5.1222-03	0.8083-01	1.2244-02
0.295	4.7804-02	26.4747-02		2.5758-01	3.5771-01	5.2122-03	0.7906-01	1.1444-02
0.300	4.8244-02	27.1897-02		2.5759-01	3.5751-01	5.3022-03	0.7729-01	1.0644-02
0.305	4.8684-02	27.9147-02		2.5760-01	3.5731-01	5.3922-03	0.7552-01	0.9844-02
0.310	4.9124-02	28.6497-02		2.5761-01	3.5711-01	5.4822-03	0.7375-01	0.9044-02
0.315	4.9564-02	29.3947-02		2.5762-01	3.5691-01	5.5722-03	0.7198-01	0.8244-02
0.320	5.0004-02	30.1497-02		2.5763-01	3.5671-01	5.6622-03	0.7021-01	0.7444-02
0.325	5.0444-02	30.9147-02		2.5764-01	3.5651-01	5.7522-03	0.6844-01	0.6644-02
0.330	5.0884-02	31.6897-02		2.5765-01	3.5631-01	5.8422-03	0.6667-01	0.5844-02
0.335	5.1324-02	32.4747-02		2.5766-01	3.5611-01	5.9322-03	0.6490-01	0.5044-02
0.340	5.1764-02	33.2697-02		2.5767-01	3.5591-01	6.0222-03	0.6313-01	0.4244-02
0.345	5.2204-02	34.0747-02		2.5768-01	3.5571-01	6.1122-03	0.6136-01	0.3444-02
0.350	5.2644-02	34.8897-02		2.5769-01	3.5551-01	6.2022-03	0.5959-01	0.2644-02
0.355	5.3084-02	35.7147-02		2.5770-01	3.5531-01	6.2922-03	0.5782-01	0.1844-02
0.360	5.3524-02	36.5497-02		2.5771-01	3.5511-01	6.3822-03	0.5605-01	0.1044-02
0.365	5.3964-02	37.3947-02		2.5772-01	3.5491-01	6.4722-03	0.5428-01	0.0244-02
0.370	5.4404-02	38.2497-02		2.5773-01	3.5471-01	6.5622-03	0.5251-01	
0.375	5.4844-02	39.1147-02		2.5774-01	3.5451-01	6.6522-03	0.5074-01	
0.380	5.5284-02	40.0897-02		2.5775-01	3.5431-01	6.7422-03	0.4897-01	
0.385	5.5724-02	41.0747-02		2.5776-01	3.5411-01	6.8322-03	0.4720-01	
0.390	5.6164-02	42.0697-02		2.5777-01	3.5391-01	6.9222-03	0.4543-01	
0.395	5.6604-02	43.0747-02		2.5778-01	3.5371-01	7.0122-03	0.4366-01	
0.400	5.7044-02	44.0897-02		2.5779-01	3.5351-01	7.1022-03	0.4189-01	
0.405	5.7484-02	45.1147-02		2.5780-01	3.5331-01	7.1922-03	0.4012-01	
0.410	5.7924-02	46.1497-02		2.5781-01	3.5311-01	7.2822-03	0.3835-01	
0.415	5.8364-02	47.1947-02		2.5782-01	3.5291-01	7.3722-03	0.3658-01	
0.420	5.8804-02	48.2497-02		2.5783-01	3.5271-01	7.4622-03	0.3481-01	
0.425	5.9244-02	49.3147-02		2.5784-01	3.5251-01	7.5522-03	0.3304-01	
0.430	5.9684-02	50.3897-02		2.5785-01	3.5231-01	7.6422-03	0.3127-01	
0.435	6.0124-02	51.4747-02		2.5786-01	3.5211-01	7.7322-03	0.2950-01	
0.440	6.0564-02	52.5697-02		2.5787-01	3.5191-01	7.8222-03	0.2773-01	
0.445	6.1004-02	53.6747-02		2.5788-01	3.5171-01	7.9122-03	0.2596-01	
0.450	6.1444-02	54.7897-02		2.5789-01	3.5151-01	8.0022-03	0.2419-01	
0.455	6.1884-02	55.9147-02		2.5790-01	3.5131-01	8.0922-03	0.2242-01	
0.460	6.2324-02	57.0497-02		2.5791-01	3.5111-01	8.1822-03	0.2065-01	
0.465	6.2764-02	58.1947-02		2.5792-01	3.5091-01	8.2722-03	0.1888-01	
0.470	6.3204-02	59.3497-02		2.5793-01	3.5071-01	8.3622-03	0.1711-01	
0.475	6.3644-02	60.5147-02		2.5794-01	3.5051-01	8.4522-03	0.1534-01	
0.480	6.4084-02	61.6897-02		2.5795-01	3.5031-01	8.5422-03	0.1357-01	
0.485	6.4524-02	62.8747-02		2.5796-01	3.5011-01	8.6322-03	0.1180-01	
0.490	6.4964-02	64.0697-02		2.5797-01	3.4991-01	8.7222-03	0.1003-01	
0.495	6.5404-02	65.2747-02		2.5798-01	3.4971-01	8.8122-03	0.0826-01	
0.500	6.5844-02	66.4897-02		2.5799-01	3.4951-01	8.9022-03	0.0649-01	

0.2700	2.11071-04	1.01744-03	6.8803E-11	2.0036E-09	3.9472E-22	3.3552E-09	1.2442E-19	2.4760E-27
1.0750	1.7444E-04	1.01744-03	3.9472E-11	1.1911E-09	1.3325E-36	1.2603E-09	6.7016E-22	4.4125E-43

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0.0700	4.4644E-01	2.2504E-10	1.0994E-38	1.0204E-17	1.1128E-12	7.9416E-04	8.6266E+00
1.0500	7.1127E-01	2.5757E-03	1.7215E-50	3.6816E-18	1.7726E-13	8.8339E-04	6.8122E+00

TIME LEFT = 2030446 MSFC

C A S E 2      53,000 FT/SEC      2.5E-7 SLUGS/CU FT

BOUNDARY CONDITIONS AT  $XI = 0.10$

BOB SHOCK CONDITIONS

SHOCK TEMPERATURE (OK)	=	14473.	NORMAL VELOCITY	=	-0.0559
SHOCK PRESSURE (ATM)	=	0.3101	TANGENTIAL VELOCITY	=	0.0996
SHOCK ANGLE, PHI (RAD)	=	0.0998	ANGLE EPS (RAD)	=	0.0002

ABLATOR SURFACE

SURFACE TEMPERATURE (OK)	=	3573.	QCOND (BTU/FT2-SEC)	=	207.
SURFACE PRESSURE (ATM)	=	0.3090	QSUBL (BTU/FT2-SEC)	=	116.
PLUMING MAT.	=	0.0582	QRR (BTU/FT2-SEC)	=	538.
MASS LOSS (LB/FT2-SEC)	=	0.0248	QTOTAL (BTU/FT2-SEC)	=	855.

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[illegible]

0.9100	0.7705E-01	3.0981E-01	1.2658E+04	4.9626E-06	7.8374E-05	3.8804E-10	2.2615E-04
0.9400	0.1658E-01	3.0392E-01	1.3119E+04	4.8519E-06	7.8999E-05	3.8330E-10	2.2709E-04
0.9700	0.3723E-01	3.1301E-01	1.3332E+04	4.8961E-06	8.0259E-05	3.7890E-10	2.2817E-04
1.0000	1.0000E+00	3.1011E-01	1.4473E+04	4.3856E-06	8.3480E-05	3.6611E-10	2.2998E-04

END PAGE 13  
QUALITY

ETA	1/2	M2	D	N	O6	M4	E-
9.9963E-11	4.9963E-02	2.5590E-09	9.9963E-11	9.9967E-11	9.9963E-11	9.9963E-11	9.9963E-11
1.0000E-09	4.9963E-02	3.5029E-07	1.0591E-07	1.3063E-06	1.3063E-06	3.2592E-05	1.3229E-09
3.1700E-02	4.9963E-02	7.6661E-07	4.9619E-07	3.0547E-06	3.0547E-06	7.6219E-05	9.0938E-09
4.9963E-02	4.9963E-02	1.6763E-06	7.6856E-07	4.2744E-06	4.2744E-06	1.0665E-04	4.3291E-09
6.0017E-06	4.9963E-02	1.5169E-06	1.3147E-06	5.9651E-06	5.9651E-06	1.4683E-04	6.0412E-09
4.9963E-02	4.9963E-02	1.6240E-06	1.7672E-06	7.0937E-06	7.0937E-06	1.7699E-04	7.1091E-09
4.9963E-02	4.9963E-02	2.1751E-06	2.7632E-06	8.3641E-06	8.3641E-06	2.0778E-04	8.4339E-09
4.9963E-02	4.9963E-02	2.5660E-06	3.0960E-06	9.5964E-06	9.5964E-06	2.3941E-04	9.7160E-09
4.9963E-02	4.9963E-02	2.5867E-06	4.0334E-06	1.0342E-05	1.0342E-05	2.7296E-04	1.1060E-08
4.9963E-02	4.9963E-02	3.6790E-06	5.2160E-06	1.2170E-05	1.2170E-05	3.0957E-04	1.2525E-08
4.9963E-02	4.9963E-02	4.0646E-06	6.7640E-06	1.3938E-05	1.3938E-05	3.4766E-04	1.4112E-08
4.9963E-02	4.9963E-02	5.5653E-06	1.0661E-05	1.7271E-05	1.7271E-05	4.3071E-04	1.7483E-08
4.9963E-02	4.9963E-02	5.9045E-06	2.1117E-05	2.2783E-05	2.2783E-05	5.6794E-04	2.3034E-08
4.9963E-02	4.9963E-02	6.4095E-06	3.9942E-05	2.9376E-05	2.9376E-05	7.3167E-04	2.9702E-08
4.9963E-02	4.9963E-02	1.3495E-05	7.9846E-05	3.7469E-05	3.7469E-05	9.3314E-04	3.7893E-08
4.9963E-02	4.9963E-02	2.2122E-05	1.3786E-04	4.7363E-05	4.7363E-05	1.1775E-04	4.7807E-08
4.9963E-02	4.9963E-02	3.7137E-05	1.9608E-04	5.4105E-05	5.4105E-05	1.3440E-04	5.4589E-08
4.9963E-02	4.9963E-02	5.1109E-05	3.2511E-04	8.4911E-05	8.4911E-05	1.6348E-04	6.4303E-08
4.9963E-02	4.9963E-02	6.2514E-05	5.1232E-04	8.0259E-05	8.0259E-05	1.9870E-04	8.0695E-08
4.9963E-02	4.9963E-02	1.3547E-04	9.5418E-04	1.0436E-04	1.0436E-04	2.5758E-04	1.0463E-07
4.9963E-02	4.9963E-02	2.4450E-04	1.4065E-04	1.4619E-04	1.4619E-04	3.3876E-04	1.3606E-07
4.9963E-02	4.9963E-02	3.3391E-04	1.9567E-04	1.7661E-04	1.7661E-04	4.3701E-04	1.7763E-07
4.9963E-02	4.9963E-02	1.1009E-03	3.1567E-03	2.3550E-04	2.3550E-04	5.7272E-03	2.3291E-07
4.9963E-02	4.9963E-02	2.2557E-03	5.2663E-03	3.1217E-04	3.1217E-04	7.3575E-03	3.0667E-07
4.9963E-02	4.9963E-02	4.7819E-03	8.3468E-03	4.1582E-04	4.1582E-04	9.9482E-03	4.0522E-07
4.9963E-02	4.9963E-02	9.4050E-03	1.2569E-02	5.5268E-04	5.5268E-04	1.3164E-02	5.3688E-07
4.9963E-02	4.9963E-02	1.3760E-02	1.0274E-02	7.4685E-04	7.4685E-04	1.7446E-02	7.1266E-07
4.9963E-02	4.9963E-02	2.8036E-02	2.7436E-02	4.8948E-02	4.8948E-02	3.0243E-02	1.2418E-06
4.9963E-02	4.9963E-02	7.6744E-02	6.5369E-02	1.8216E-03	1.8216E-03	4.0152E-02	1.6575E-06
4.9963E-02	4.9963E-02	5.1589E-02	8.4657E-02	2.4746E-03	2.4746E-03	5.2921E-02	2.1799E-06
4.9963E-02	4.9963E-02	1.1794E-02	1.1264E-01	3.3340E-03	3.3340E-03	8.8944E-02	2.4940E-06
4.9963E-02	4.9963E-02	1.4115E-01	1.4386E-01	4.4340E-03	4.4340E-03	8.8944E-02	3.4750E-06
4.9963E-02	4.9963E-02	1.5655E-01	1.7884E-01	5.8475E-03	5.8475E-03	1.1117E-01	4.7818E-06
4.9963E-02	4.9963E-02	1.7615E-01	2.1563E-01	7.9474E-03	7.9474E-03	1.3625E-01	5.9388E-06
4.9963E-02	4.9963E-02	1.7698E-01	2.8807E-01	1.1694E-02	1.1694E-02	1.9052E-01	8.4684E-06
4.9963E-02	4.9963E-02	1.8506E-01	3.6976E-01	1.7940E-02	1.7940E-02	2.6758E-01	1.1824E-05
4.9963E-02	4.9963E-02	1.9044E-01	4.8948E-01	2.7404E-02	2.7404E-02	3.2758E-01	1.4711E-05
4.9963E-02	4.9963E-02	1.9791E-01	6.5369E-01	3.7416E-02	3.7416E-02	4.7693E-01	1.5947E-05
4.9963E-02	4.9963E-02	1.8136E-01	3.9500E-01	4.7466E-02	4.7466E-02	3.6036E-01	1.7269E-05
4.9963E-02	4.9963E-02	1.7191E-01	3.7846E-01	5.4209E-02	5.4209E-02	4.1466E-01	1.8294E-05
4.9963E-02	4.9963E-02	1.6253E-01	3.6153E-01	6.7365E-02	6.7365E-02	4.2735E-01	1.9104E-05
4.9963E-02	4.9963E-02	1.5417E-01	3.4920E-01	7.5112E-02	7.5112E-02	4.3855E-01	1.9778E-05
4.9963E-02	4.9963E-02	1.4655E-01	3.3761E-01	8.1981E-02	8.1981E-02	4.4736E-01	2.0360E-05
4.9963E-02	4.9963E-02	1.3678E-01	3.2878E-01	8.8409E-02	8.8409E-02	4.5492E-01	2.0873E-05
4.9963E-02	4.9963E-02	1.2638E-01	3.1224E-01	9.4167E-02	9.4167E-02	4.6166E-01	2.1331E-05
4.9963E-02	4.9963E-02	1.1674E-01	3.1471E-01	9.9413E-02	9.9413E-02	4.6719E-01	2.1743E-05
4.9963E-02	4.9963E-02	1.0737E-01	3.0897E-01	1.0469E-01	1.0469E-01	4.7227E-01	2.2166E-05
4.9963E-02	4.9963E-02	1.0174E-01	3.0385E-01	1.0943E-01	1.0943E-01	4.7691E-01	2.2457E-05
4.9963E-02	4.9963E-02	1.1233E-01	2.9360E-01	1.1447E-01	1.1447E-01	4.8007E-01	2.2749E-05
4.9963E-02	4.9963E-02	1.0737E-01	2.8540E-01	1.1803E-01	1.1803E-01	4.8453E-01	2.3055E-05
4.9963E-02	4.9963E-02	1.0373E-01	2.7164E-01	1.2192E-01	1.2192E-01	4.8749E-01	2.3319E-05
4.9963E-02	4.9963E-02	1.0147E-01	2.6433E-01	1.2556E-01	1.2556E-01	4.9082E-01	2.3567E-05
4.9963E-02	4.9963E-02	9.6174E-02	2.4534E-01	1.2900E-01	1.2900E-01	4.9336E-01	2.3786E-05
4.9963E-02	4.9963E-02	9.2757E-02	2.2626E-01	1.3219E-01	1.3219E-01	4.9599E-01	2.3992E-05
4.9963E-02	4.9963E-02	8.9268E-02	2.0018E-01	1.3516E-01	1.3516E-01	4.9818E-01	2.4181E-05
4.9963E-02	4.9963E-02	8.6596E-02	2.7798E-01	1.3787E-01	1.3787E-01	5.0012E-01	2.4352E-05
4.9963E-02	4.9963E-02	8.3841E-02	2.7604E-01	1.3787E-01	1.3787E-01	5.0012E-01	2.4352E-05
4.9963E-02	4.9963E-02	8.1414E-02	2.7437E-01	1.4034E-01	1.4034E-01	5.0181E-01	2.4508E-05

0.9700	1.1103E-05	2.1504E-05	7.9377E-02	2.7509E-01	1.4238E-01	5.0307E-01	2.4628E-05
1.0000	1.1103E-05	2.8914E-05	7.7799E-02	2.7231E-01	1.4395E-01	5.0384E-01	2.4717E-05

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ETA	C	H2	C3-G	CN	C2H	C2H2
0.0	2.102E-02	2.202E-02	9.133E-02	9.962E-11	1.3071E-01	5.688E-02
0.0050	2.177E-02	2.453E-02	1.754E-01	9.503E-05	1.636E-01	5.802E-02
0.0111	2.264E-02	2.655E-02	3.691E-01	2.221E-04	1.710E-01	5.902E-02
0.0150	2.331E-02	2.731E-02	3.687E-01	3.106E-04	1.721E-01	5.913E-02
0.0200	2.401E-02	2.807E-02	3.684E-01	4.332E-04	1.718E-01	5.907E-02
0.0231	2.474E-02	2.874E-02	3.683E-01	5.148E-04	1.713E-01	5.900E-02
0.0263	2.547E-02	2.947E-02	3.681E-01	6.038E-04	1.705E-01	5.896E-02
0.0294	2.620E-02	3.020E-02	3.679E-01	6.917E-04	1.696E-01	5.893E-02
0.0325	2.693E-02	3.093E-02	3.677E-01	7.917E-04	1.686E-01	5.890E-02
0.0356	2.766E-02	3.166E-02	3.675E-01	8.927E-04	1.674E-01	5.887E-02
0.0388	2.839E-02	3.239E-02	3.673E-01	1.003E-03	1.661E-01	5.884E-02
0.0450	3.011E-02	3.411E-02	3.668E-01	1.241E-03	1.631E-01	5.858E-02
0.0538	3.300E-02	3.700E-02	3.660E-01	1.625E-03	1.578E-01	5.725E-02
0.0625	3.684E-02	4.084E-02	3.650E-01	2.071E-03	1.509E-01	4.256E-02
0.0713	4.068E-02	4.468E-02	3.638E-01	2.593E-03	1.419E-01	3.902E-02
0.0800	4.452E-02	4.852E-02	3.624E-01	3.202E-03	1.307E-01	3.517E-02
0.0888	4.836E-02	5.236E-02	3.608E-01	3.945E-03	1.232E-01	3.262E-02
0.0975	5.220E-02	5.620E-02	3.591E-01	4.807E-03	1.105E-01	2.912E-02
0.1060	5.604E-02	6.004E-02	3.574E-01	5.746E-03	9.633E-02	2.526E-02
0.1100	5.777E-02	6.177E-02	3.562E-01	6.516E-03	9.539E-02	2.000E-02
0.1150	6.061E-02	6.461E-02	3.547E-01	7.065E-03	9.530E-02	1.866E-02
0.1200	6.345E-02	6.745E-02	3.532E-01	7.518E-03	9.530E-02	1.866E-02
0.1250	6.629E-02	7.029E-02	3.517E-01	7.971E-03	9.530E-02	1.866E-02
0.1300	6.913E-02	7.313E-02	3.502E-01	8.424E-03	9.530E-02	1.866E-02
0.1350	7.197E-02	7.597E-02	3.487E-01	8.877E-03	9.530E-02	1.866E-02
0.1400	7.481E-02	7.881E-02	3.472E-01	9.330E-03	9.530E-02	1.866E-02
0.1450	7.765E-02	8.165E-02	3.457E-01	9.783E-03	9.530E-02	1.866E-02
0.1500	8.049E-02	8.449E-02	3.442E-01	1.0236E-02	9.530E-02	1.866E-02
0.1550	8.333E-02	8.733E-02	3.427E-01	1.0689E-02	9.530E-02	1.866E-02
0.1600	8.617E-02	9.017E-02	3.412E-01	1.1142E-02	9.530E-02	1.866E-02
0.1650	8.901E-02	9.301E-02	3.397E-01	1.1595E-02	9.530E-02	1.866E-02
0.1700	9.185E-02	9.585E-02	3.382E-01	1.2048E-02	9.530E-02	1.866E-02
0.1750	9.469E-02	9.869E-02	3.367E-01	1.2501E-02	9.530E-02	1.866E-02
0.1800	9.753E-02	1.015E-01	3.352E-01	1.2954E-02	9.530E-02	1.866E-02
0.1850	1.0037E-01	1.043E-01	3.337E-01	1.3407E-02	9.530E-02	1.866E-02
0.1900	1.0321E-01	1.071E-01	3.322E-01	1.3860E-02	9.530E-02	1.866E-02
0.1950	1.0605E-01	1.099E-01	3.307E-01	1.4313E-02	9.530E-02	1.866E-02
0.2000	1.0889E-01	1.127E-01	3.292E-01	1.4766E-02	9.530E-02	1.866E-02
0.2050	1.1173E-01	1.155E-01	3.277E-01	1.5219E-02	9.530E-02	1.866E-02
0.2100	1.1457E-01	1.183E-01	3.262E-01	1.5672E-02	9.530E-02	1.866E-02
0.2150	1.1741E-01	1.211E-01	3.247E-01	1.6125E-02	9.530E-02	1.866E-02
0.2200	1.2025E-01	1.239E-01	3.232E-01	1.6578E-02	9.530E-02	1.866E-02
0.2250	1.2309E-01	1.267E-01	3.217E-01	1.7031E-02	9.530E-02	1.866E-02
0.2300	1.2593E-01	1.295E-01	3.202E-01	1.7484E-02	9.530E-02	1.866E-02
0.2350	1.2877E-01	1.323E-01	3.187E-01	1.7937E-02	9.530E-02	1.866E-02
0.2400	1.3161E-01	1.351E-01	3.172E-01	1.8390E-02	9.530E-02	1.866E-02
0.2450	1.3445E-01	1.379E-01	3.157E-01	1.8843E-02	9.530E-02	1.866E-02
0.2500	1.3729E-01	1.407E-01	3.142E-01	1.9296E-02	9.530E-02	1.866E-02
0.2550	1.4013E-01	1.435E-01	3.127E-01	1.9749E-02	9.530E-02	1.866E-02
0.2600	1.4297E-01	1.463E-01	3.112E-01	2.0202E-02	9.530E-02	1.866E-02
0.2650	1.4581E-01	1.491E-01	3.097E-01	2.0655E-02	9.530E-02	1.866E-02
0.2700	1.4865E-01	1.519E-01	3.082E-01	2.1108E-02	9.530E-02	1.866E-02
0.2750	1.5149E-01	1.547E-01	3.067E-01	2.1561E-02	9.530E-02	1.866E-02
0.2800	1.5433E-01	1.575E-01	3.052E-01	2.2014E-02	9.530E-02	1.866E-02
0.2850	1.5717E-01	1.603E-01	3.037E-01	2.2467E-02	9.530E-02	1.866E-02
0.2900	1.6001E-01	1.631E-01	3.022E-01	2.2920E-02	9.530E-02	1.866E-02
0.2950	1.6285E-01	1.659E-01	3.007E-01	2.3373E-02	9.530E-02	1.866E-02
0.3000	1.6569E-01	1.687E-01	2.992E-01	2.3826E-02	9.530E-02	1.866E-02
0.3050	1.6853E-01	1.715E-01	2.977E-01	2.4279E-02	9.530E-02	1.866E-02
0.3100	1.7137E-01	1.743E-01	2.962E-01	2.4732E-02	9.530E-02	1.866E-02
0.3150	1.7421E-01	1.771E-01	2.947E-01	2.5185E-02	9.530E-02	1.866E-02
0.3200	1.7705E-01	1.799E-01	2.932E-01	2.5638E-02	9.530E-02	1.866E-02
0.3250	1.7989E-01	1.827E-01	2.917E-01	2.6091E-02	9.530E-02	1.866E-02
0.3300	1.8273E-01	1.855E-01	2.902E-01	2.6544E-02	9.530E-02	1.866E-02
0.3350	1.8557E-01	1.883E-01	2.887E-01	2.6997E-02	9.530E-02	1.866E-02
0.3400	1.8841E-01	1.911E-01	2.872E-01	2.7450E-02	9.530E-02	1.866E-02
0.3450	1.9125E-01	1.939E-01	2.857E-01	2.7903E-02	9.530E-02	1.866E-02
0.3500	1.9409E-01	1.967E-01	2.842E-01	2.8356E-02	9.530E-02	1.866E-02
0.3550	1.9693E-01	1.995E-01	2.827E-01	2.8809E-02	9.530E-02	1.866E-02
0.3600	1.9977E-01	2.023E-01	2.812E-01	2.9262E-02	9.530E-02	1.866E-02
0.3650	2.0261E-01	2.051E-01	2.797E-01	2.9715E-02	9.530E-02	1.866E-02
0.3700	2.0545E-01	2.079E-01	2.782E-01	3.0168E-02	9.530E-02	1.866E-02
0.3750	2.0829E-01	2.107E-01	2.767E-01	3.0621E-02	9.530E-02	1.866E-02
0.3800	2.1113E-01	2.135E-01	2.752E-01	3.1074E-02	9.530E-02	1.866E-02
0.3850	2.1397E-01	2.163E-01	2.737E-01	3.1527E-02	9.530E-02	1.866E-02
0.3900	2.1681E-01	2.191E-01	2.722E-01	3.1980E-02	9.530E-02	1.866E-02
0.3950	2.1965E-01	2.219E-01	2.707E-01	3.2433E-02	9.530E-02	1.866E-02
0.4000	2.2249E-01	2.247E-01	2.692E-01	3.2886E-02	9.530E-02	1.866E-02
0.4050	2.2533E-01	2.275E-01	2.677E-01	3.3339E-02	9.530E-02	1.866E-02
0.4100	2.2817E-01	2.303E-01	2.662E-01	3.3792E-02	9.530E-02	1.866E-02
0.4150	2.3101E-01	2.331E-01	2.647E-01	3.4245E-02	9.530E-02	1.866E-02
0.4200	2.3385E-01	2.359E-01	2.632E-01	3.4698E-02	9.530E-02	1.866E-02
0.4250	2.3669E-01	2.387E-01	2.617E-01	3.5151E-02	9.530E-02	1.866E-02
0.4300	2.3953E-01	2.415E-01	2.602E-01	3.5604E-02	9.530E-02	1.866E-02
0.4350	2.4237E-01	2.443E-01	2.587E-01	3.6057E-02	9.530E-02	1.866E-02
0.4400	2.4521E-01	2.471E-01	2.572E-01	3.6510E-02	9.530E-02	1.866E-02
0.4450	2.4805E-01	2.499E-01	2.557E-01	3.6963E-02	9.530E-02	1.866E-02
0.4500	2.5089E-01	2.527E-01	2.542E-01	3.7416E-02	9.530E-02	1.866E-02
0.4550	2.5373E-01	2.555E-01	2.527E-01	3.7869E-02	9.530E-02	1.866E-02
0.4600	2.5657E-01	2.583E-01	2.512E-01	3.8322E-02	9.530E-02	1.866E-02
0.4650	2.5941E-01	2.611E-01	2.497E-01	3.8775E-02	9.530E-02	1.866E-02
0.4700	2.6225E-01	2.639E-01	2.482E-01	3.9228E-02	9.530E-02	1.866E-02
0.4750	2.6509E-01	2.667E-01	2.467E-01	3.9681E-02	9.530E-02	1.866E-02
0.4800	2.6793E-01	2.695E-01	2.452E-01	4.0134E-02	9.530E-02	1.866E-02
0.4850	2.7077E-01	2.723E-01	2.437E-01	4.0587E-02	9.530E-02	1.866E-02
0.4900	2.7361E-01	2.751E-01	2.422E-01	4.1040E-02	9.530E-02	1.866E-02
0.4950	2.7645E-01	2.779E-01	2.407E-01	4.1493E-02	9.530E-02	1.866E-02
0.5000	2.7929E-01	2.807E-01	2.392E-01	4.1946E-02	9.530E-02	1.866E-02
0.5050	2.8213E-01	2.835E-01	2.377E-01	4.2399E-02	9.530E-02	1.866E-02
0.5100	2.8497E-01	2.863E-01	2.362E-01	4.2852E-02	9.530E-02	1.866E-02
0.5150	2.8781E-01	2.891E-01	2.347E-01	4.3305E-02	9.530E-02	1.866E-02
0.5200	2.9065E-01	2.919E-01	2.332E-01	4.3758E-02	9.530E-02	1.866E-02
0.5250	2.9349E-01	2.947E-01	2.317E-01	4.4211E-02	9.530E-02	1.866E-02
0.5300	2.9633E-01	2.975E-01	2.302E-01	4.4664E-02	9.530E-02	1.866E-02
0.5350	2.9917E-01	3.003E-01	2.287E-01	4.5117E-02	9.530E-02	1.866E-02
0.5400	3.0201E-01	3.031E-01	2.272E-01	4.5570E-02	9.530E-02	1.866E-02
0.5450	3.0485E-01	3.059E-01	2.257E-01	4.6023E-02	9.530E-02	1.866E-02
0.5500	3.0769E-01	3.087E-01	2.242E-01	4.6476E-02	9.530E-02	1.866E-02
0.5550	3.1053E-01	3.115E-01	2.227E-01	4.6929E-02	9.530E-02	1.866E-02
0.5600	3.1337E-01	3.143E-01	2.212E-01	4.7382E-02	9.530E-02	1.866E-02
0.5650	3.1621E-01	3.171E-01	2.197E-01	4.7835E-02	9.530E-02	1.866E-02
0.5700	3.1905E-01	3.199E-01	2.182E-01	4.8288E-02	9.530E-02	1.866E-02
0.5750	3.2189E-01	3.227E-01	2.167E-01	4.8741E-02	9.530E-02	1.866E-02
0.5800	3.2473E-01	3.255E-01	2.152E-01	4.9194E-02	9.530E-02	1.866E-02
0.5850	3.2757E-01	3.283E-01	2.137E-01	4.9647E-02	9.530E-02	1.866E-02
0.5900	3.3041E-01	3.311E-01	2.122E-01	5.0100E-02	9.530E-02	1.866E-02
0.5950	3.3325E-01	3.339E-01	2.107E-01	5.0553E-02	9.530E-02	1.866E-02
0.6000	3.3609E-01	3.367E-01	2.092E-01	5.1006E-02	9.530E-02	1.866E-02
0.6050	3.3893E-01	3.395E-01	2.077E-01	5.1459E-02	9.530E-02	1.866E-02
0.6100	3.4177E-01	3.423E-01	2.062E-01	5.1912E-02		

0.2700	2.2304E-04	1.0151E-03	9.1073E-11	2.1662E-09	4.2163E-22	3.5966E-09	1.3434E-19	2.7225E-27
1.0000	1.3014E-04	1.0151E-03	3.5775E-11	1.1942E-09	1.3359E-38	1.2716E-09	8.7242E-22	4.6246E-43

ORIGINAL PAGE IS  
OF POOR QUALITY

## -SPECIES MASS FRACTIONS-

ETA	CF	C3H	C4H	HCN	C2	C+	ANH
0.0	6.957E-11	9.965E-11	9.995E-11	9.963E-11	4.218E-02	9.993E-11	1.334E+01
0.0030	6.971E-01	1.297E-06	2.240E-09	1.724E-06	1.852E-02	2.462E-08	1.321E+01
0.0111	6.987E-01	2.903E-06	5.090E-09	3.632E-06	1.553E-02	5.762E-08	1.304E+01
0.0153	6.999E-01	3.940E-06	7.050E-09	4.763E-06	1.060E-02	8.063E-08	1.292E+01
0.0200	6.997E-01	5.277E-06	9.748E-09	6.375E-06	1.104E-02	1.125E-07	1.278E+01
0.0231	7.002E-01	6.112E-06	1.151E-08	7.852E-06	1.168E-02	1.340E-07	1.264E+01
0.0263	7.007E-01	6.774E-06	1.342E-08	9.161E-06	1.254E-02	1.572E-07	1.259E+01
0.0294	7.013E-01	7.813E-06	1.535E-08	1.048E-05	1.350E-02	1.818E-07	1.250E+01
0.0325	7.018E-01	8.653E-06	1.737E-08	1.185E-05	1.457E-02	2.077E-07	1.241E+01
0.0356	7.023E-01	9.495E-06	1.947E-08	1.328E-05	1.575E-02	2.354E-07	1.231E+01
0.0388	7.029E-01	1.026E-05	2.173E-08	1.480E-05	1.709E-02	2.669E-07	1.222E+01
0.0420	7.034E-01	1.204E-05	2.437E-08	1.791E-05	2.060E-02	3.322E-07	1.202E+01
0.0451	7.039E-01	1.481E-05	3.340E-08	2.263E-05	2.522E-02	4.451E-07	1.175E+01
0.0482	7.044E-01	1.671E-05	4.150E-08	2.757E-05	3.167E-02	5.870E-07	1.147E+01
0.0513	7.049E-01	1.844E-05	5.025E-08	3.270E-05	3.769E-02	7.736E-07	1.119E+01
0.0544	7.054E-01	2.101E-05	5.963E-08	3.766E-05	4.456E-02	1.019E-06	1.092E+01
0.0575	7.059E-01	2.410E-05	6.530E-08	4.317E-05	5.529E-02	1.201E-06	1.076E+01
0.0606	7.064E-01	2.760E-05	7.413E-08	4.891E-05	6.594E-02	1.543E-06	1.053E+01
0.0637	7.069E-01	2.888E-05	8.310E-08	4.667E-05	7.667E-02	2.002E-06	1.031E+01
0.0668	7.074E-01	2.617E-05	9.527E-08	4.864E-05	9.042E-02	2.879E-06	1.002E+01
0.0699	7.079E-01	2.686E-05	1.065E-07	4.786E-05	1.018E-01	4.255E-06	9.759E+00
0.0730	7.084E-01	2.649E-05	1.152E-07	4.390E-05	1.076E-01	6.484E-06	9.513E+00
0.0761	7.089E-01	2.621E-05	1.213E-07	3.696E-05	1.060E-01	1.019E-05	9.291E+00
0.0792	7.094E-01	2.482E-05	1.217E-07	2.813E-05	9.622E-02	1.648E-05	9.093E+00
0.0823	7.099E-01	2.278E-05	1.156E-07	1.915E-05	7.970E-02	2.747E-05	8.914E+00
0.0854	7.104E-01	2.015E-05	1.023E-07	1.168E-05	5.977E-02	4.678E-05	8.748E+00
0.0885	7.109E-01	1.702E-05	8.321E-08	6.581E-06	4.347E-02	8.076E-05	8.587E+00
0.0916	7.114E-01	9.779E-06	3.445E-08	1.924E-06	1.375E-02	2.517E-04	8.276E+00
0.0947	7.119E-01	6.490E-06	1.944E-08	1.025E-06	7.506E-03	4.959E-04	8.144E+00
0.0978	7.124E-01	3.768E-06	8.657E-09	5.525E-07	4.458E-03	9.382E-04	8.037E+00
0.1009	7.129E-01	1.987E-06	6.542E-09	2.540E-07	2.788E-03	1.722E-03	7.979E+00
0.1040	7.134E-01	7.167E-07	7.839E-10	1.039E-07	2.140E-03	3.038E-03	7.931E+00
0.1071	7.139E-01	2.305E-07	1.528E-10	3.666E-08	1.270E-03	5.001E-03	8.084E+00
0.1102	7.144E-01	4.015E-08	1.571E-11	1.203E-08	6.272E-04	7.578E-03	8.245E+00
0.1133	7.149E-01	1.702E-09	3.517E-13	1.060E-09	1.162E-04	1.324E-02	8.712E+00
0.1164	7.154E-01	2.049E-11	2.004E-15	1.185E-10	6.847E-06	1.454E-02	9.434E+00
0.1195	7.159E-01	1.321E-13	9.853E-18	7.174E-12	3.211E-07	8.710E-03	9.652E+00
0.1226	7.164E-01	4.494E-16	6.126E-21	2.866E-13	1.052E-08	3.163E-03	9.932E+00
0.1257	7.169E-01	4.397E-19	5.408E-24	1.247E-14	2.004E-10	7.052E-04	9.807E+00
0.1288	7.174E-01	2.770E-22	8.952E-28	2.751E-15	2.662E-11	3.532E-04	9.643E+00
0.1319	7.179E-01	5.432E-25	7.573E-31	1.065E-15	2.384E-11	3.990E-04	9.507E+00
0.1350	7.184E-01	7.064E-27	8.692E-34	1.452E-15	2.027E-11	4.204E-04	9.397E+00
0.1381	7.189E-01	4.118E-27	1.789E-34	1.142E-15	1.802E-11	4.396E-04	9.304E+00
0.1412	7.194E-01	4.076E-27	1.213E-34	9.256E-16	1.621E-11	4.554E-04	9.224E+00
0.1443	7.199E-01	3.603E-27	8.742E-35	7.491E-16	1.478E-11	4.695E-04	9.154E+00
0.1474	7.204E-01	2.917E-27	6.465E-35	6.492E-16	1.357E-11	4.823E-04	9.091E+00
0.1505	7.209E-01	4.268E-27	4.892E-35	5.551E-16	1.253E-11	4.942E-04	9.035E+00
0.1536	7.214E-01	1.411E-27	3.738E-35	4.785E-16	1.161E-11	5.034E-04	8.982E+00
0.1567	7.219E-01	1.444E-27	2.870E-35	4.146E-16	1.078E-11	5.165E-04	8.939E+00
0.1598	7.224E-01	1.203E-27	2.206E-35	3.599E-16	9.249E-12	5.276E-04	8.898E+00
0.1629	7.229E-01	9.671E-28	1.676E-35	3.110E-16	9.234E-12	5.394E-04	8.860E+00
0.1660	7.234E-01	7.615E-28	1.243E-35	2.667E-16	8.467E-12	5.524E-04	8.825E+00
0.1691	7.239E-01	7.767E-28	8.738E-36	2.234E-16	7.632E-12	5.681E-04	8.794E+00
0.1722	7.244E-01	4.974E-28	5.761E-36	1.814E-16	6.720E-12	5.874E-04	8.764E+00
0.1753	7.249E-01	2.521E-28	3.121E-36	1.362E-16	5.617E-12	6.132E-04	8.738E+00
0.1784	7.254E-01	1.274E-28	1.333E-36	9.259E-17	4.316E-12	6.513E-04	8.714E+00
0.1815	7.259E-01	3.255E-29	2.379E-37	4.247E-17	2.591E-12	7.155E-04	8.693E+00



0.2700 6.0137E-01 3.5194E-10 1.0168E-32 1.2054E-17 1.1748E-12 7.9266E-04 9.6765E+00  
1.0000 7.0921E-01 2.5224E-03 1.7260E-50 3.6911E-18 1.7772E-13 8.9568E-04 8.6644E+00

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## VITA

John Frederick Balhoff was born Thursday, October 14, 1948 in Howell, Michigan about 30 miles west of Detroit to his proud but poor parents Jack and Kaye. Shortly after accepting his discharge from the hospital he made the acquaintance of his two older brothers, Mike and Tom. Little did he realize he was to be followed by six more brothers and two sisters - Bob, Don, Bill, Dick, George, Margaret, Kathleen, and Danny. At this writing it appears that no more siblings will be announced.

Elementary education was achieved at Sacred Heart and St. Aloysius parochial schools in Baton Rouge where a somewhat dubious career of academia commenced. In the Fall of 1962, secondary education was initiated at St. Joseph's Seminary in St. Benedict, Louisiana where an attempt was made to become a man of the cloth. Three years was sufficient to be convinced that the calling was no longer there and another profession would have to be pursued. The last year of secondary education was accomplished at Catholic High School in Baton Rouge and graduation came in May of 1966.

In June of 1966 the spurious career of a new chemical engineer began when the author enrolled at Louisiana State University in Baton Rouge. In May of 1970 he was awarded a Bachelor of Science in chemical engineering. Continuing his education by attending graduate school at L.S.U. in August, 1970, the author recieved his Master of Science in chemical engineering in May of 1972. Being ambitious and

somewhat foolish, pursuit of a doctorate was initiated immediately following his M.S. graduation. On the 29th of October 1973, prior to completion of all the requirements of the doctoral program, he started employment at Ethyl Corporation, Baton Rouge naively assuming that in a couple of months the work on the doctorate would be finished.

Presently the author is still working for Ethyl Corporation as a Process Evaluation Engineer and plans to attend graduation in May of 1975. At this time the author is not married with no immediate prospects in sight.